

BMR [3] and BFMRB [2] use Euclidean space time; this introduces a factor $(-1)^{p'}$ compared to our definition of Feynman diagrams, where p' is the weighted number of propagators and numerators in the corresponding diagram.

The conventions are such that the tadpoles become:

$$\begin{aligned} T^{BMR} &= - \left(\frac{\mu_0^2}{a} \right)^\epsilon \frac{a}{\epsilon(1-\epsilon)}, \\ T^{BFMRB} &= - \frac{m^2}{4\epsilon(1-\epsilon)}. \end{aligned}$$

For an l -loop master integral with p propagators the relations between the integrals become (assuming here $\mu_0^2 = m^2 = a = 1$):

$$\begin{aligned} M_{l,p'}^{BMR} &= \frac{(-1)^{p'}}{[e^{\gamma\epsilon}\Gamma(1+\epsilon)]^l} M_{l,p'}^{CGR}, \\ M_{l,p'}^{BFMRB} &= \frac{(-1)^{p'}}{[4e^{\gamma\epsilon}\Gamma(1+\epsilon)]^l} M_{l,p'}^{CGR}. \end{aligned}$$

One has also to take into account the various definitions of the series expansions:

$$\begin{aligned} M_{l,p'}^{BMR} &= \left(\frac{\mu_0^2}{a} \right)^{2\epsilon} \sum_k M_k^{BMR}(l, p', \dots) \epsilon^k, \\ M_{l,p'}^{BFMRB} &= \sum_k M_k^{BFMRB}(l, p', \dots) (-2)^k \epsilon^k. \end{aligned}$$