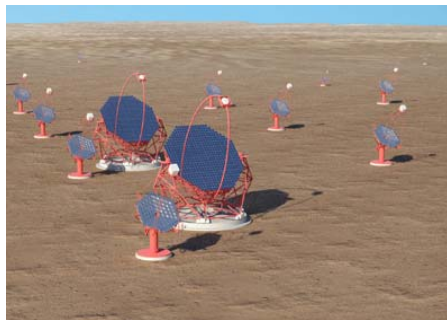


Time Synchronization and Array Trigger in CTA with WhiteRabbit : Ready to go ...

Ralf Wischnewski (DESY)
Martin Brueckner (HumboldtUniv Berlin)

CTA-Collaboration Meeting,
Rome / Italy
23.10.2012



Outline

- > Array Time Synchronization for large distributed arrays
 - A generic job: for modern astroparticle experiments
 - Going for a *standardized system*

- > WhiteRabbit for Time Synchronization
 - Synchronization & TimeStamping – WRabbit features & architecture
 - WRabbit is HiSCORE's top candidate

- > WhiteRabbit at DESY: Tests and Installation
 - Started January-2012: “Node \leftrightarrow Node” @ field
 - August-2012: “Switch \leftrightarrow Multi-Node”
 - 23rdOctober-2012: \rightarrow Rabbits are running in Siberia .

- > Summary
 - WhiteRabbit – Clock distribution: ready to go.
 - Trigger-time stamping @nsec - ready (in the DESY-implementation)



Array Time Synchronization : The Task

- > Time Synchronization of large distributed arrays
 - Experiments like HiSCORE (10km²), CTA (few km²), KM3Net, AUGERnext, LHAASO
 - Detectors (Cameras, Stations) generate a “local trigger signal” , which needs to be time-stamped

- > Two Tasks:
 - (1) **Build a time distribution system**

Synchronize clocks in all detector units to a common central clock with nsec precision;
Focus: relative timing between detector units must be correct at any moment (not the absolute time)
 - (2) **Verify the time distribution system**

Long-term monitor / control in-situ, for at least a subset of the installed array.
Verification by an *independent* system is as important as the synchronization;
and a comparable in technical challenge

- > White Rabbit
 - WR-Specifications: An excellent candidate
 - In reality: a system still under development - so needs a detailed verification.



White Rabbit (1)



White Rabbit – a new CERN-based extension to Ethernet / PTP for :

- > Synchronous mode (clock synchronization)
- > Deterministic routing (package latency guaranteed)

Performance :

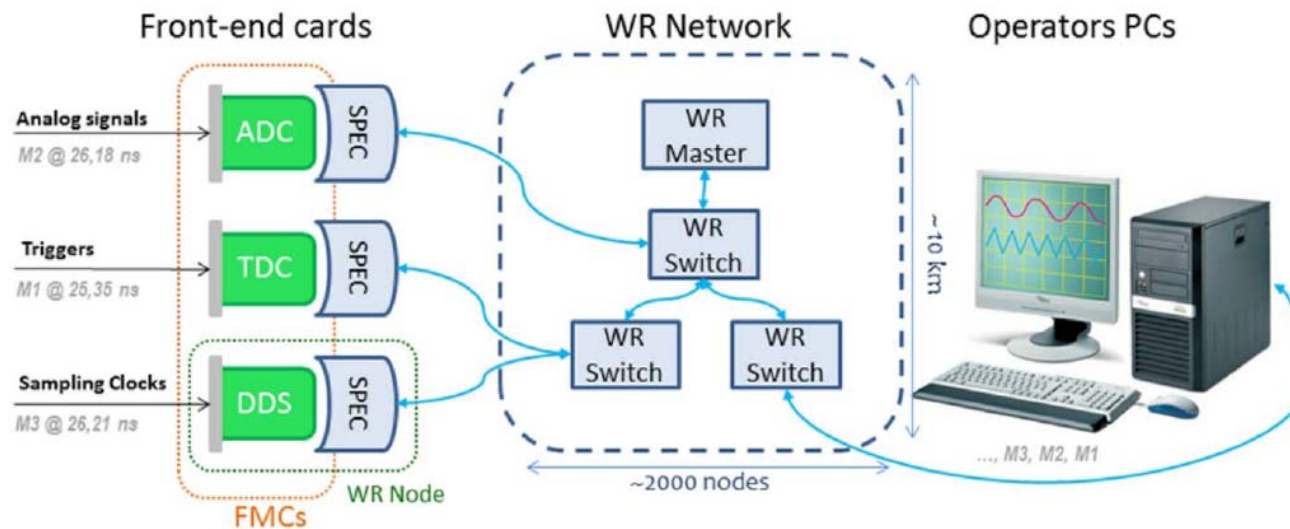
- > ~ 1 ns precision, 20 ps jitter
- > 10 km fiber links
- > Up to 2000 nodes

Important :

- > Development for CERN & GSI accelerator complex; much external interest
- > Open Hardware & SW Project w/ peer review (ie. open for extensions)
- > Hardware is commercially available
- > Standardization planned (IEEE...)
- > A guaranteed large user community: it will be a well debugged system... !!



White Rabbit (2)



■ For concept + details, see eg:

Cern-WR-site <http://www.ohwr.org/projects/white-rabbit/wiki> ;

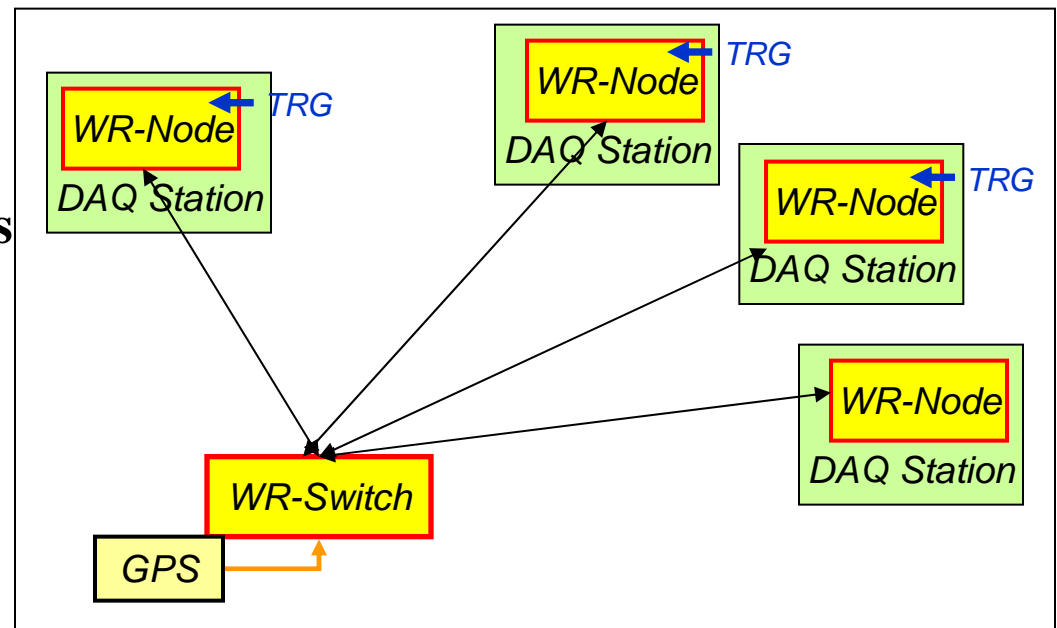
(also <http://znwiki3.ifh.de/TUNKA/>)



Basic Layout of a WhiteRabbit based timing DAQ

White Rabbit DAQ architecture:

- (1) **WR-network**
= **WRSwitches + GPS/RbCl**
to distribute the clock to Nodes
- (2) **WR-Nodes** ('endpoints')
for time-stamping, ...



Components

- WhiteRabbit Switches (WRS)
- White-Rabbit Nodes
eg. SPEC (Simple PCIe FMC Carrier)
(or build your own board: OWHR)



since July/2012



since 2011

The WR Node : SPEC Card

- **SPEC (“Simple PCIe FMC Carrier”)** is the WR node currently available.
- **For tests, it can also be configured as WR-Master** (used for April/2012 tests).
- **It carries the mezzanine-card for your DAQ:**
available/planned
 - **Digital InpOutFMC / FMC DEL / FMC FADC(100MHz) / TDC;**
 - **eg. possible: design a DRS4-based mezzanine for HiSCORE/CTA**

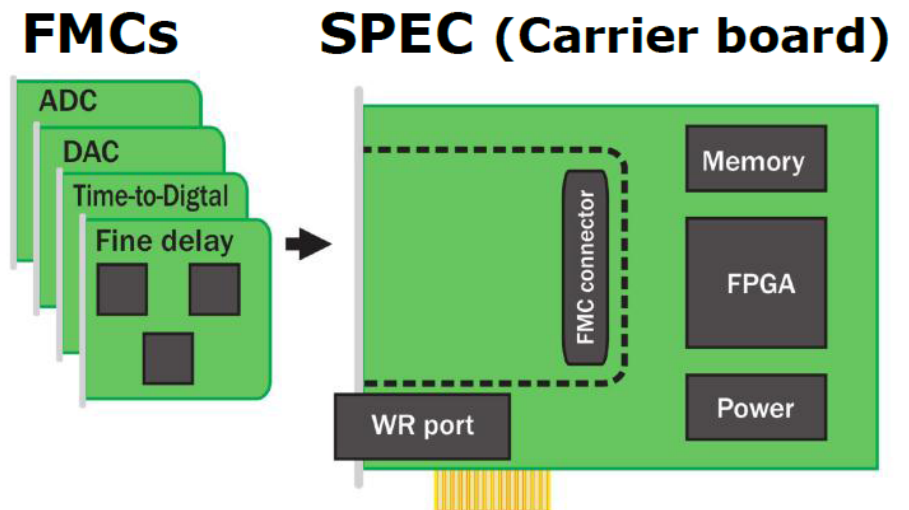
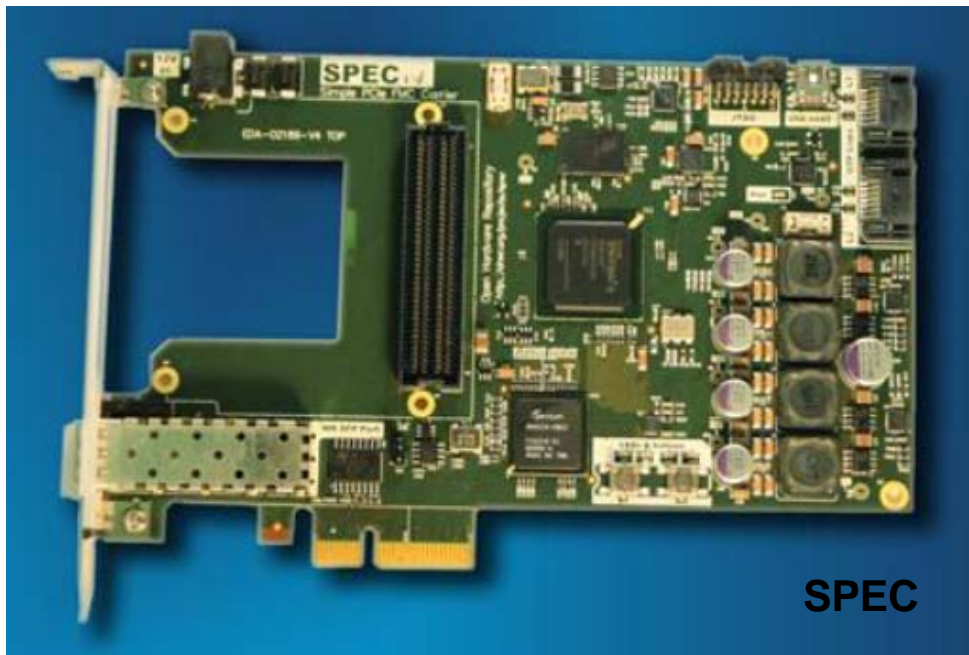
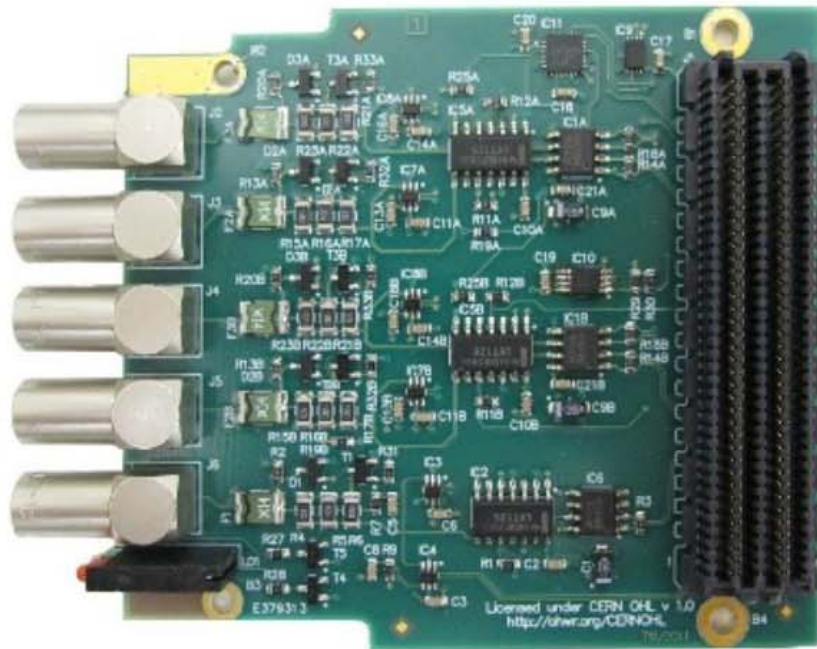


Illustration of the connection of SPEC with FMCs

FMC-DIO-5CHTTLA FMC 5-CHANNEL DIGITAL I/O MODULE

The *fmc-dio-5chttla* 5-channel digital I/O module is a simple board for digital I/O on LEMO connectors.

It has been designed for testing White Rabbit functionality as part of the SPEC Demonstration Package for White Rabbit ([manual](#)), and it can be used for other applications too.



FUNCTIONAL SPECIFICATIONS

- 5 input/output ports (Lemo 00 connectors)
- Output levels: LVTTTL, capable of driving +3.3 V over a 50-Ohm load. At power-up the outputs should be in Hi-Z state
- Input levels: any logic standard from $V_{ih} = 1$ V to $V_{ih} = 5$ V (programmable threshold)
- Output Rise/fall times: max. 2 ns
- Input bandwidth: min. 200 MHz
- Programmable 50-Ohm input termination in each channel
- LVDS I/O on the carrier side
- One of the inputs shall be capable of driving a global clock net in the carrier's FPGA
- Inputs need to be protected against +15V pulses with a pulse width of at least 10us @ 50Hz (with protection diodes if possible)
- Withstands a continuous short-circuit on all the outputs at the same time

Example Layout for HiSCORE clock / trigger (~CTA)

HiSCORE / EA : 20-40 DAQ-Stations (km² prototype)
CTA : 50-100 Telescopes

> *Array center:*

WR-switches (18 x out) - with N=4 → 67 Stations/Telescopes
(+ 1x GPS / GPS-disciplined Rub.Clock)



> *Dedicated SM-fiber* to every Station / Telescope

> *Every Station / Telescope:*

Houses one WR-SPEC time-latching unit : SPEC



> Price: $O(1100\text{Eur})$ per station

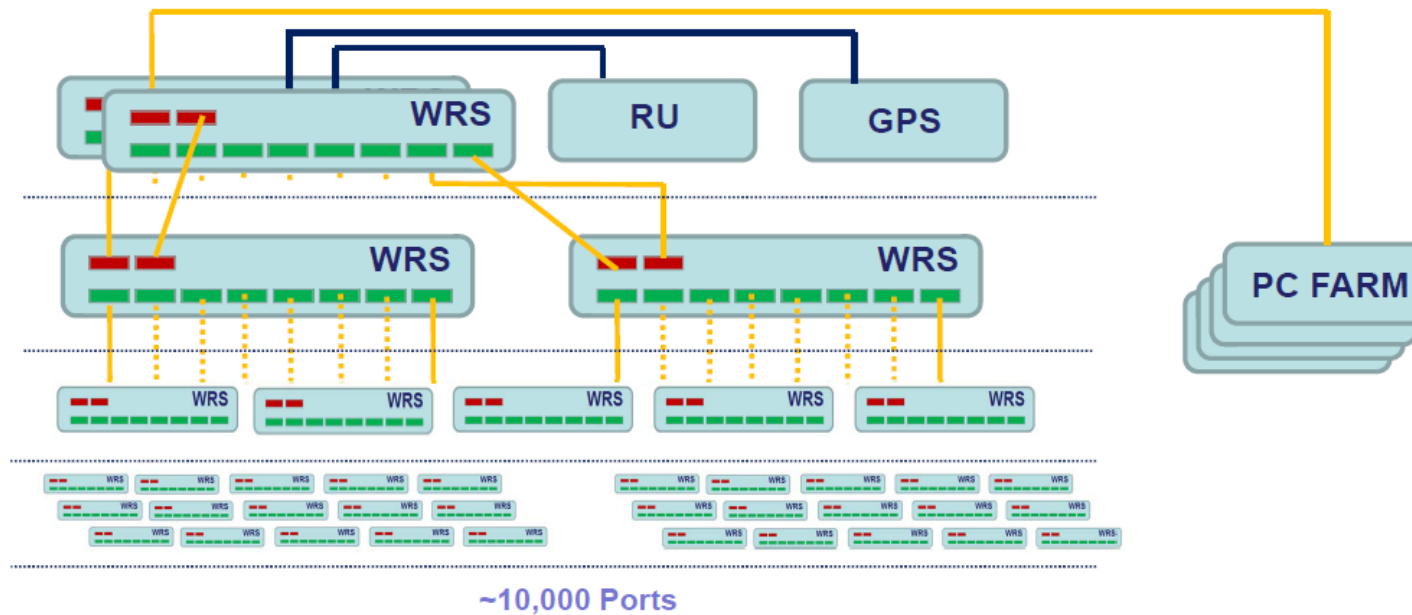
= 900 (SPEC+DIO) + $1/17 \times 3500$ (WRS) (very conservative)



Example: LHAASO - WhiteRabbit Layout

- > Other WR-application proposals:
 - **LHAASO : ~10000 nodes to synchronize**

Design study
G.Gong, ICALEPCS, 2011.



WRabbit Tests @ DESY - Overview

> Started January-2012:

Basic functional & field test “Node \leftrightarrow Node”

> August-2012:

Full architecture setup “Switch \leftrightarrow Multi-Node”

> 23rdOctober-2012:

WR@HiSCORE commissioned “Switch + 5 Nodes”

→ ***Rabbits are running in Siberia***

> The DESY-setup: Each Node (ie. the SPEC-PCIe card) contains

(1) the standard WR-Clock

(2) TimeStamping with 125MHz (8ns) resolution /DESY/

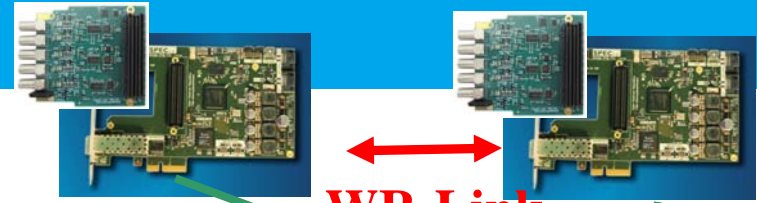
(3) Precision TDC 2 ns (future: 1ns) for ns time stamping /DESY/

(4) TriggerTime UDP-transport over WR-link /DESY/

(5) Various Test-signal INPUT/OUTPUT via the FMC-5CH-DIO mezzanine



Laboratory Tests: March 2012



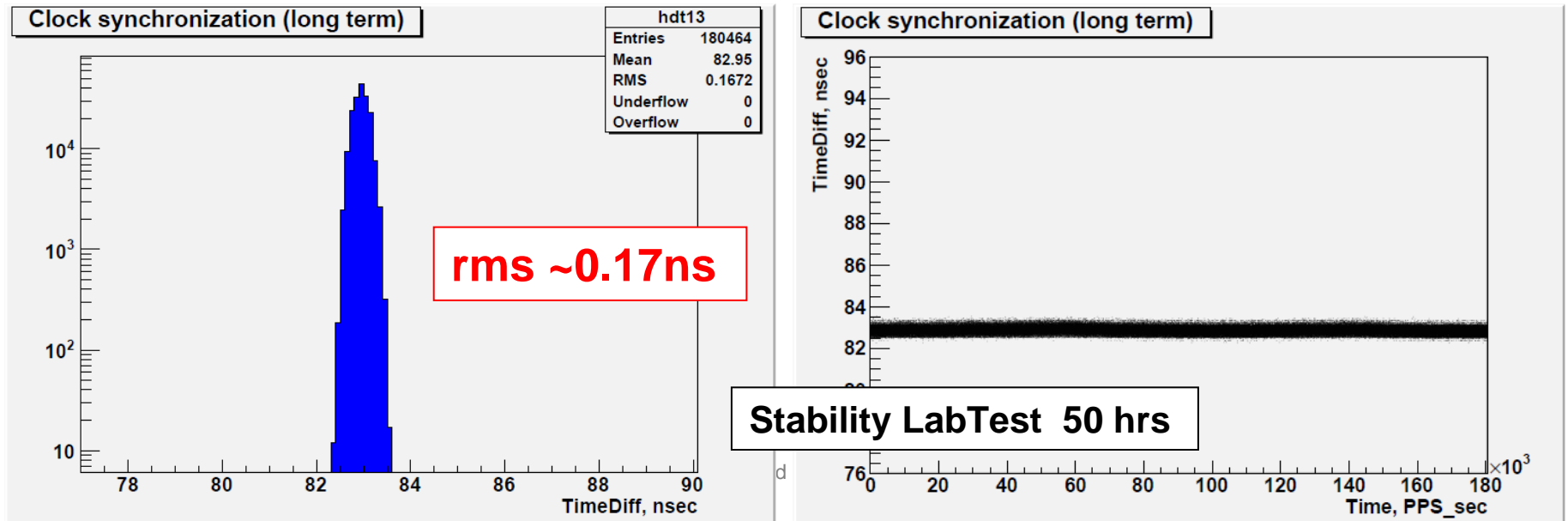
> Setup:

- Two WR-SPECs: Master-Slave WR-link + fiber
- Time tests: comparing the master/slave 1-PPS output
- Setup emulates the basic element: WRS+SPEC

1 PPS timing
(DRS4 5GS/s)

> Result: **TimeDiff (Master-Slave) rms < 0.2 ns !!**

(reaching time-measurement precision)



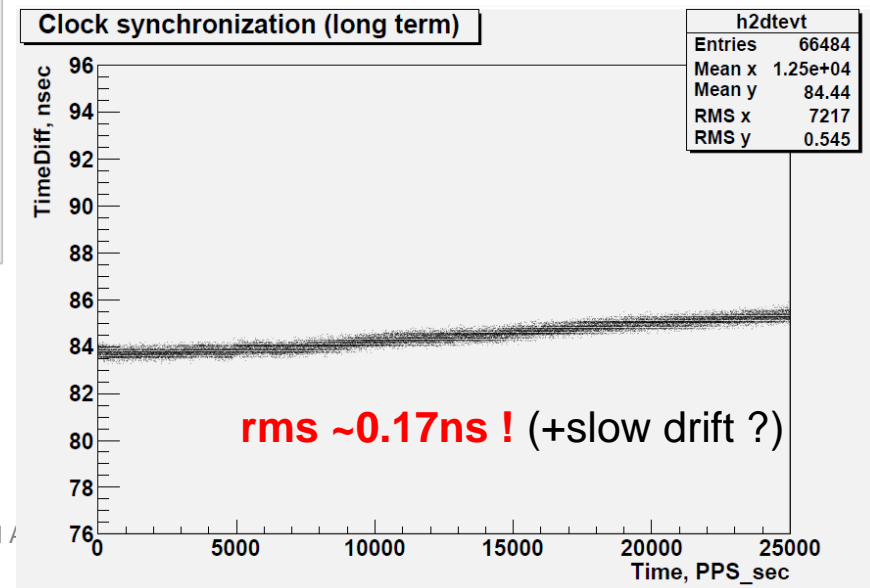
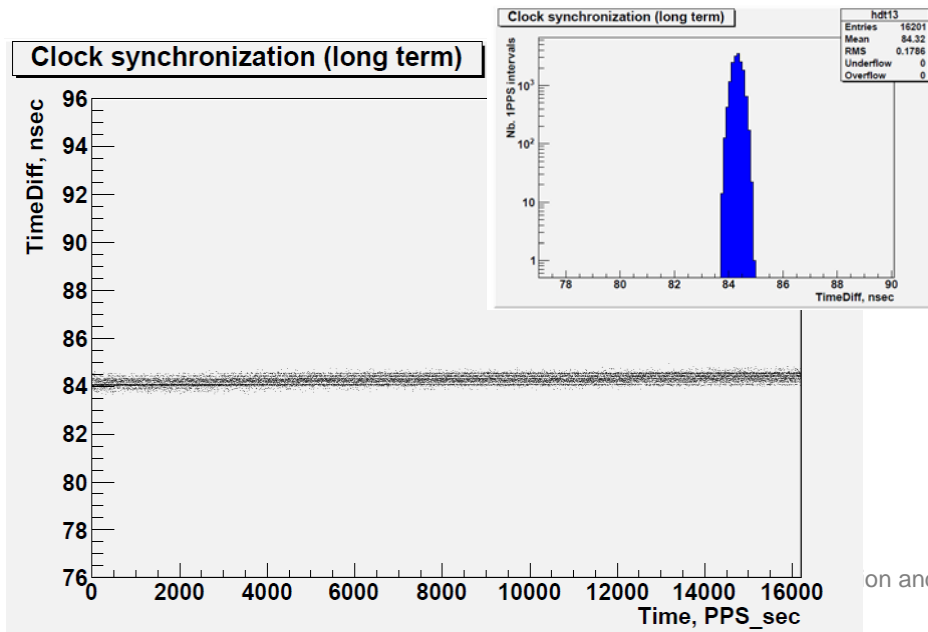
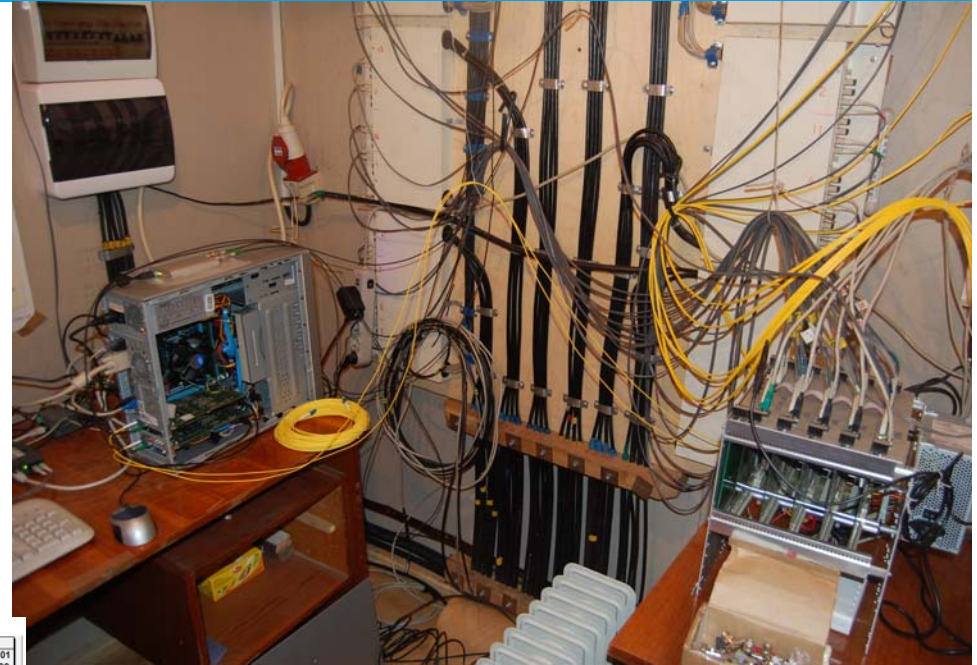
Test at Tunka/HiSCORE : April 2012 2km-Loopback

> Field tests at Tunka/HiSCORE site

- April 2012: use DESY-Lab-setup
- use the real 1+1 km Tunka-fibers

> Confirmed Lab-result: rms <0.2 ns !

- Quick setup; rare occasions of slow drifts (artefacts or temperature?)



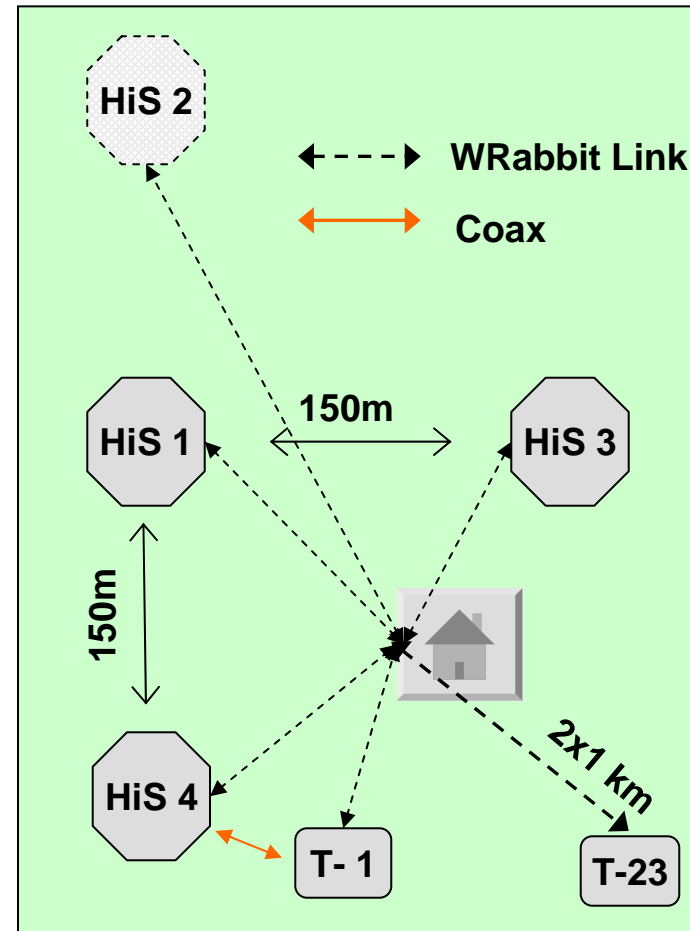
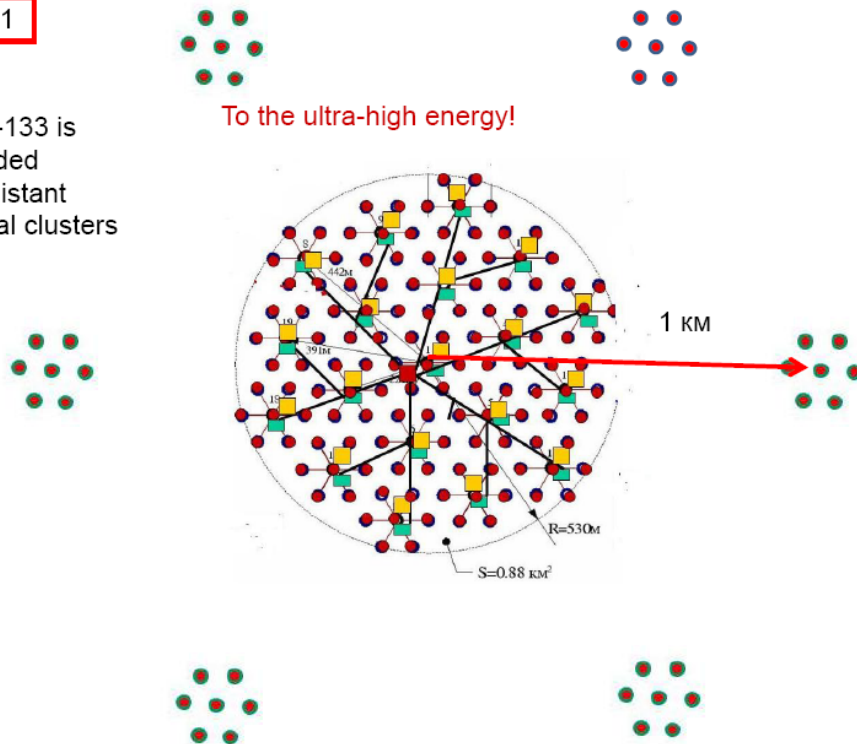
The HiSCORE prototype setup: Oct. 2012

- > Three HiSCORE Prototype Stations deployed in Oct.2012

Tunka-133 Array (running independently)

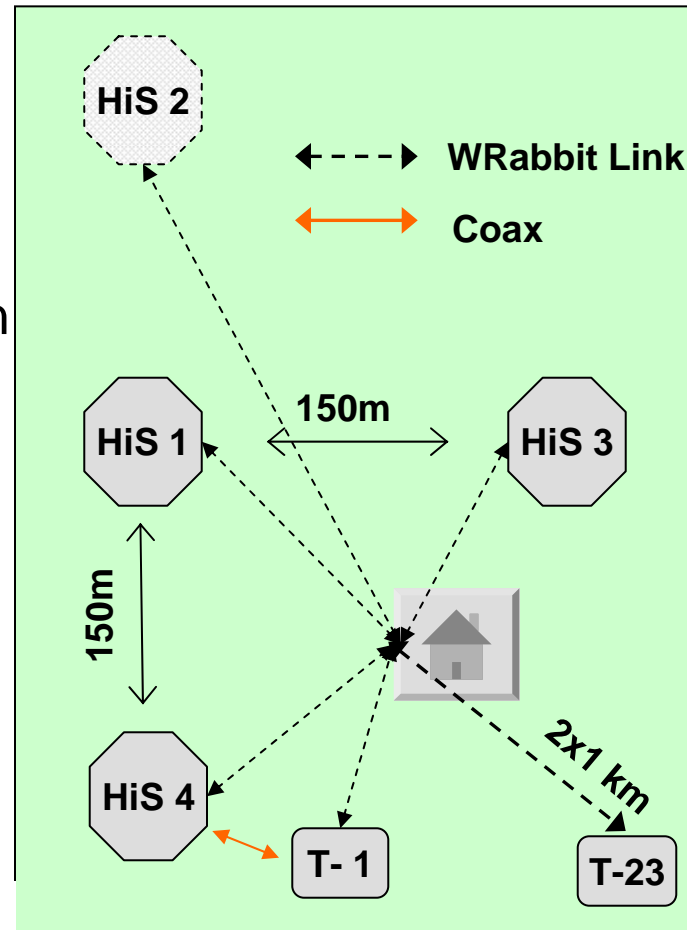
2011

Tunka-133 is extended by 6 distant external clusters



The HiSCORE prototype and WhiteRabbit

- > Three HiSCORE Prototype Stations deployed in Oct.2012
- > WhiteRabbit is TriggerLatch of HiSCORE
 - 1 WR-Switch + 6 WR-Nodes
 - Various cross calibration tools foreseen
 - Twin – WR-Nodes
 - Coax – WR-Nodes connection
 - Tunka-timing system (10ns)
 - Fiber-Loopback to Laboratory



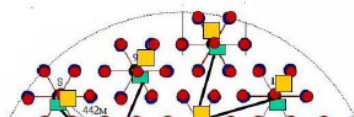
Tunka-133 Array (running independently)

2011



To the ultra-high energy!

Tunka-133 is extended by 6 distant external clusters



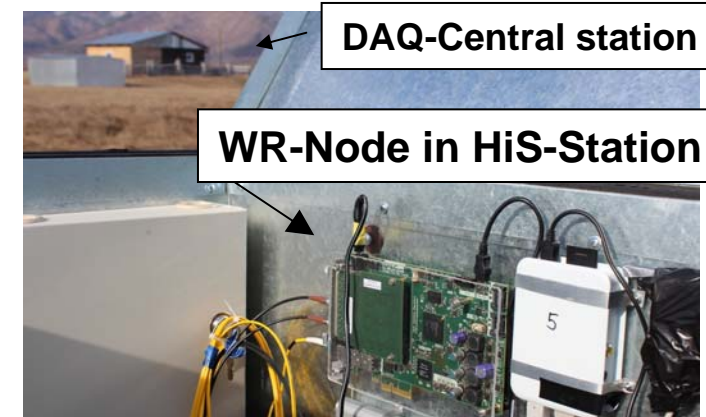
Test at Tunka/HiSCORE

- > A full WhiteRabbit system is running
 - WR-Switch
 - 6 WR-Nodes
- > Confirmed Prototype Result: rms <0.2 ns
 - Analysis is in progress, stay tuned

WR-Switch



DAQ-Central station



WR-Node in HiS-Station

HiS-Station



Test at Tunka/HiSCORE

- > A full WhiteRabbit system commissioned
 - 1x WR-Switch & 6x WR-Nodes
- > Confirmed Prototype Result: rms <0.2 ns
 - Analysis is in progress, stay tuned

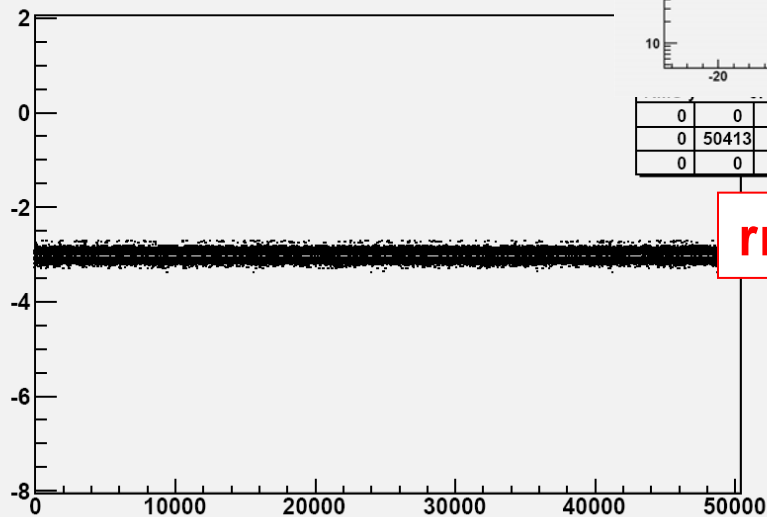
WR-Switch



1. Test :

**LoopBack Fiber over 2x1km
back to Laboratory
(5GHz DRS4-measurement)**

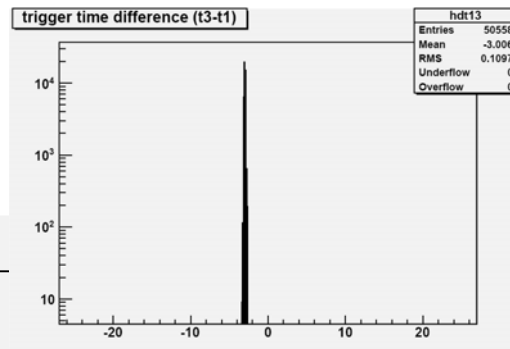
time difference (t3-t1) % event number (PPS sec)



rms <0.2 ns

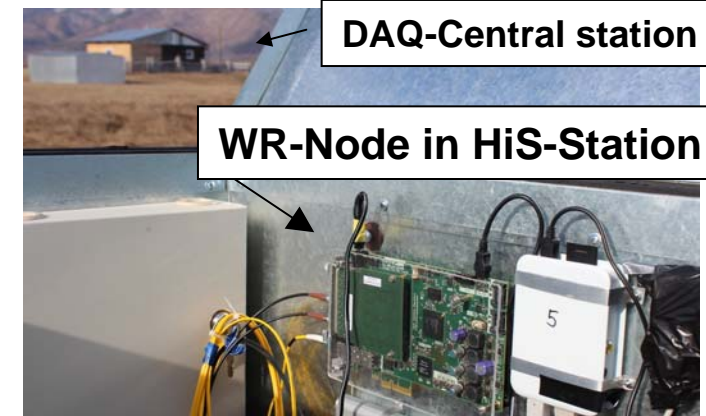
14hr run

ization and Array Trigger in CTA



DAQ-Central station

WR-Node in HiS-Station



HiS-Station



Test at Tunka/HiSCORE

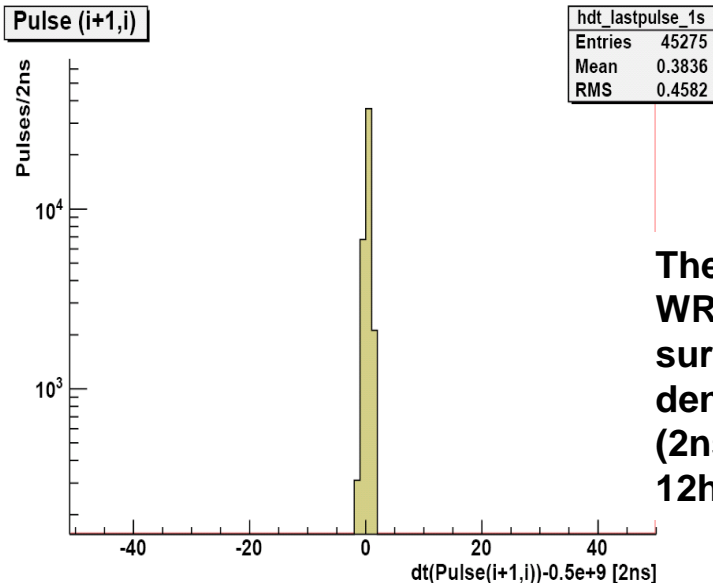
- > A full WhiteRabbit system commissioned
 - 1x WR-Switch & 6x WR-Nodes
- > Confirmed Prototype Result: rms <0.2 ns
 - Analysis is in progress, stay tuned

WR-Switch



2. Test :

2 WR-Nodes in neighboured stations, connected by 95m CoaxCable (sending 1PPS signals), TimeStamping with SPEC-TDC like for trigger. Excellent time stability !



fwhm <2 ns

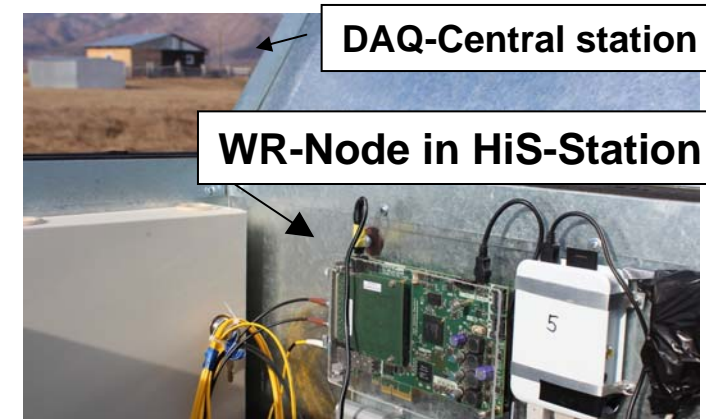
The time difference of WR-1PPS pulses, measured with an independent WR-station (2ns TDC resolution, 12hr run).

2ns/bin; average == 1.00sec

ynchronization and Array Trigger in CTA

DAQ-Central station

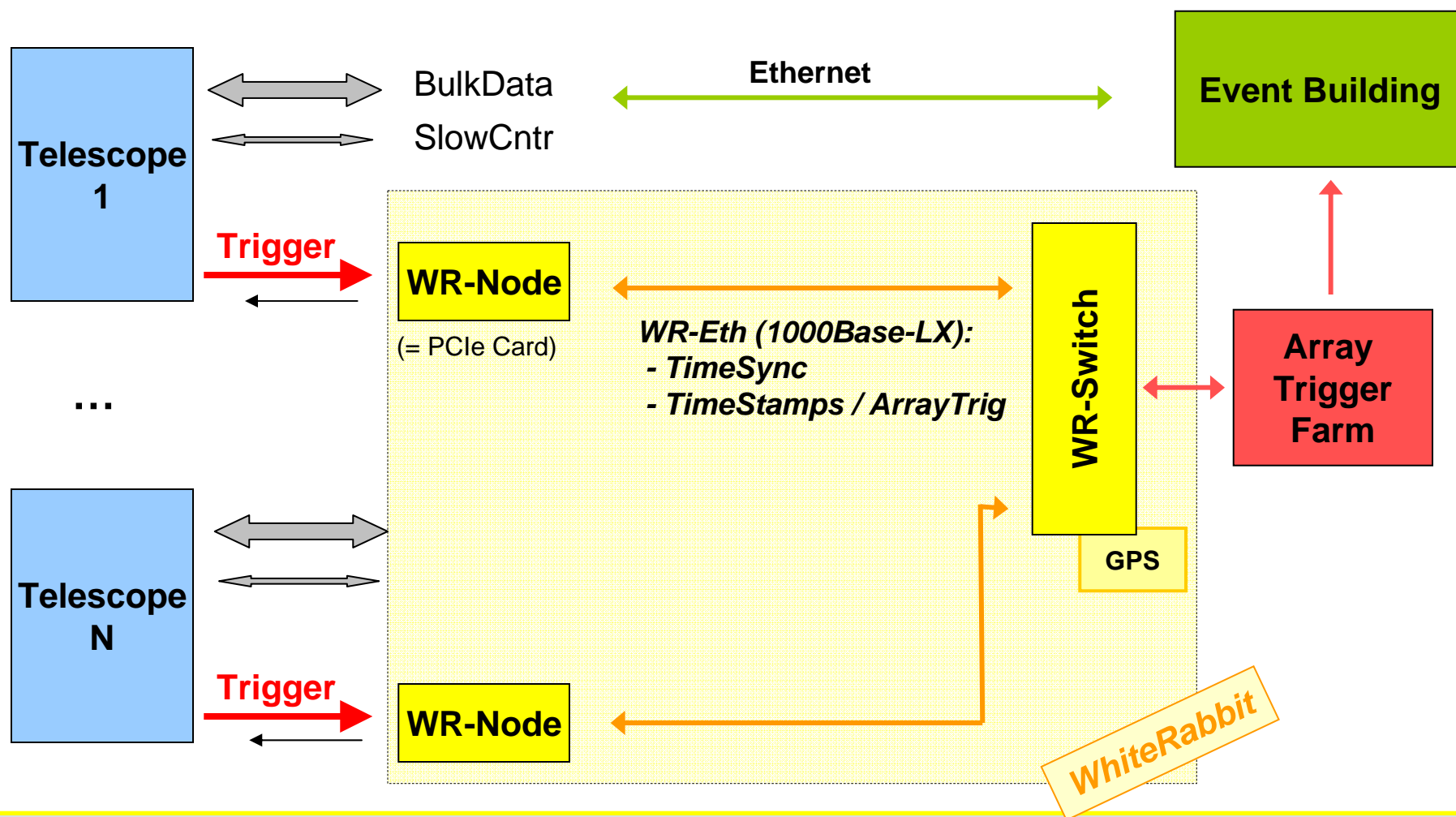
WR-Node in HiS-Station



HiS-Station



White Rabbit : Layout for Array-Timing, TimeStamping & ArrayTrigger

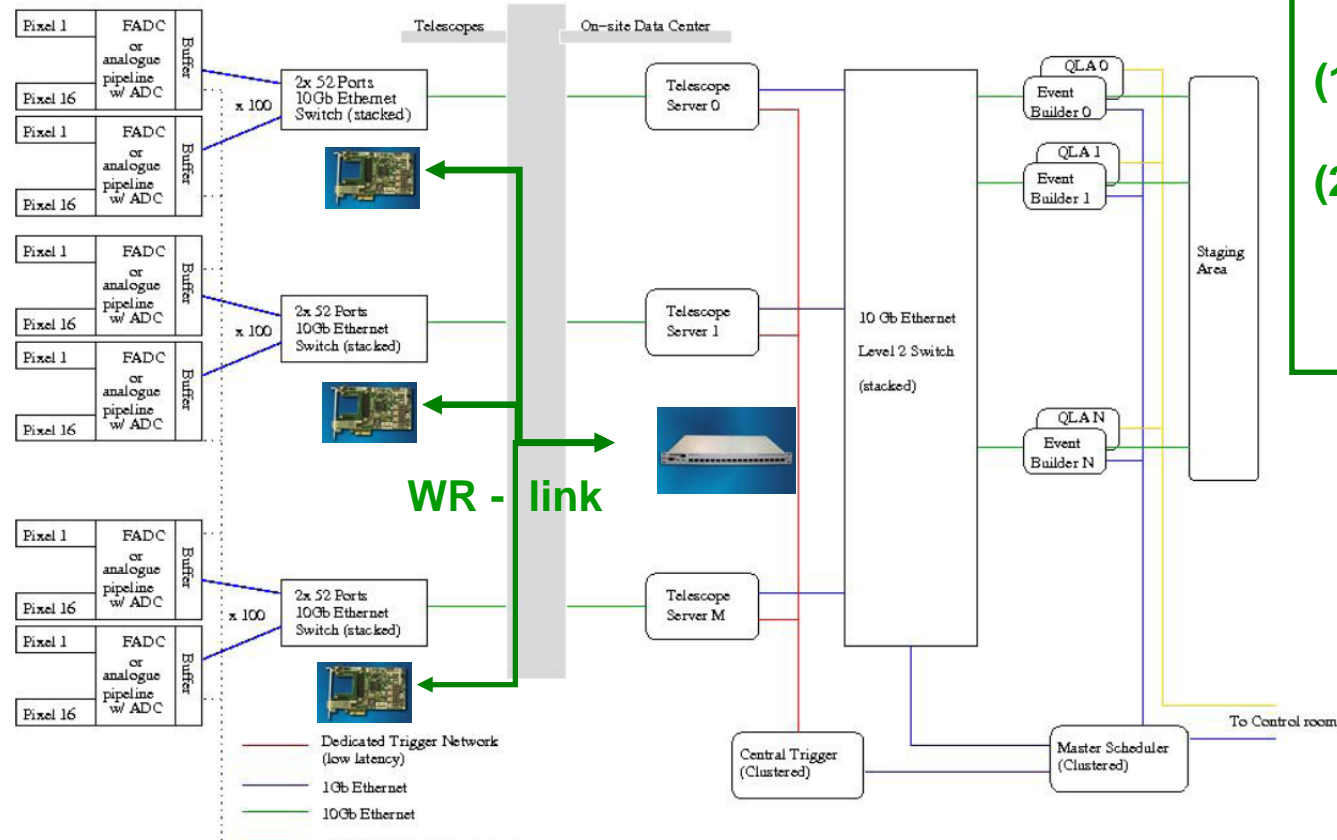


Only a few components are needed:

- per telescope : 1 x WhiteRabbit Node (PCIe) + 1 standard fiber (SM;1390/1510nm)
- per array : 1...n WhiteRabbit Switches

CTA : DAQ with White Rabbit - Timing & ArrayTrigger

DAQ Architecture



Use the WR-link for

(1) nsec-timing, and

(2) Array-trigger request/
confirm messages
(1Gbs) with guaranteed
latency.

Here, we considered just the baseline functionality: Time-stamping of Camera trigger. There many are more options, including development of CTA-WR boards; Which could integrate FE/DAQ components with WR-cores. E.g.: a WR-Mezzanine with DRS4 (or NecTar).



Summary

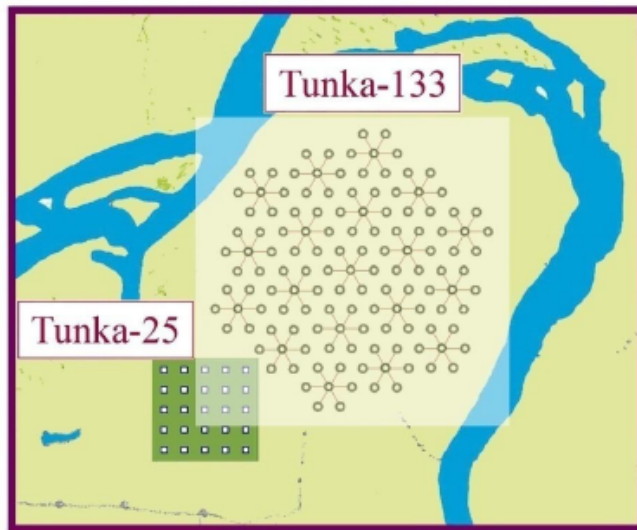
- > **White Rabbit: Advantages, that make WR a top candidate for CTA:**
 - **Real standard, commercial support, open source HW & SW**
 - **Reliability, easy maintenance, cost effective, scalability, ...**
 - **Big user-community. Eg. HiSCORE, LHAASO, ...**
- > **Thus, for CTA, White Rabbit is an excellent candidate for**
 - **Time-synchronization: local clocks - specs are fulfilled**
 - **Array trigger : favourable network architecture for time-stamp based coincidences**
- > **DESY build and operates a realistic WR Trigger TimeStamping setup**
Functionality added: 8ns and 1(2) ns latching, WR-data transport, ...
- > **Commissioned for routine DataTaking at HiSCORE/Tunka**
 - **Clock stability (longterm) < 0.2 ns rms**
- > **Need: Independent performance verification in the field (GPS, Radio, MUTIN...)**
- > **DESY is committed to HiSCORE WR-DAQ; thinks about a related CTA initiative. ASTRI ?**

> Thanks.

> Additional slides....



The Tunka-133 Cherenkov-EAS Array

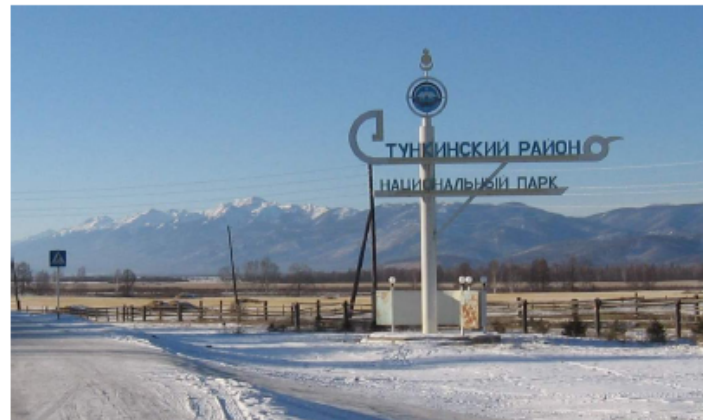


Tunka-133 – 1 km² “dense” EAS
Cherenkov light array

Energy threshold 10¹⁵ eV

Accuracy: core location ~ 10 m
energy resolution ~ 15%
 $\delta X_{\max} < 25 \text{ g}\cdot\text{cm}^{-2}$

51° 48' 35" N
103° 04' 02" E
675 m a.s.l.



The Tunka-133 Cherenkov-EAS Array

2011

Tunka-133 is extended by 6 distant external clusters

To the ultra-high energy!

