

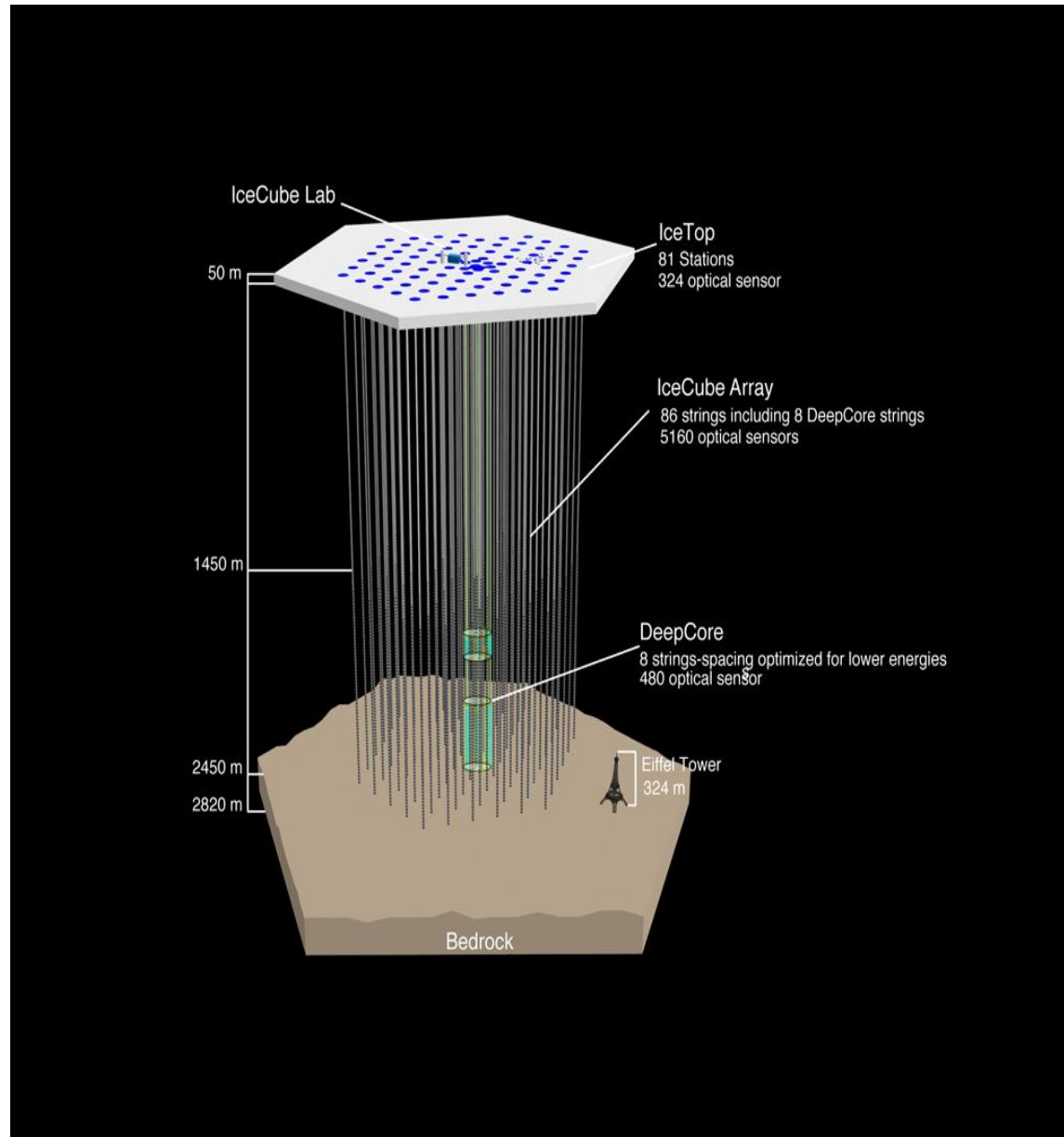
Identifying Weak Cosmic Ray Signatures in Cascade-Like Events in IceCube.

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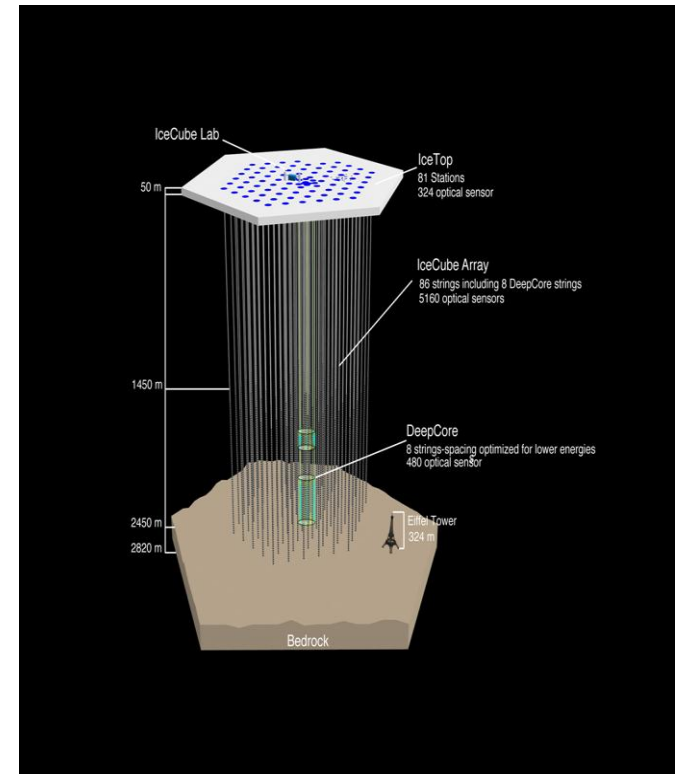
THE ICECUBE DETECTOR

- Neutrino telescope ,
1450 to 2450m in the
Antarctic Ice
- 86 strings, each with 60
DOMs. 125m horizontal
spacing, and 17m vertical
- High energy astrophysical
neutrinos $>100\text{Gev}$.
- Low cross section of
neutrino interactions; large
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- Deep Core section
- DOMs detect Cherenkov
radiation



THE ICECUBE DETECTOR

- Neutrino telescope ,buried between 1450 to 2450 meters deep in the Antarctic Ice
- 86 strings, each with 60 DOMs. 125m horizontal spacing, and 17m vertical
- Sensitive to high energy astrophysical neutrinos $>100\text{GeV}$.
- Low cross-section of neutrino interaction hence large detector volume
- DOMs detect Cherenkov radiation emitted by secondary charged particles produced by neutrino interaction with the ice
- Deep Core section , sensitive to low energy neutrinos, where the DOMs are more closely spaced.



- Each DOM has a photomultiplier tube.
- The voltage drop over the resistor is the recorded signal.

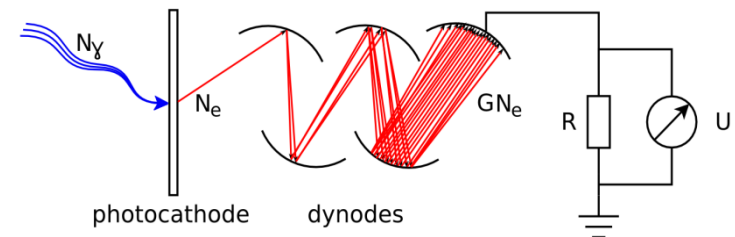


Photo-multiplier

EVENT SIGNATURES

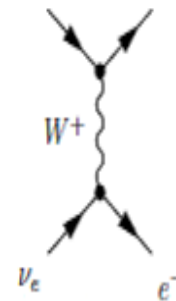
CASCADE-LIKE EVENTS

- Electron neutrino charged current; electromagnetic cascade
- Neutral Current produces hadronic cascade. Sensitive to all 3 neutrino flavors.
- Good energy resolution.
- Launch time is color-coded from red to blue.

CC Example:

$Nuclei \Rightarrow d$

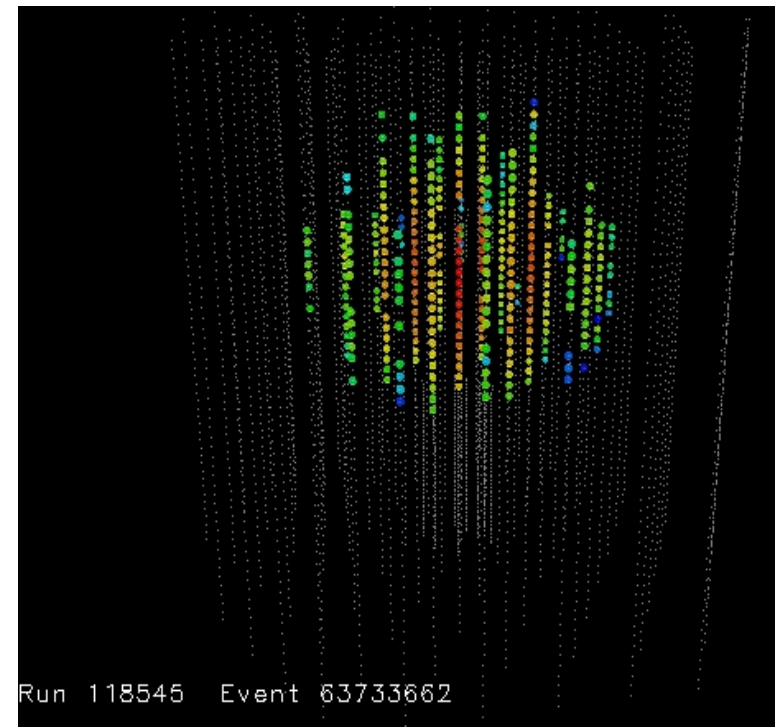
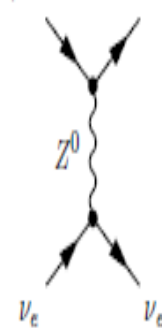
$u \Rightarrow HadronicCas.$



NC Example:

$Nuclei \Rightarrow d$

$d \Rightarrow HadronicCas. \nu_{Cas.}$

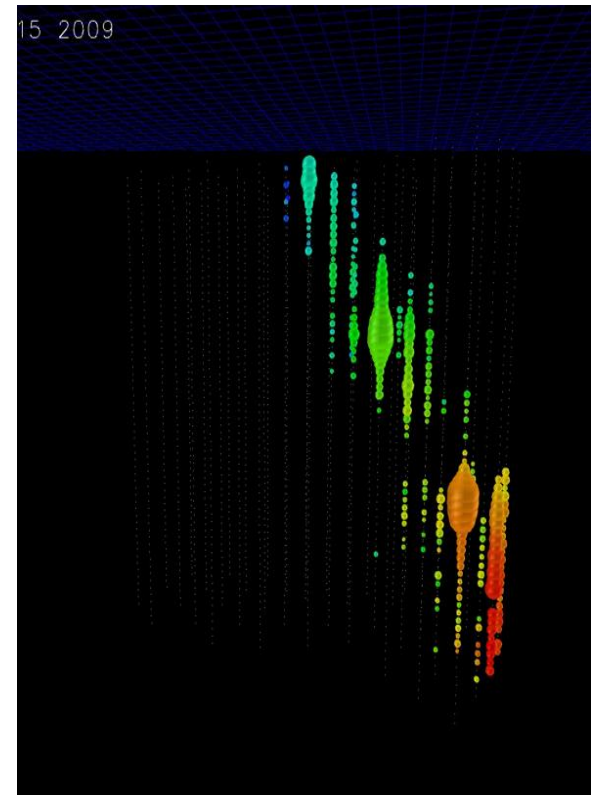


Run 118545 Event 63733662

EVENT SIGNATURES

TRACK LIKE EVENTS

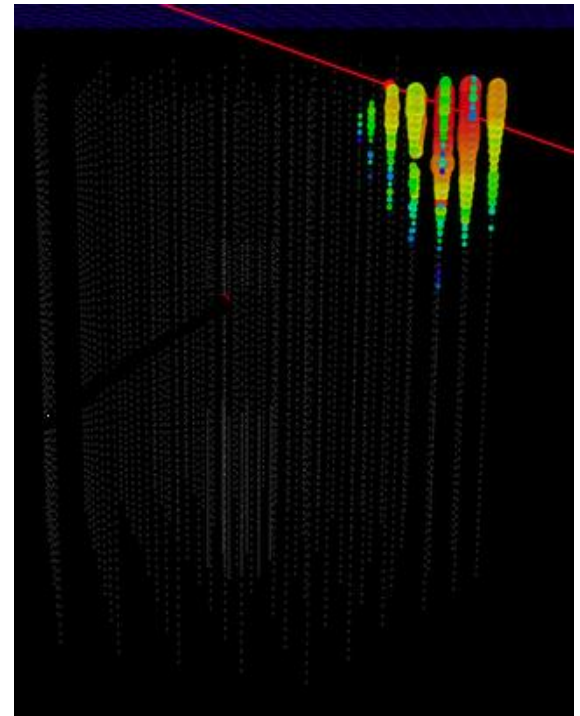
- charged current ν_μ interaction and cosmic ray μ
- Muon emits mainly Cherenkov without significant energy loss.
- good direction resolution; point back to sources.
- Primary cosmic rays produce muons which are the dominant background in the cascade channel.



Size denotes collected charge

WEAK COSMIC RAY SIGNATURES

- cosmic μ may create a signature almost identical to a cascade event
 - ❖ Catastrophic energy loss along μ track
 - ❖ Corner clipper
- AIM: Evaluate different cut strategies to remove this class of background events.
- By studying differences in waveforms, charge and arrival times.



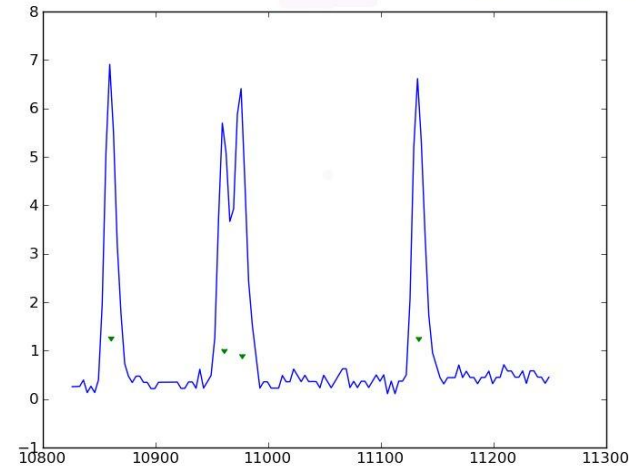
Example of a Corner Clipper.

ANALYSIS

Studying early waveforms and pulses

- **Waveform:** Time resolved recording of the voltage drop over the resistor.
- **pulses:** response of Photomultiplier Tube to a single photo electron. Reconstructed from the waveform.
- For cascade event, first peak is expected to be highest

Waveform and Pulses for string 58, 10 in event 92032

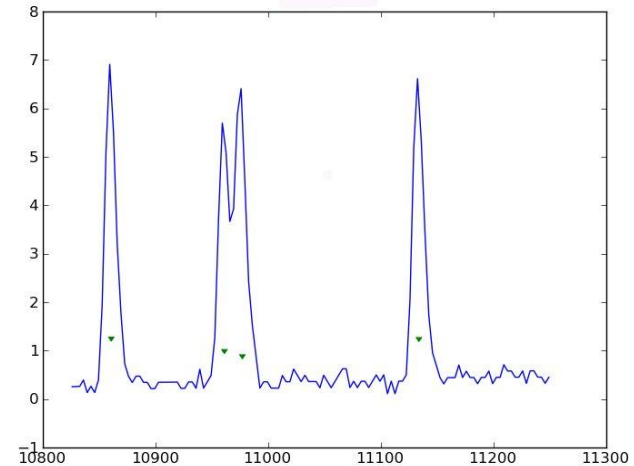


ANALYSIS

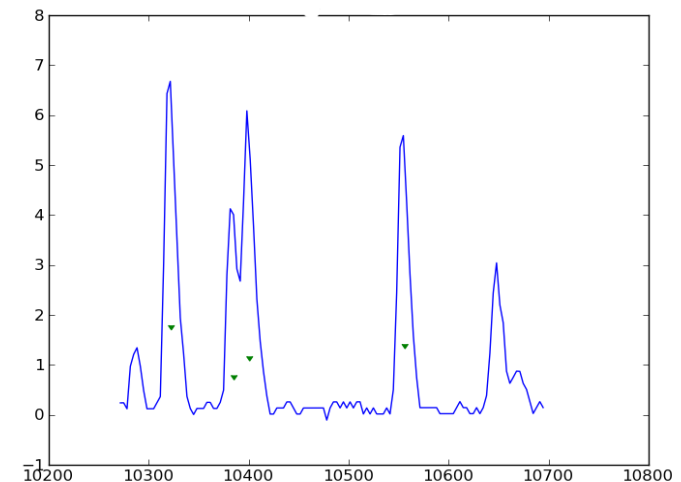
Studying early waveforms and pulses

- **Waveform:** The time resolved recording of the voltage drop over the resistor.
- **pulses:** response of the Photomultiplier Tube to a single photo electron. This is reconstructed from the waveform.
- For cascade event, the first peak is expected to be the highest
 - ☐ Some waveforms did not have pulses in the small features!
 - ☐ This needs to be investigated further.

Waveform and Pulses for string 58, 10 in event 92032



Waveform and Pulses for string 58, 12 in event 92032

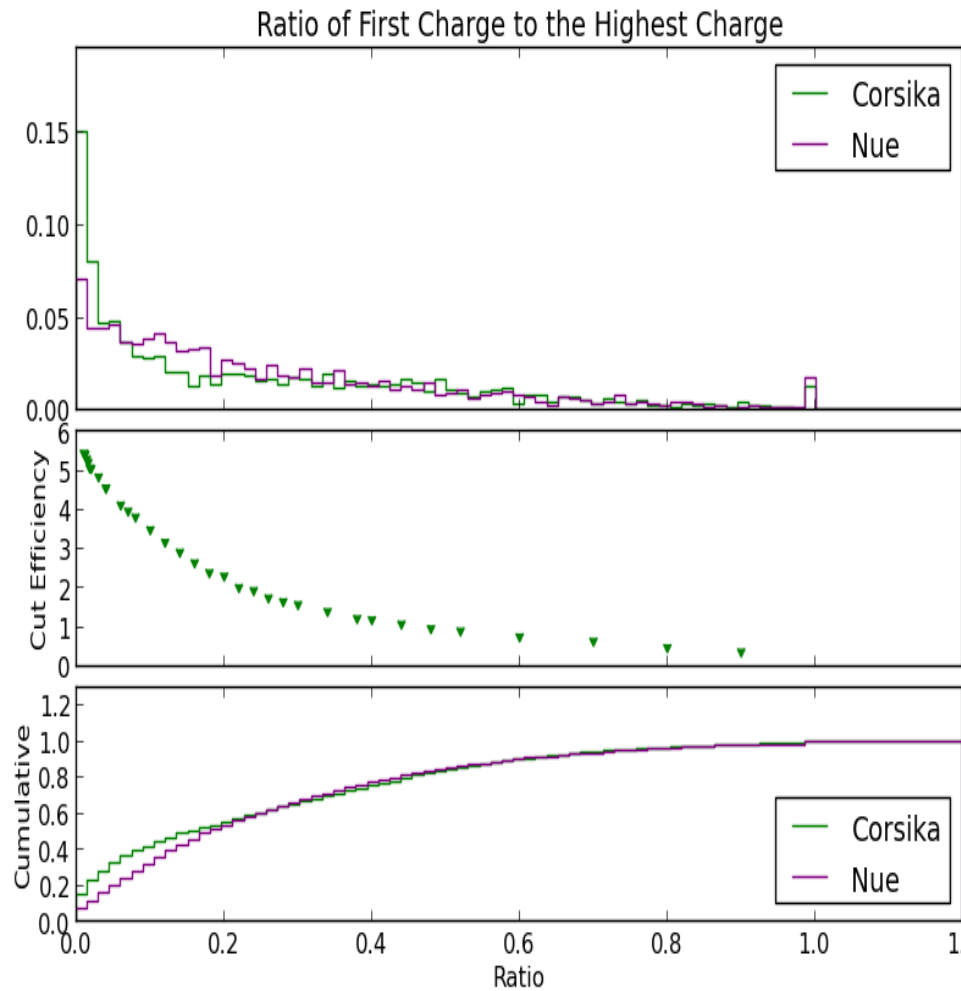


- Data sets used:
 - Pre-filtered Simulated cosmic ray muons (μ that mimic cascade events)
 - Simulated electron neutrinos
- An efficient variable makes a clear distinction between background and signal
- Goal is to keep as much signal while removing as much background as possible.

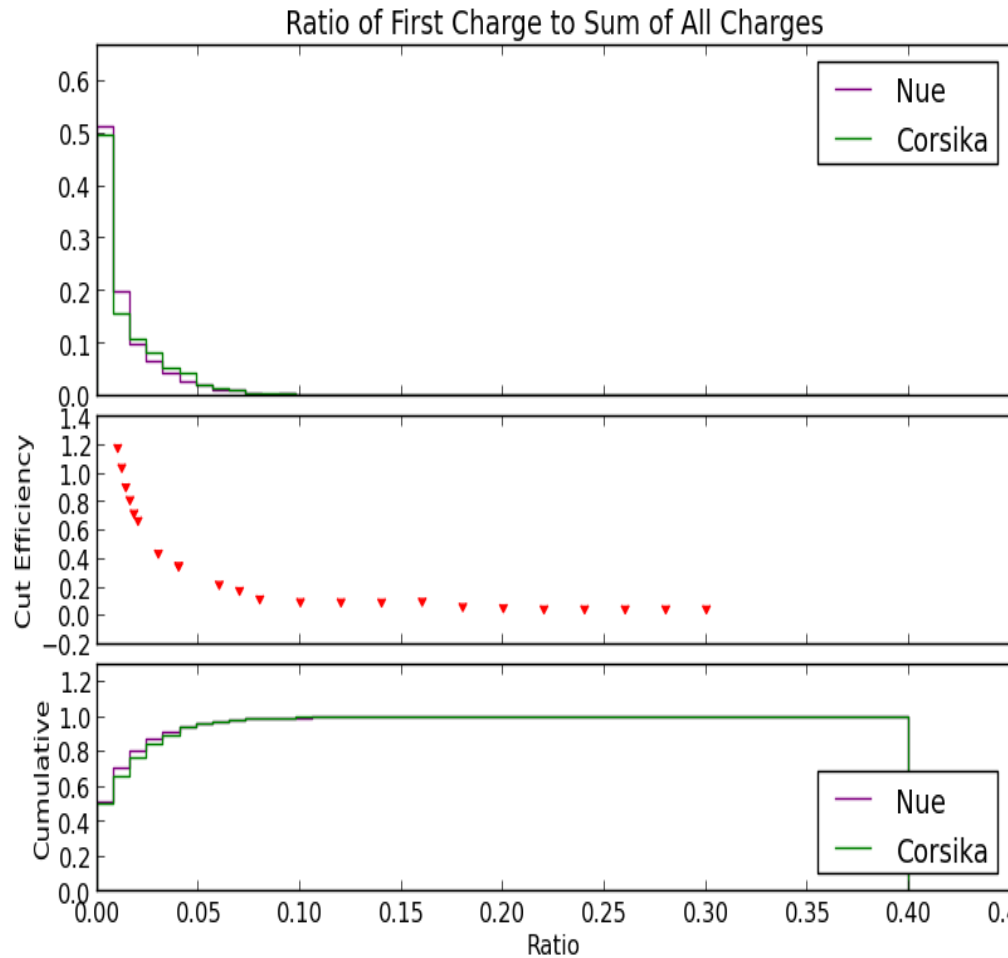
$$\text{Cut Efficiency} = \frac{\text{Number of remaining signal events}}{\sqrt{\text{Number of remaining background events}}}$$

- Cumulative distribution shows signal efficiency

A) Ratio of the First Charge to the Highest Charge.



B) Ratio of the First Charge to the Sum of all Charges

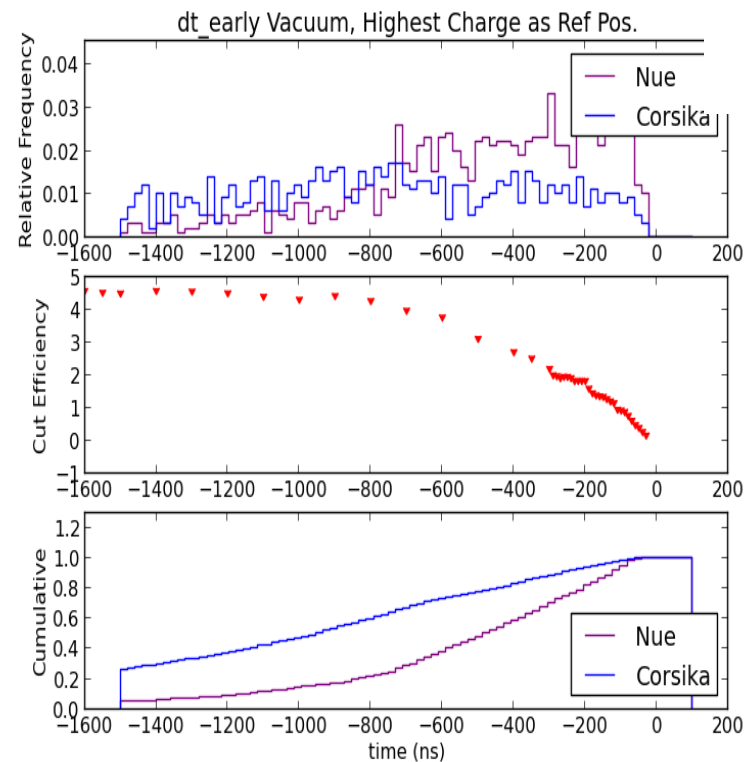
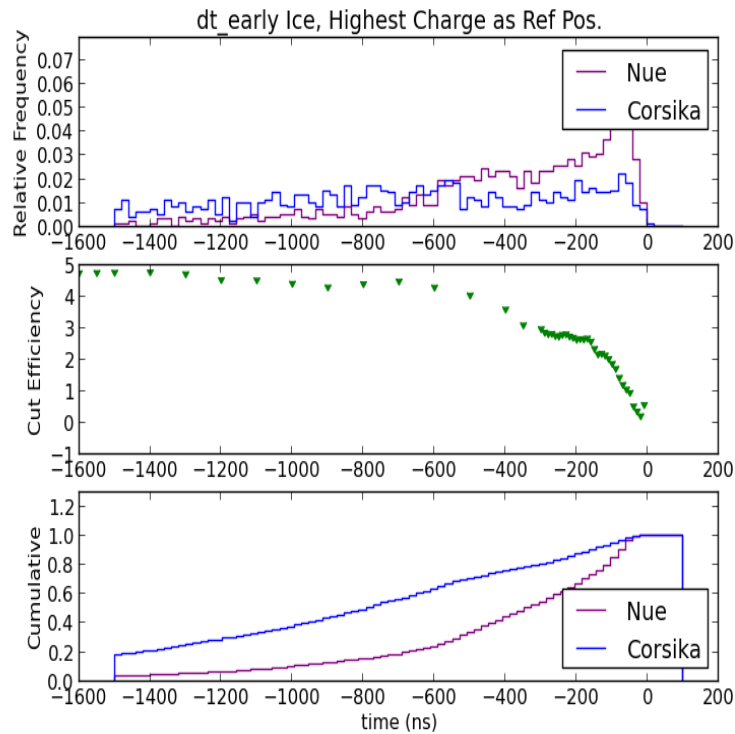
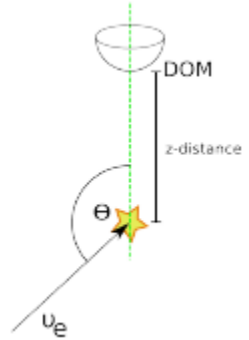


- Low signal efficiency

C) Delay time window, dt_nearly

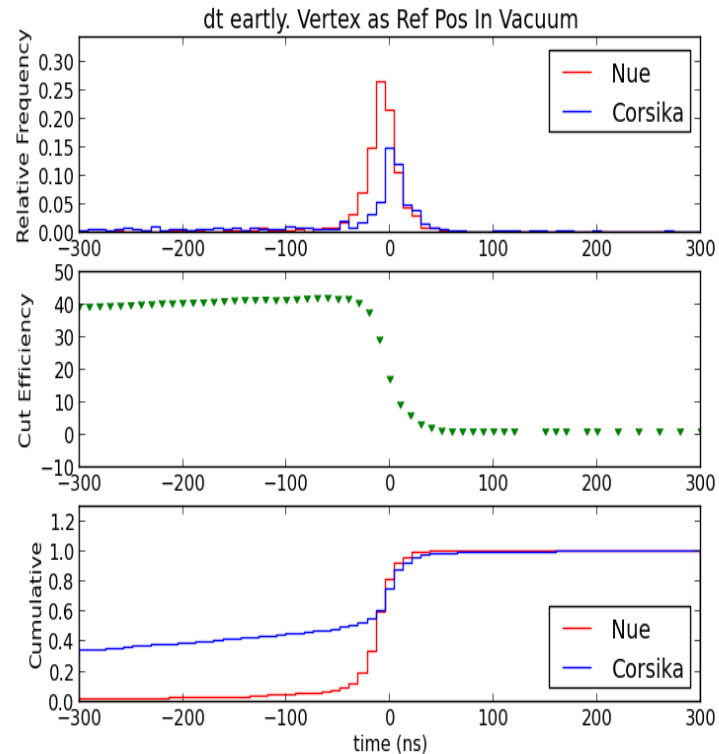
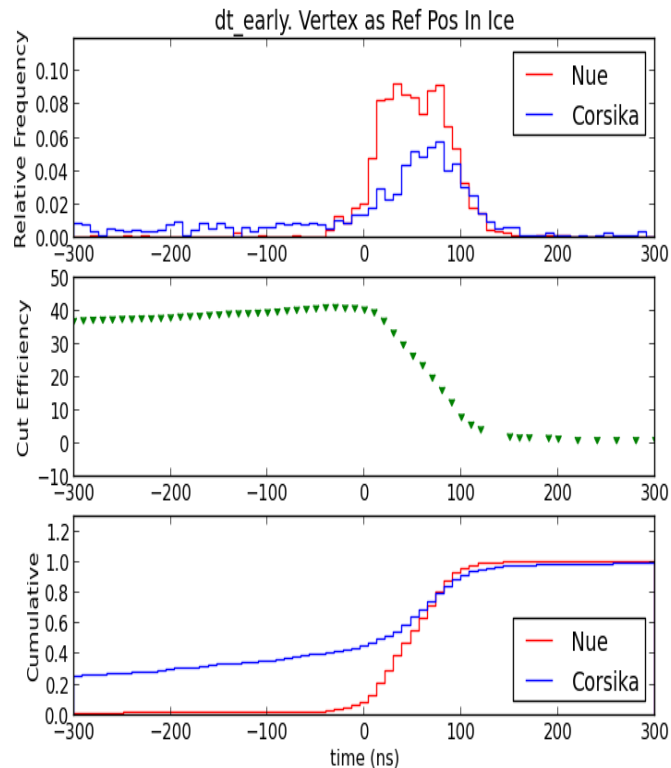
dt_nearly = Expected arrival time - Actual arrival time, with respect to a reference position.

i) DOM with the highest charge as reference position



) Delay time window, dt_nearly

ii) Vertex position as reference position.



SUMMARY

- The problem of the reconstruction of pulses of small features in the waveforms needs further investigation
- dt_nearly significantly better in ice with vertex position as reference position
- The assumption that the earliest pulse is the highest may not be ideal.

Thank you!

This has been an invaluable learning
experience.