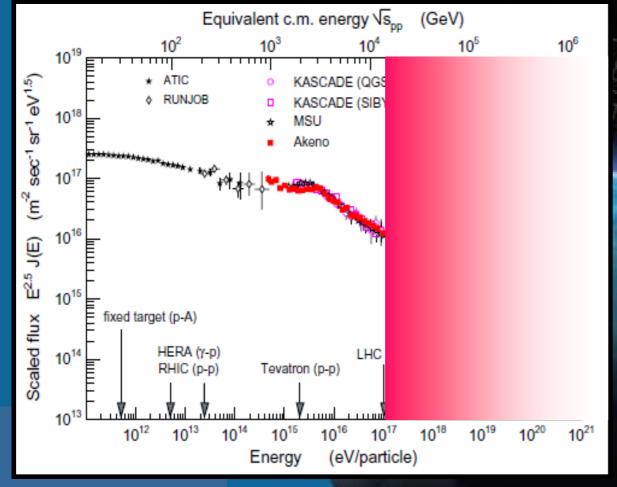
Astroparticle Physics and the LHC

Christian Spiering, DESY

EPS-ECFA 2009 Cracow, July 2009

http://www.aspera-eu.org

Astroparticle Physics helps LHC



No black holes from LHC which would eat the Earth !

http://www.aspera-eu.org

Appec Astroparticle Physics for Europe

CERN & Cosmic Rays

- Cross section @ high energies
 - TOTEM @ CMS
 - ATLAS forward (ALFA)
- Inelasticity and spectra
 - Zero Degree Calorimeters ATLAS, CMS, ALICE

LHCf

- dedicated CR-study experiment
- individual particle, particle ID, imaging calorimetry,

NA61:

Detailed inspection of forward region

Auger:

Composition seems to become heavier towards higher energies

- Astrophysics ?
- "New" physics ?
- Details of air shower development ?

Improve air shower simulations and understanding of CR spectra and composition at highest energy

- knee
- towards end of galactic spectrum
- GZK region

SUSY and Dark Matter

http://www.aspera-eu.org

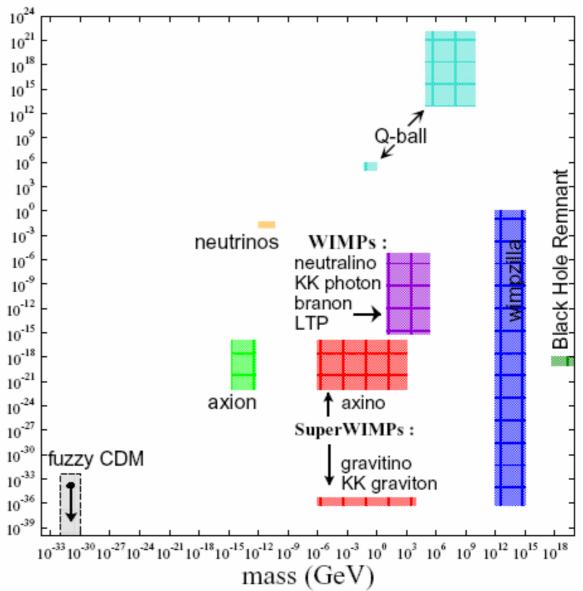
Dark Matter Searches

DM candidates: • WIMPs

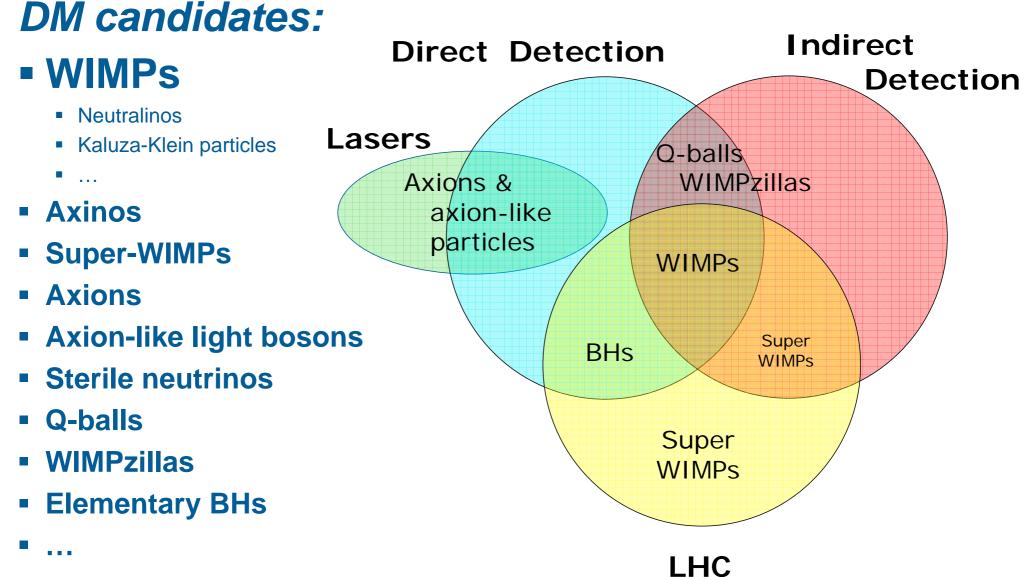
- Neutralinos
- Kaluza-Klein particles
- • •
- Axinos
- Super-WIMPs
- Axions
- Axion-like light bosons

σ_{int} (pb)

- Sterile neutrinos
- Q-balls
- WIMPzillas
- Elementary BHs



Dark Matter Searches



Dark Matter Searches

Dark matter search strategies 1. Direct detection > Milky way 2. Indirect detection > γ, e+, p WIMP WIMP Sun Earth < 3. Production at the Large Hadron Collider

DM candidates:

WIMPs

- Neutralinos
- Kaluza-Klein particles
- • • •
- Axinos
- Super-WIMPs
- Axions
- Axion-like light bosons
- Sterile neutrinos
- Q-balls
- WIMPzillas
- Elementary BHs

Direct Dark Matter Searches

DM candidates: • WIMPs

Astroparticle Physics for Europe

Neutralinos

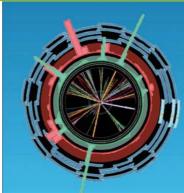
Appec

Kaluza-Klein particles

Dark matter search strategies

1. Direct detection

~ 20 experiments worldwide
→ Need of convergence



Milky way

WIMP



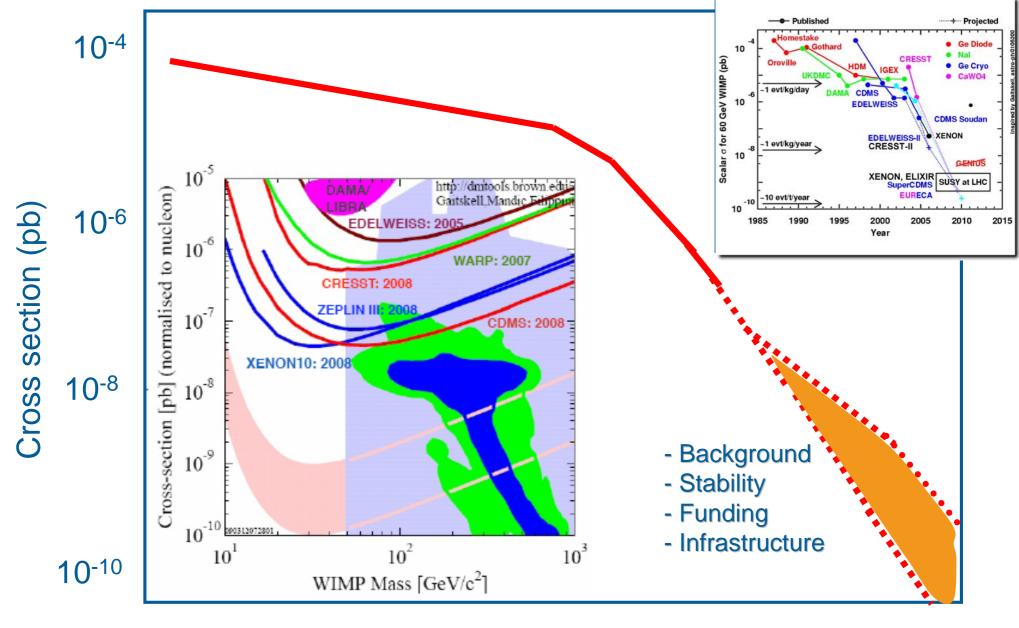
< 3. Production at the Large Hadron Collider

WIMP

Sun

WIMP

Direct Dark Matter Searches



1985

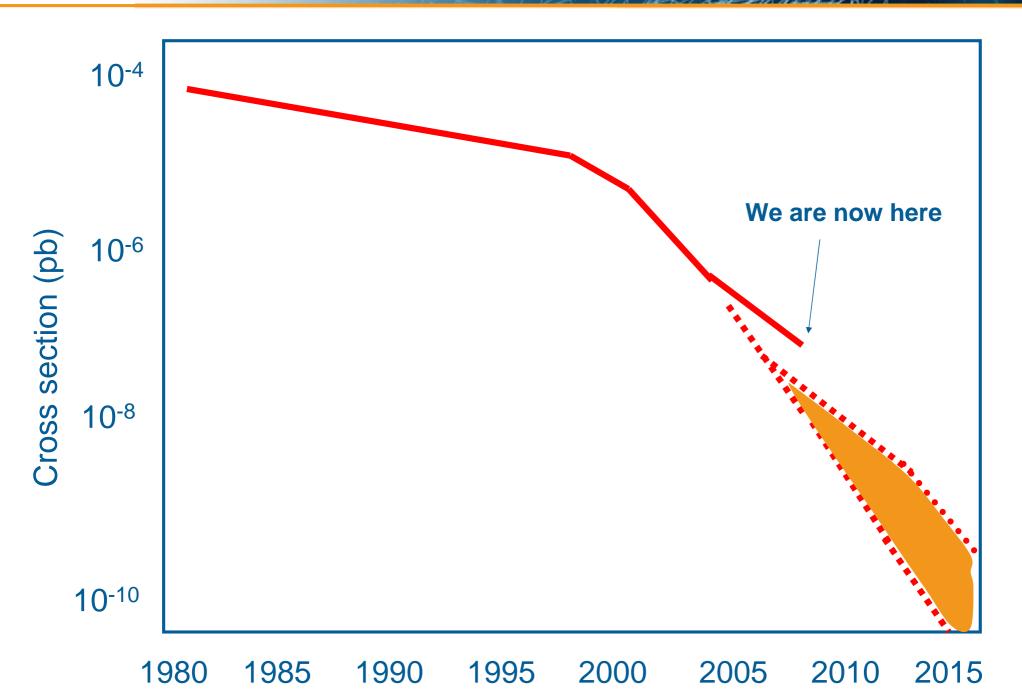
1990

1995

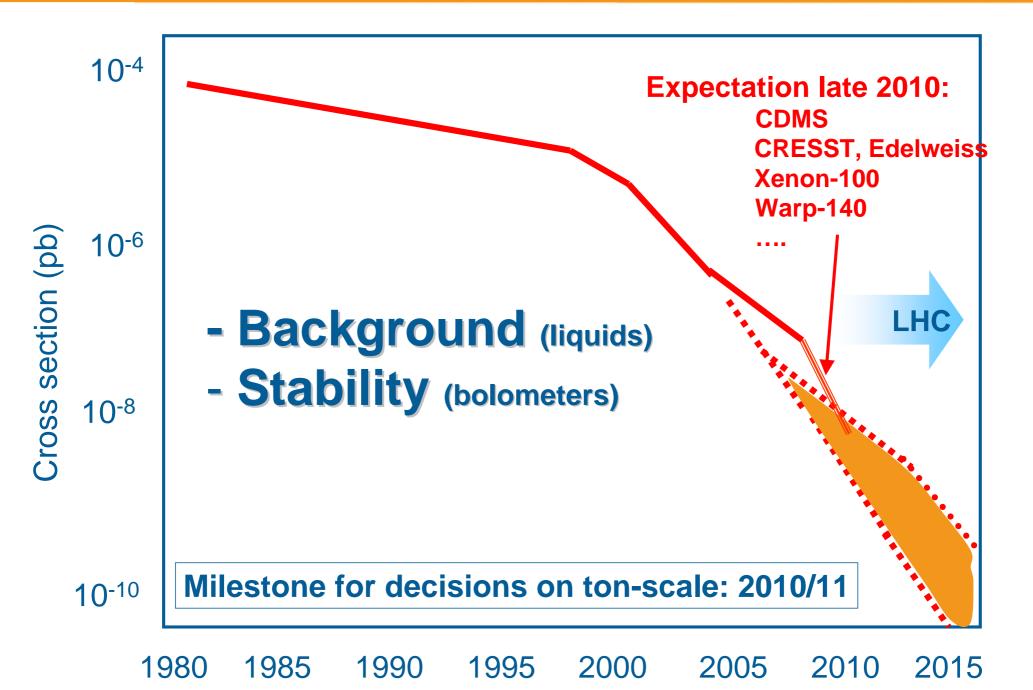
2000

2005 2010 2015

Direct Dark Matter Searches



Direct Dark Matter Searches





Indirect Dark Matter Searches

DM candidates: • WIMPs

- Neutralinos
- Kaluza-Klein particles
- • • •



0.02

0.01

10

20

50

100

 E_{e} [GeV]

200

0.010Fermi prelim. 10°<|b|<20' I FI FI FI FI E₇³Φ₇[MeV cm⁻²x⁻¹s⁻¹] 0 EG PM only Pulsars only 10 10^{2} 0.110E_r [GeV] 5001.000.50 200 PAMELA (²0.20 ¹0.10 ¹0.05 ²0.05 100Galactic bkg

500

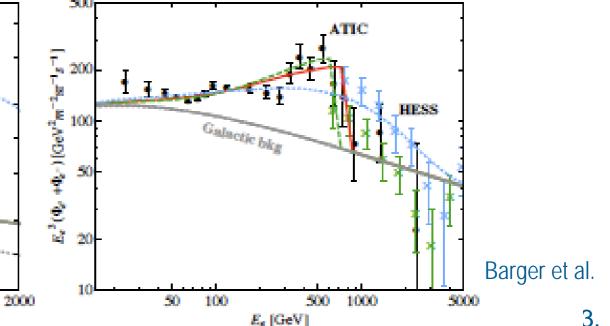
1000

Dark Matter annihilation ?)

Indirect Searchest Pamela & Co.

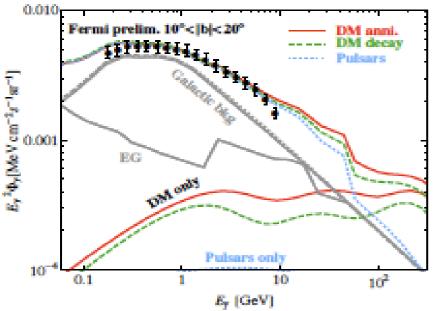
- Dark Matter Decay?
- Nearby Pulsars?
 - Remember: Many discoveries in astroparticle and astrophysics started as puzzle !

– Neutrino oscillations, quasars,



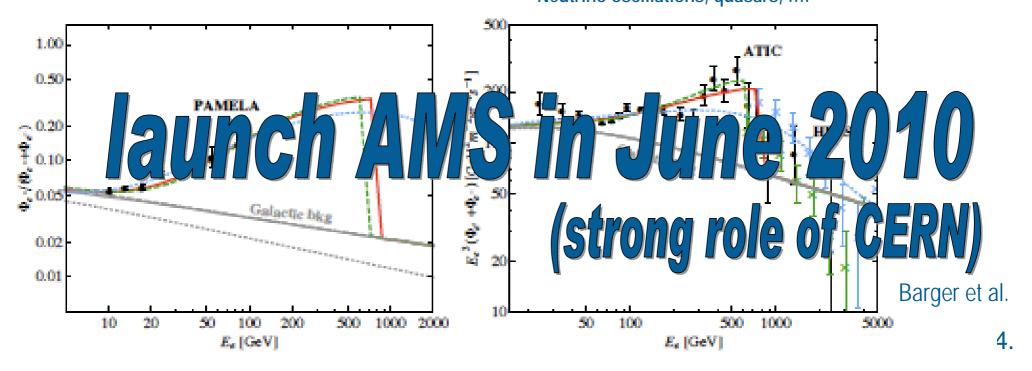
• New Physics

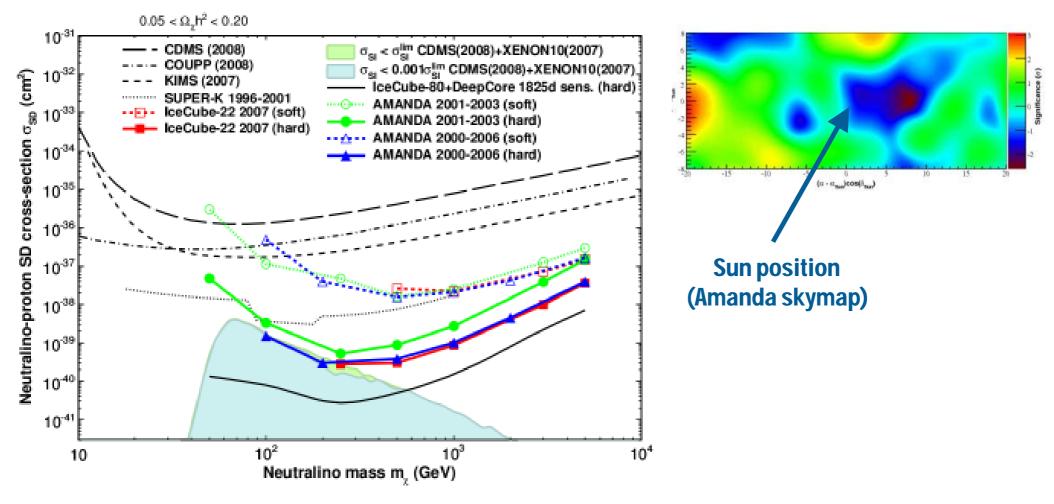
Astrophysics



- Dark Matter annihilation ?
 - Dark Matter Decay?
 - Nearby Pulsars ?
 - Remember: Many discoveries in astroparticle and astrophysics started as puzzle !

 Neutrino oscillations, quasars,





Indirect Searchest Example leec

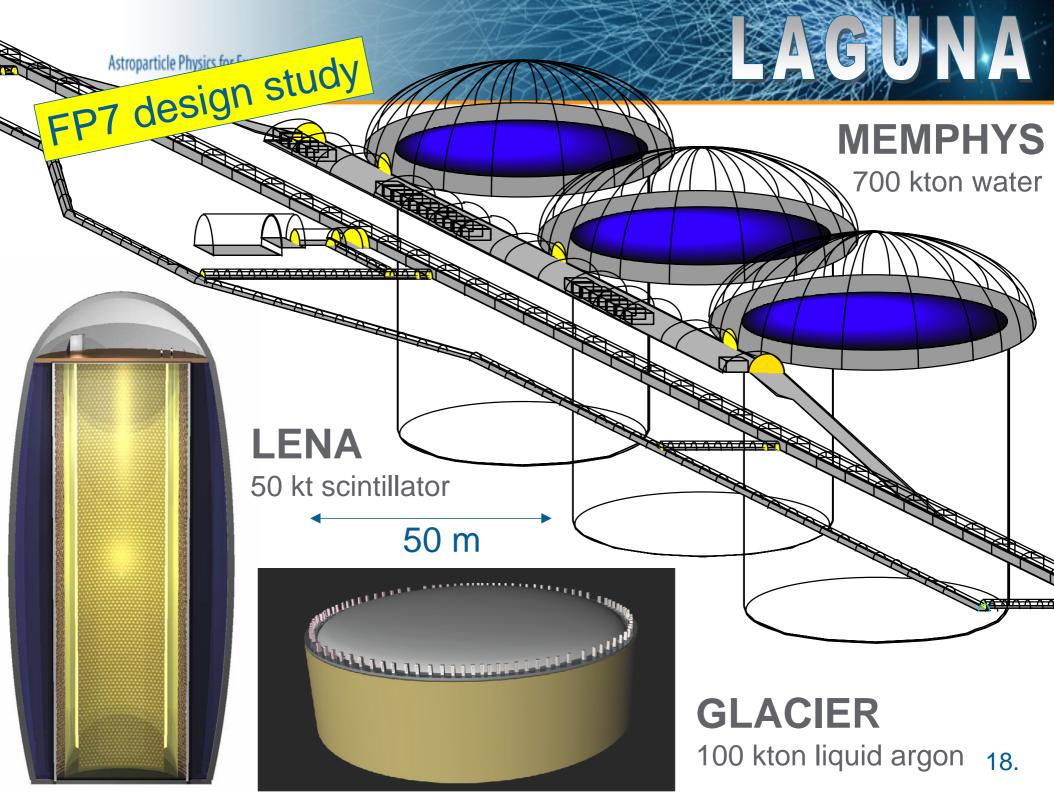
- Models with strong spin-dependent coupling are the least constrained by direct DM searches.
- W.r.t. spin-dependent coupling, underice/underwater detectors are ~100 times more sensitive than direct search experiments (Sun is mostly hydrogen)

SUSY and Proton Decay

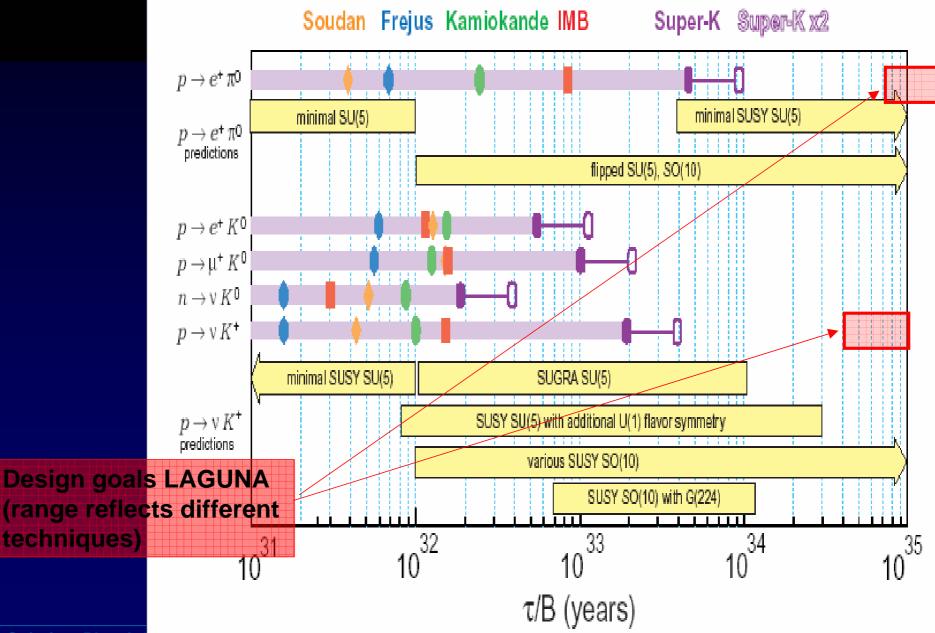
http://www.aspera-eu.org

Soudan Frejus Kamiokande IMB Super-K x2 Super-K $p \to e^{\star} \, \pi^{\rm O}$ minimal SUSY SU(5) minimal SU(5) $p \rightarrow e^{+} \pi^{0}$ 111 predictions flipped SU(5), SO(10) $p \to e^{\rm +} \, K^{\rm 0}$ $p \rightarrow \mu^+ K^0$ 1 1 1 1 1 1 $n \mathop{\rightarrow} \mathsf{v} \, K^{\mathbf{0}}$ 1.1.1.1.1.1 $p \to \mathrm{v}\, K^{\mathrm{+}}$ minimal SUSY SU(5) SUGRA SU(5) $p \rightarrow v K^{+}$ SUSY SU(5) with additional U(1) flavor symmetry predictions various SUSY SO(10) SUSY SO(10) with G(224) 1 1 1 1 1 1 10³² 10³⁴ 10³⁵ 33 10³¹ 10 τ/B (years)

Proton Deca



Proton Deca



APPEC Astroparticle Physics at the Negaton Scale

- Proton decay: improve sensitivity by > factor 10 and test a new class of Supersymmetry models
- Galactic Supernova: 10⁴- 10⁵ events
 Incredibly detailed information on the early SN phase
- Diffuse flux from past SN: probe cosmological star formation rate
- Solar neutrinos: details of the Standard Solar Model determined with percent accuracy
- Atmospheric neutrinos: high statistics would improve knowledge neutrino mixing and provide unique information on the neutrino mass hierarchy
- Geo-neutrinos: improve understanding of the Earth interior
- Indirect WIMP search
- Neutrinos from accelerators over a long baseline (also with dedicated smaller detectors): neutrino properties

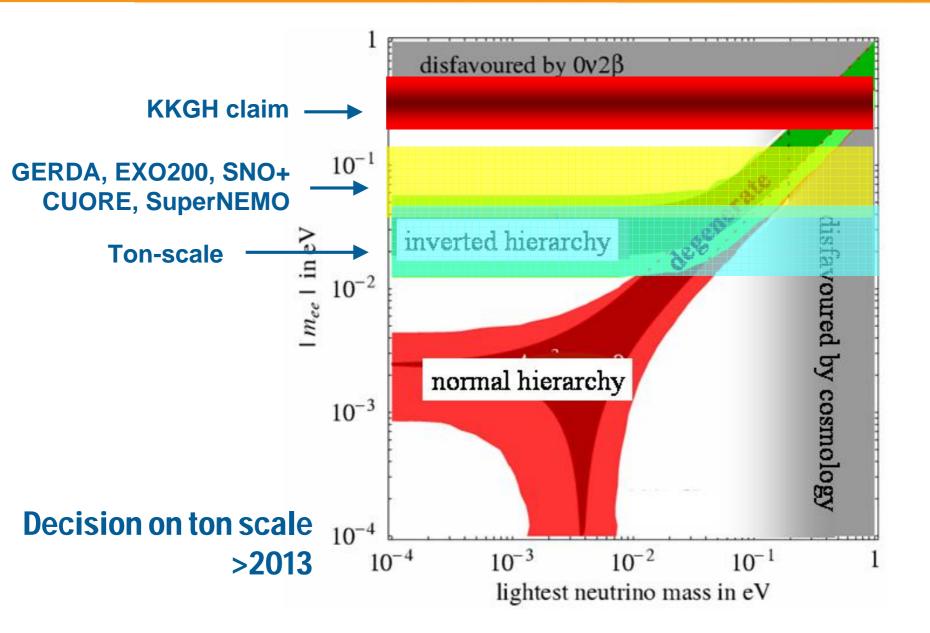
Properties of Neutrinos

Majorana nature, mass & mass hierarchy

.. may be related to GUT scale via see-saw mechanism and to matter-antimatter asymmetry via Majorana decay

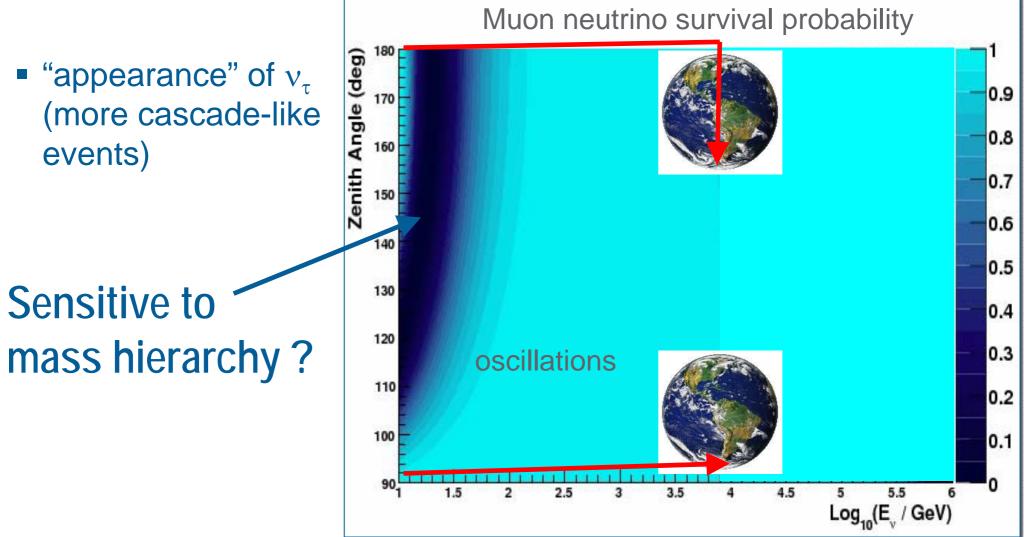
http://www.aspera-eu.org

Double Beta Decay



Oscillation physics with atmospheric neutrinos

 v_µ disappearance (less muon tracks)



Effects of Quantum Gravity

cosmic neutrinos, gamma rays, charged cosmic rays

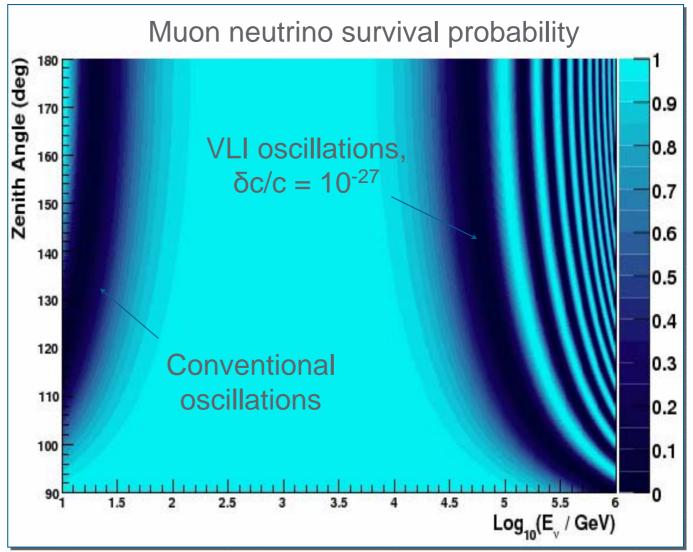
http://www.aspera-eu.org

Exotic Oscillations

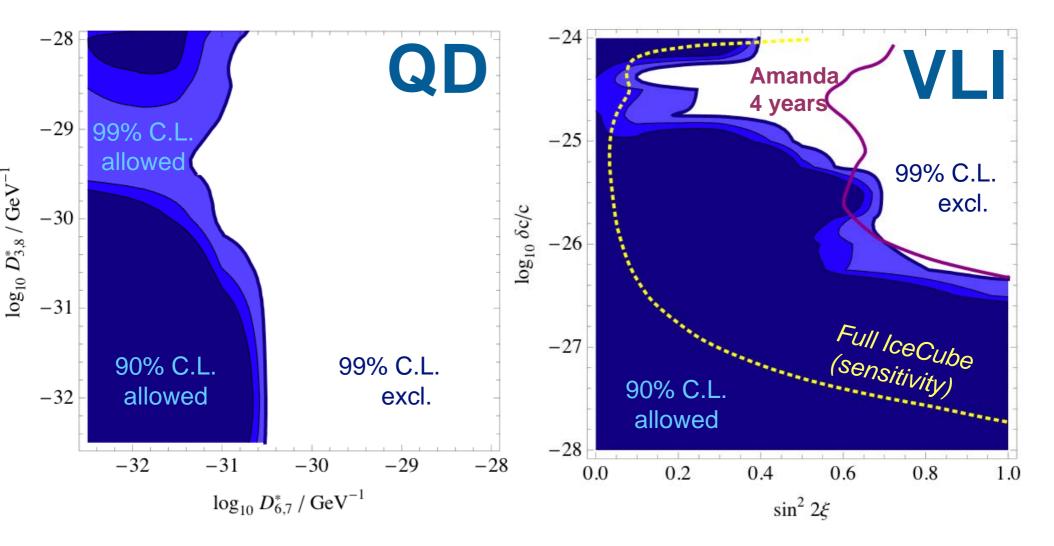
- Violation of Lorentz invariance
- Quantum
 decoherence
- (both appear in quantum gravity theories)

parameters of interest: VLI: $\delta c/c$, sin 2 ξ , phase η QD: D₃ and D₈, D₆ and D₇

Different to standard oscillations (~ 1/E), effects of QG oscillations go ~ E



Appec Astroparticle Physics for Europe



Limits from 7 years AWANDA

(2000-2006)

The ApPEC Roadmap and the Magnificent Seven

http://www.aspera-eu.org

The Magnificent Seven

Font size corresponds to realization time (not importance!)

Einstein Telescope E.T.

CTA

Ton-scale Double Beta Megaton (LAGUNA)

Auger-Nord KM3NeT

Ton-scale Dark Matter

Astrophysics

Particle Physics

Double Beta Dark Matter LAGUNA KM3NeT Auger CTA E.T.

The Magnificent Seven

Majorana nature (→ leptogenesis) v mass

> Cosmology SUSY, KK, ... Q.C indition proton decay v astronomy dark matter accelerator beam v oscillations (hierarchy?)

v astronomy v oscillations Q.G., TD, ... indirect DM and other exotic particles

Astrophys. Indirect DM Q.G, ...

Astrophysics Q.G., TD, ... GZK physics cross sections

Test Relativity Extreme objects

Astrophysics

Particle Physics

Double Beta Dark Matter LAGUNA KM3NeT Auger CTA E.T.

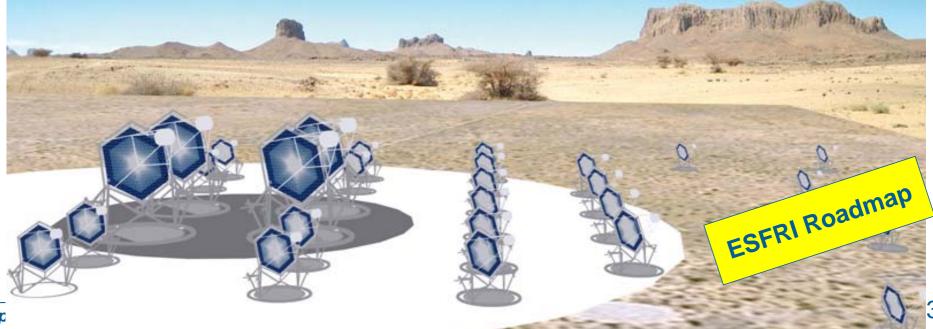
Astroparticle Theory & CERN

• A European Centre for Astroparticle Theory could be established at CERN. Given the synergy between LHC physics and astroparticle physics, CERN would be a natural host, particularly in view of several astroparticle experiments being CERN recognized experiments.



Cherenkov Telescope Array

- Increased sensitivity
- Extended energy range
- Improved angular resolution
- 50 to 100 large, medium and small telescopes
- Observatory with flexible and robotic operation
 Arrays in North and South for full sky coverage



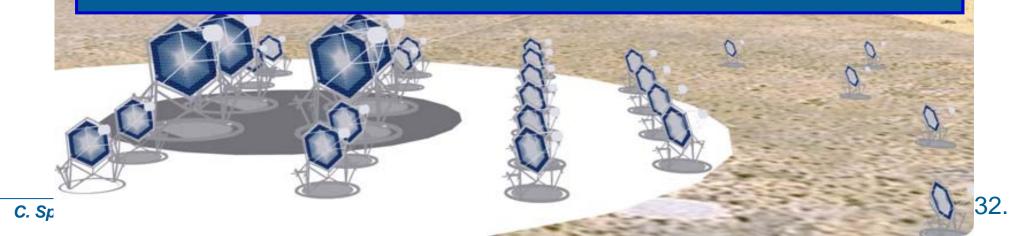
Cherenkov Telescope Array

Increased sensitivityExtended energy range

50 to 100 large, medium

S

The priority project of VHE gamma astrophysics is CTA. We recommend design and prototyping of CTA and selection of site(s), and proceeding decidedly towards start of deployment in 2012.

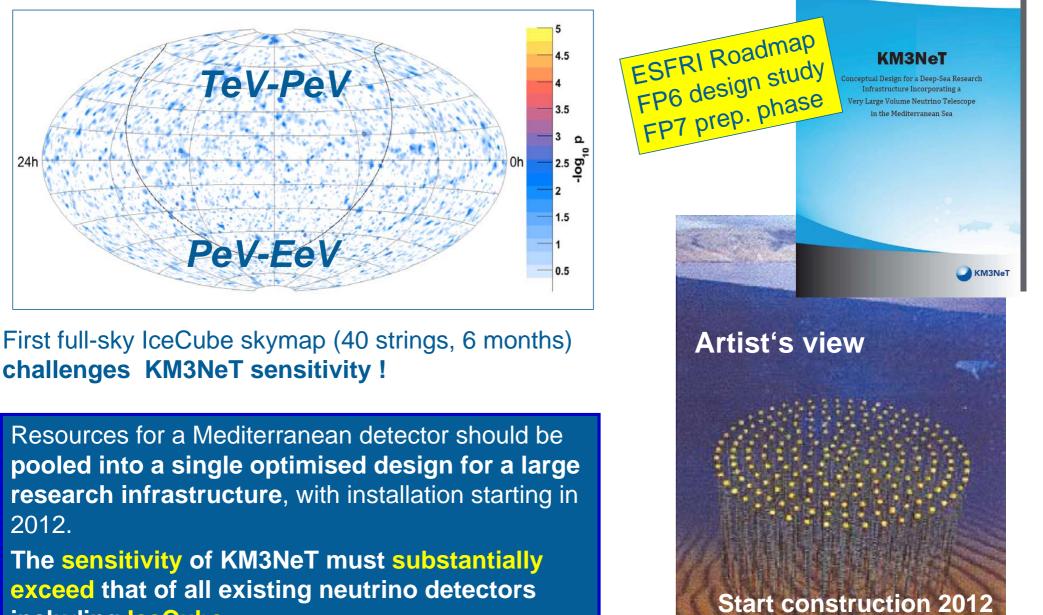


Pierre Auger Observatory

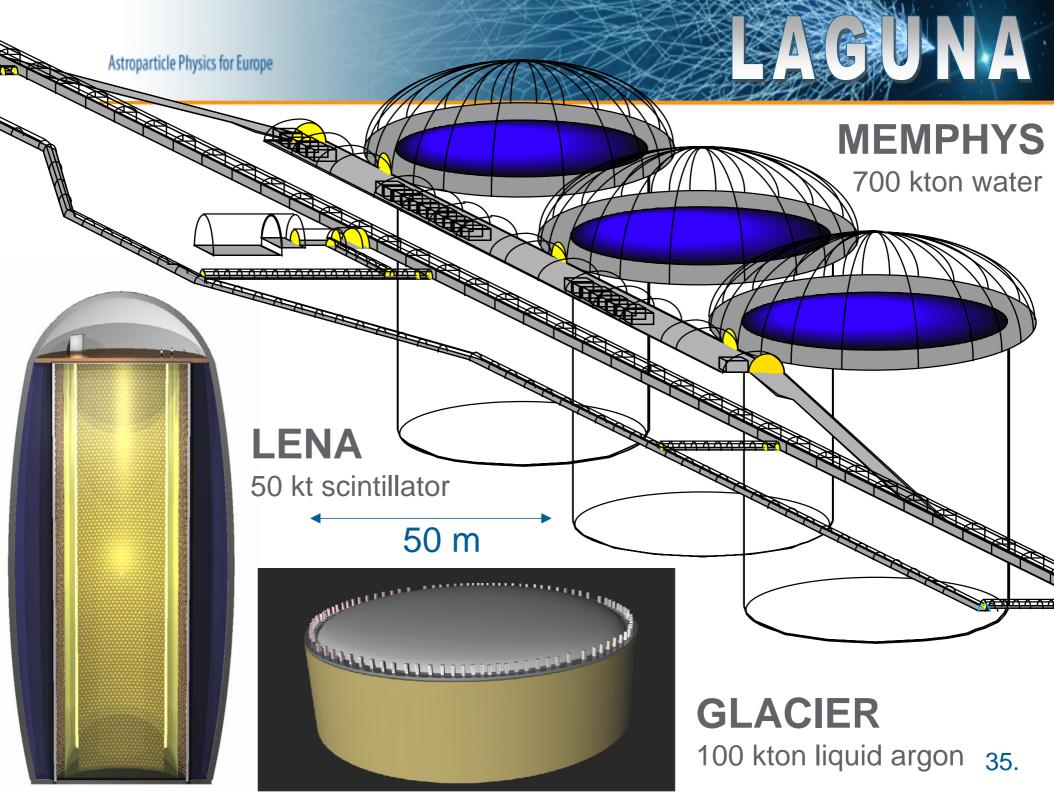


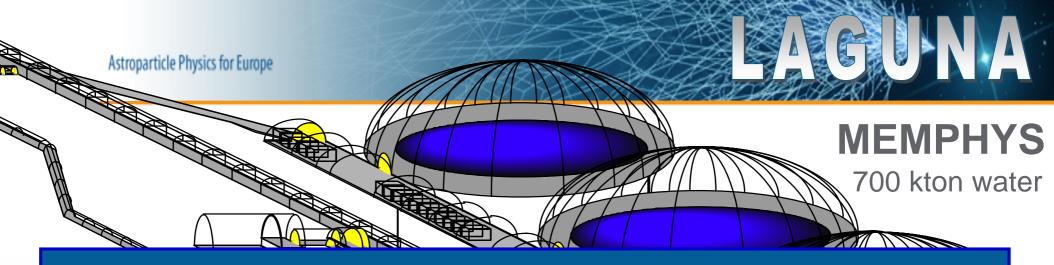
Auger-North: high statistics astronomy with reasonably fast data collection calls for a substantially larger array than Auger South, full sky coverage calls for a Northern site. A larger array would also allow a more detailed inspection of the high energy cut-off of the particle spectrum, which recently has been firmly established by Auger-South.

Appec Astroparticle Physics for Europe



including lceCube.

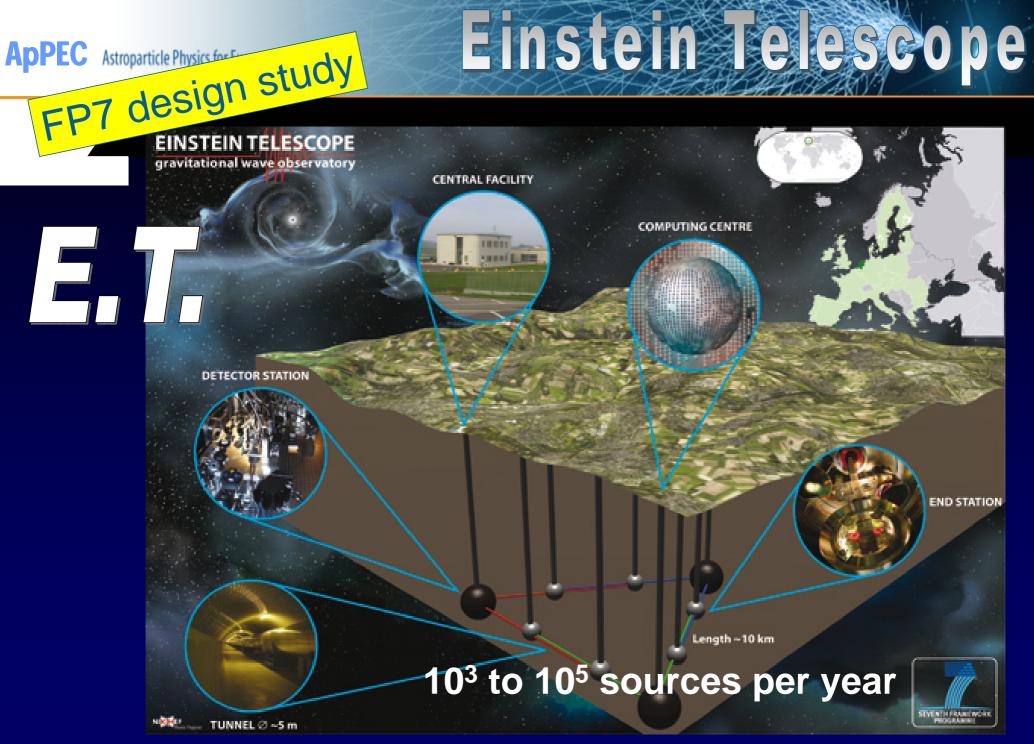




We recommend supporting the work towards a large infrastructure for proton decay and low energy neutrino astronomy, possibly also accelerator neutrinos in long baseline experiments, in a worldwide context (coherent approach with efforts in USA and Japan). **Results of the LAGUNA FP7 design study are expected around 2010 and should be followed by work towards a technical design report. Depending on technology, site and worldwide cost sharing, construction could start between 2012 and 2015.**



GLACIER 100 kton liquid argon 36.





E.T. is the long-term future project of ground-based gravitational wave astronomy in Europe.

A decision on funding for the construction of E.T. will earliest after first detections with enhanced LIGO/Virgo but is most likely after collecting about a year of data with advanced LIGO/Virgo in approximately 2014/15.

Targeted start of E.T. construction ~ 2017.

TUNNEL @ ~S m

Conclusions

- Astroparticle and LHC physics are closely linked.
- Several questions can be convincingly answered only if astroparticle searches are combined with LHC results.

Astroparticle Physics:

- Moving into regions with fair/high discovery potential.
- Accelerated increase of sensitivity in nearly all fields.
- ApPEC has defined milestones for decisions on best technologies and has initiated a process of convergence.
- Need a substantial increase in funding to make the possibilities a reality
- ApP could enhance the scientific portfolio of CERN