

Elementary Particles

and their

Interactions

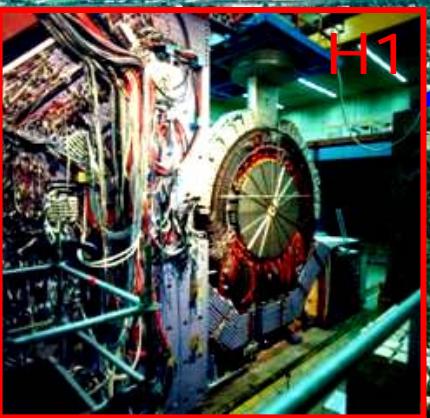
Thomas Naumann



DEUTSCHES ELEKTRONEN-SYNCHROTRON

Thomas.Naumann@desy.de

http://www-zeuthen.desy.de/~naumann/lectures/summer_lecture.pdf



H1

Protons 920 GeV

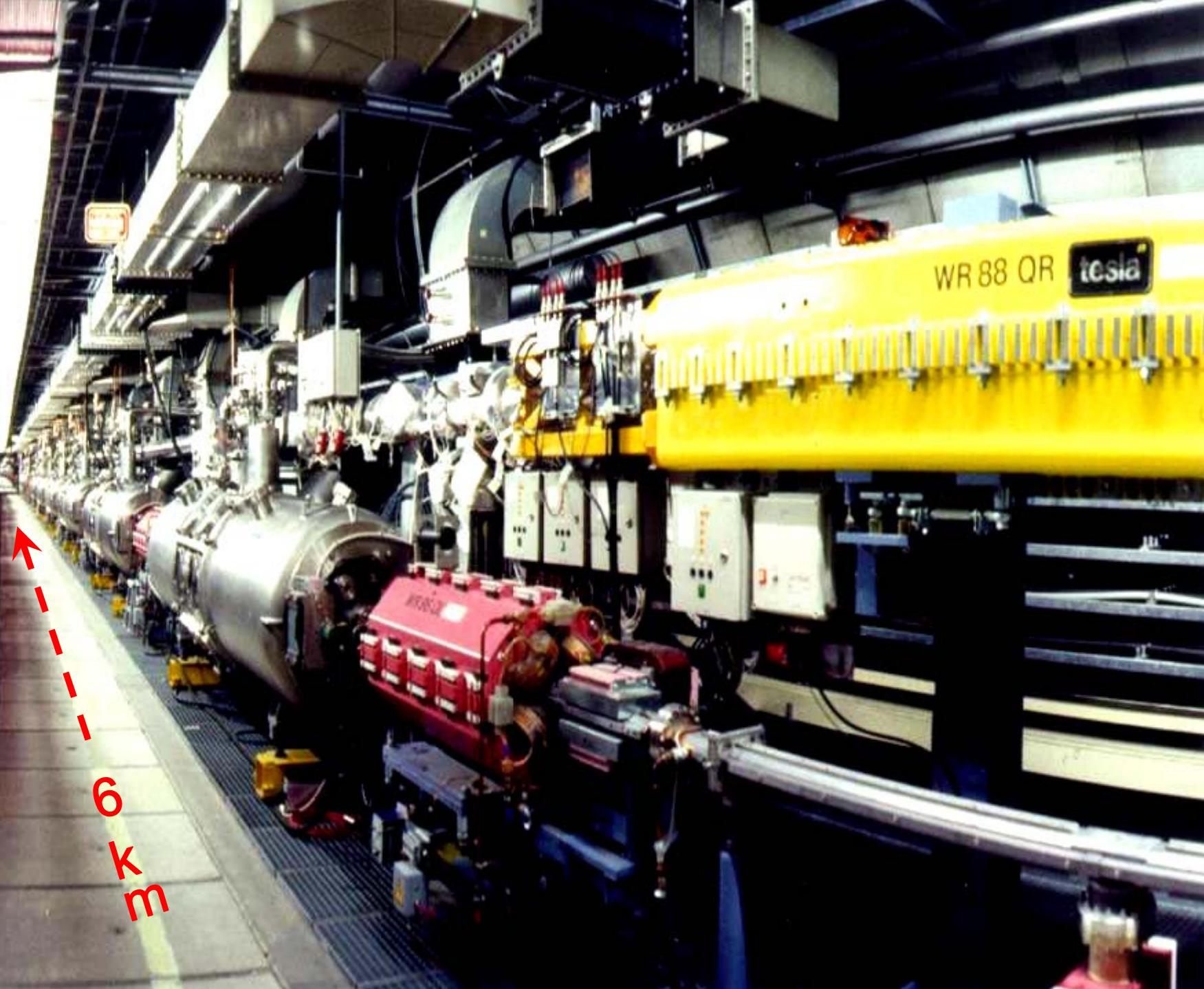


ZEUS



Electrons 27.6 GeV

PETRA



WR 88 QR

tesla

6
km

1. Introduction

1. Leptons
2. Hadrons

2. The Quark Model

1. Isospin
2. Multiplets and SU(N)
3. Spectroscopy of Light Quarks
4. Spectroscopy of Heavy Quarks
5. Measurement of Quark Charges

3. Symmetries and Conservation Laws

1. Charge Conjugation C
2. Parity P
3. *CP Violation*

4. Phenomenology of Weak Interactions

1. Neutral and Charged Currents
2. Neutrino Physics
3. Cross section and Fermi Constant
4. W- and Z-Bosons

5. Gauge Theory of Weak Interactions

(W. Lohmann)

1. Gauge Theories
2. Spontaneous Symmetry Breaking and Higgs Mechanism
3. Lagrangian of Weak Interactions
4. Experimental Tests of the Standard Model
5. Higgs Search

(Th. Naumann)

6. Gauge Theory of Strong Interactions

(J. Blümlein)

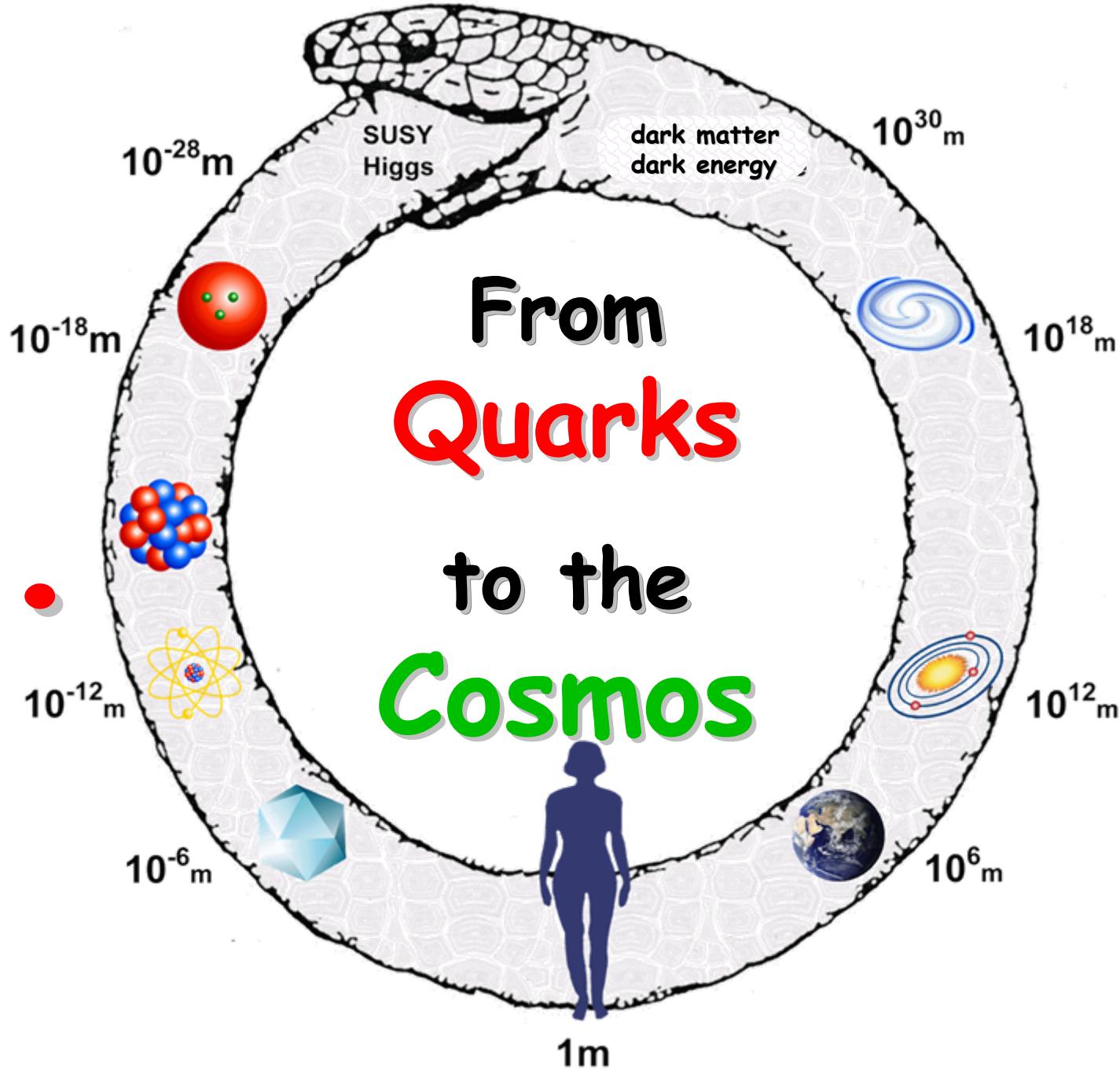
1. Quantum Chromodynamics
2. Hadron Structure

7. Grand Unification and Cosmology

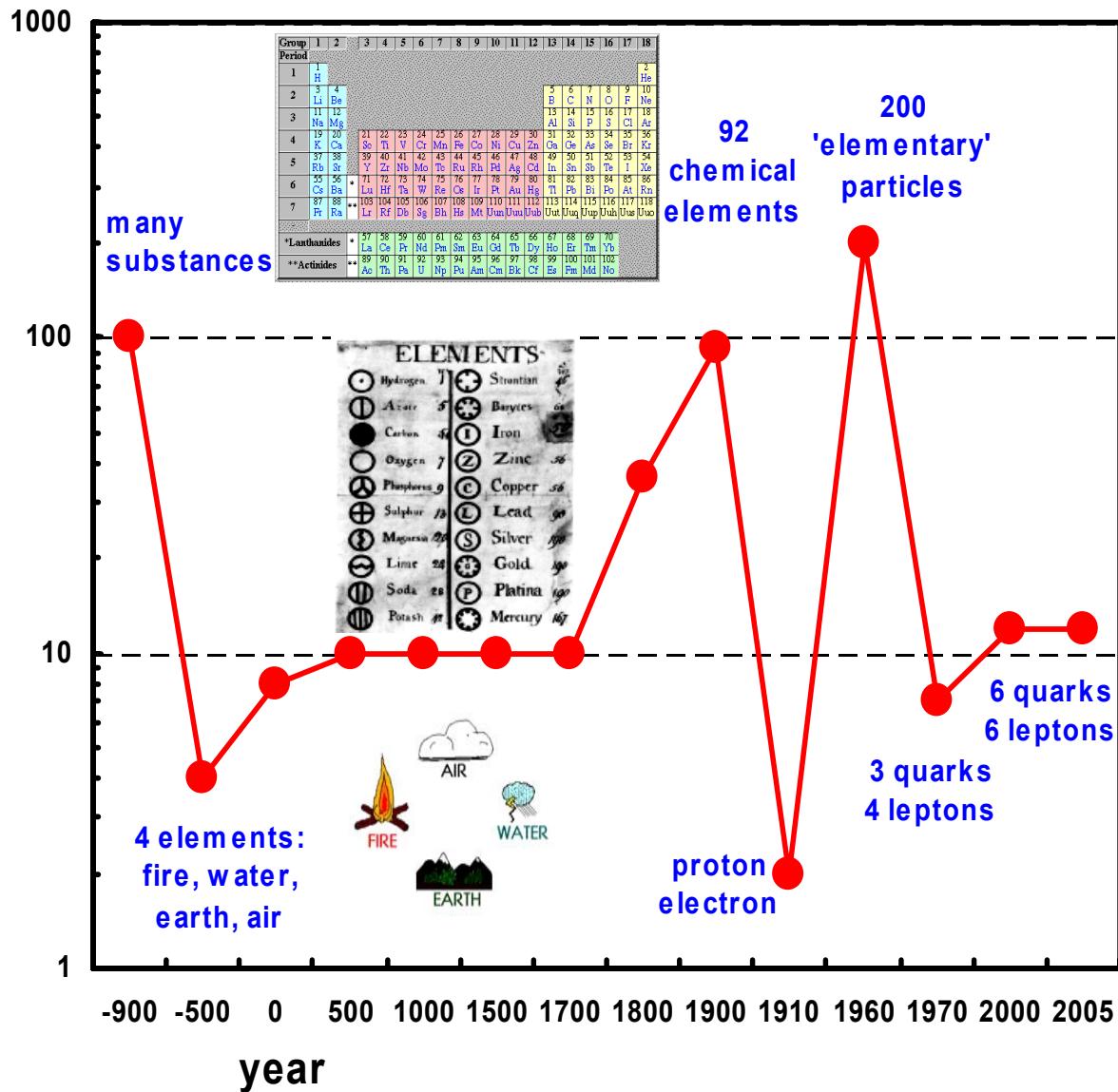
1. Unified Gauge Theories
2. Proton Decay and Baryon Asymmetry of the Universe
3. Supersymmetry and Superstrings
4. Cosmology and Particle Physics

(H. Nowak)

1.



The Building Blocks of Matter



Structure of Matter

subatomic units:
electron-volt

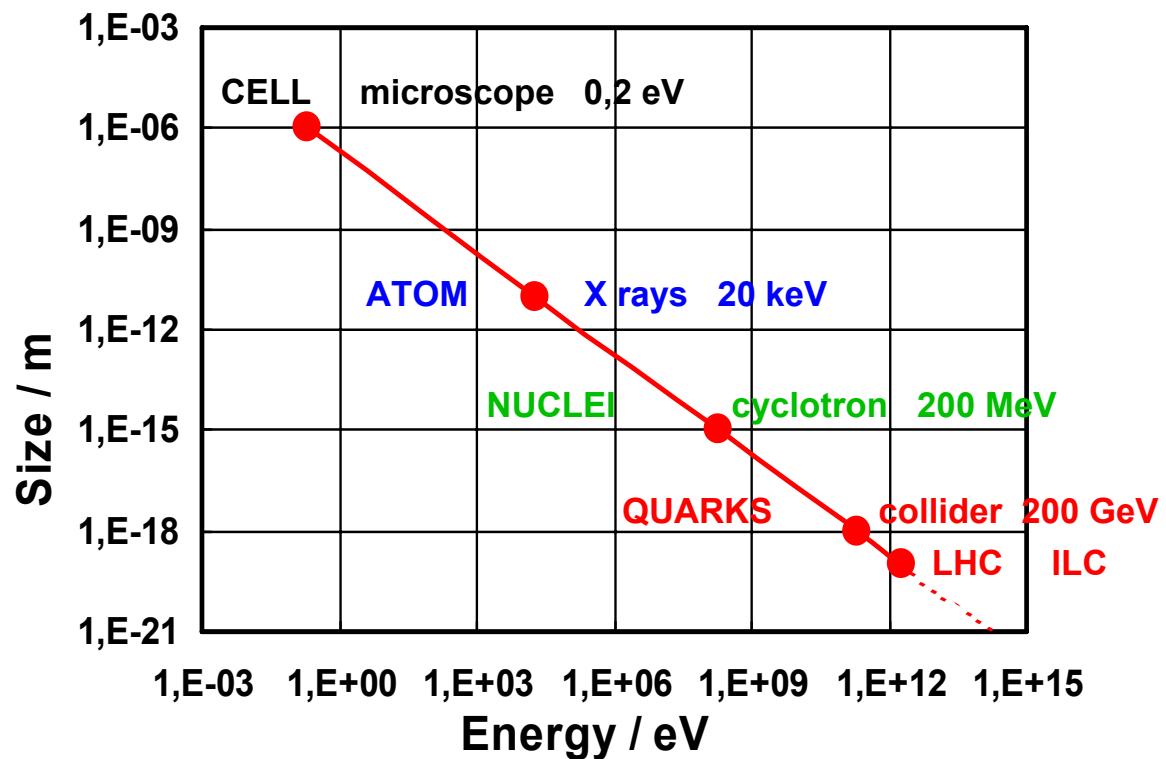
uncertainty
relation

$$1 \text{ eV} \approx k \cdot 10.000 \text{ K}$$

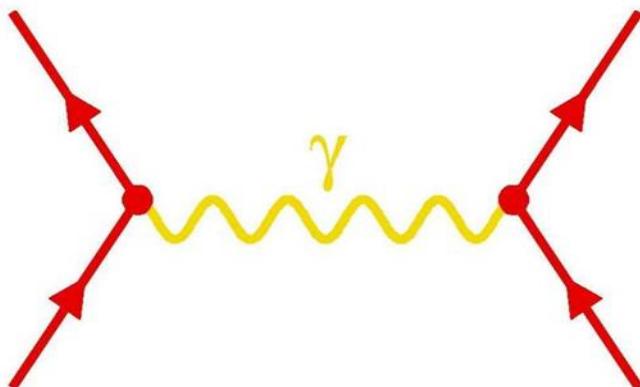
$$\Delta E \Delta x = \hbar c$$

$\approx 200 \text{ MeV fm}$

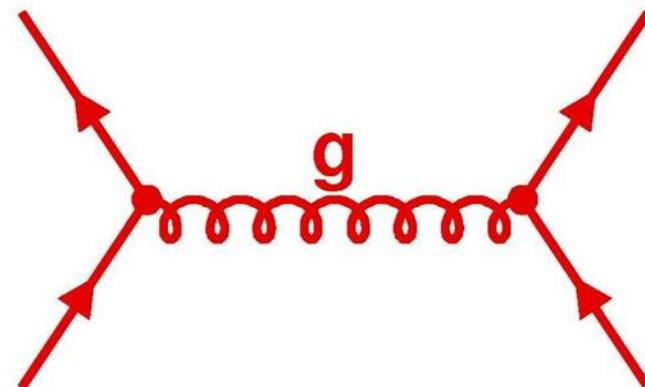
Energy	Size	Device	Object	Year
0.2 eV	10^{-6} m	microscope	cell	1600
20 keV	10^{-11} m	X rays	atom	1910
200 MeV	10^{-15} m	cyclotron	nuclei	1946
200 GeV	10^{-18} m	collider	quarks	1998



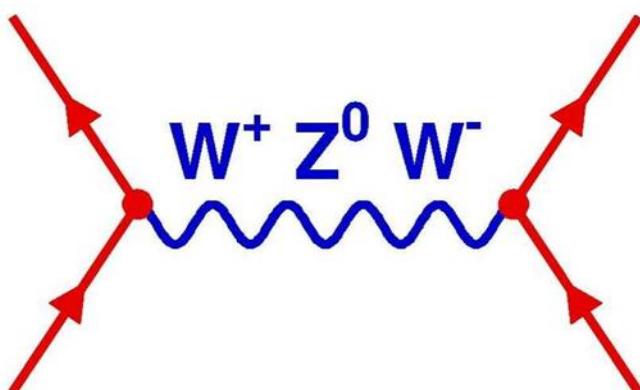
The Forces



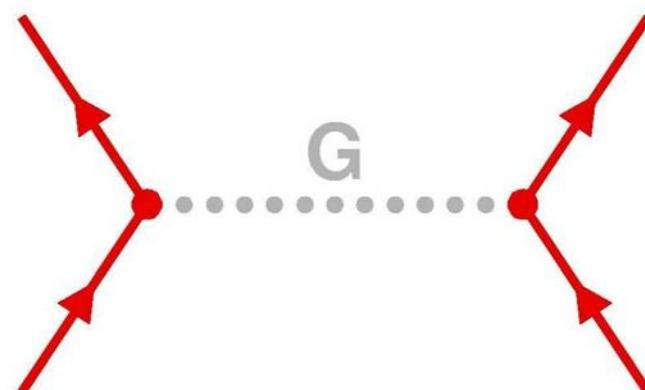
elektromagn. Kraft



starke Kraft



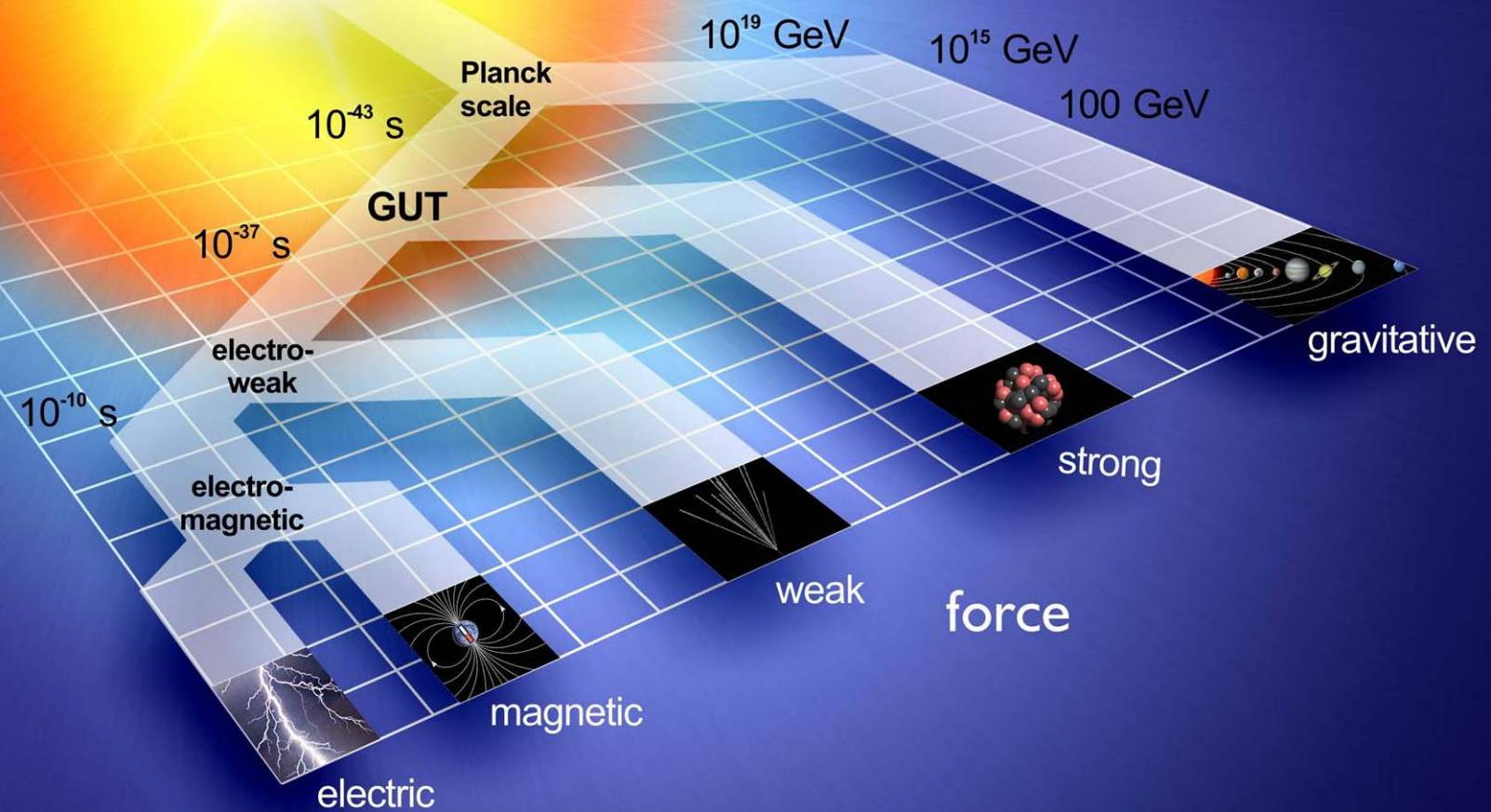
schwache Kraft



Gravitation

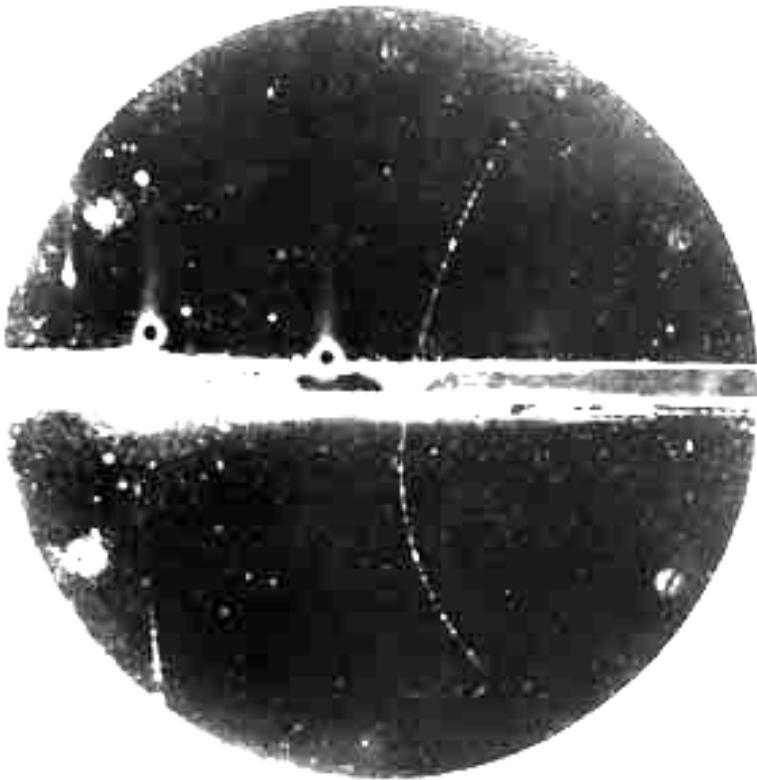
Unification of Forces

Big Bang



The Positron

- **curvature** R in magnetic field B: $R \sim p \sim 1/B$
- **ionization** I ~ velocity $\beta = v/c$: $I \sim 1/\beta \sim 1/(p/m)$



C. Anderson,

The
Positive
Electron

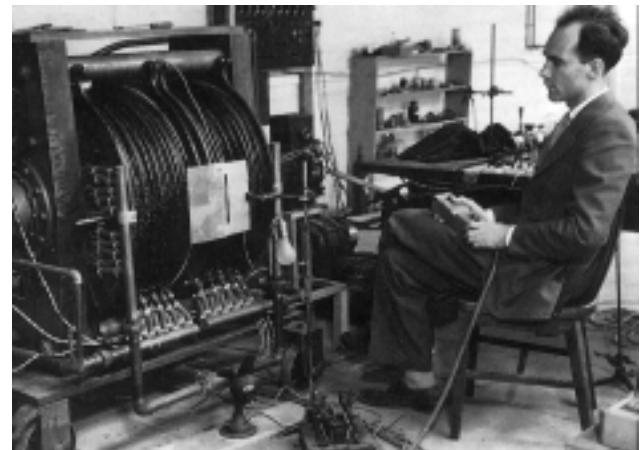
Phys. Rev.
43, 491
(1933)

Nobel prize 1936
Nobel prize 1927

C. Anderson
Ch. Wilson



Wilson's
cloud
chamber



for the discovery of the positron
for the cloud chamber

The Muon

1935: H. Yukawa:
carrier of nuclear force
mass ~ 200 MeV (between e + p)



Nobel prize 1949

J.C. Street, E.C. Stevenson,
New Evidence for the Existence of a Particle of Mass
Intermediate between the Proton and the Electron,
Phys. Rev. 52, 1003 (1937).

Too penetrating - NOT the Yukawa particle !

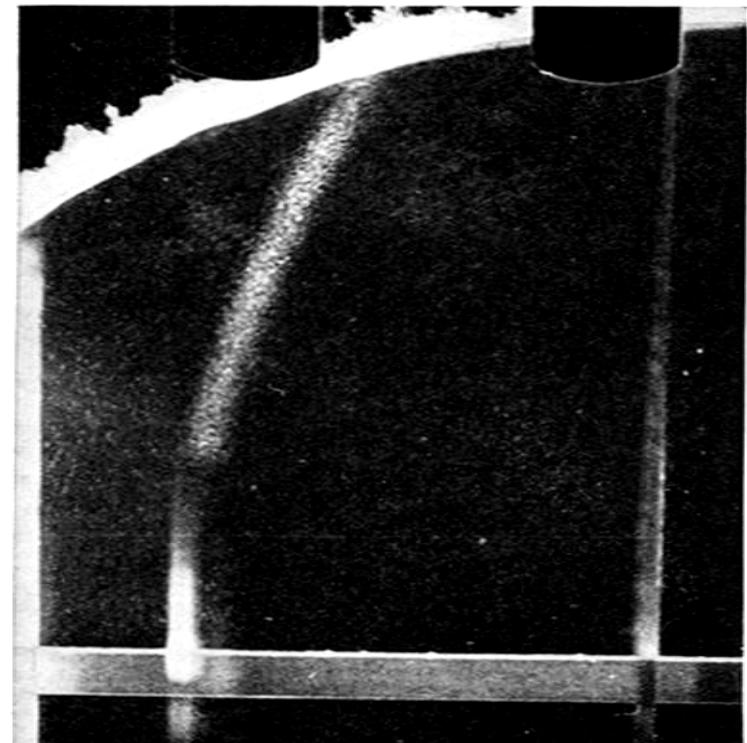


I.I. Rabi

Nobel 1944

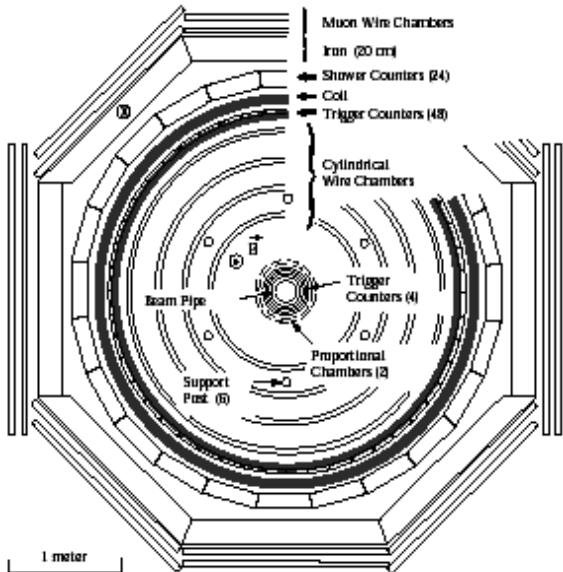
curvature R in magnetic field B:
 $R \sim p \sim 1/B$

ionization $I \sim$ velocity $\beta = v/c$:
 $I \sim 1/\beta \sim 1/(p/m)$



Who ordered that ?

τ lepton



MARK-I detector
at SPEAR $e^+ e^-$ ring
SLAC, USA.

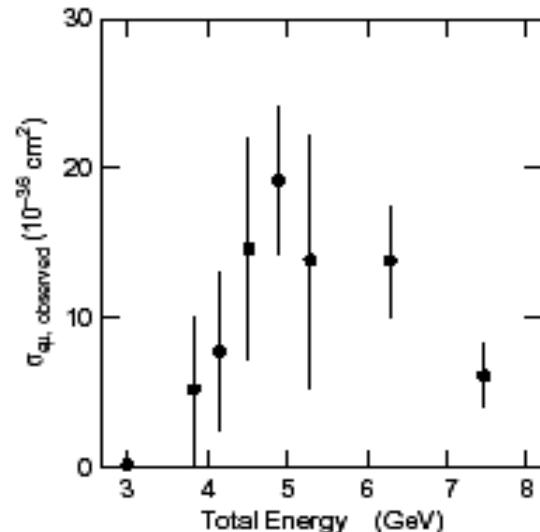
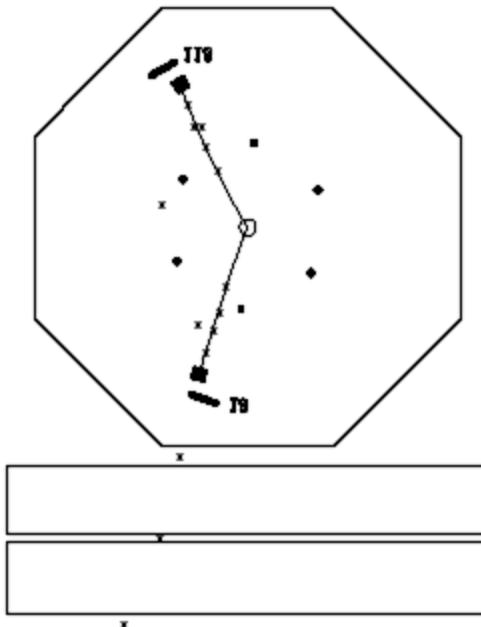


FIG. 2. The *observed* cross section for the signature $e\mu$ events.

M. Perl et al., 1975-77:

$$e^+ e^- \rightarrow L^+ L^-$$

above 3.5 GeV threshold: unlike lepton pair production:

$$\begin{aligned} L^+ &\rightarrow e^+ + \text{unseen } \nu's \text{ carrying off energy} + L_e \\ L^- &\rightarrow \mu^- + \text{unseen } \nu's \text{ carrying off energy} + L_\mu \end{aligned}$$



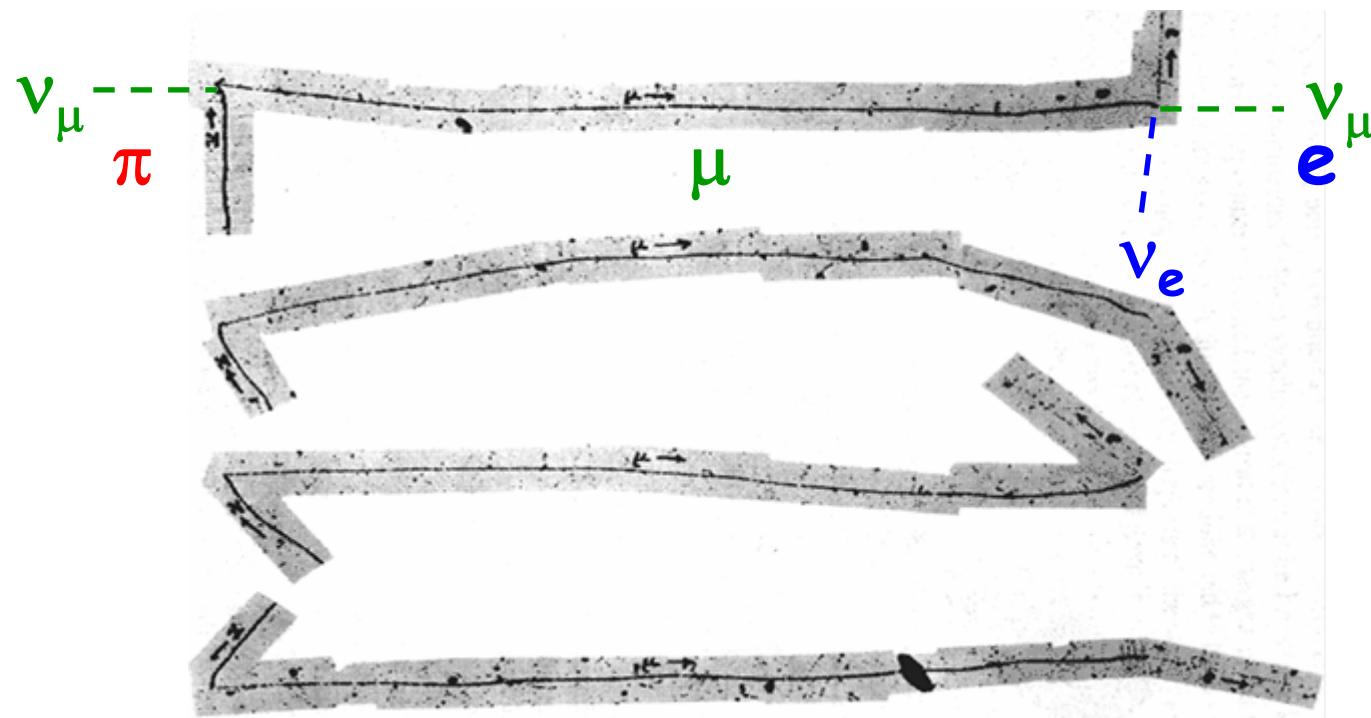
Nobel prize



1995

The Pion

- stopping track + typical decay cascade
- decay product fixed range in nuclear emulsion = always same E
- two body decay to muon + neutrino



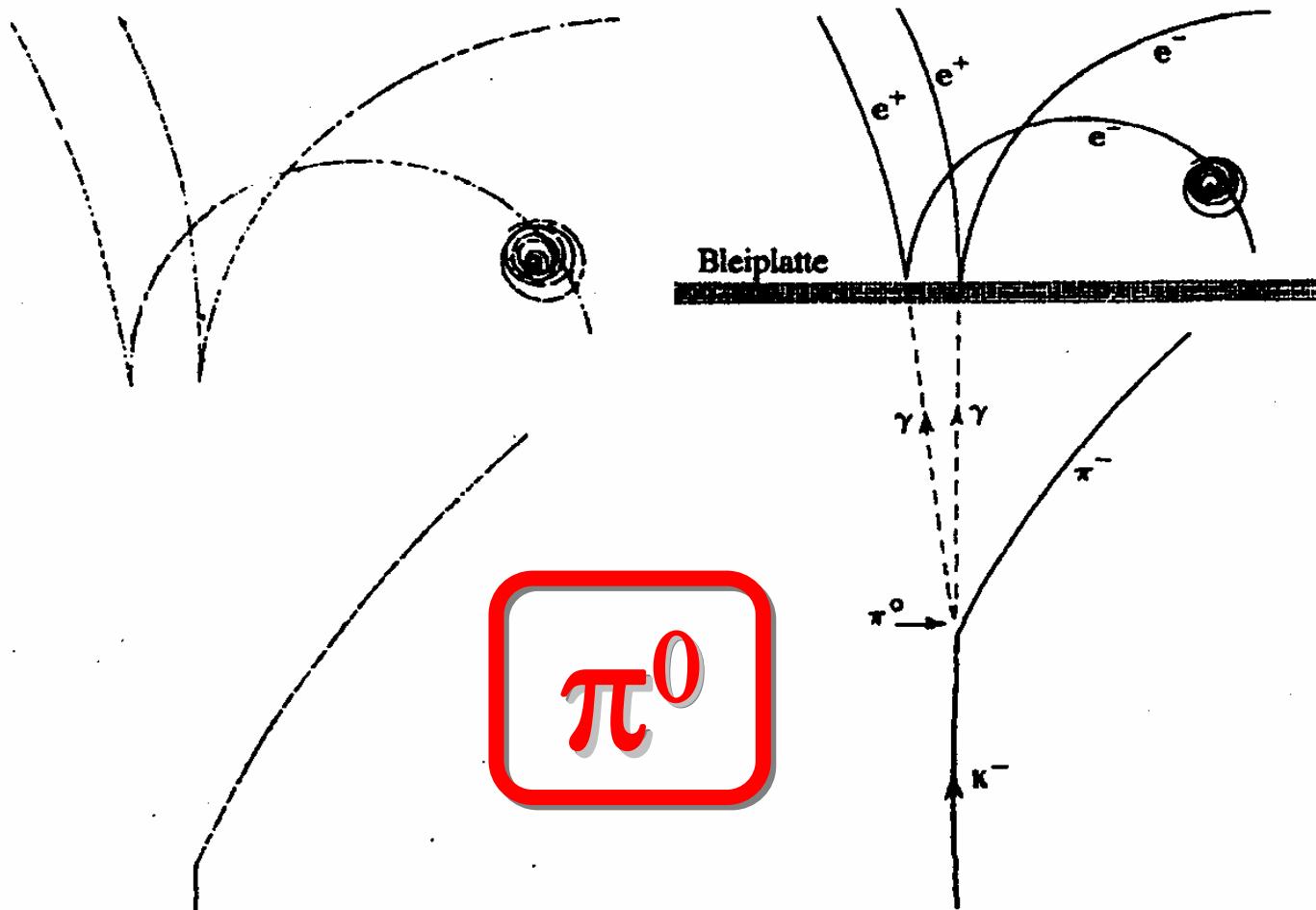
Nobel prize 1948:
Nobel prize 1950:



P. Blackett, Use of cloud chambers in cosmic radiation
C.F. Powell, Discoveries on mesons with emulsions

(still used, τ in OPERA, Gran Sasso)





Die beiden Bilder zeigen eine Fotografie (bei der alle für den Vorgang unwichtigen Linien gelöscht wurde) und eine Zeichnung zur Erklärung des Zerfalls eines negativen Kaons in ein negatives und ein neutrales Pion. Die enge Spur rechts oben stammt von einem Elektron, das aus einem Atom der Blasenkammerflüssigkeit herausgeschlagen wurde. Die Aufnahme stammt aus der Blasenkammer des LBL.

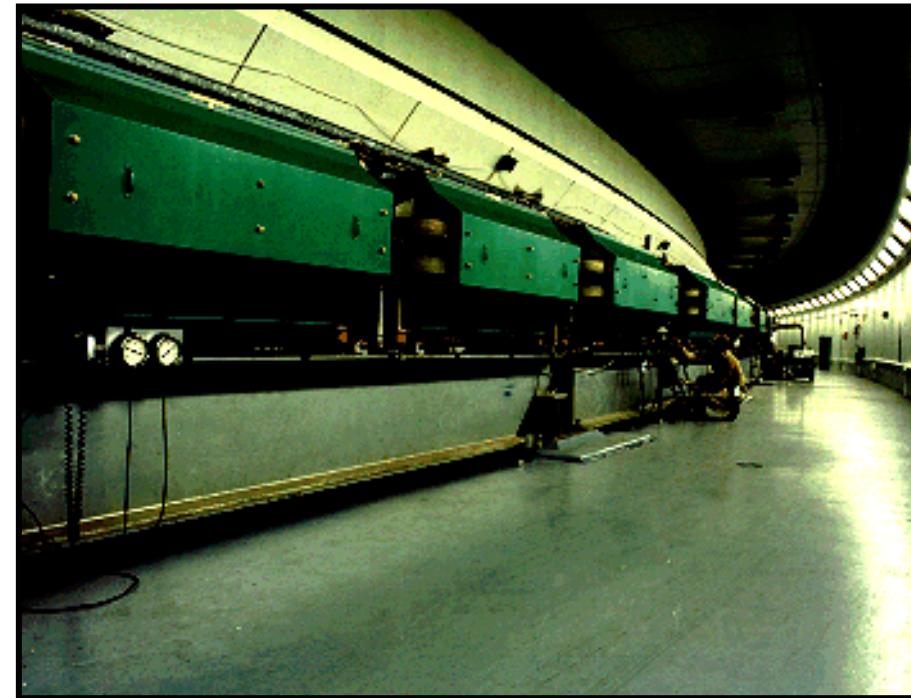
Accelerators

find particle zoo:

BNL Brookhaven National Lab., USA

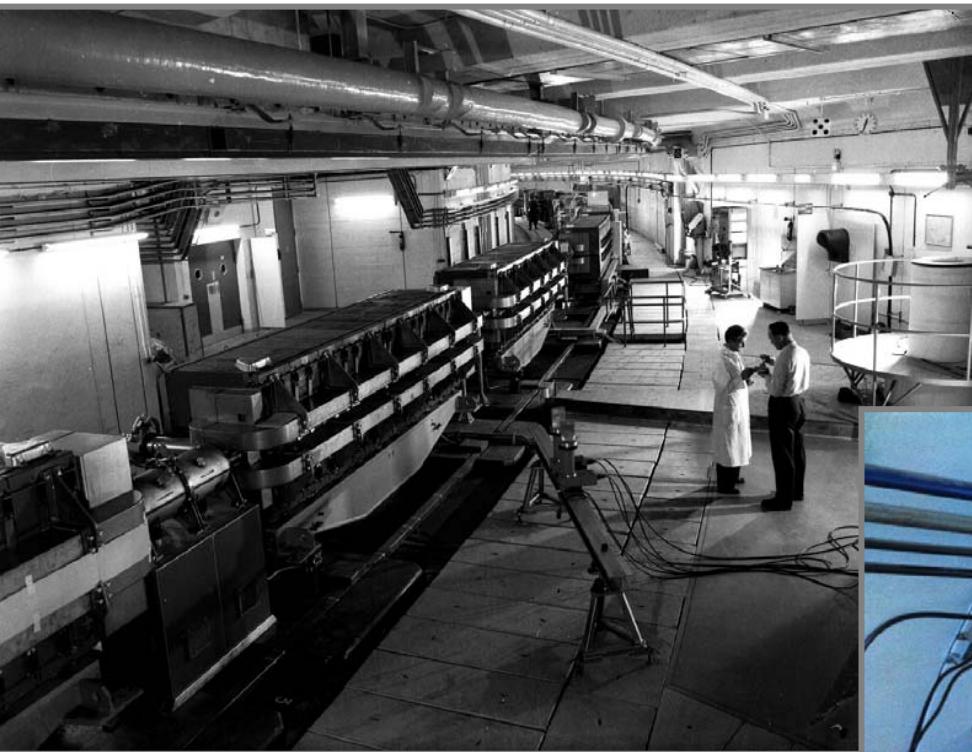


Cosmotron: 1st proton synchrotron.
1953: 3.3 GeV



AGS: Alternating Gradient Synchrotron
1960: 33 GeV

CERN Accelerators



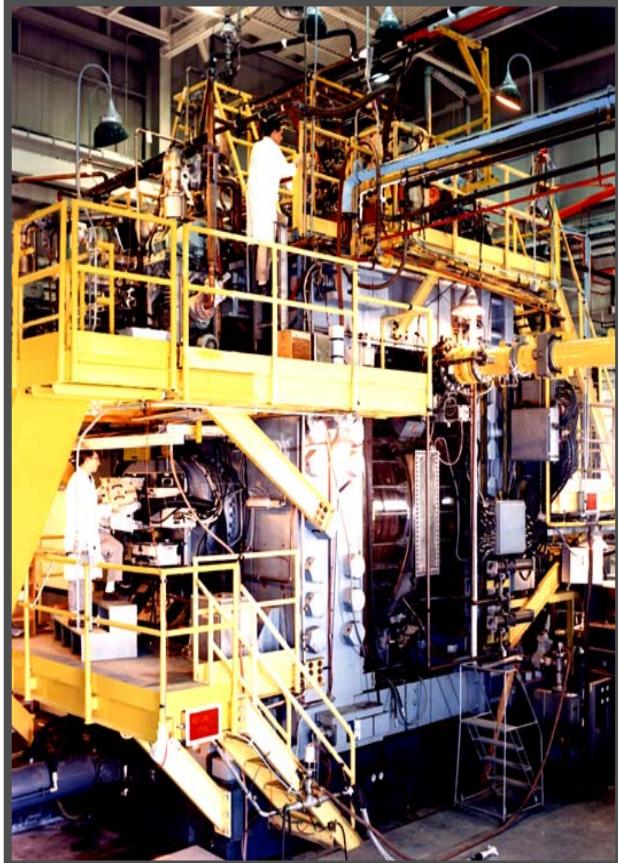
CERN 1959:
26 GeV
Proton Synchrotron

CERN 1976:
400 GeV
Super Proton Synchrotron
7 km



Bubble Chambers

Particle

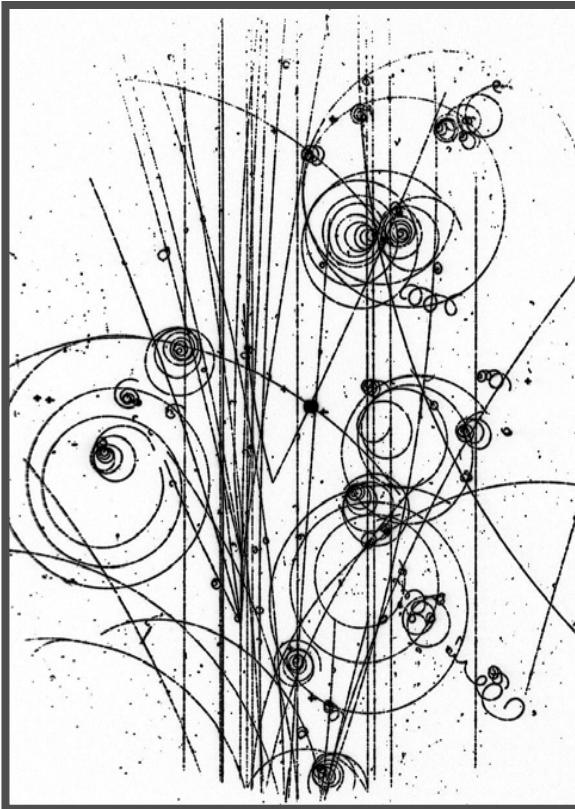


BNL 80" Bubble Chamber

1963-74

Ω^-

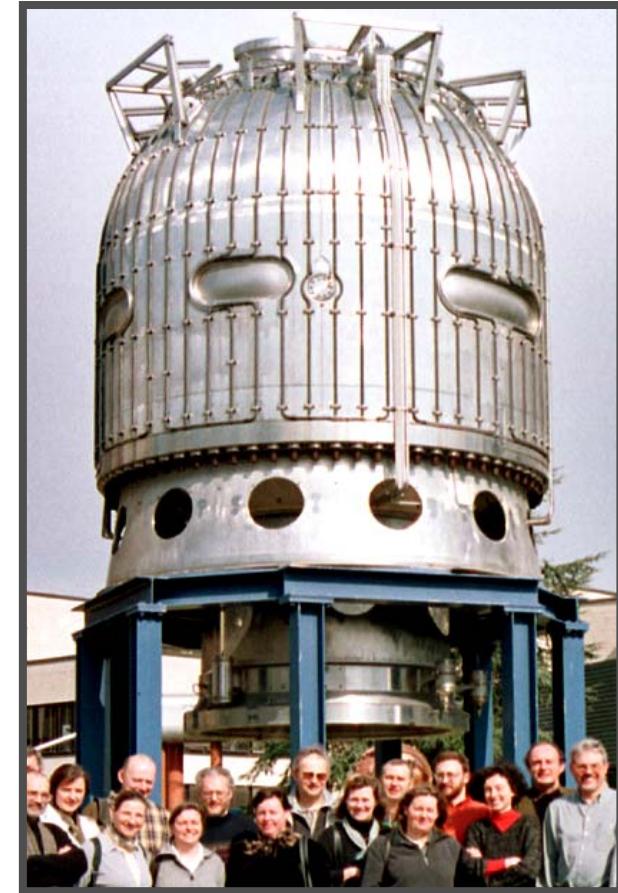
Nobel prize



Zoo

Quark model !

~100 million photos
US + EU



CERN: BEBC

Big European Bubble Chamber

3.7 m, 20 m³, 3.5 T

6 million photos 1974-84

Strangeness

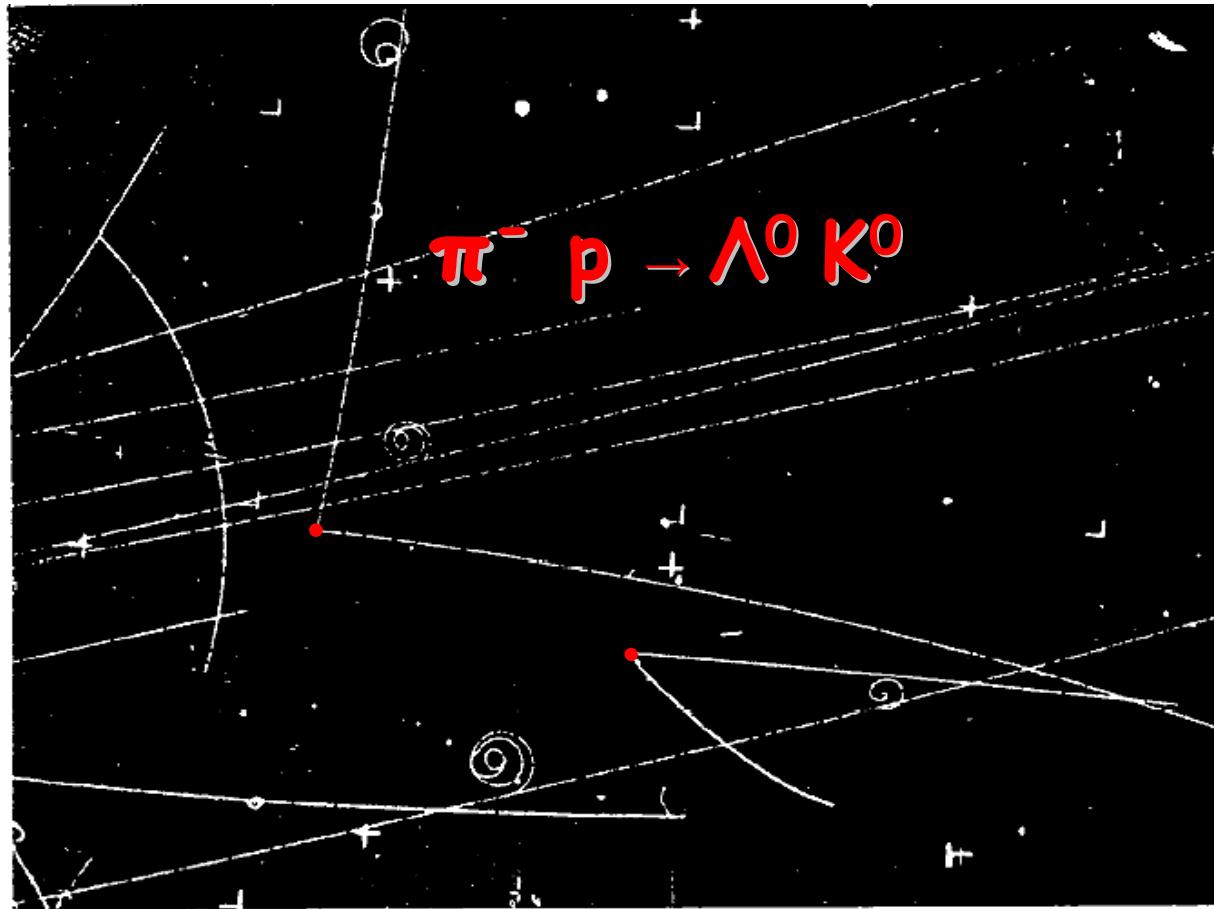


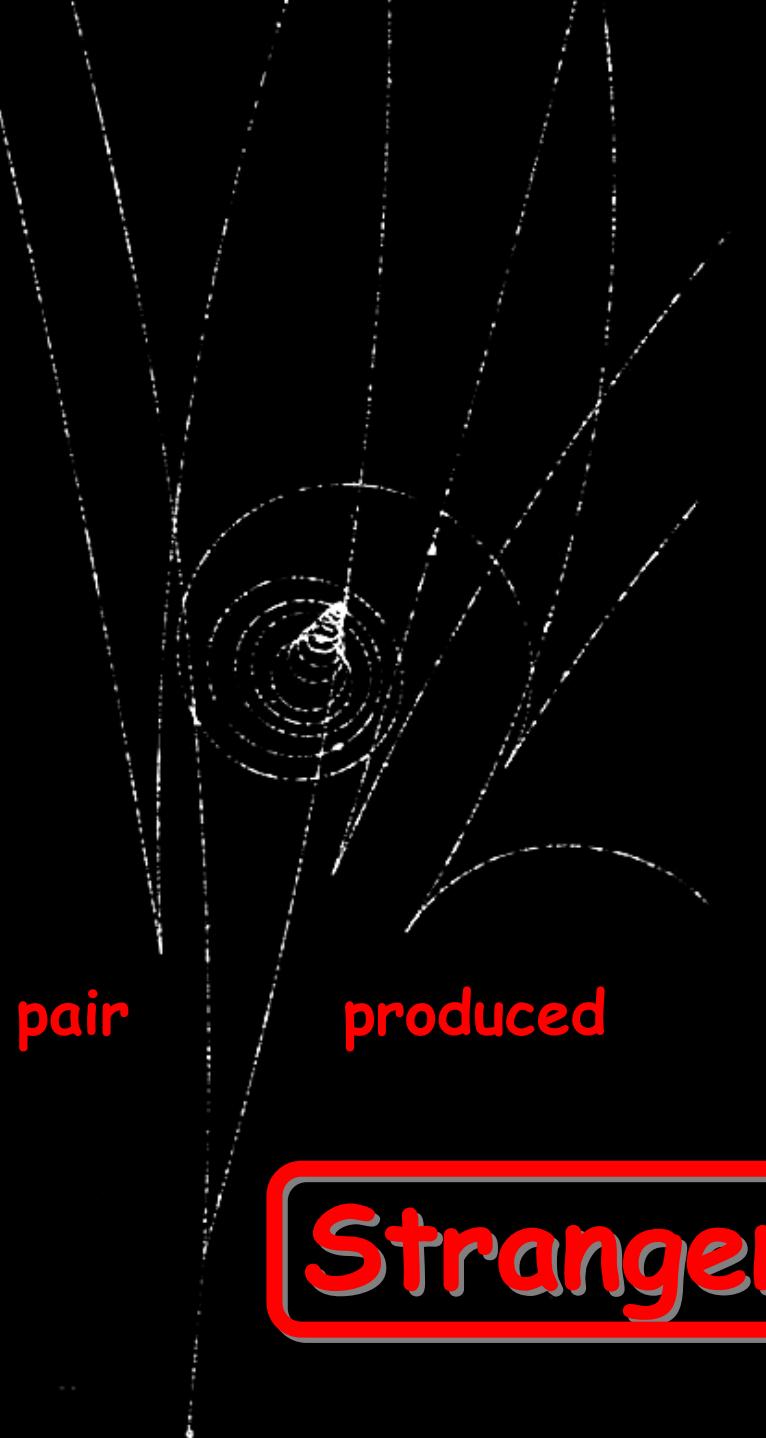
Fig. 12. Associated production, $\pi^- + p \rightarrow \Lambda^0 + K^0$ at about 1 GeV with subsequent decays in Alvarez's hydrogen bubble chamber.

D. Glaser, Nobel prize 1960
L. Alvarez,
1968



invention of bubble chambers
use of bubble chambers

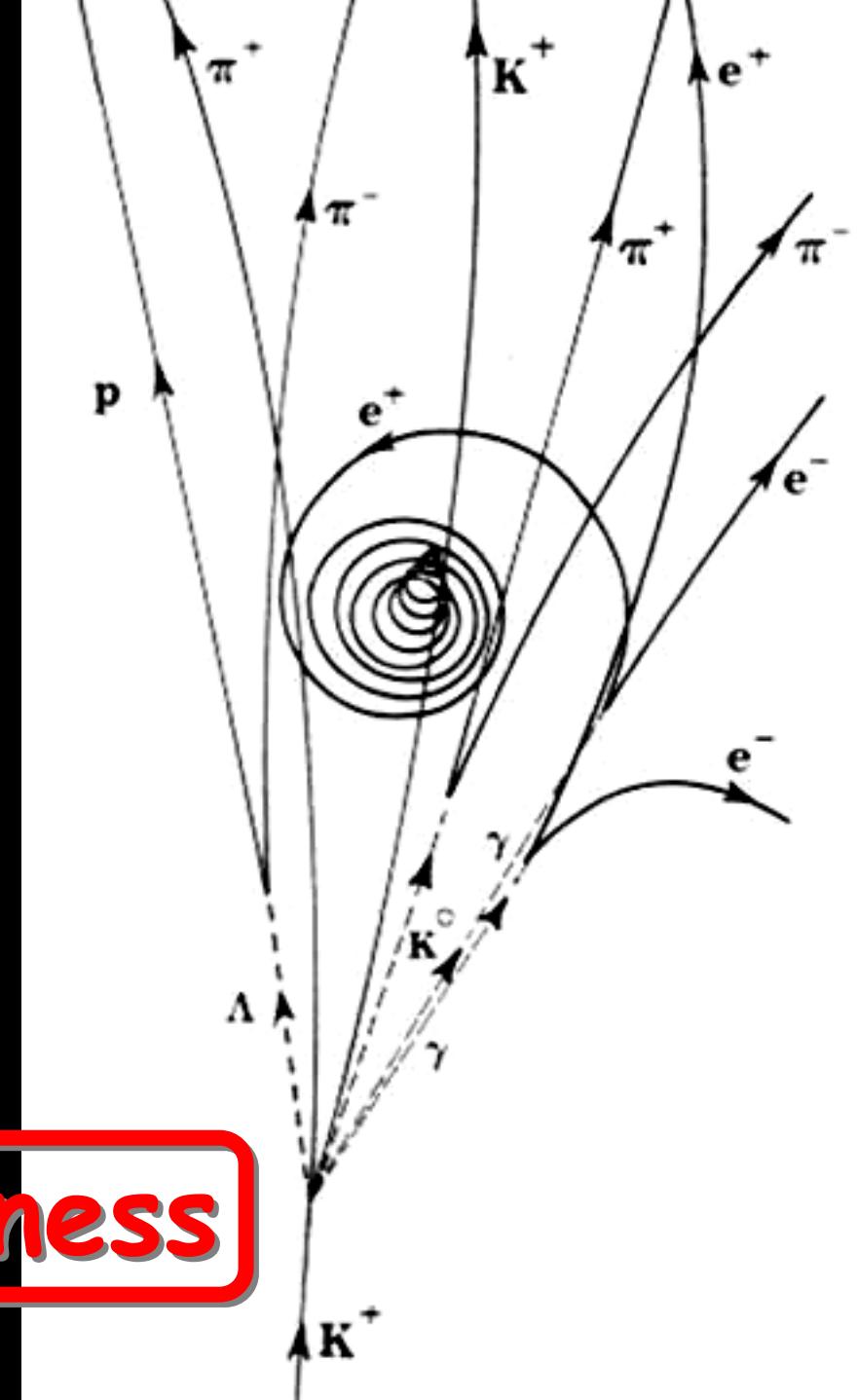
Strange particles:
always pair produced in strong interactions:
new conservation law !
life time $\tau \sim 10^{-8} \dots 10^{-10}$ s
decay length $c\tau \sim \text{cm} \dots \text{m}$



pair

produced

Strangeness



Particle Zoo

weak
decays

lifetime
 $\tau \sim 10^{-8} \dots 10 \text{ s}$

decay length
 $c\tau \sim \text{cm...m}$

π^+ \bar{p}

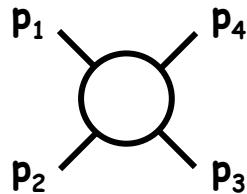
Λ

$\Lambda \rightarrow p \pi^-$

$\Xi^- \rightarrow \Lambda \pi^-$

$\bar{\nu} p$

Strongly bound hadrons



four-momentum
 $p_i = (E_i, \vec{p}_i)$

invariant mass
of a particle:

$$m^2 = p^2 = s = (p_1 + p_2)^2$$

width:

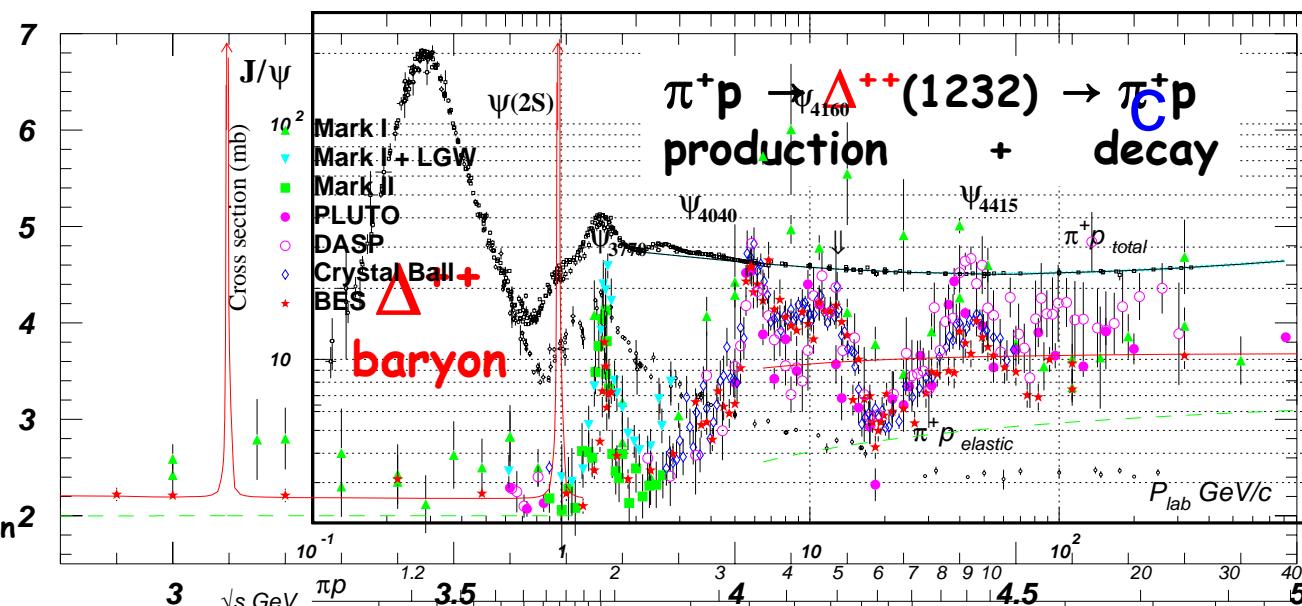
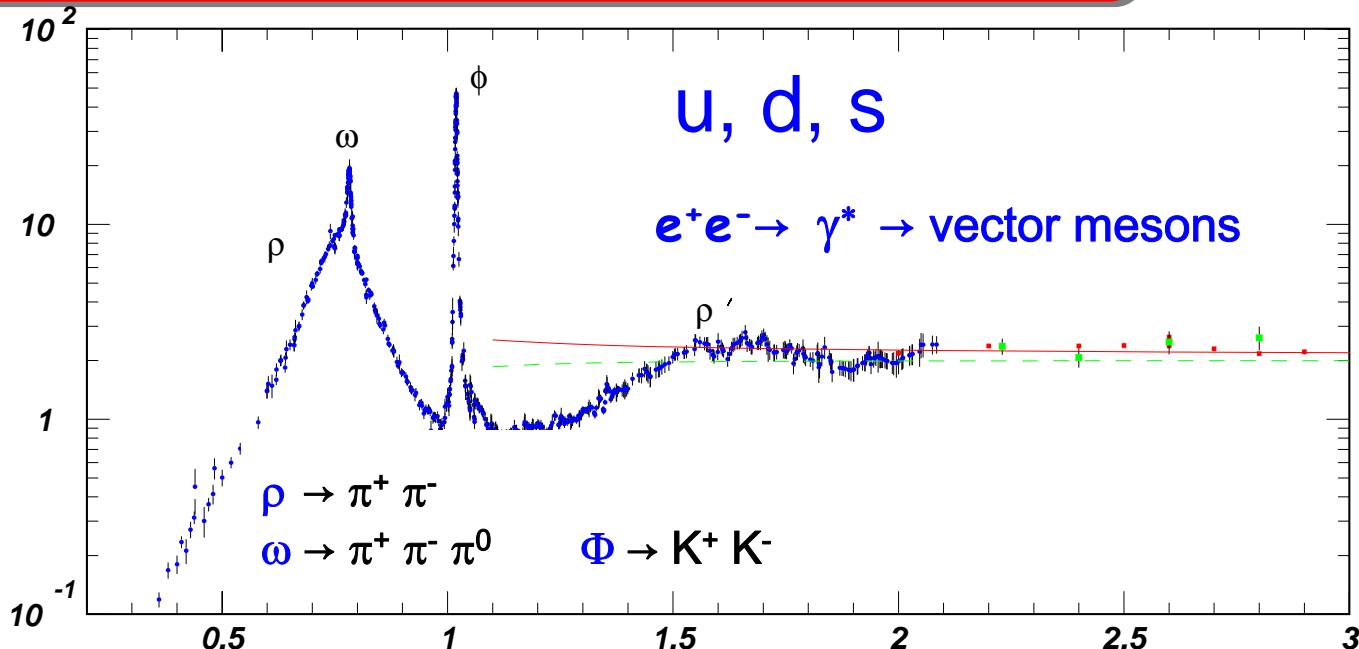
$$\Gamma \sim 120 \text{ MeV}$$

lifetime:

$$\tau \sim 10^{-23} \text{ s}$$

decay length:

$$c\tau \sim \text{fm} \sim r_{\text{proton}}^2$$



The Particle Zoo

Light Mesons

$n^{2s+1}\ell_J$	J^{PC}	$ = 1$ $ud, \bar{u}d, \frac{1}{\sqrt{2}}(d\bar{d} - u\bar{u})$	$ = \frac{1}{2}$ $u\bar{s}, d\bar{s}, \bar{d}s, -\bar{u}s$	$ = 0$ f'	$ = 0$ f	θ_{quad} [°]	θ_{lin} [°]
1^1S_0	0^-+	π	K	η	$\eta'(958)$	-11.5	-24.6
1^3S_1	1^{--}	$\rho(770)$	$K^*(892)$	$\phi(1020)$	$\omega(782)$	38.7	36.0
1^1P_1	1^{+-}	$b_1(1235)$	K_{1B}^\dagger	$h_1(1380)$	$h_1(1170)$		
1^3P_0	0^{++}	$a_0(1450)$	$K_0^*(1430)$	$f_0(1710)$	$f_0(1370)$		
1^3P_1	1^{++}	$a_1(1260)$	K_{1A}^\dagger	$f_1(1420)$	$f_1(1285)$		
1^3P_2	2^{++}	$a_2(1320)$	$K_2^*(1430)$	$f_2'(1525)$	$f_2(1270)$	29.6	28.0
1^1D_2	2^{-+}	$\pi_2(1670)$	$K_2(1770)^\dagger$	$\eta_2(1870)$	$\eta_2(1645)$		
1^3D_1	1^{--}	$\rho(1700)$	$K^*(1680)^\dagger$		$\omega(1650)$		
1^3D_2	2^{--}		$K_2(1820)^\dagger$				
1^3D_3	3^{--}	$\rho_3(1690)$	$K_3^*(1780)$	$\phi_3(1850)$	$\omega_3(1670)$	32.0	31.0
1^3F_4	4^{++}	$a_4(2040)$	$K_4^*(2045)$		$f_4(2050)$		
1^3G_5	5^{--}	$\rho_5(2350)$					
1^3H_6	6^{++}	$a_6(2450)$			$f_6(2510)$		
2^1S_0	0^-+	$\pi(1300)$	$K(1460)$	$\eta(1475)$	$\eta(1295)$	-22.4	-22.6
2^3S_1	1^{--}	$\rho(1450)$	$K^*(1410)^\dagger$	$\phi(1680)$	$\omega(1420)$		

LIGHT UNFLAVORED ($S = C = B = 0$)		STRANGE ($S = \pm 1, C = B = 0$)		BOTTOM ($B = \pm 1$)	
$J^G(J^{PC})$	$J^G(J^{PC})$	$J^G(J^{PC})$	$J^G(J^{PC})$	$J^G(J^{PC})$	$J^G(J^{PC})$
• π^\pm	$1^-(0^-)$	• $\pi_2(1670)$	$1^-(2-+)$	• K^\pm	$1/2(0^-)$
• π^0	$1^-(0^-+)$	• $\phi(1680)$	$0^-(1^-)$	• K^0	$1/2(0^-)$
• η	$0^+(0^-+)$	• $\rho_3(1690)$	$1^+(3^-)$	• K_2^0	$1/2(0^-)$
• $\delta_0(600)$	$0^+(0^++)$	• $\rho(1700)$	$1^+(1^-)$	• K_1^0	$1/2(0^-)$
• $\rho(770)$	$1^+(1^-)$	• $\omega_3(1700)$	$1^-(2++)$	• $K_2^*(892)$	$1/2(1^-)$
• $\omega(782)$	$0^-(1^-)$	• $\delta_2(1710)$	$0^+(0++)$	• $K_1(1270)$	$1/2(1^-)$
• $\eta'(958)$	$0^+(0^-+)$	• $\eta(1760)$	$0^+(0^-+)$	• $K_1(1400)$	$1/2(1^-)$
• $\delta_0(980)$	$0^+(0^++)$	• $\pi(1800)$	$1^-(0^-+)$	• $K^*(1410)$	$1/2(1^-)$
• $a_0(980)$	$1^-(0^++)$	• $\delta_2(1810)$	$0^+(2++)$	• $K_2^*(1430)$	$1/2(0^*)$
• $\phi(1020)$	$0^-(1^-)$	• $\phi_3(1850)$	$0^-(3^-)$	• $K_2^*(1430)$	$1/2(2^+)$
• $h_1(1170)$	$0^-(1^-+)$	• $\eta_2(1870)$	$0^+(2-+)$	• $K_2^*(1430)$	$1/2(2^+)$
• $b_1(1235)$	$1^+(1^-+)$	• $\rho(1900)$	$1^+(1^-)$	• $K_1(1460)$	$1/2(0^-)$
• $a_1(1260)$	$1^-(1^++)$	• $\delta_2(1910)$	$0^+(2++)$	• $K_2(1580)$	$1/2(2^-)$
• $\delta_2(1270)$	$0^+(2^++)$	• $\delta_2(1950)$	$0^+(2++)$	• $K_1(1630)$	$1/2(2^?)$
• $\delta_1(1285)$	$0^+(1^++)$	• $\rho_3(1990)$	$1^+(3^-)$	• $K_1(1650)$	$1/2(1^+)$
• $\eta(1295)$	$0^+(0^-+)$	• $\delta_2(2010)$	$0^+(2++)$	• $K^*(1680)$	$1/2(1^-)$
• $\pi(1300)$	$1^-(0^-+)$	• $\delta_2(2020)$	$0^+(0++)$	• $K_2(1770)$	$1/2(2^-)$
• $a_2(1320)$	$1^-(2++)$	• $a_4(2040)$	$1^-(4++)$	• $K_2^*(1780)$	$1/2(3^-)$
• $\delta_0(1370)$	$0^+(0^++)$	• $\delta_2(2050)$	$0^+(4++)$	• $K_2^*(1820)$	$1/2(2^?)$
• $h_1(1380)$	$1^-(1^-+)$	• $\pi_2(2100)$	$1^-(2-+)$	• $K_1(1830)$	$1/2(0^-)$
• $\pi_1(1400)$	$1^-(1^-+)$	• $\pi_2(2100)$	$0^+(0^+)$	• $K_1^*(1850)$	$1/2(0^+)$
• $\eta(1405)$	$0^+(0^-+)$	• $\delta_2(2150)$	$0^+(2++)$	• $K_2^*(1880)$	$1/2(2^+)$
• $\delta_1(1420)$	$0^+(1^++)$	• $\rho_1(2150)$	$1^+(1^-)$	• $K_1^*(2045)$	$1/2(0^+)$
• $\omega(1420)$	$0^-(1^-)$	• $\delta_2(2200)$	$0^+(0++)$	• $K_2(2250)$	$1/2(2^-)$
• $\delta_2(1430)$	$0^+(2++)$	• $\delta_2(2220)$	$0^+(2++)$	• $K_3(2320)$	$1/2(3^+)$
• $a_0(1450)$	$1^-(0^++)$	• $\alpha_4(2240)$	$\alpha_4(4++)$	• $K_2(2380)$	$1/2(5^-)$
• $\rho(1450)$	$1^+(1^-)$	• $\eta(2225)$	$0^+(0-+)$	• $K_4(2500)$	$1/2(4^-)$
• $\eta(1475)$	$0^+(0^-+)$	• $\rho_3(2250)$	$1^+(3^-)$	• $K(3100)$	$7^?(???)$
• $\delta_0(1500)$	$0^+(0^++)$	• $\delta_2(2300)$	$0^+(2++)$		
• $\delta_1(1510)$	$0^+(1^++)$	• $\delta_2(2340)$	$0^+(2++)$		
• $\rho_2(1525)$	$0^+(2++)$	• $\rho_2(2350)$	$1^+(5^-)$		
• $\delta_2(1565)$	$0^+(2++)$	• $\rho_2(2350)$	$\alpha_4(4++)$		
• $h_1(1595)$	$0^-(1^+)$	• $\delta_2(2450)$	$0^+(0-+)$		
• $\pi_1(1600)$	$1^-(1^-+)$	• $\delta_2(2510)$	$1^-(6++)$		
• $a_1(1640)$	$1^-(1^++)$	• $\delta_2(2510)$	$0^+(6++)$		
• $\pi_2(1645)$	$0^+(2++)$	• $\delta_2(2510)$	$0^+(2++)$		
• $\omega(1650)$	$0^-(1^-+)$	• $\delta_2(2510)$	$0^+(2++)$		
• $\omega_3(1670)$	$0^-(3^-)$	• $\delta_2(2510)$	$0^+(2++)$		
OTHER LIGHT					
Further States					
CHARMED (C = ± 1)					
• D^\pm			$1/2(0^-)$		
• D^0			$1/2(0^-)$		
• $D^*(2007)^0$			$1/2(1^-)$		
• $D^*(2010)^\pm$			$1/2(1^+)$		
• $D_1(2420)^\pm$			$1/2(1^+)$		
• $D_2(2460)^0$			$1/2(2^+)$		
• $D_2(2460)^\pm$			$1/2(2^+)$		
• $D_2(2640)^\pm$			$1/2(2^?)$		
CHARMED, STRANGE (C = S = ± 1)					
• D_s^\pm			$0(0^-)$		
• D_s^0			$0(7^?)$		
• $D_{sJ}(2317)^\pm$			$0(0^+)$		
• $D_{sJ}(2460)^\pm$			$0(1^+)$		
• $D_{s1}(2536)^\pm$			$0(1^+)$		
• $D_{s2}(2573)^\pm$			$0(7^?)$		
NON-q-q CANDIDATES					
NON-q-q CANDIDATES					

Particle Types

LEPTONS
B=0 L=1

fundamental fermions
no strong interaction
no substructure
spin = 1/2

ELECTRONS
Q=1

e m = 511 keV
 μ 106 MeV
 τ 1.8 GeV

NEUTRINOS
Q=0

ν_e m \lesssim 1 eV
 ν_μ
 ν_τ

HADRONS
L=0

B baryon nr
L lepton nr
S strangeness

BARYONS
B=1

MESONS
B=0

S=0

NUCLEONS
S=0

N = (n, p)

m_p = 938 MeV
 $m_n - m_p$ = 1 MeV

p stable
 $n \rightarrow p e^- \bar{\nu}_e$ $\tau = 886$ s

HYPERONS
S=1

$\Sigma^+(1189) \rightarrow N \pi$
 $\Lambda^0(1116) \rightarrow p \pi^-$
 $\Xi^-(1321) \rightarrow \Lambda \pi^-$

$\tau \approx 10^{-10}$ s $c\tau \approx$ cm

Pions = $(\pi^+ \pi^0 \pi^-)$

$m(\pi^\pm) = 140$ MeV
 $m(\pi^0) = 135$ MeV
 $\pi^+ \rightarrow \mu^+ \nu_\mu$ $\tau = 26$ ns $c\tau = 8$ m

Kaons = $(K^+ K^0)$, $(K^- \bar{K}^0)$

$m(K^\pm) = 494$ MeV
 $m(K^0) = 498$ MeV
 $K^+ \rightarrow \mu^+ \nu_\mu$ $\tau = 12$ ns $c\tau = 3$ m

S=1

2.

Quark Model

- **observation:** level scheme of **mirror nuclei** similar:

$$^{15}_7N \equiv ^{15}_8O \quad ^{14}_6C \ ^{14}_7N \ ^{14}_8O \quad ^{13}_7N \equiv ^{13}_6C$$

- **strong interactions symmetric w.r.t. exchange $n \leftrightarrow p$?**

can we say: $n \equiv p$???

- **but:** $^3_1H \neq ^3_2He$

$$\begin{array}{lll} nn & \neq & np = ^2_1H \\ \text{not binding} & & \text{deuterium} \end{array} \neq pp \quad \text{repulsive}$$

- **symmetry broken** by electromagnetic interaction:

$$(m_n - m_p) / m_p = 1 \text{ MeV} / 1 \text{ GeV} = 1 \% \quad \text{small}$$

$p e = H$ atom bound system

$n e =$ not bound

- and by weak interaction: $n \rightarrow p e^- \nu_e$

- **Heisenberg 1932:** isotopic spin:

rotations in 3D real vector space **$SO(3)$** (spin algebra) isomorphic to

rotations in 2D complex vector space **$SU(2)$** (isospin algebra)

- nucleons (n, p) $I = 1/2$ $(2I+1) = 2$ iso-doublet

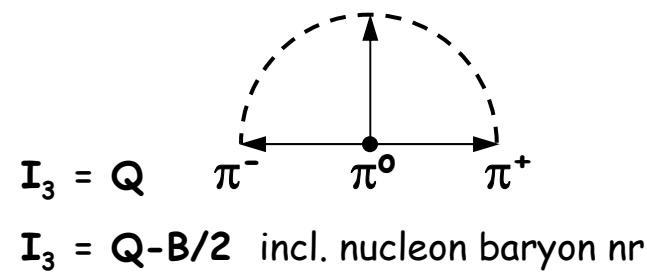
- deuterium ($n p$) $I = 0$ $(2I+1) = 1$ iso-singlet

- pions ($\pi^+ \pi^0 \pi^-$) $I = 1$ $(2I+1) = 3$ iso-triplet

- same **algebra** as for spin:

Clebsch-Gordon coefficients give cross section and decay branching ratios. See exercises !

Iso-Spin



Clebsch-Gordon coefficients

- Quantum Mechanics: vector addition of angular momenta $j_1 \oplus j_2$:
- triangular condition: $|j_1 - j_2| < J < j_1 + j_2$ $M = m_1 + m_2$

$$|JM\rangle = \sum_{m_1, m_2} |j_1 m_1 j_2 m_2\rangle \langle j_1 m_1 j_2 m_2 |JM\rangle$$

- multiplicity: $\sum_{j=|j_1-j_2|}^{j_1+j_2} (2j+1) = (2j_1+1)(2j_2+1)$
- normalization: $\sum_{m_1, m_2} |\langle j_1 m_1 j_2 m_2 |JM\rangle|^2 = 1 \quad =>$

$$\begin{aligned} \langle j_1 m_1 j_2 m_2 | jm \rangle &= \delta_{m_1+m_2, m} \sqrt{\frac{(j_1 + j_2 - j)! (j + j_1 - j_2)! (j + j_2 - j_1)! (2j+1)}{(j + j_1 + j_2 + 1)!}} \\ &\times \sum_k \frac{(-1)^k \sqrt{(j_1 + m_1)! (j_1 - m_1)! (j_2 + m_2)! (j_2 - m_2)! (j + m)! (j - m)!}}{k! (j_1 + j_2 - j - k)! (j_1 - m_1 - k)! (j_2 + m_2 - k)! (j - j_2 + m_1 + k)! (j - j_1 - m_2 + k)!} \end{aligned}$$

Wigner 1931
Racah 1942

- Abramowitz, M. and Stegun, I.A. (Eds.). "Vector-Addition Coefficients." § 27.9 in "Handbook of Mathematical Functions with Formulas, Graphs, and Mathematical Tables", 9th printing. New York: Dover, pp. 1006-1010, 1972.
- Condon, E.U. and Shortley, G. § 3.6-3.14 in "The Theory of Atomic Spectra." Cambridge, England: Cambridge University Press, pp. 56-78, 1951.
- Messiah, A. "Clebsch-Gordan Coefficients and 3j Symbols." Appendix C.I in "Quantum Mechanics", Vol. 2. Amsterdam, Netherlands: North-Holland, pp. 1054-1060, 1962.

$J_1 \times J_2$

$1/2 \times 1/2$	$\begin{matrix} 1 \\ +1 \\ +1/2+1/2 \\ +1/2-1/2 \\ -1/2+1/2 \\ -1/2-1/2 \end{matrix}$	$\begin{matrix} 1 & 0 \\ 0 & 0 \\ 1/2 & 1/2 \\ 1/2 & -1/2 \\ 1/2 & +1/2 \\ 1/2 & -1/2 \end{matrix}$	$\begin{matrix} 1 \\ 0 \\ 1 \\ 1 \\ -1 \\ -1 \end{matrix}$
------------------	---	---	--

$1 \times 1/2$	$\begin{matrix} 3/2 \\ +3/2 \\ +1+1/2 \end{matrix}$	$\begin{matrix} 3/2 & 1/2 \\ 1 & +1/2+1/2 \end{matrix}$	$\begin{matrix} 3/2 & 1/2 \\ 1/3 & 2/3 \\ 2/3 & -1/3 \end{matrix}$	$\begin{matrix} -1/2-1/2 \\ 0-1/2 \\ -1+1/2 \end{matrix}$	$\begin{matrix} 1 \\ 1/3 \\ 2/3 \\ 1/3-2/3 \\ -3/2 \end{matrix}$
2×1	$\begin{matrix} 3 \\ +3 \\ +2+1 \end{matrix}$	$\begin{matrix} 3 & 2 \\ 1 & +2+2 \end{matrix}$	$\begin{matrix} 3 & 2 & 1 \\ 1/3 & 2/3 & 1/3-2/3 \end{matrix}$	$\begin{matrix} -1-1/2 \\ 1+1 \\ +1 \end{matrix}$	$\begin{matrix} 1 \\ 1/3 \\ 2/3 \\ 1/3-2/3 \\ +1 \end{matrix}$
1×1	$\begin{matrix} 2 \\ +2 \\ +1+1 \end{matrix}$	$\begin{matrix} 2 & 1 \\ 1 & +1+1 \end{matrix}$	$\begin{matrix} 2 & 1 \\ 1/15 & 1/3 \\ 8/15 & 1/6-3/10 \\ 6/15 & -1/2 \\ 1/10 & \end{matrix}$	$\begin{matrix} 3 & 2 & 1 \\ 1/5 & 1/2 & 3/10 \\ 0 & 0 & 3/5 \\ 0 & -2/5 & \\ -1+1 & 1/5-1/2 & 3/10 \end{matrix}$	$\begin{matrix} 0 \\ 0+1 \\ 0 \\ 0 \\ 0 \end{matrix}$
	$\begin{matrix} +1 & 0 \\ 0 & +1 \end{matrix}$	$\begin{matrix} 1/2 & 1/2 \\ 1/2 & -1/2 \end{matrix}$	$\begin{matrix} 2 & 1 & 0 \\ 0 & 0 & 0 \end{matrix}$	$\begin{matrix} +1-1 \\ 0 \\ -1+1 \end{matrix}$	$\begin{matrix} 1/6 & 1/2 & 1/3 \\ 2/3 & 0-1/3 & \\ 1/6 & -1/2 & 1/3 \end{matrix}$
	$\begin{matrix} +1 & -1 \\ 0 & 0 \\ -1 & +1 \end{matrix}$	$\begin{matrix} 1/2 & 1/2 \\ 0-1/3 & \\ 1/6 & -1/2 \\ 1/3 & \end{matrix}$	$\begin{matrix} 2 & 1 \\ -1 & -1 \end{matrix}$	$\begin{matrix} 0-1 \\ -1 \\ -1 \end{matrix}$	$\begin{matrix} 1/2 & 1/2 \\ 1/2 & -1/2 \\ -1/2 & 1/2 \\ -1/2 & -1/2 \end{matrix}$
	$\begin{matrix} -1 & 0 \\ -1 & -1 \end{matrix}$	$\begin{matrix} 1/2 & 1/2 \\ 1/2 & -1/2 \end{matrix}$	$\begin{matrix} 2 \\ -2 \end{matrix}$	$\begin{matrix} 1 \\ 1 \end{matrix}$	$\begin{matrix} 2 \\ -2 \end{matrix}$

Clebsch-Gordon coefficients

$$J = J_1 \otimes J_2$$

$$M = m_1 + m_2$$

$2 \times 1/2$	$\begin{matrix} 5/2 \\ +5/2 \\ +2+1/2 \end{matrix}$	$\begin{matrix} 5/2 & 3/2 \\ 1 & 3/2+3/2 \end{matrix}$	$\begin{matrix} 5/2 & 3/2 \\ 1/5 & 4/5 \\ 4/5 & -1/5 \end{matrix}$	$\begin{matrix} j_1 \times j_2 & '2 \\ +1/2 & +1/2 \end{matrix}$	$\begin{matrix} 1/2-1/2 \\ 2/5 & 3/5 \\ 3/5 & -2/5 \end{matrix}$
$3/2 \times 1/2$	$\begin{matrix} 2 \\ +2 \\ +3/2+1/2 \end{matrix}$	$\begin{matrix} 2 & 1 \\ 1 & +1+1 \end{matrix}$	$\begin{matrix} 1/4 & 3/4 \\ 3/4 & -1/4 \end{matrix}$	$\begin{matrix} 5/2 & 3/2 \\ 1/2 & 1/2 \\ 1/2 & -1/2 \end{matrix}$	$\begin{matrix} 4/5 & 1/5 \\ 1/5 & -4/5 \\ -5/2 & \end{matrix}$
$3/2 \times 1$	$\begin{matrix} 5/2 \\ +5/2 \\ +3/2+1 \end{matrix}$	$\begin{matrix} 5/2 & 3/2 \\ 1 & +3/2+3/2 \end{matrix}$	$\begin{matrix} 5/2 & 3/2 & 1/2 \\ 2/5 & 3/5 & +1/2 \\ 3/5 & -2/5 & +1/2 \end{matrix}$	$\begin{matrix} 1/2-1/2 \\ 1/2 & 1/2 \\ 1/2 & -1/2 \end{matrix}$	$\begin{matrix} 2 & 1 \\ -1 & -1 \end{matrix}$
$3/2 \times 3/2$	$\begin{matrix} 3 \\ +3 \\ +3/2+3/2 \end{matrix}$	$\begin{matrix} 3 & 2 \\ 1 & +2+2 \end{matrix}$	$\begin{matrix} 1/2 & 1/2 \\ 1/2 & -1/2 \end{matrix}$	$\begin{matrix} 3/2 & 1/2 \\ 1/5 & 1/2 \\ 3/5 & 0 \end{matrix}$	$\begin{matrix} 3/2 & 1/2 \\ 1/4 & 1/4 \\ 1/4-3/4 & -2 \end{matrix}$
	$\begin{matrix} +1 & -1 \\ -1 & 0 \\ -2 & +1 \end{matrix}$	$\begin{matrix} 1/2 & 1/2 \\ 1/2 & -1/2 \\ 1/2 & -1/2 \end{matrix}$	$\begin{matrix} 3 & 2 \\ -1 & -2 \end{matrix}$	$\begin{matrix} 1/10 & 2/5 & 1/2 \\ 3/5 & 1/15 & -1/3 \\ 3/10 & -8/15 & 1/6 \end{matrix}$	$\begin{matrix} 2 & 1 \\ -3/2 & 1/2 \\ -3/2 & -1/2 \end{matrix}$
	$\begin{matrix} -1 & 0 \\ -3/2 & +1 \end{matrix}$	$\begin{matrix} 1/2 & 1/2 \\ 1/2 & -1/2 \\ 1/2 & -1/2 \end{matrix}$	$\begin{matrix} 3 & 2 \\ -2 & -2 \end{matrix}$	$\begin{matrix} 1/10 & 8/15 & 1/6 \\ 3/5 & -1/15 & -1/3 \\ 1/10 & -2/5 & 1/2 \end{matrix}$	$\begin{matrix} 5/2 & 3/2 \\ 5/2 & 3/2 \\ -3/2 & -3/2 \end{matrix}$
	$\begin{matrix} -3/2 & -1 \end{matrix}$	$\begin{matrix} 1/2 & 1/2 \\ 1/2 & -1/2 \\ 1/2 & -1/2 \end{matrix}$	$\begin{matrix} 3 & 2 \\ -2 & -1 \end{matrix}$	$\begin{matrix} 3/5 & 2/5 \\ 2/5 & -3/5 \\ -5/2 & \end{matrix}$	$\begin{matrix} 5/2 & 3/2 \\ 5/2 & 3/2 \\ -3/2 & -1 \end{matrix}$

$2 \times 3/2$	$\begin{matrix} 7/2 \\ +7/2 \\ +2+3/2 \end{matrix}$	$\begin{matrix} 7/2 & 5/2 \\ 1 & +5/2+5/2 \end{matrix}$	$\begin{matrix} 7/2 & 5/2 & 3/2 \\ 3/7 & 4/7 & 7/2 \\ 4/7 & 3/7 & 5/2 \\ 1/2 & 1/2 & 3/2 \end{matrix}$
----------------	---	---	--

$3/2 \times 3/2$	$\begin{matrix} 3 \\ +3 \\ +3/2+3/2 \end{matrix}$	$\begin{matrix} 3 & 2 \\ 1 & +2+2 \end{matrix}$	$\begin{matrix} 1/2 & 1/2 \\ 1/2 & -1/2 \end{matrix}$
	$\begin{matrix} +3/2 & +1/2 \\ +1/2 & +3/2 \end{matrix}$	$\begin{matrix} 3 & 2 & 1 \\ 1/5 & 1/2 & 3/10 \\ 3/5 & 0 & -2/5 \end{matrix}$	$\begin{matrix} 3 & 2 & 1 \\ 1/5 & 1/2 & 3/10 \\ 3/5 & 0 & -2/5 \end{matrix}$

Notation:

J	J	\dots
M	M	\dots

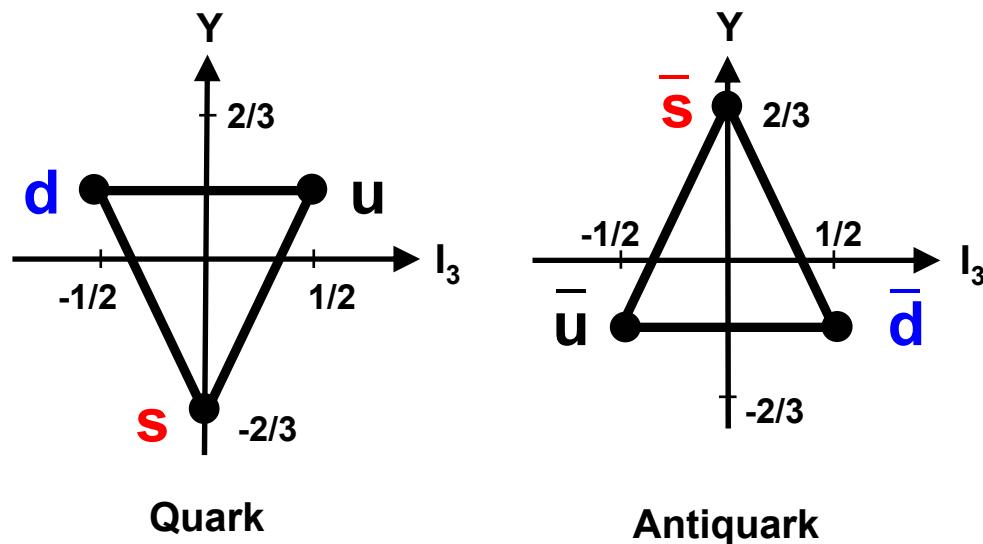
m_1	m_2
m_1	m_2
\vdots	\vdots
m_1	m_2

Coefficients

Quarks

Quark	Q	I_3	B	S	Y
u	+2/3	+1/2		0	+1/3
d	-1/3	-1/2	1/3	0	+1/3
s	-1/3	0		-1	-2/3

spin
 $J = 1/2$



$$Y = B + S$$

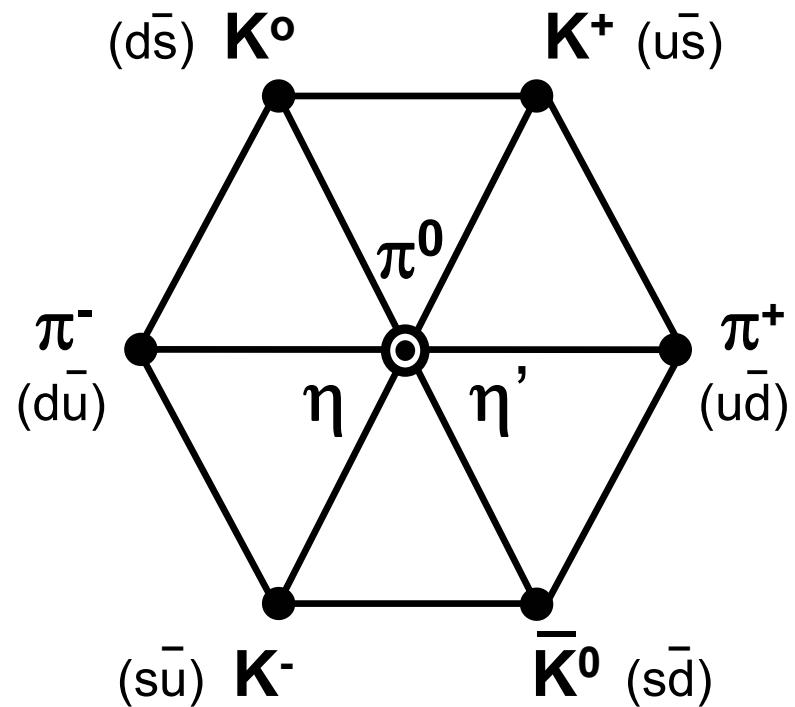
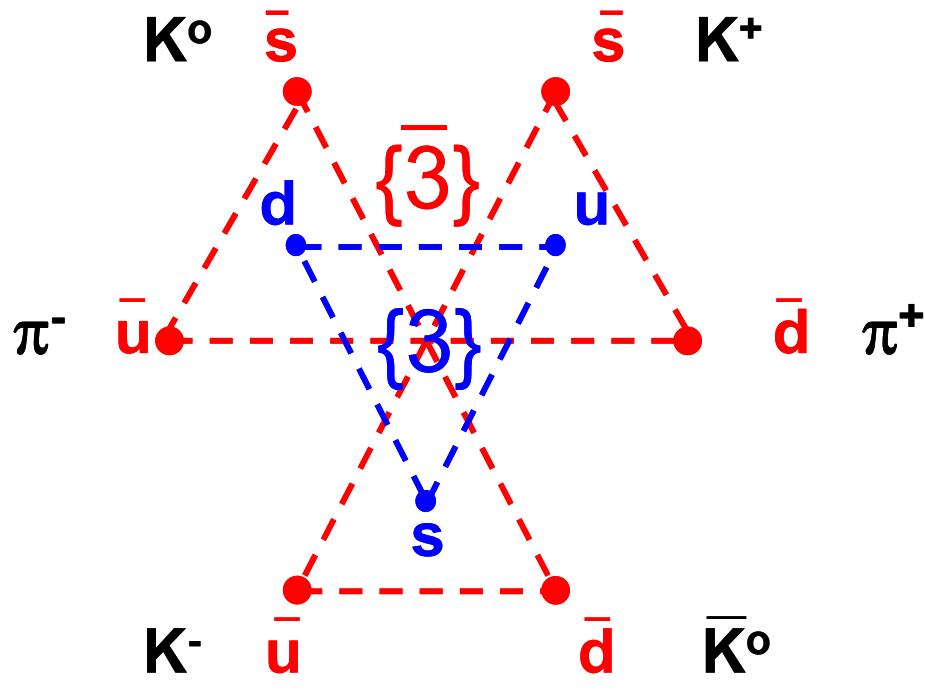
$$I_3 = Q - Y/2$$

hypercharge
 Gell-Mann - Nishijima rule

Meson Octet

$$\{q\} \otimes \{\bar{q}\} = \nabla \otimes \Delta = \{2\} \otimes \{\bar{2}\} = \{1\} \oplus \{3\} \text{ in } \text{SU}(2)$$

$$= \{3\} \otimes \{\bar{3}\} = \{1\} \oplus \{8\} \text{ in } \text{SU}(3)$$



Meson Octet

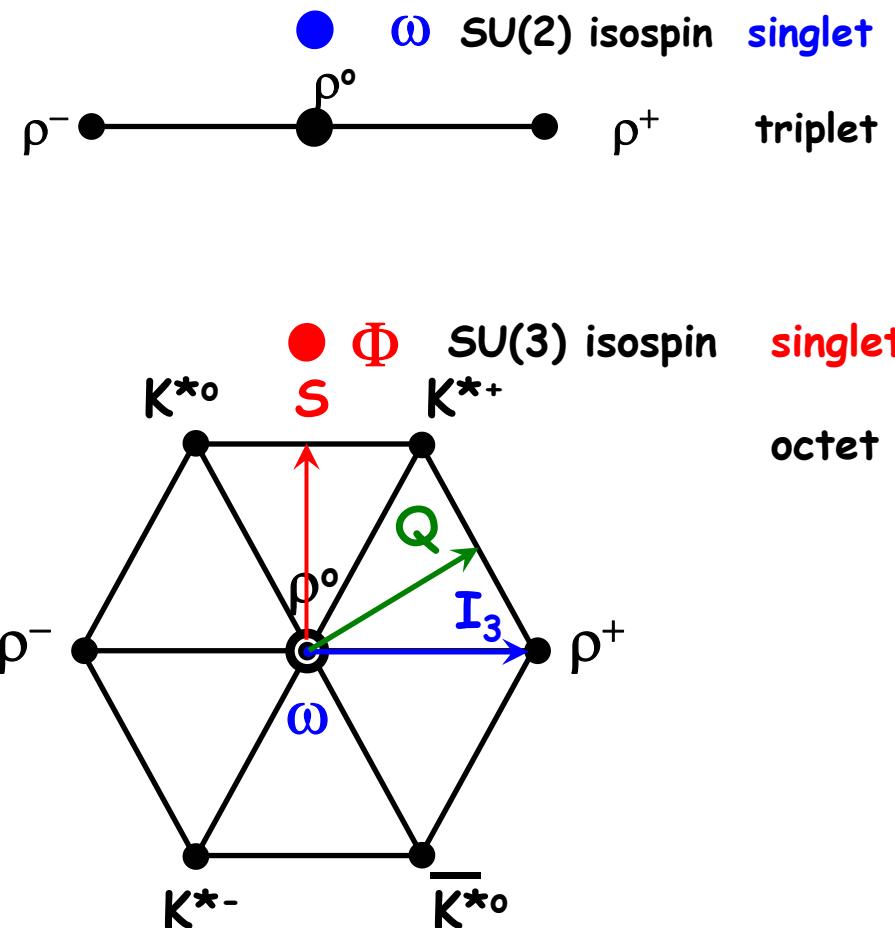
$$= \{q\} \otimes \{\bar{q}\} = \nabla \otimes \Delta$$

$$= \{2\} \otimes \{\bar{2}\} = \{1\} \oplus \{3\} \text{ in } \mathbf{SU(2)}$$

$$= \{3\} \otimes \{\bar{3}\} = \{1\} \oplus \{8\} \text{ in } \mathbf{SU(3)}$$

hyper-charge
 $Y = B + S$

Gell-Mann - Nishijima:
 $I_3 = Q - Y/2$

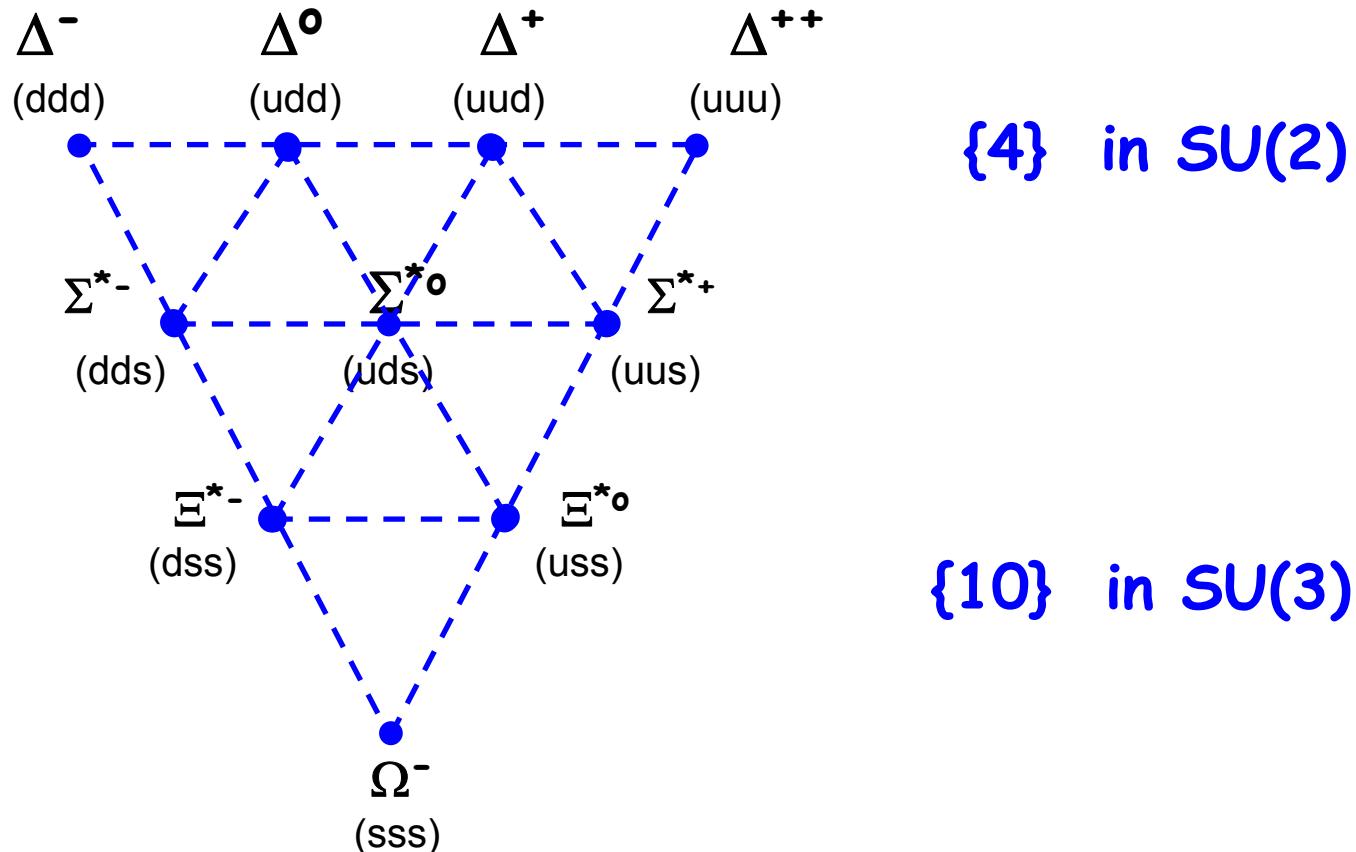


Baryon Decuplet

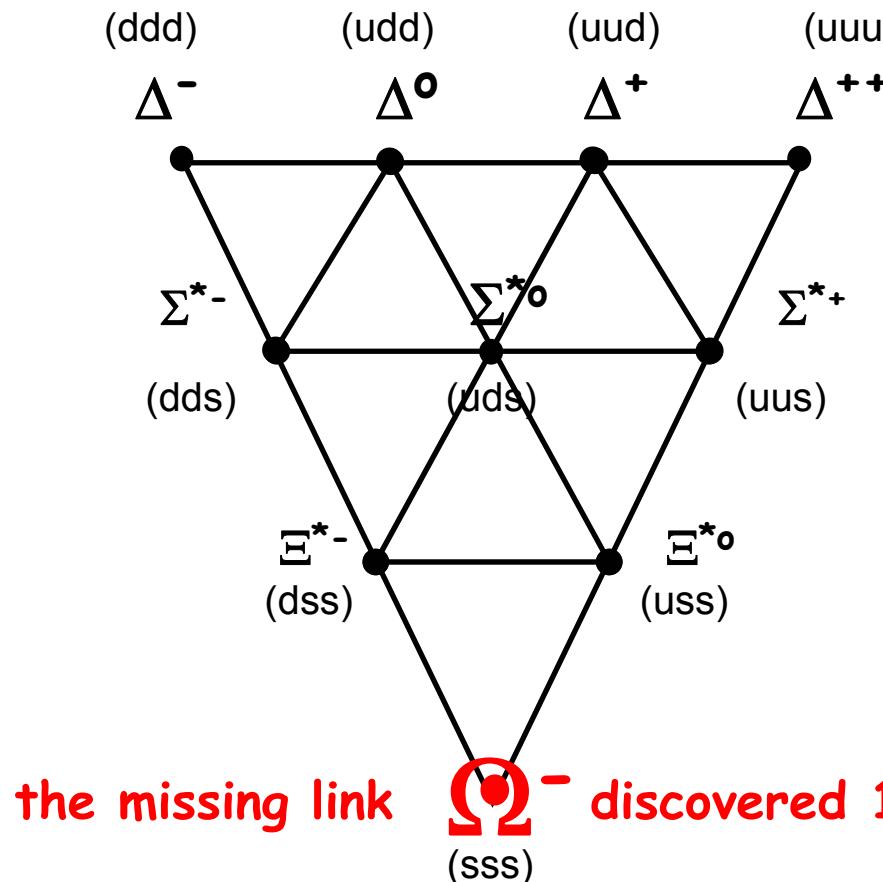
$$= \{q\} \otimes \{q\} \otimes \{q\} = \nabla \otimes \nabla \otimes \nabla$$

$$= \{2\} \otimes \{2\} \otimes \{2\} = \{2\} \oplus \{2\} \oplus \{4\} \quad \text{in } SU(2)$$

$$= \{3\} \otimes \{3\} \otimes \{3\} = \{1\} \oplus \{8\} \oplus \{8\} \oplus \{10\} \quad \text{in } SU(3)$$



Baryon Decuplet

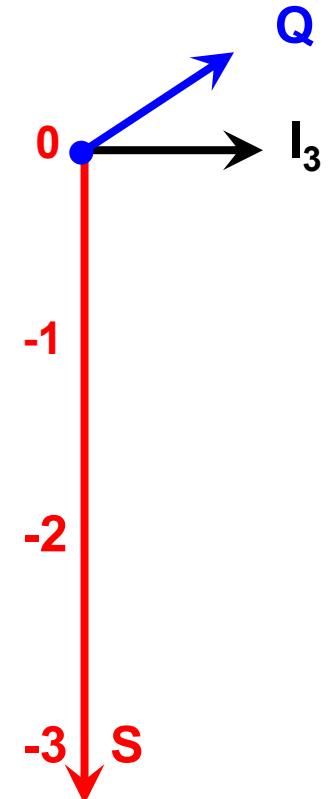


in Brookhaven 80 inch bubble chamber

Nobel prizes:



1968 Alvarez for bubble chamber
1969 Gell-Mann for quark model



Quarks

Three quarks for Muster Mark!
Sure he hasn't got much of a bark
And sure any he has it's all beside the mark.

James Joyce, Finnegans Wake, Book 2, Episode 4, Page 383.

Poem against King Mark, cuckolded husband in the Tristan legend.



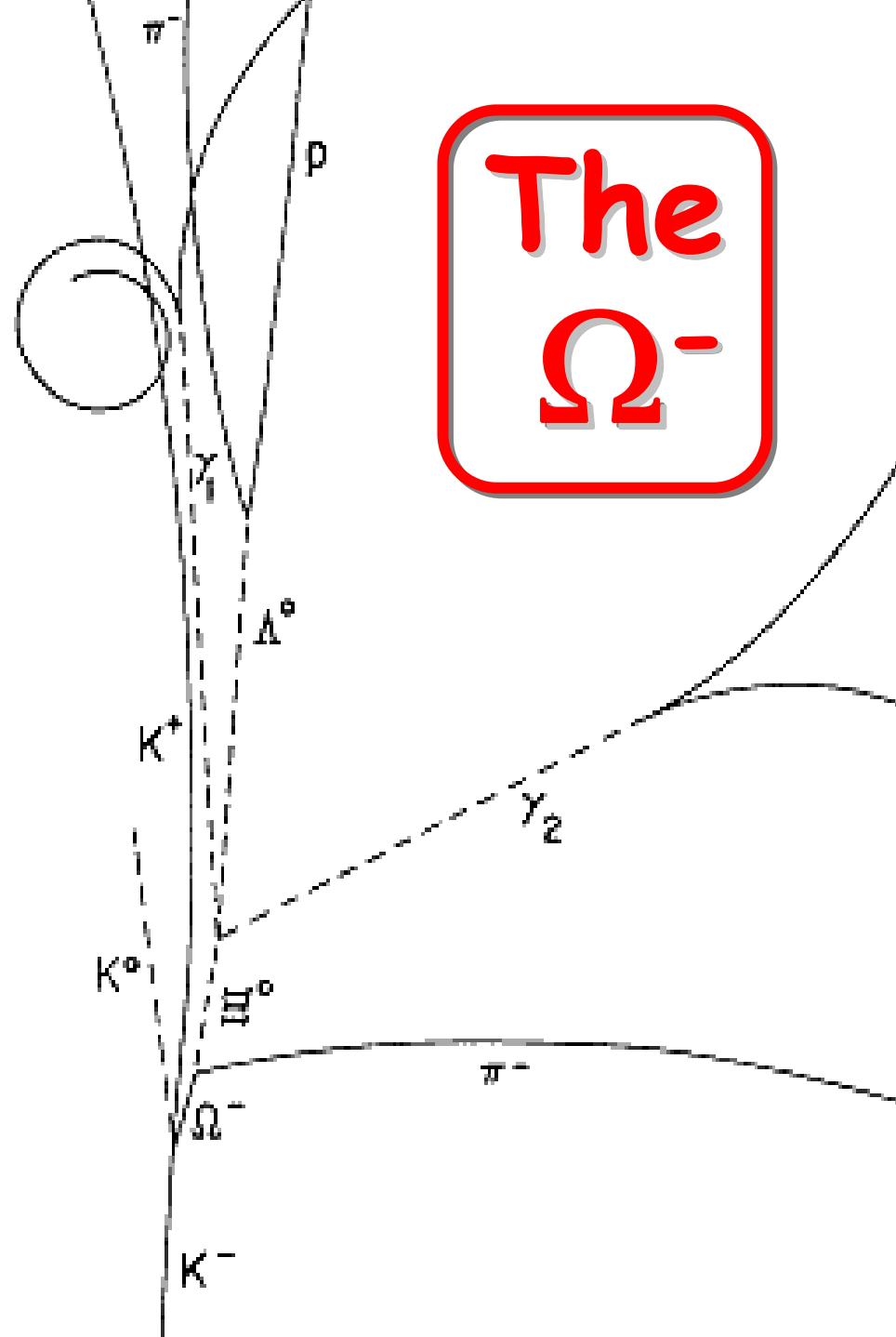
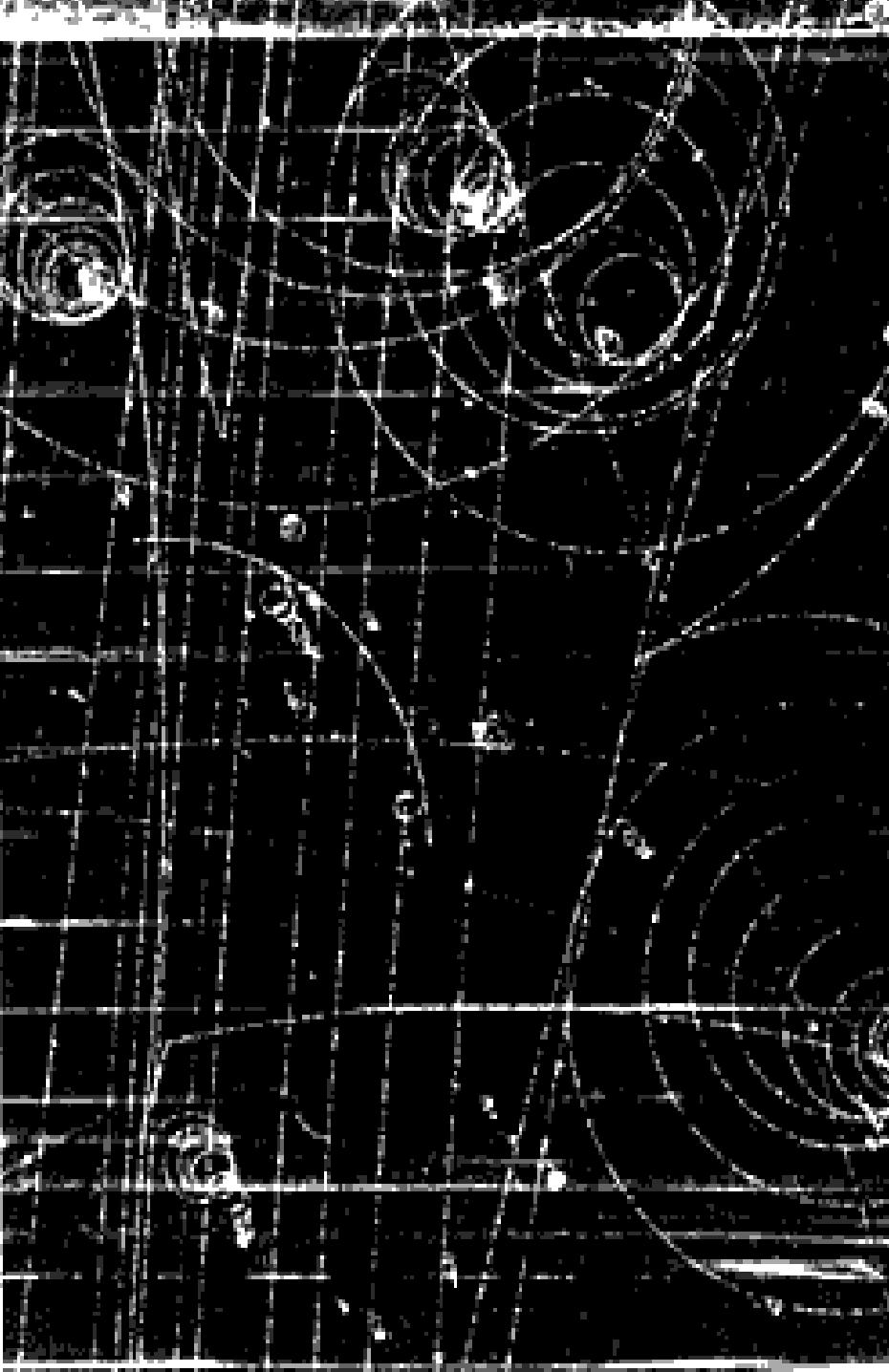
Nobel prize 1969

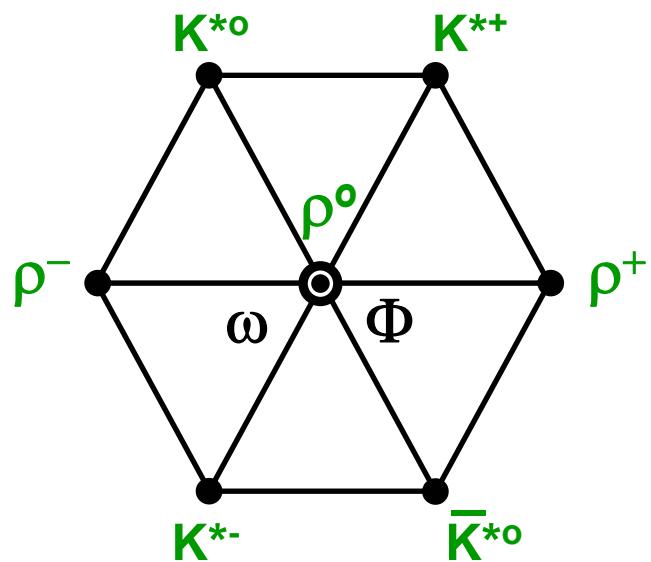
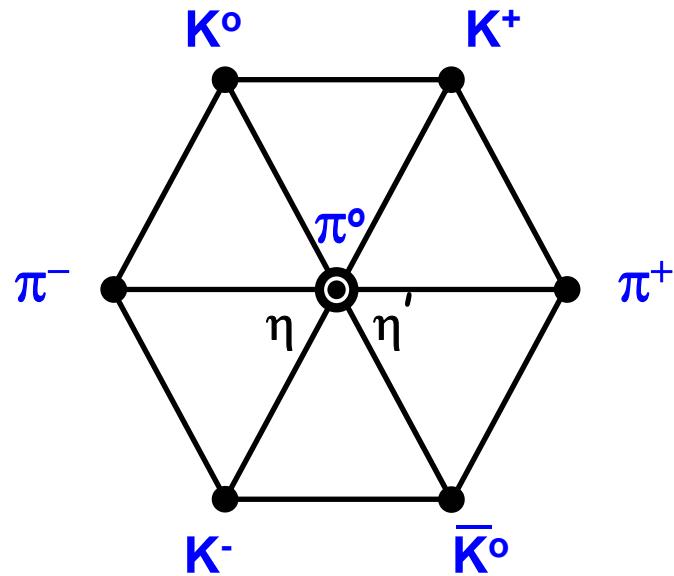
The allusion to **three quarks** seemed perfect ...
by supposing that one ingredient of the line
'Three quarks for Muster Mark'
was a cry of 'Three quarts for Mister Mark'
heard in H.C. Earwicker's pub.



M. Gell-Mann
for quark model

Murray Gell-Mann, private Letter to the Editor
of the Oxford English Dictionary, June 27, 1978 .





PSEUDO-SCALAR

$\uparrow \downarrow \quad J^P = 0^{-+}$

weak / elm. decay

The Spin

VECTOR

$\uparrow \uparrow \quad J^P = 1^-$

strong decay

The Quark Model

MESONS

Mass MeV	$J^P = 0^-$ $(\uparrow \downarrow)$	$J^P = 1^-$ $(\uparrow \uparrow)$	Mass MeV
494/498	$(d\bar{s}) K^0$ $K^+ (us^-)$ π^- π^0 π^+ (du) η η' (ud)	$(d\bar{s}) K^{*0}$ $K^{*+} (us^-)$ ρ^- ρ^0 ρ^+ (du) ω Φ (ud)	892
135 140 548/958	(su) K^- $\bar{K}^0 (s\bar{d})$	(su) K^{*-} $\bar{K}^{*0} (s\bar{d})$	770 782 / 1020
494/498			892

BARYONS

	$J^P = 1/2^+$ $(\uparrow \downarrow \uparrow)$	$J^P = 3/2^+$ $(\uparrow \uparrow \uparrow)$	
939 / 938	$(udd) n$ $p (uud)$ Σ^- Σ^0 Σ^+ (dds) Λ $\Xi^- (dss)$ $\Xi^0 (uss)$	$(ddd) \Delta^-$ $\Delta^0 (udd)$ $\Delta^+ (uud)$ $\Delta^{++} (uuu)$ $\Sigma^{*-} (dds)$ $\Sigma^*0 (uds)$ $\Sigma^{*+} (uus)$ $\Xi^{*-} (dss)$ $\Xi^0 (uss)$	1232
1197/1193/1189			1385
1116			1532
1321 / 1315			1672
	OCTET	DECU $\Omega^- (sss)$ PLET	

Diagram illustrating the quark model for baryons. The left side shows the Octet (8 particles) and the right side shows the Decu and Plet (10 particles). The vertical axis represents the total spin quantum number S , ranging from -3 to +1. The horizontal axis represents the isospin quantum number I_3 , ranging from -3 to +3. The diagonal axis represents the charge quantum number Q .

Quark Masses

$$1. \quad M_{\Sigma^-} + M_{\Sigma^0} - 2 M_{\Sigma^+} = 3m_d - 3m_u = 11 \text{ MeV}$$

$$(dds) + (uds) - 2(uus) = 3d - 3u$$

$$m_d - m_u \sim \text{MeV}$$

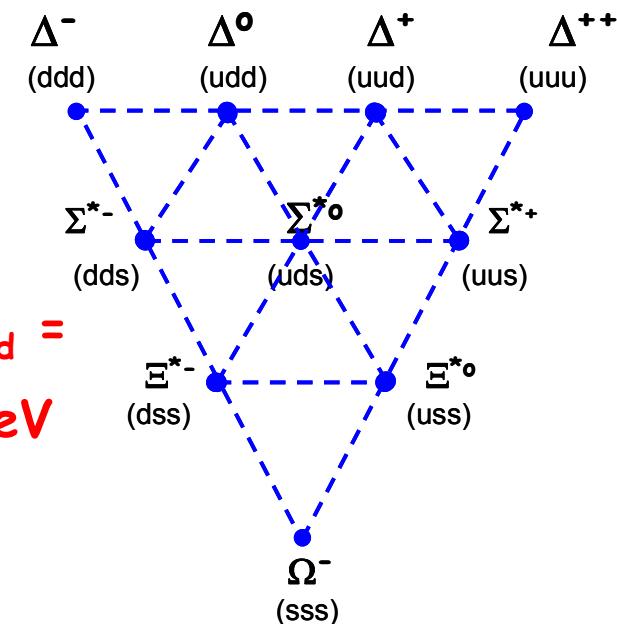
2. decuplet equal spacing rule:

$$M_\Omega - M_{\Xi^*} = M_{\Xi^*} - M_{\Sigma^*} = M_{\Sigma^*} - M_\Delta = m_s - m_d = \\ 142 \sim 145 \sim 153 \sim 145 \text{ MeV}$$

3. Gell-Mann-Okubo mass relation:

$$3 M_\Lambda + M_\Sigma = 2 (M_N + M_\Xi)$$

$$3(uds) + (uds) = 2 [(uud) + (dss)]$$



Hadron Spectroscopy

$SU(N)_F \otimes O(3)$: N flavors \otimes spatial excitations

1. S hyperfine splitting: energy of spin flip:

	$\uparrow\uparrow$	$\uparrow\downarrow$	mass split / GeV
K^*	-	K	~ 0.4
	$\uparrow\uparrow\uparrow$	$\uparrow\downarrow\uparrow$	
Δ	-	N	~ 0.3
Σ^*	-	Σ	~ 0.2
			$\sim E^* - E$

$$M_{\Sigma^*} - M_\Sigma = M_{E^*} - M_E = M_\Delta - M_N$$

$$196 \quad \sim \quad 214 \quad \sim \quad 294 \quad \text{MeV}$$

2. L orbital momentum:  $J = L + S \sim M^2$

Regge theory:

3. N radial excitations: 

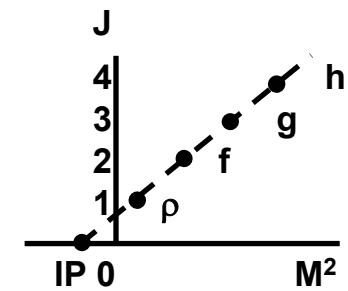
$$\rho' - \rho \quad dm \sim 0.5 \text{ GeV}$$

heavy quarks: non-relativistic potential: harmonic oscillator:

$$\Psi' - \Psi \quad dm \sim 0.6 \text{ GeV}$$

$$\Psi'' - \Psi'$$

$$Y' - Y$$



Spectroscopy of light mesons

$n^{2s+1}\ell_J$	J^{PC}	$ = 1$ $ud, \bar{u}d, \frac{1}{\sqrt{2}}(d\bar{d} - u\bar{u})$	$ = \frac{1}{2}$ $u\bar{s}, d\bar{s}; \bar{d}s, -\bar{u}s$	$ = 0$ f'	$ = 0$ f	θ_{quad} [°]	θ_{lin} [°]
1^1S_0	0^{-+}	π	K	η	$\eta'(958)$	-11.5	-24.6
1^3S_1	1^{--}	$\rho(770)$	$K^*(892)$	$\phi(1020)$	$\omega(782)$	38.7	36.0
1^1P_1	1^{+-}	$b_1(1235)$	K_{1B}^\dagger	$h_1(1380)$	$h_1(1170)$		
1^3P_0	0^{++}	$a_0(1450)$	$K_0^*(1430)$	$f_0(1710)$	$f_0(1370)$		
1^3P_1	1^{++}	$a_1(1260)$	K_{1A}^\dagger	$f_1(1420)$	$f_1(1285)$		
1^3P_2	2^{++}	$a_2(1320)$	$K_2^*(1430)$	$f'_2(1525)$	$f_2(1270)$	29.6	28.0
1^1D_2	2^{-+}	$\pi_2(1670)$	$K_2(1770)^\dagger$	$\eta_2(1870)$	$\eta_2(1645)$		
1^3D_1	1^{--}	$\rho(1700)$	$K^*(1680)^\ddagger$		$\omega(1650)$		
1^3D_2	2^{--}		$K_2(1820)^\ddagger$				
1^3D_3	3^{--}	$\rho_3(1690)$	$K_3^*(1780)$	$\phi_3(1850)$	$\omega_3(1670)$	32.0	31.0
1^3F_4	4^{++}	$a_4(2040)$	$K_4^*(2045)$		$f_4(2050)$		
1^3G_5	5^{--}	$\rho_5(2350)$					
1^3H_6	6^{++}	$a_6(2450)$			$f_6(2510)$		
2^1S_0	0^{-+}	$\pi(1300)$	$K(1460)$	$\eta(1475)$	$\eta(1295)$	-22.4	-22.6
2^3S_1	1^{--}	$\rho(1450)$	$K^*(1410)^\ddagger$	$\phi(1680)$	$\omega(1420)$		

Heavy Quarks

flavor =

property to be up, down, strange, ... quark

N flavors

SU(N) flavor space =

unitary N-dim. complex vector space

from SU(3) to SU(N):



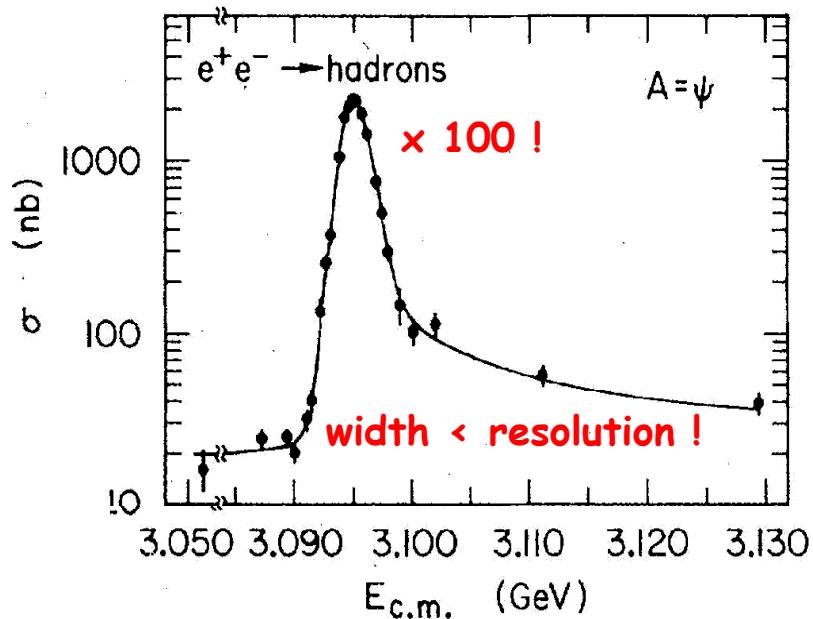
B. Richter

Charm

discovered
1974



S. Ting



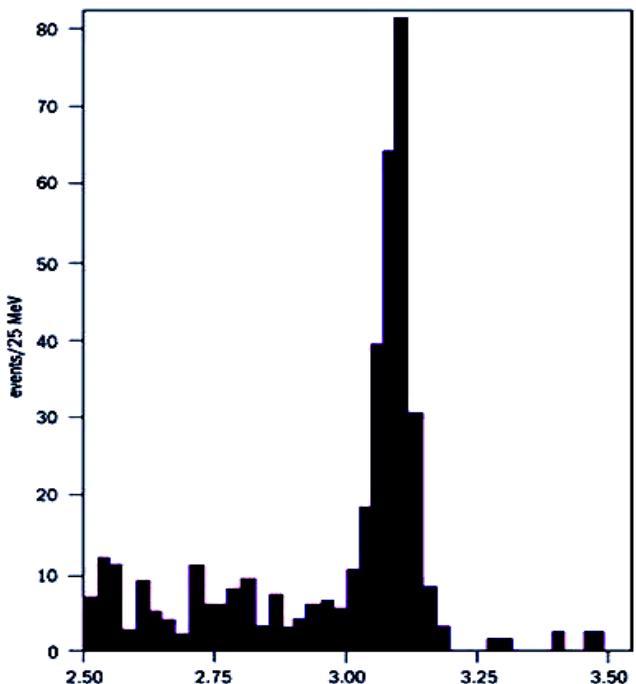
$e^+e^- \rightarrow \Psi \rightarrow \text{hadrons}$
3.1 GeV

MARK-I @ SPEAR
Stanford Linear Accelerator Lab.

Nobel

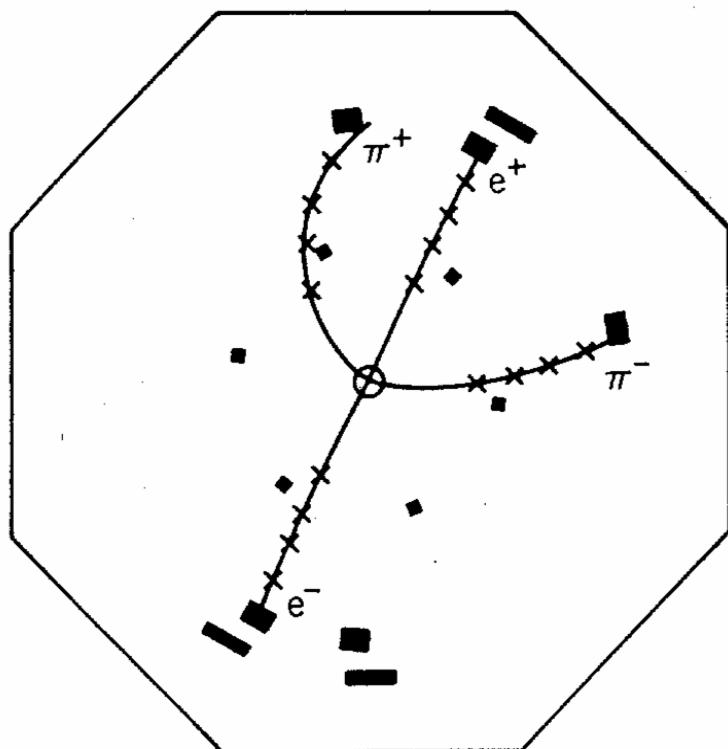


Prize
1976



$p\text{Be} \rightarrow JX \rightarrow e^+e^-X$
28.5 GeV
double arm spectrometer @ AGS
Brookhaven National Lab.

Charmonium

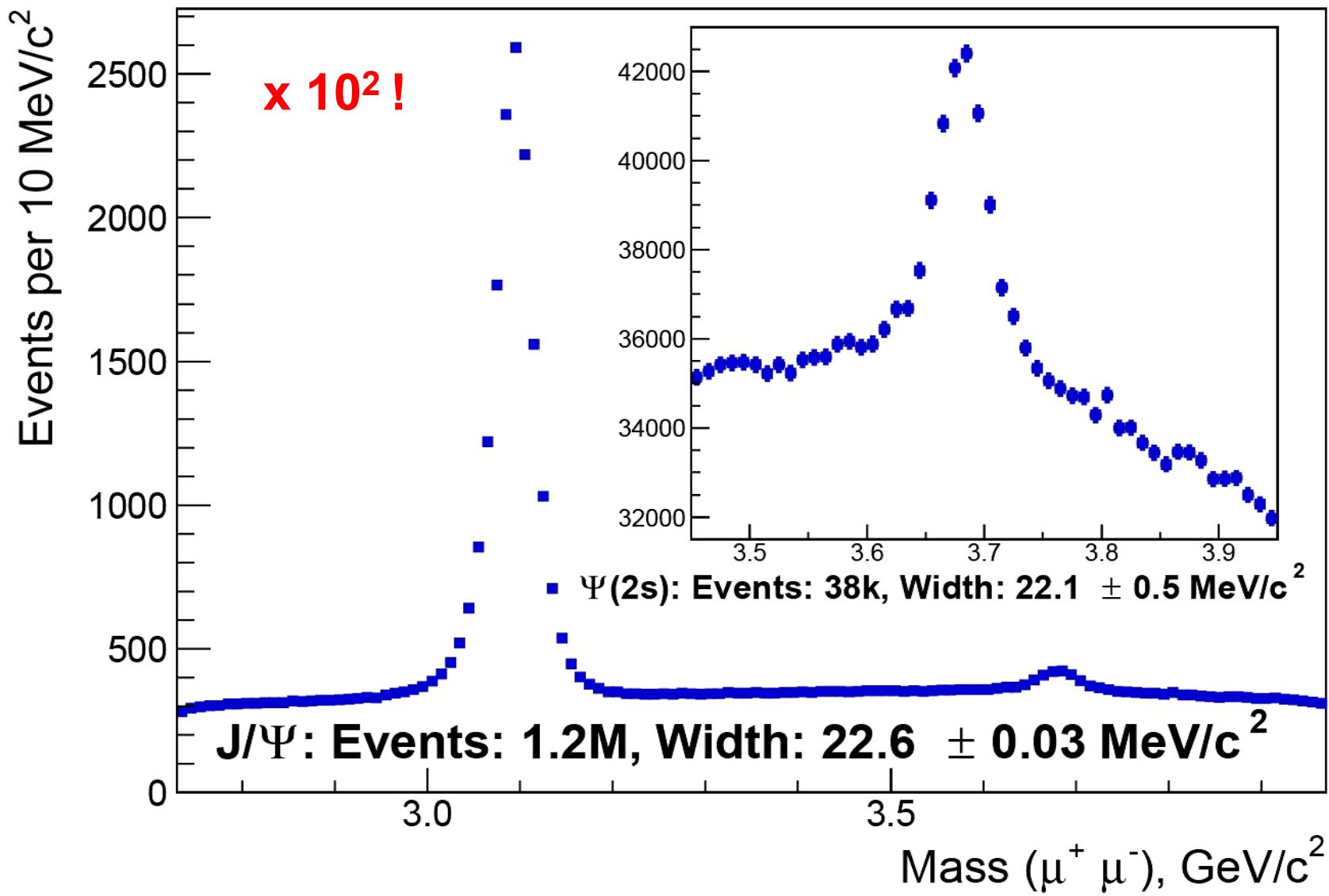


$$e^+e^- \rightarrow \Psi' \rightarrow \Psi \pi^+ \pi^- \rightarrow e^+ e^-$$

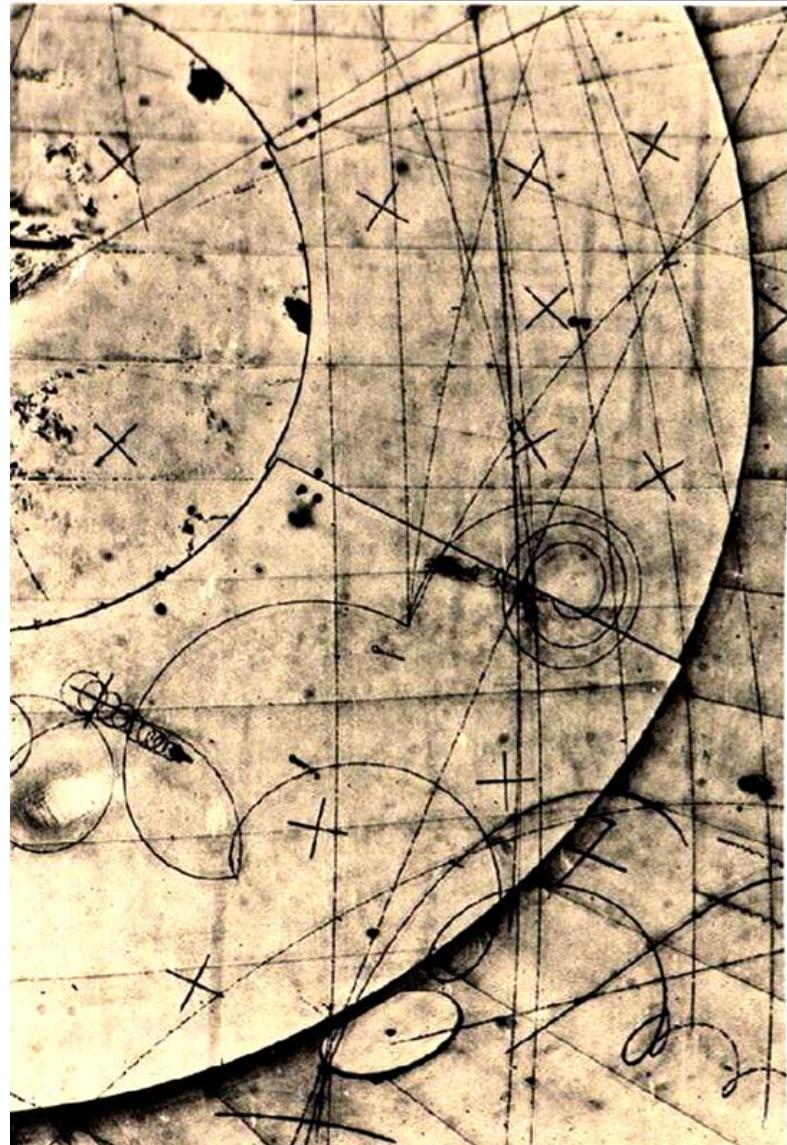
$$\begin{aligned} m_\Psi &= 3.1 \text{ GeV} \\ m_c &= 1.5 \text{ GeV} \end{aligned}$$

MARK-I @ SPEAR 1975
Stanford Linear Accelerator Lab.

Charm

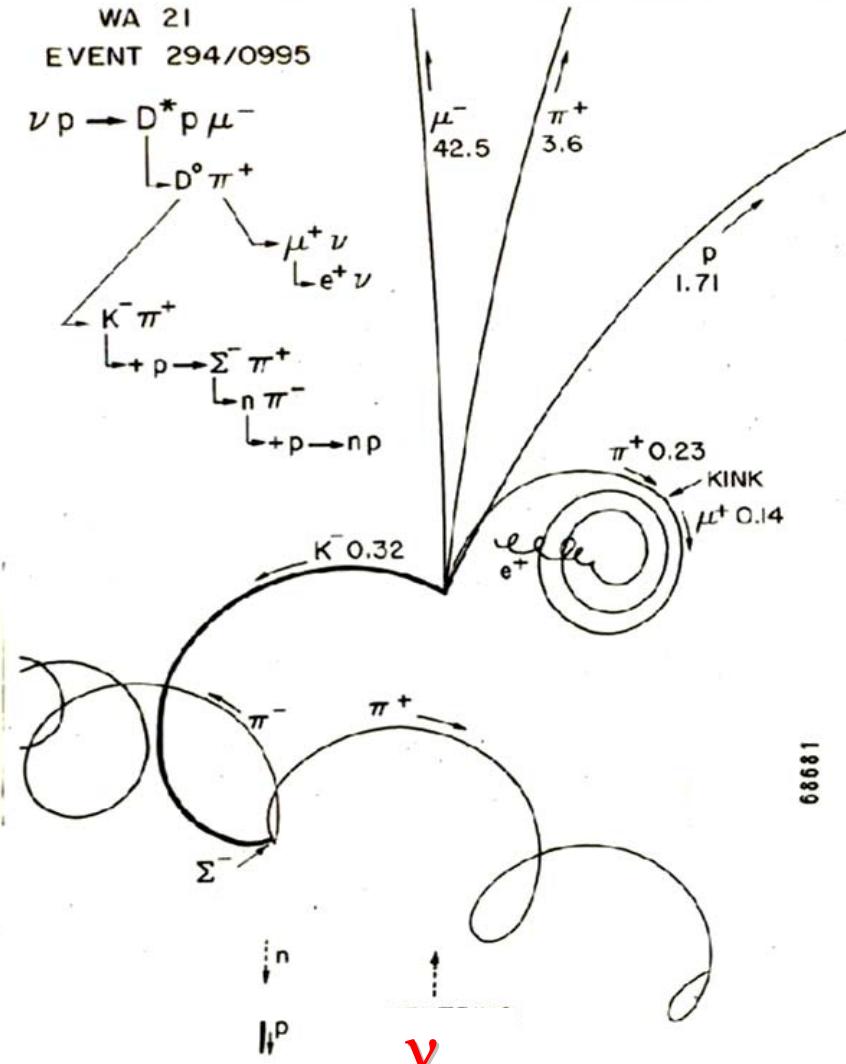
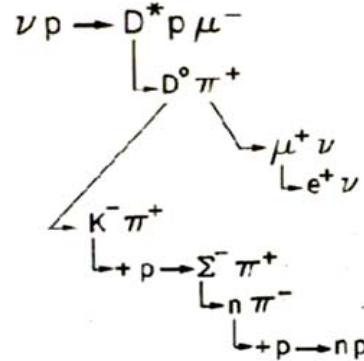


Open Charm Decay



AACHEN-BONN-CERN-MUNICH-OXFORD COLLABORATION

WA 21
EVENT 294/0995



production flavor violating = weak

Open Charm Decay

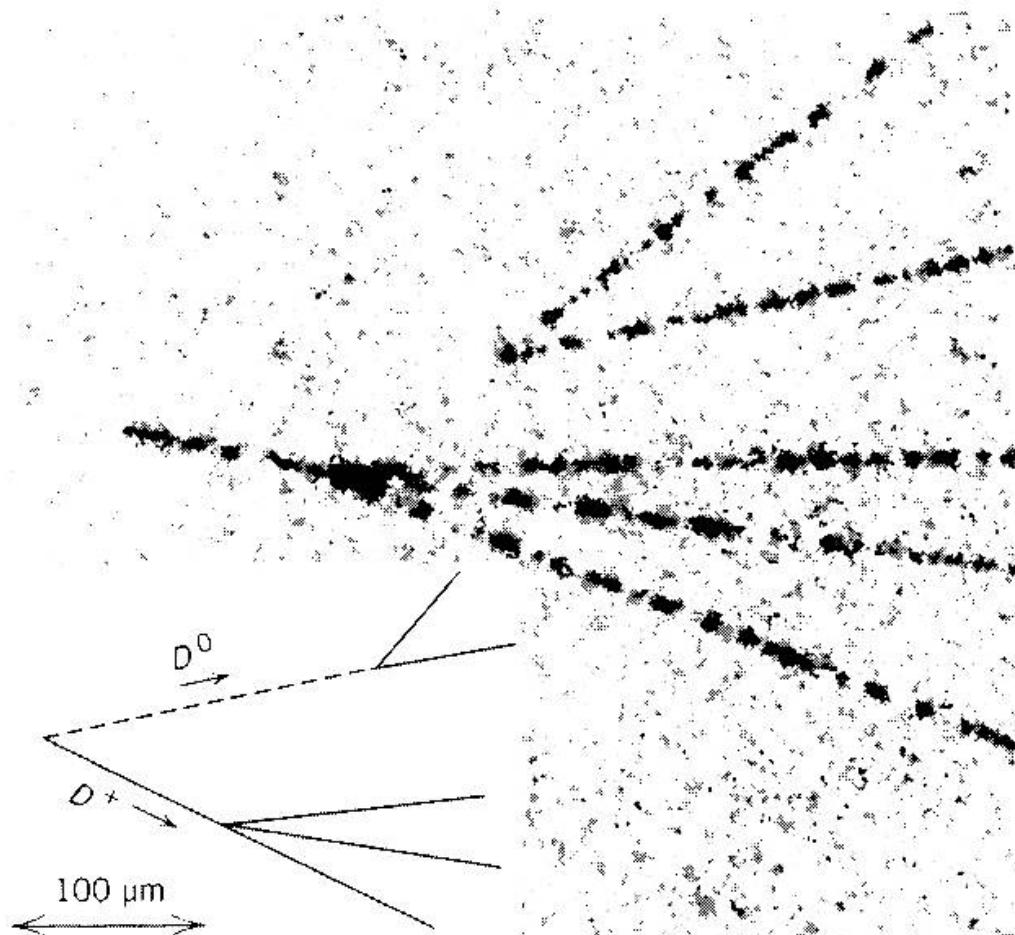


FIGURE 18-11 Example of D meson decays observed in a bubble chamber.

From K. Abe et al., *Phys. Rev. Lett.* **48**, 1526 (1982).



$$\tau = 0.4 \text{ ps}$$

$$c\tau = 123 \mu\text{m}$$

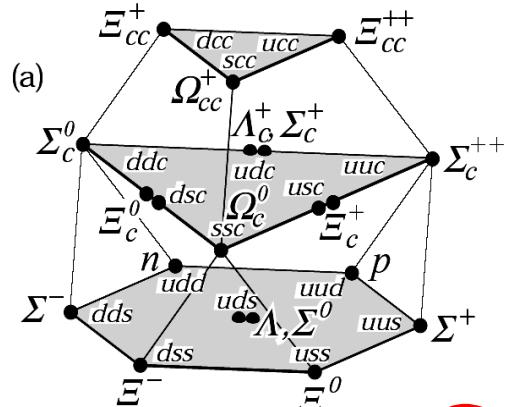


$$\tau = 1.0 \text{ ps}$$

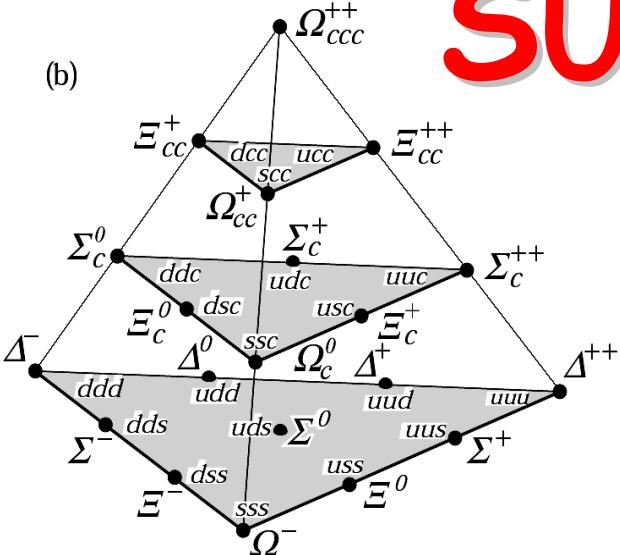
$$c\tau = 315 \mu\text{m}$$

Charm Spectroscopy

$J=1/2$
 $\uparrow \downarrow \uparrow$

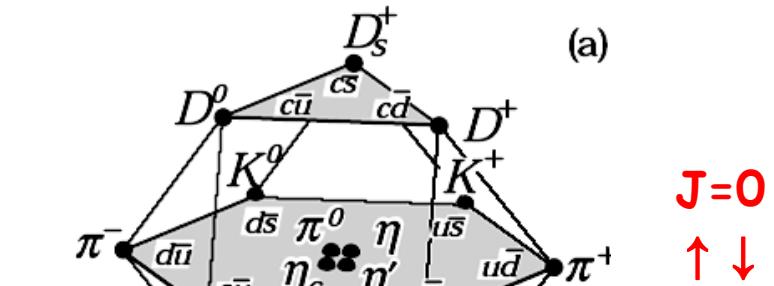


$J=3/2$
 $\uparrow \uparrow \uparrow$

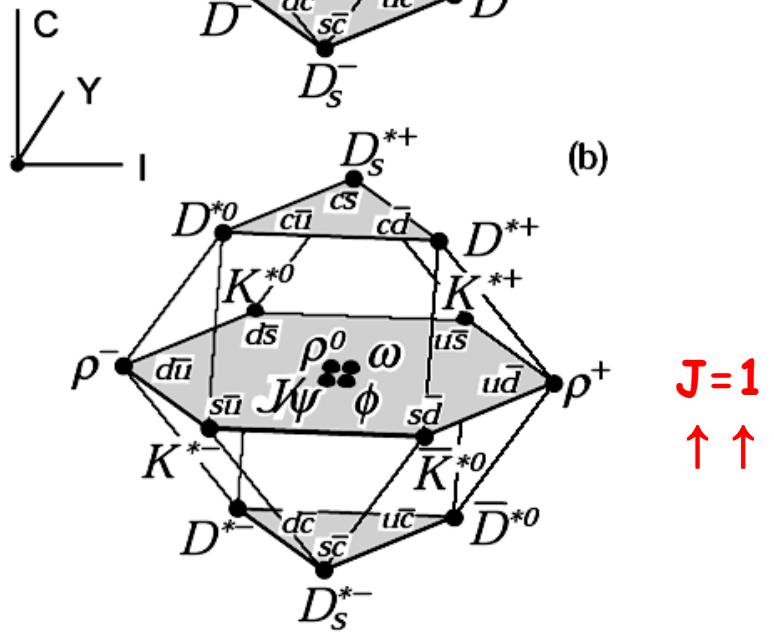


Baryons

SU(4)



$J=0$
 $\uparrow \downarrow$



$J=1$
 $\uparrow \uparrow$

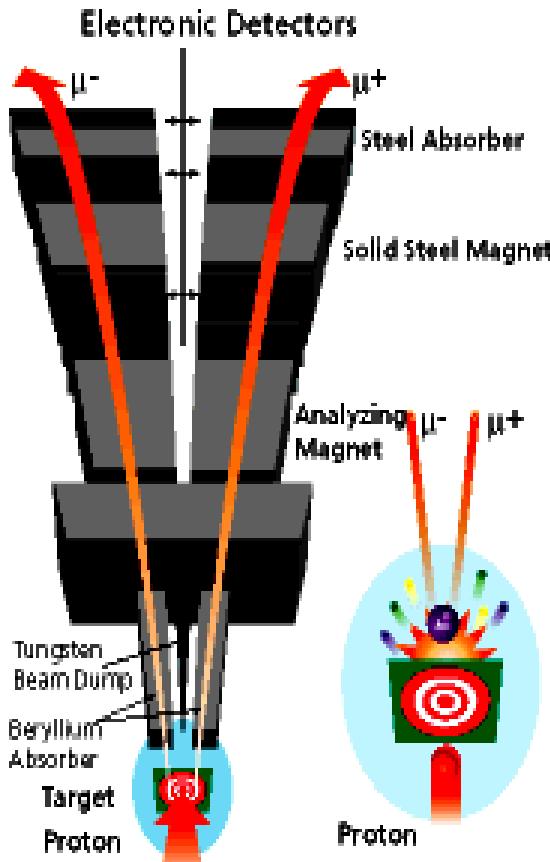
Mesons

discovered 1977 by

L. Lederman et al.

$p + N \rightarrow \mu^+ \mu^- + X$

Fermilab, USA

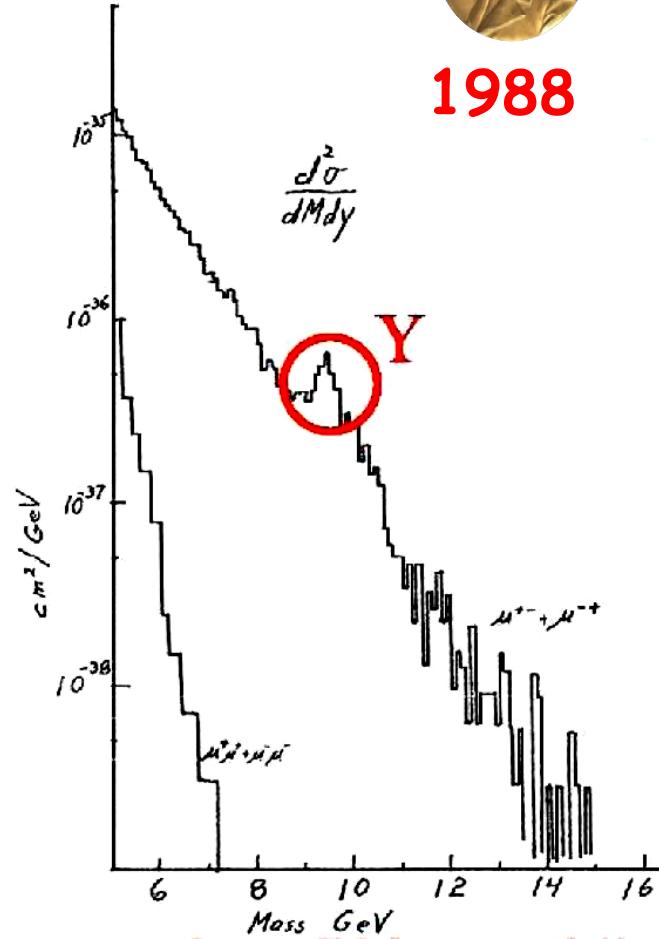


Beauty

Nobel prize



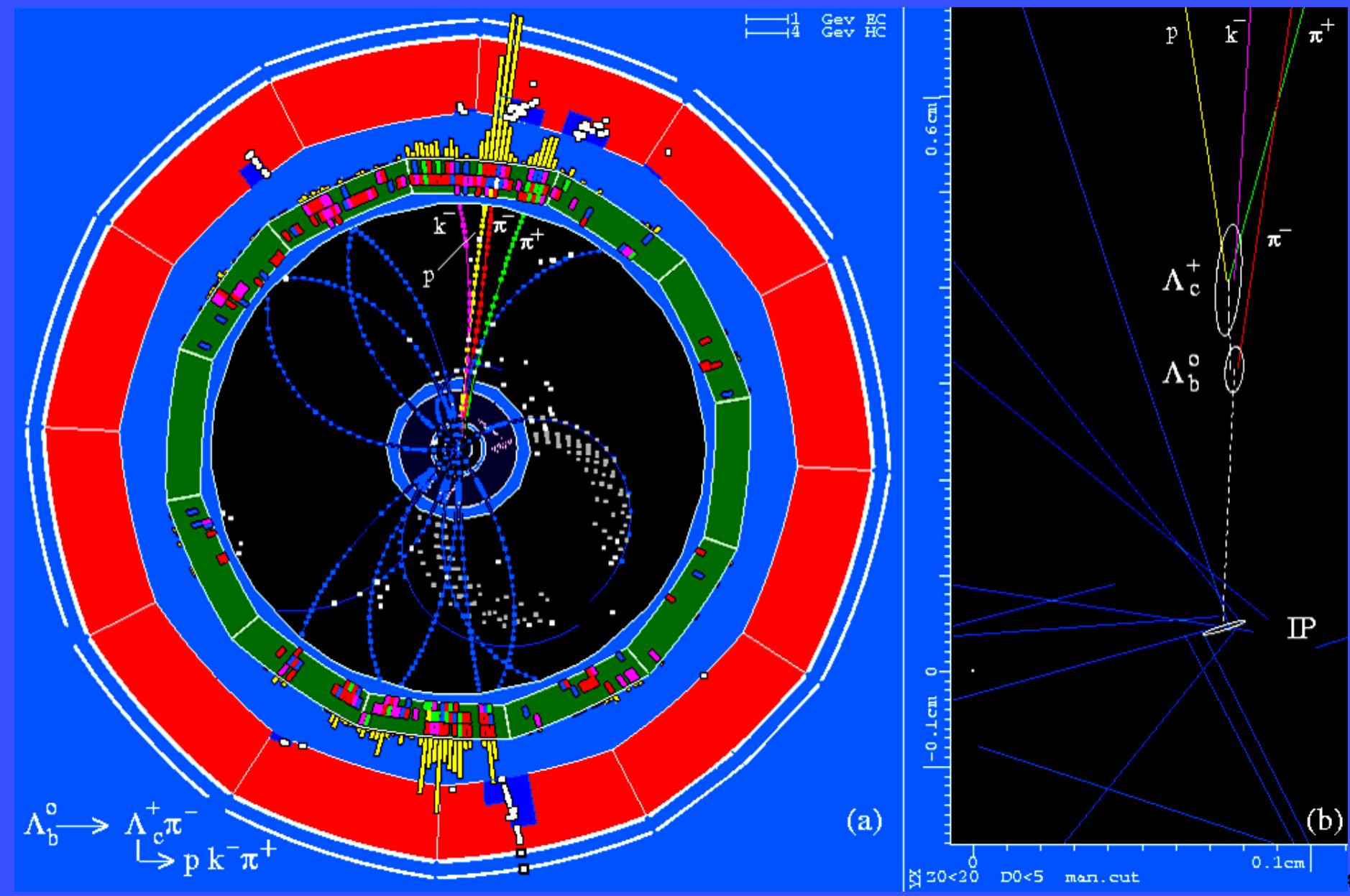
1988



L.Lederman



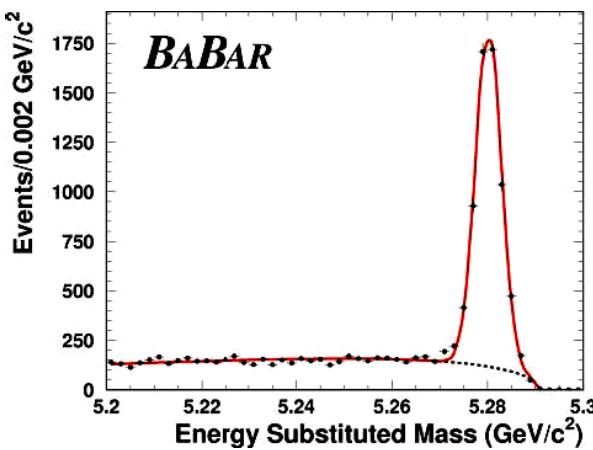
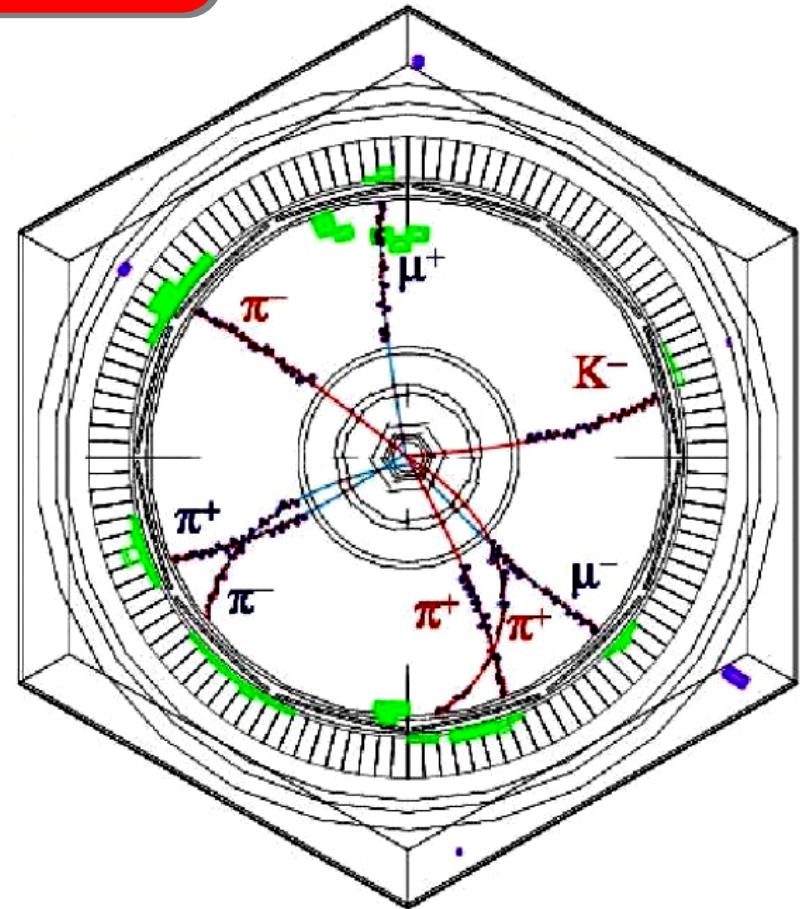
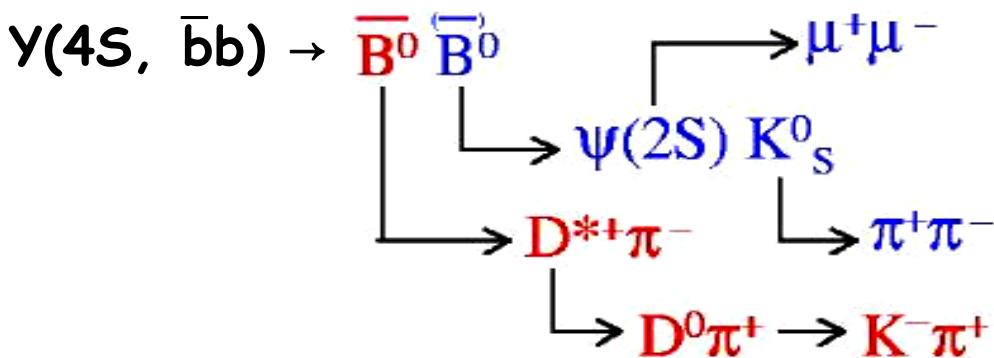
Heavy Quark Decays



Bottom quark

BaBar @ PEP,

SLAC: $e^+e^- \rightarrow$

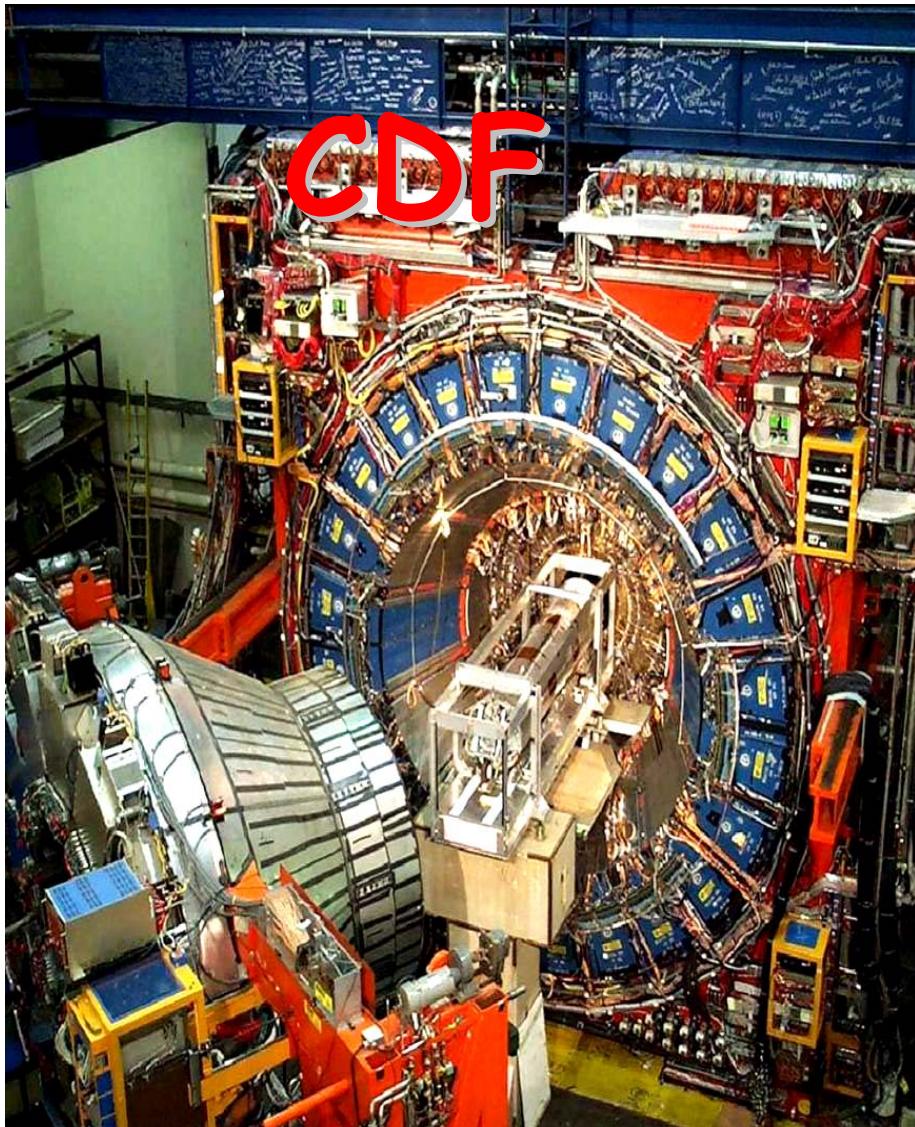


BaBar @ PEP, SLAC, Stanford, USA
Belle @ KEK, Tsukuba, Japan:

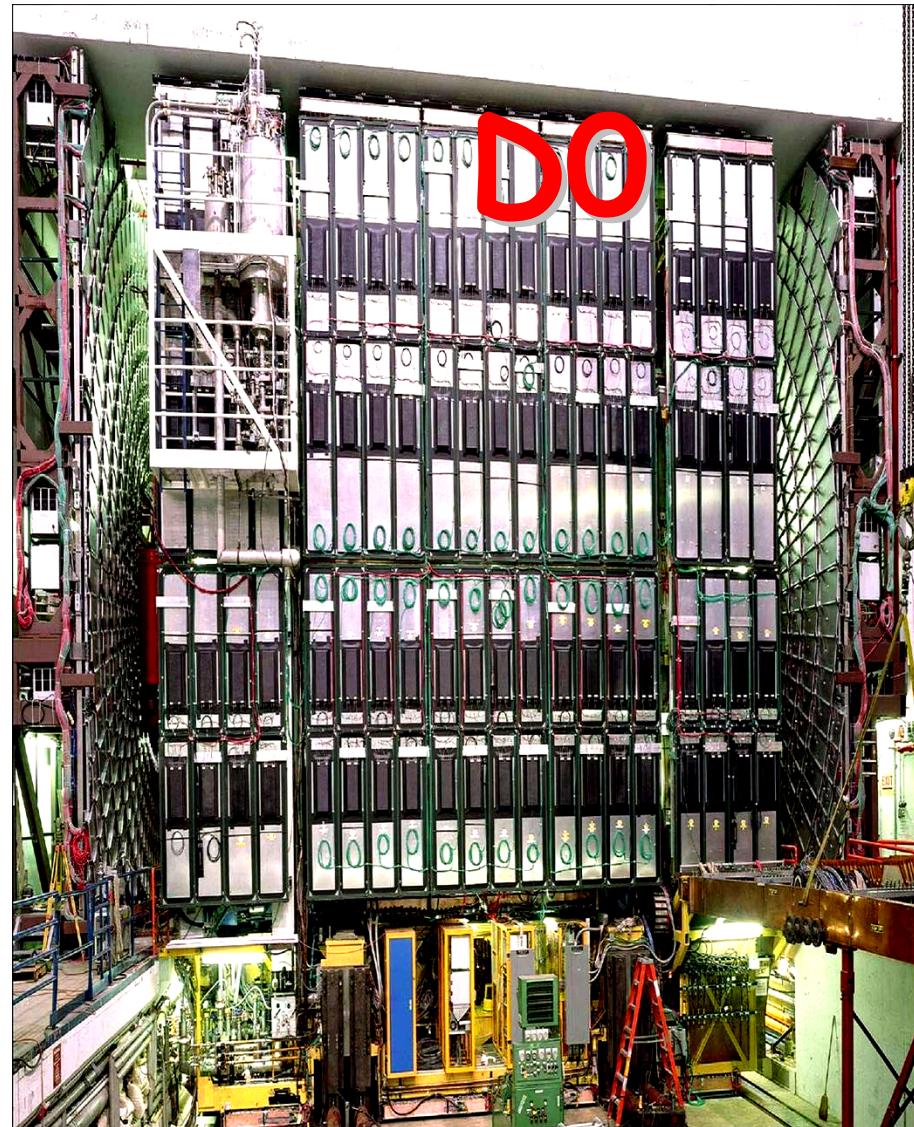
Belle: 500 million $\bar{B}B$ pairs in 500 fb^{-1} in 6 years

Tevatron

CDF



DO

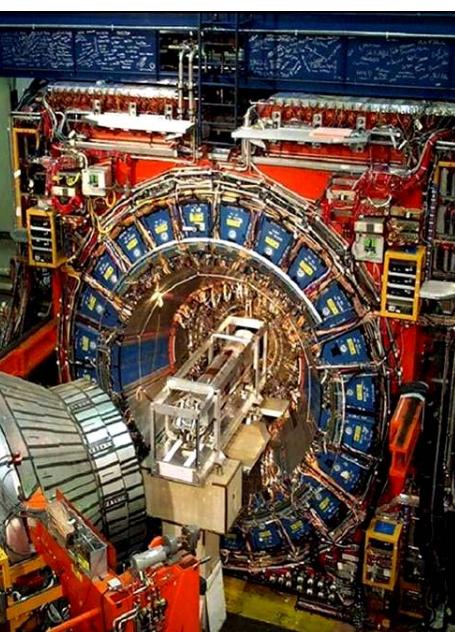


Fermi National Lab., Chicago: $\bar{p} p$ @ 1x1 TeV

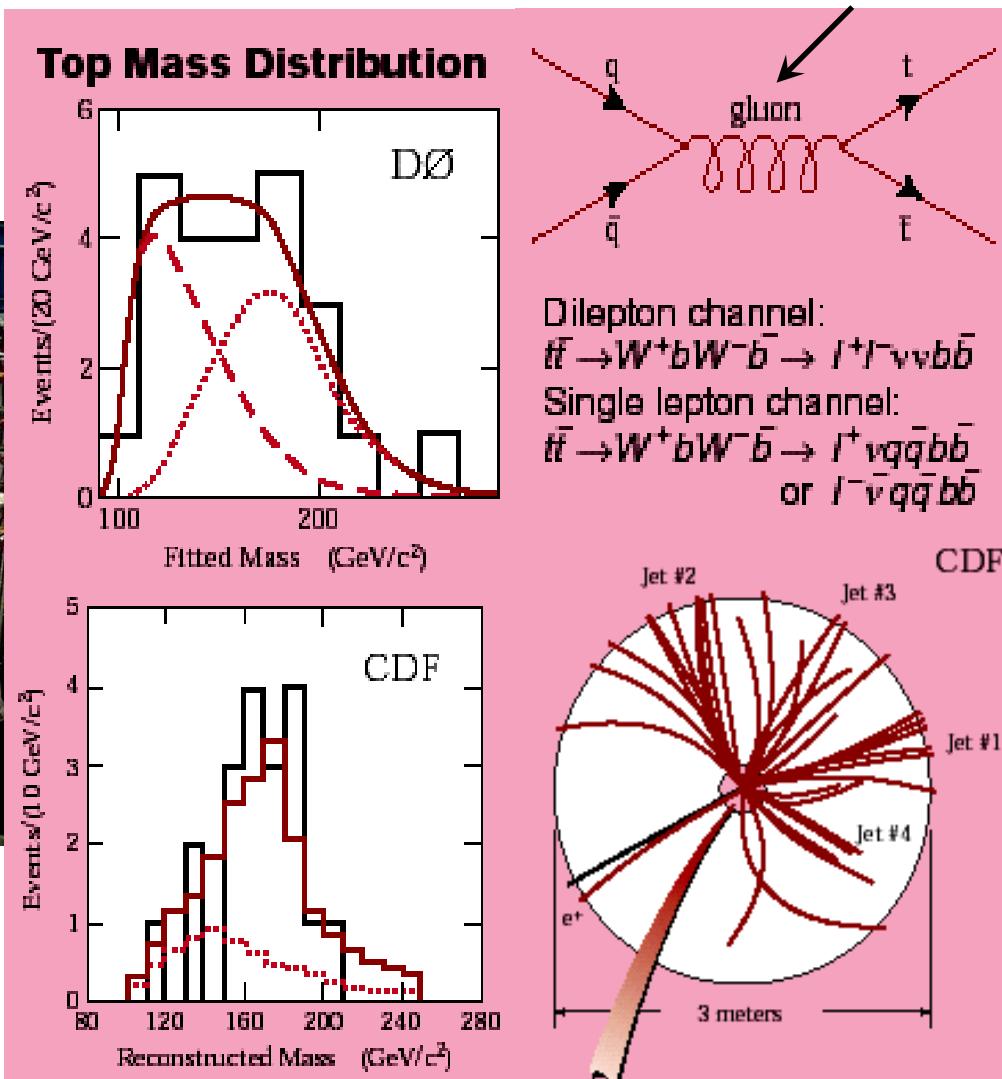
Top

discovered
1994

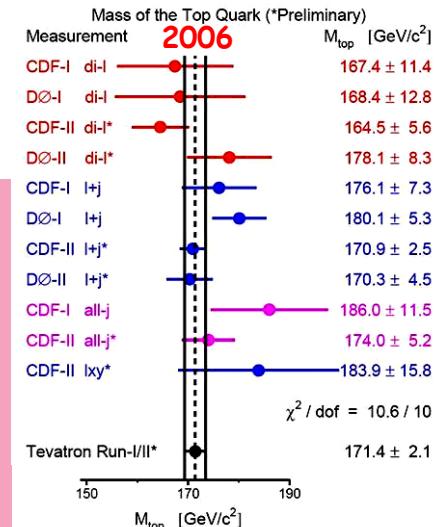
CDF + DØ



Tevatron
p p @ 2x1 TeV
Fermi Natl. Lab.
USA



quantum of
strong int.



$m_t = 172 \pm 2 \text{ GeV} !$

prediction
from
radiative corrections
to Z mass + width:
 $172 \pm 4 \text{ GeV} !$

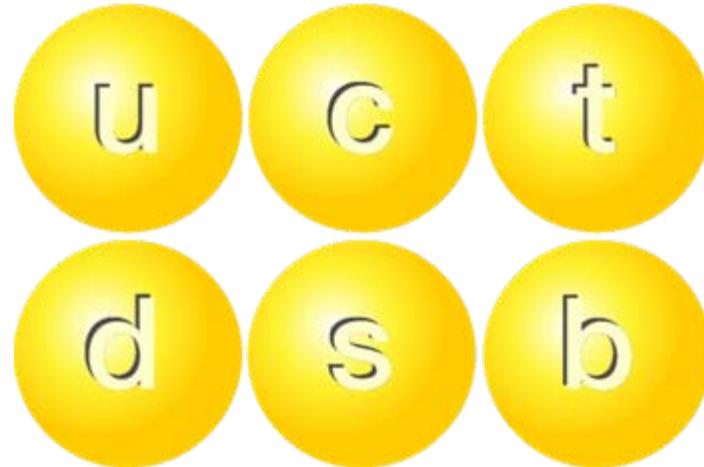
weak decay
within 10^{-26} s -
faster than
hadronic time scale

too fast
to find a partner -

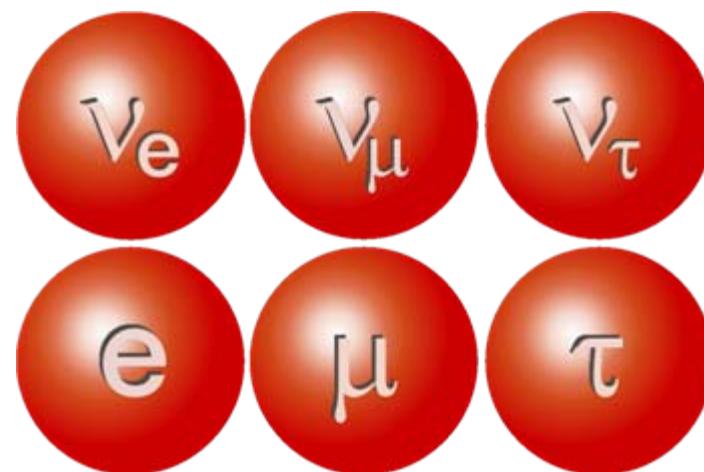
top quark
never becomes
a dressed hadron

The Building Blocks

Quarks



Leptons



Fermions

QUARKS			Q	I_3	B	L
u	c	t	+2/3	+1/2		
d	s	b	-1/3	-1/2	1/3	0
LEPTONS						
e ⁻	μ^-	τ^-	-1	-1/2	0	1
ν_e	ν_μ	ν_τ	0	+1/2		

Spin
 $J=1/2$

Questions:

- why no free quarks ?
=> confinement, QCD
- fractional charge 1/3 ?

Masses

u	c	t	e ⁻	μ^-	τ^-
2-5 MeV	1.25 GeV	172 GeV	511 keV	106 MeV	1.78 GeV
d	s	b	ν_e	ν_μ	ν_τ
5-8 MeV	120 MeV	4.25 GeV	<2 eV	<190 keV	<18.2 MeV

The Building Blocks

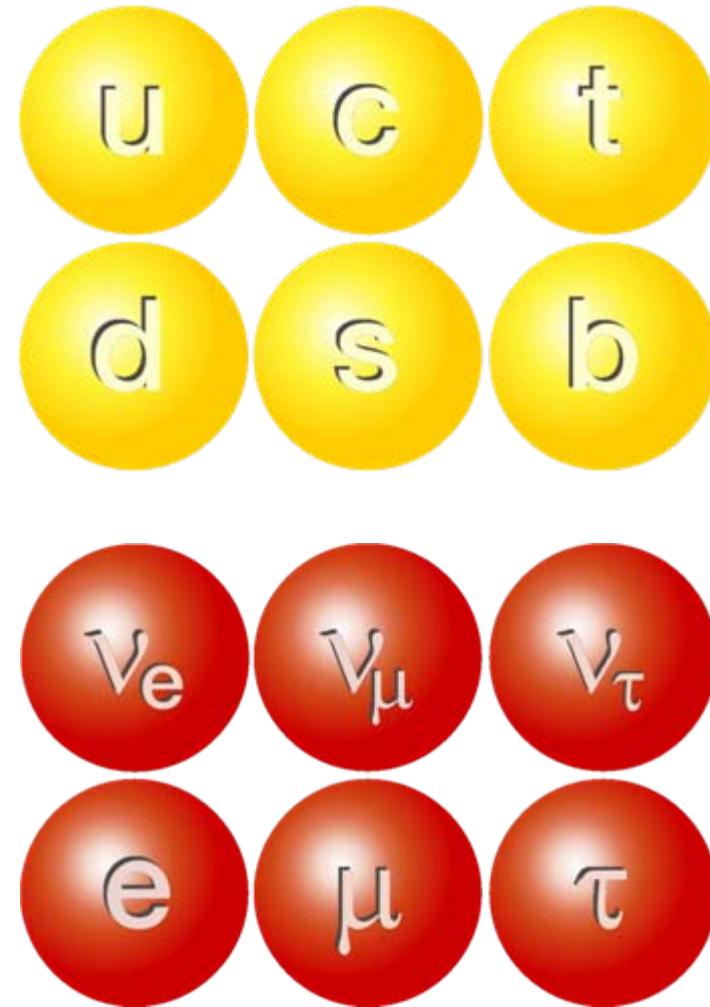
Great scheme,

BUT:

all 3 symmetries

mysterious:

- up-down ((weak) isospin)
- lepton-quark
- 3 families

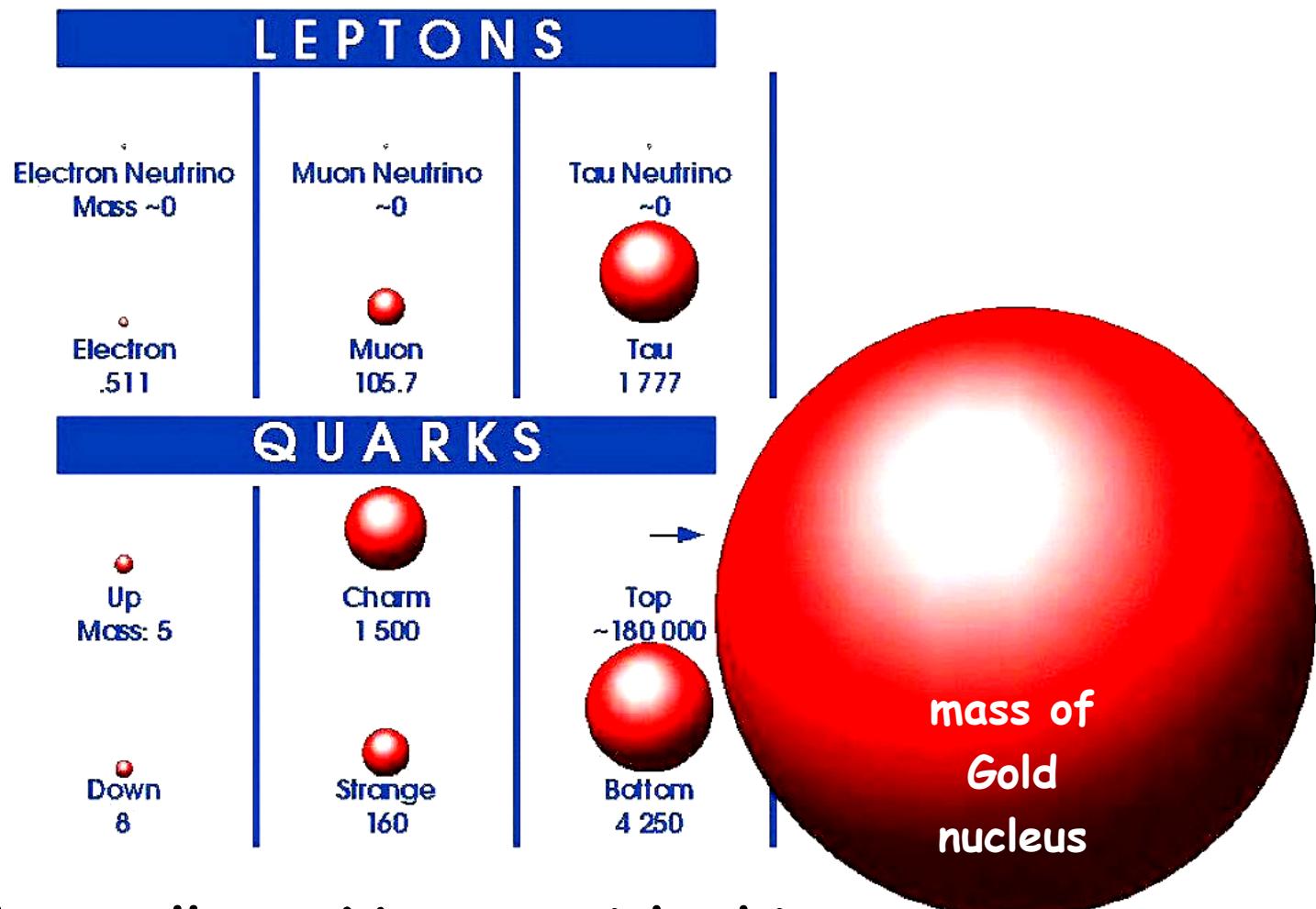


PARTICLES

FORCES

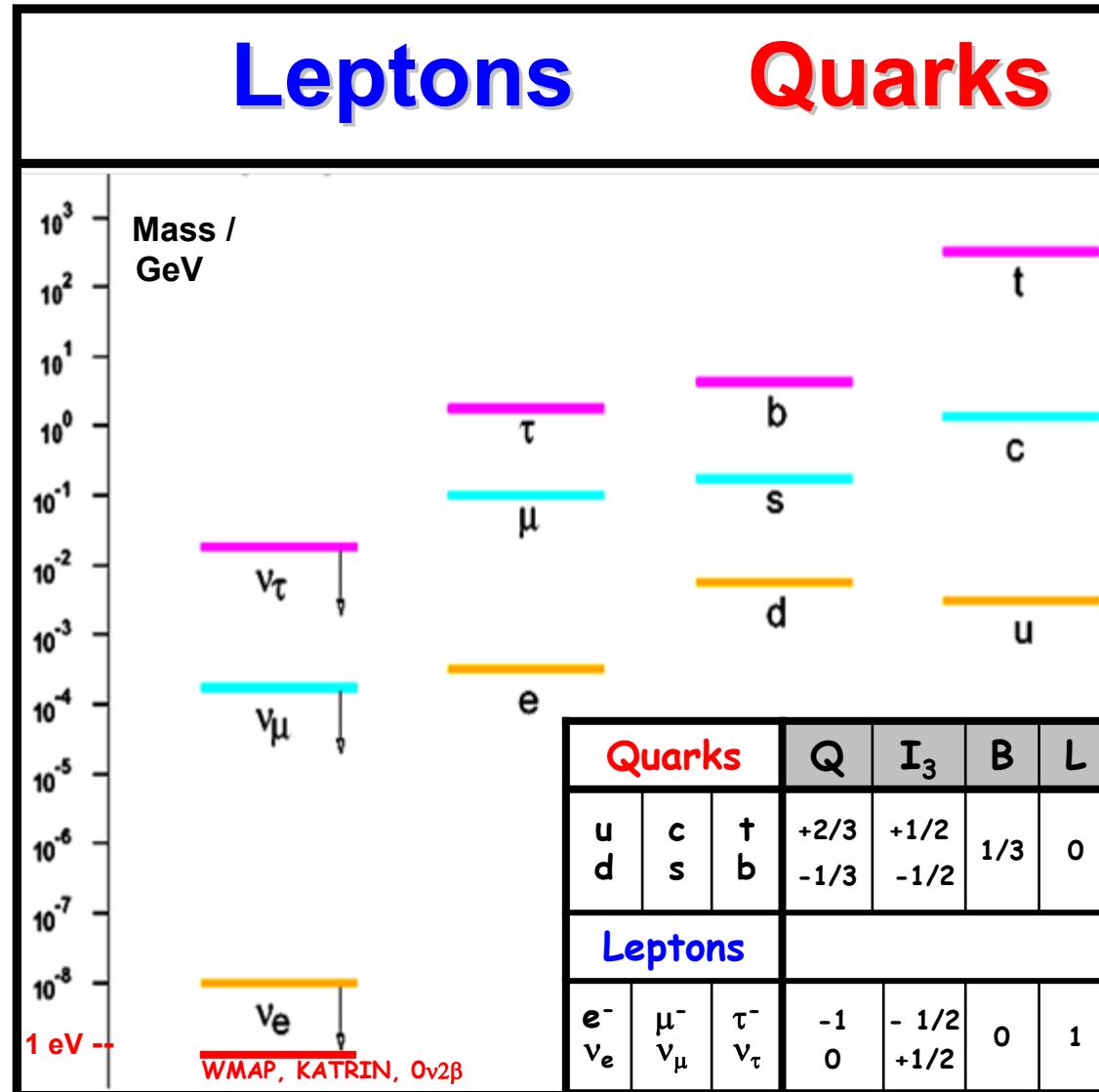
			Electro-Magnet.	Weak	Nuclear	Gravitation
	Charge Symmetry	Electric	Weak	Color	Mass	
Matter Particles		Fermions J=1/2				
Quarks	Up	u	c	+2/3	I _w , Y _w	r g b
	Down	d	s	-1/3		
Leptons	Electrons	e	μ	-1	I _w , Y _w	
	Neutrinos	ν _e	ν _μ	0		
Force Particles		Bosons J=1				
Photon	γ					
Weak Bosons	W ⁺ , Z ⁰ , W ⁻					
Gluons	8 g _{ij}					
Graviton (J=2)	G					

Fermion Mass Spectrum



What tells us Nature with this
new spectroscopy ?

Fermion Mass Spectrum

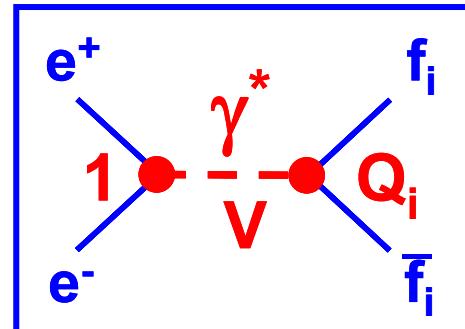
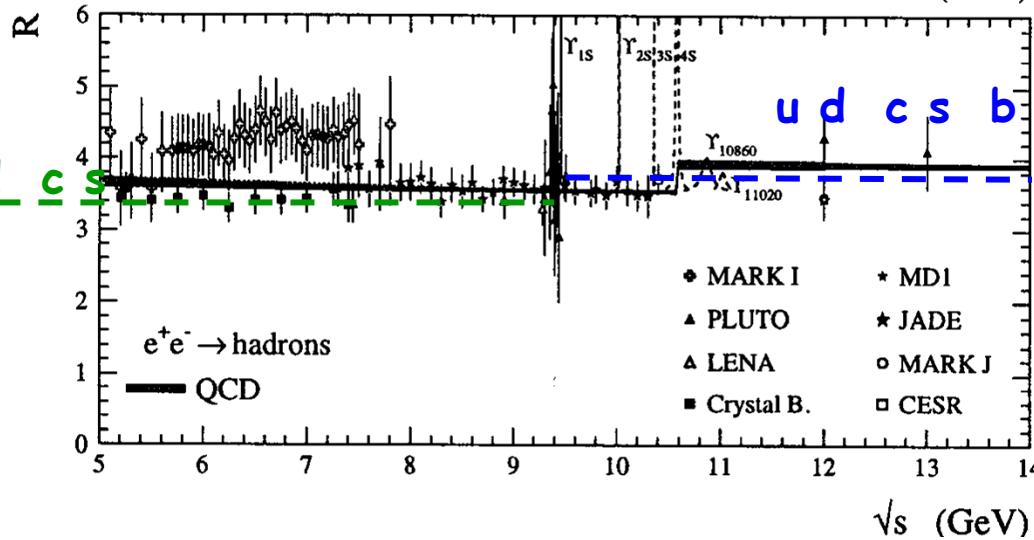
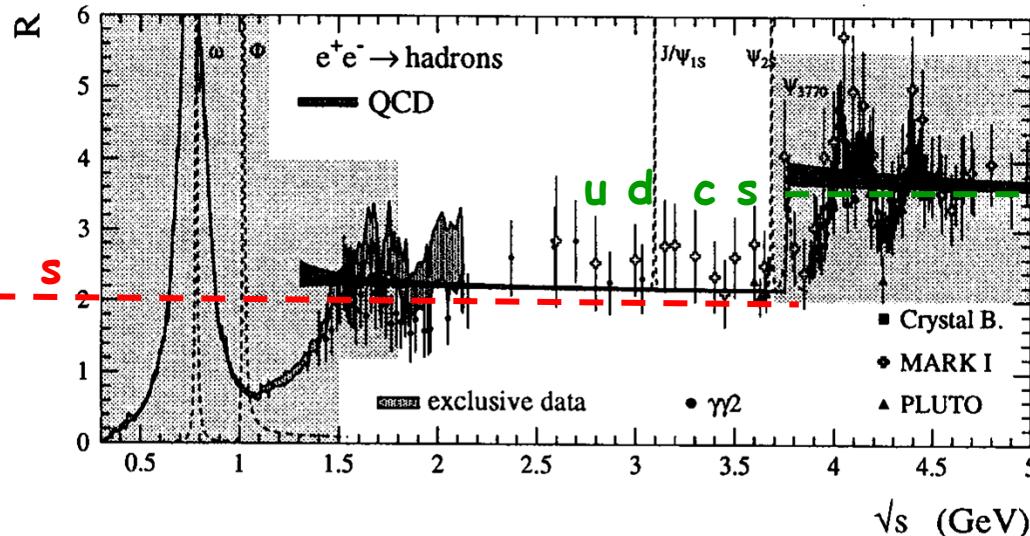


What
tells us
Nature
with this
spectrum

?

Hadron
spectroscopy
+
mixing

Quark Charge

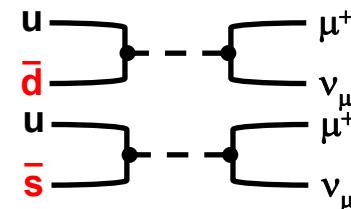


$$R = \frac{\sigma(e^+e^- \rightarrow q\bar{q} \rightarrow \text{hadrons})}{\sigma(e^+e^- \rightarrow \mu^+\mu^-)} = \sum_{\text{flavors}} Q_i^2 = \sum_{\text{colors}}$$

N_F	Q_i^2
3	$\left(\frac{2}{3}\right)^2 + \left(\frac{1}{3}\right)^2 + \left(\frac{1}{3}\right)^2 = \frac{6}{9} * 3 = \frac{6}{3} \quad \begin{pmatrix} u \\ d \\ s \end{pmatrix}$
4	$\left(\frac{2}{3}\right)^2 + \left(\frac{1}{3}\right)^2 + \left(\frac{2}{3}\right)^2 + \left(\frac{1}{3}\right)^2 = \frac{10}{9} * 3 = \frac{10}{3} \quad \begin{pmatrix} u \\ d \\ c \\ s \end{pmatrix}$
5	$\left(\frac{2}{3}\right)^2 + \left(\frac{1}{3}\right)^2 + \left(\frac{2}{3}\right)^2 + \left(\frac{1}{3}\right)^2 + \left(\frac{1}{3}\right)^2 = \frac{11}{9} * 3 = \frac{11}{3} \quad \begin{pmatrix} u \\ d \\ c \\ s \\ b \end{pmatrix}$

Quark Mixing

Decay	$\Delta M/\text{MeV}$	τ/ns
$\pi^+ \rightarrow \mu^+ \nu_\mu$	$139 - 105 = 34$	26
$K^+ \rightarrow \mu^+ \nu_\mu$	$494 - 105 = 389$	8



- coupling to d,s could be identical or zero, but different ?
- N.Cabbibo 1963:
universal couplings, but weak interaction
does not see eigenstates **d** and **s** of mass + strong interaction,
but mixed states **d'**, **s'** with mixing angle θ_C :

$$d' = d \cos \theta_C + s \sin \theta_C$$

$$s' = -d \sin \theta_C + s \cos \theta_C$$

with Cabbibo angle: $\sin \theta_C = 0.220 \pm 0.002$

- $\sin \theta_C$: mixing strange - normal world or probability of family change
- SU(2) doublet: **(u d')** , $\Gamma(K^+ \rightarrow \mu^+ \nu_\mu) / \Gamma(\pi^+ \rightarrow \mu^+ \nu_\mu) \sim \sin^2 \theta_C / \cos^2 \theta_C$

GIM Mechanism

- **Cabbibo theory:** weak neutral current

$$= (ud') (ud')^T = \bar{u}u + d'd' =$$

$$= \bar{u}u + \cos^2 \theta_C \bar{d}d + \sin^2 \theta_C \bar{s}s + \dots \quad \Delta S = 0$$

$$\sin \theta_C \cos \theta_C (\bar{s}d + \bar{d}s) \quad \Delta S = 1$$

- $\Delta S = 1$ flavor changing neutral current not observed

- 1970: Glashow, Iliopoulos, Maiani: 2. quark doublet

$$d' = d \cos \theta_C + s \sin \theta_C$$

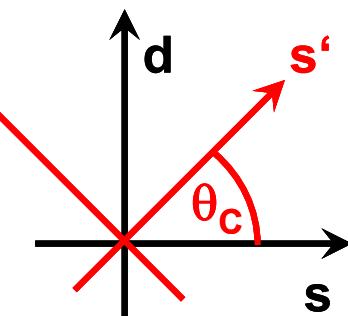
$$s' = -d \sin \theta_C + s \cos \theta_C$$

$$\text{neutral current} = \bar{u}u + d'd' + \bar{c}c + s's'$$

$$= \bar{u}u + d\bar{d} + \bar{c}\bar{c} + s\bar{s} \quad \Delta S = 0$$

Decay	Current	Γ_i / Γ
$K^+ \rightarrow \mu^+ \nu_\mu$	charged	60 %
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	neutral	$< 10^{-9}$
$K^0 \rightarrow l^+ l^-$	neutral	$< 10^{-8}$

$$\begin{pmatrix} u \\ d' \end{pmatrix} \quad \begin{pmatrix} c \\ s' \end{pmatrix}$$



- 1974: Charm discovered by S.Ting (SLAC) and B.Richter (BNL)

- 1973: Kobayashi + Maskawa postulate 3rd quark family mixing matrix with 1 complex phase \rightarrow CP violation

Quark Mixing

1973: Cabibbo-Kobayashi-Maskawa postulate 3rd quark family

CKM matrix V_{CKM} transforms

mass eigen states (d s b) to
weak eigen states (d's'b')

$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix} \equiv \hat{V}_{CKM} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

represent by 3 Euler angles in (dsb) space

$$c_{ij} = \cos \theta_{ij}, \quad s_{ij} = \sin \theta_{ij} :$$

$$\hat{V}_{CKM} = \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \cdot \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \cdot \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & e^{i\delta} \end{pmatrix} \cdot \begin{pmatrix} c_{13} & 0 & s_{13} \\ 0 & 1 & 0 \\ -s_{13} & 0 & c_{13} \end{pmatrix}$$

+ complex phase δ (allows CP violation)

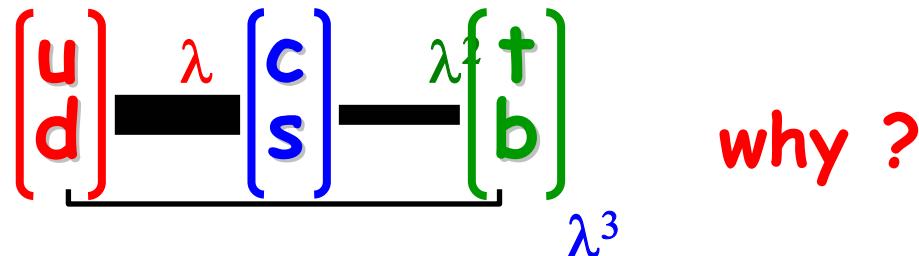
$$\hat{V}_{CKM} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -s_{23}c_{12} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

Quark Mixing

Experiment:

$$\begin{pmatrix} 0.9741 \text{ to } 0.9756 & 0.219 \text{ to } 0.226 & 0.0025 \text{ to } 0.0048 \\ 0.219 \text{ to } 0.226 & \underline{0.9732 \text{ to } 0.9748} & \underline{0.038 \text{ to } 0.044} \\ 0.004 \text{ to } 0.014 & 0.037 \text{ to } 0.044 & \underline{0.9990 \text{ to } 0.9993} \end{pmatrix}$$

hierarchic suppression of family change :



Wolfenstein parameterization :

$$V_{CKM} = \begin{pmatrix} 1 & -\lambda^2/2 & A\lambda^3(\rho-i\eta) \\ -\lambda & 1 & A\lambda^2 \\ A\lambda^3(1-\rho-i\eta) & -A\lambda^2 & 1 \end{pmatrix} = \begin{pmatrix} \text{gray} & \text{red} & \cdot \\ \text{red} & \text{gray} & \text{green} \\ \cdot & \text{green} & \text{gray} \end{pmatrix}$$

mixing angles:

$s_{12} = V_{us}$	$= \lambda$	$= 0.223 \pm 0.003$	Cabbibo
$s_{23} = V_{cb}$	$= A\lambda^2$	$= 0.041 \pm 0.002$	
$s_{13} = V_{ub} e^{i\delta}$	$= A\lambda^3 (\dots)$	$= 0.004 \pm 0.001 \pm 0.001$	
	A	$= 0.83 \pm 0.02$	

3. C+P Symmetries

C Parity

Dirac, QFT:

matter-antimatter: each particle has antiparticle !

C operator

inverts all charge like additive quantum numbers:

C

L,B → -L, -B

lepton + baryon nr.

F → -F

flavor: F = u, d, s, c, b, t

y,Q → -y, -Q

((weak) hyper) charge

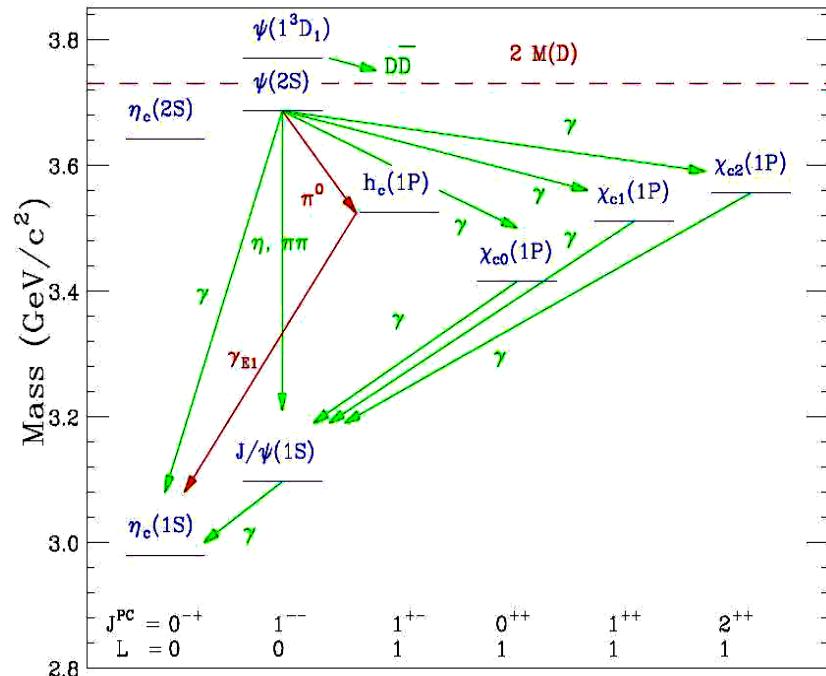
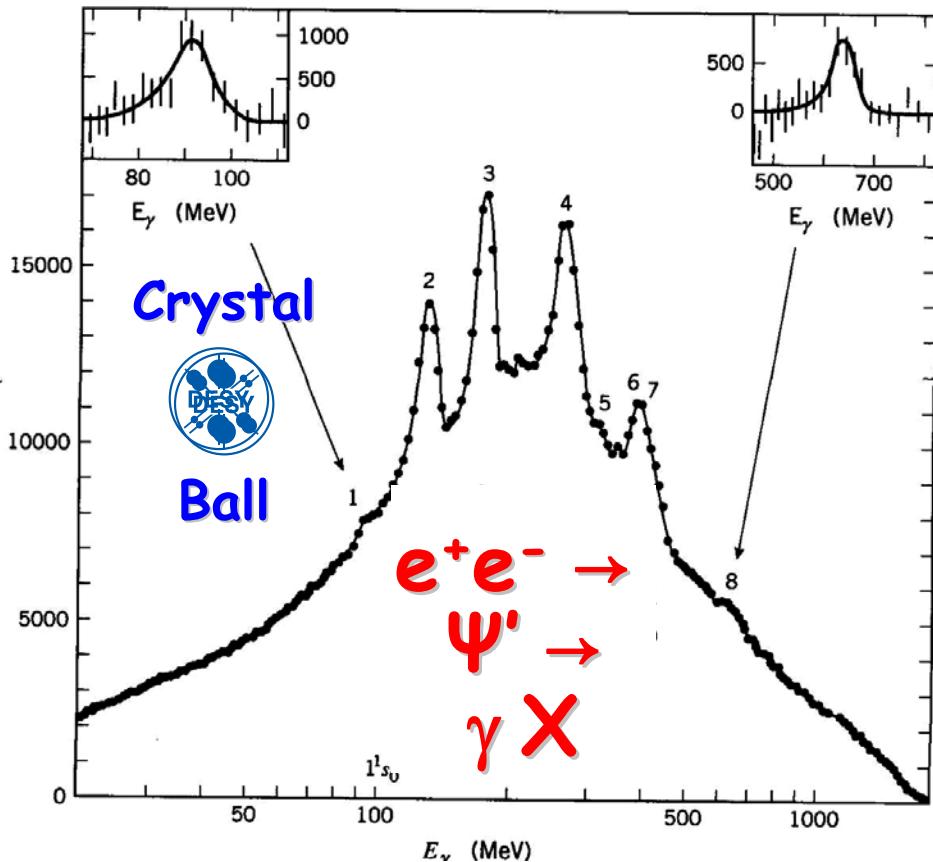
E,B → -E, -B

electric + magnetic field

$$\Rightarrow \textcolor{blue}{C} |\gamma\rangle = -|\gamma\rangle$$

C: multiplicative, not charge like

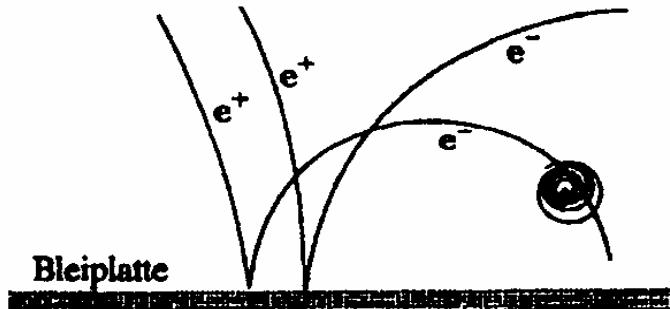
Charmonium Spectrum



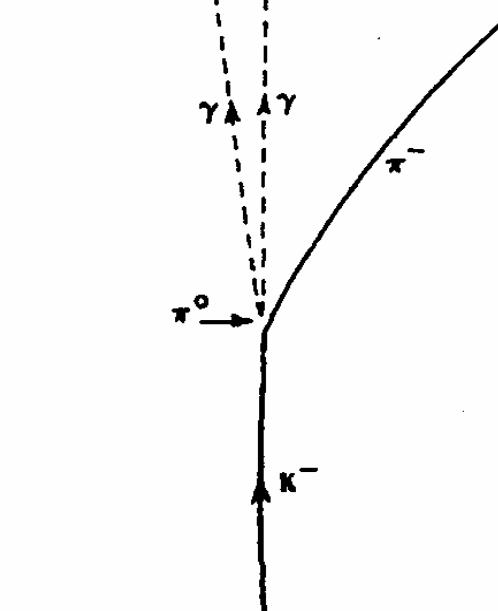
$$N^{2S+1} \quad L_{J=L+S}$$

- input state fixed by Ψ' , measure γ in NaI (Tl) crystals =>
- missing mass spectrum
- 1S_0 and 3P states n and X only in Ψ' decays, not in e^+e^-
- $\Psi' \not\rightarrow \Psi \gamma : 1^{--} \not\rightarrow 1^{--} 1^{--}$ (C violated)

$\Psi' - \Psi$	radial	N
$\Psi - n$	$\uparrow \uparrow - \uparrow \downarrow$	S
$\Psi - X$	orbital	L



C Parity

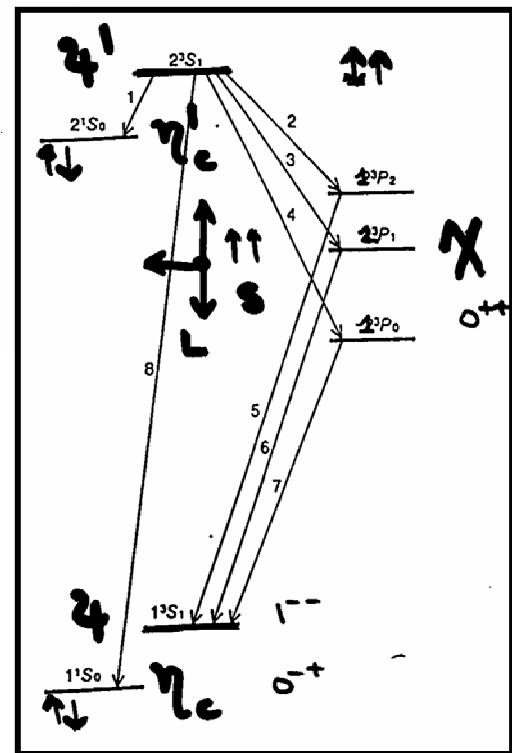


$$\Psi' \not\rightarrow \Psi \gamma$$

$$JPC \quad 1^- \not\rightarrow 1^- 1^-$$

$$\pi^0 \rightarrow \gamma \gamma$$

$$JPC \quad 0^{+} \rightarrow 1^{-} 1^{-}$$



C Parity

$$\pi^0 \not\rightarrow \gamma\gamma\gamma$$

$$\Gamma(\pi^0 \rightarrow 3\gamma) / \Gamma(\pi^0 \rightarrow 2\gamma) < 3 \cdot 10^{-8} \ll \alpha = 1/137$$

$$\pi^0 \rightarrow \gamma\gamma$$

$$C |\pi^0\rangle = C_\gamma^2 |\pi^0\rangle = + |\pi^0\rangle$$

$$J^{PC} = 0^{-+} \quad \text{pseudo-scalar mesons}$$

$$V \rightarrow P \gamma$$

$$\rho^0 \rightarrow \pi^0 \gamma$$

$\rho^0 \not\rightarrow 2$ identical particles :

$$\rho^0 \not\rightarrow \pi^0 \pi^0$$

$$\rho^0 \not\rightarrow \gamma\gamma$$

$$C |\rho^0\rangle = +C |\gamma\rangle = - |\rho^0\rangle$$

$$J^{PC} = 1^{--} \quad \text{vector mesons}$$

quarkonium spectroscopy:

$$\Psi' \not\rightarrow \Psi \gamma \quad 1^{--} \not\rightarrow 1^{--} 1^{--}$$

$$\rightarrow \eta_C \gamma \quad 1^{--} \rightarrow 1^{--} 0^{++}$$

$$\rightarrow \chi \gamma \quad 1^{--} \rightarrow 1^{--} 0,1,2^{++}$$

$\Phi, \Psi, Y \not\rightarrow 2$ gluons

OZI rule

$\rightarrow 3$ gluons

C Parity

- not eigenstates of C operator !
only eigenstates have eigenvalue !
- only totally neutral states with
all additive quantum numbers zero:
 $Q, B, L, S, C = 0$

$$\begin{array}{lcl} C \mid \pi^+ \rangle = \mid \pi^- \rangle & (Q) \\ C \mid n \rangle = \mid \bar{n} \rangle & (B) \\ C \mid K^0 \rangle = \mid \bar{K}^0 \rangle & (S) \\ C \mid \nu \rangle = \mid \bar{\nu} \rangle & (L) \end{array}$$

$\gamma, \pi^0, \eta^0, \rho^0, \omega^0, \Phi^0, \Psi, Y, Z$

are eigenstates of C operator !

- spin like magnetic field = rotating charge:
negative C parity

$$C \mid f \bar{f} \rangle = C \mid \text{Meson} \rangle = \mid f \bar{f} \rangle (-1)^{L+S}$$

L	S	J	C	Multiplet	Expl.
L=S			+		
0	0	0	+	Pseudo-Scalar	π^0
0	0	0	-	FORBIDDEN !	
0	1	1	-	Vector (like γ)	ρ^0
1	1	0	+	Scalar L=1 rare	A_0
1	1	1	+	Axial-Vector	A_1
1	1	2	+	Tensor	f^0

P Parity

P Operation	Type	Dirac	J ^P	Example
$P \ \vec{r}\rangle = - \vec{r}\rangle$	vector	γ_μ	1 ⁻	
$P \ \vec{p}\rangle = - \vec{p}\rangle$	vector	γ_μ	1 ⁻	$\vec{p} = d\vec{r}/dt$
$P \ t\rangle = t\rangle$	scalar	1	0 ⁺	
$P \ E\rangle = E\rangle$	scalar	1	0 ⁺	
$P \ \vec{B}\rangle = \vec{B}\rangle$	axialvector	$\gamma_5 \gamma_\mu$	1 ⁺	$\vec{B} = \vec{v} \times \vec{E} = \vec{\nabla} \times \vec{A}$
$P \ \vec{\sigma}\rangle = \vec{\sigma}\rangle$	axialvector	$\gamma_5 \gamma_\mu$	1 ⁺	$\vec{L} = \vec{v} \times \vec{r}$
$P \ \vec{p} \vec{\sigma}\rangle = - \vec{p} \vec{\sigma}\rangle$	pseudoscalar	γ_5	0 ⁻	$H = \vec{p} \vec{\sigma} / \vec{p} \vec{\sigma} $
$P \ F_{\mu\nu}\rangle = F_{\mu\nu}\rangle$	tensor	$\gamma_\mu \gamma_\nu$	2 ⁺	$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$
$T \ t\rangle = - t\rangle$	time reversal			

H = helicity
helix = screw

$$P^2 = 1 \quad P \dots \text{unitary}$$

combined CPT conserved in field theory

$\theta - \tau$ paradox

1955: $P |\pi\rangle = -|\pi\rangle$

τ : $K^\pm \rightarrow \pi^\pm \pi^+ \pi^-$ $P |3\pi\rangle = -|3\pi\rangle$

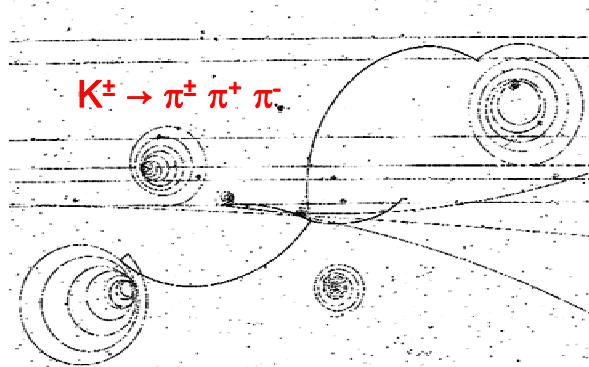
θ : $K^\pm \rightarrow \pi^\pm \pi^0$ $P |2\pi\rangle = +|2\pi\rangle$

Are θ and τ different?

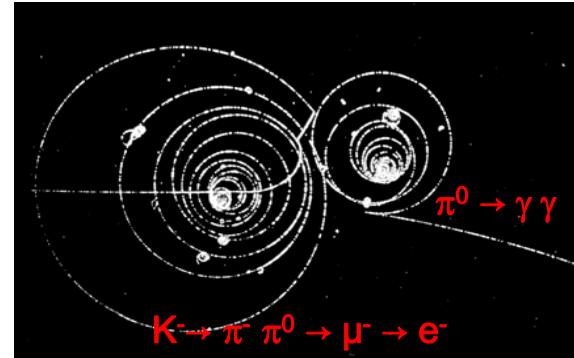
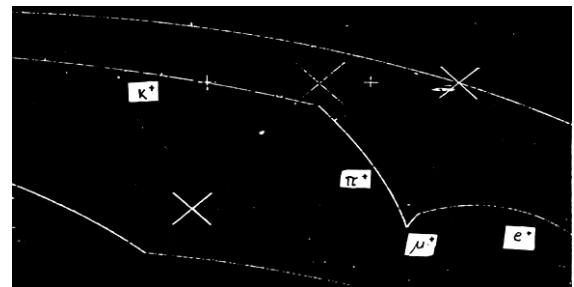
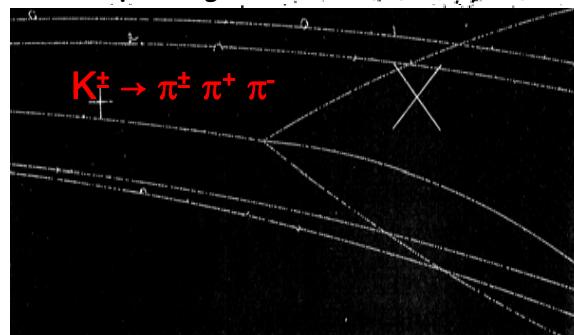
But their masses are equal!

Solution: $\theta = \tau$

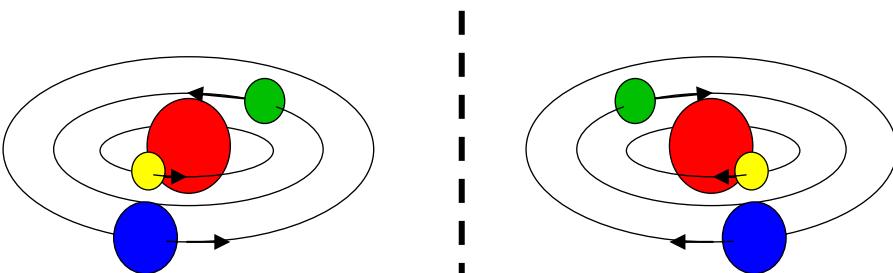
Parity VIOLATED in weak decays!



decay in flight: uneven nr of tracks!



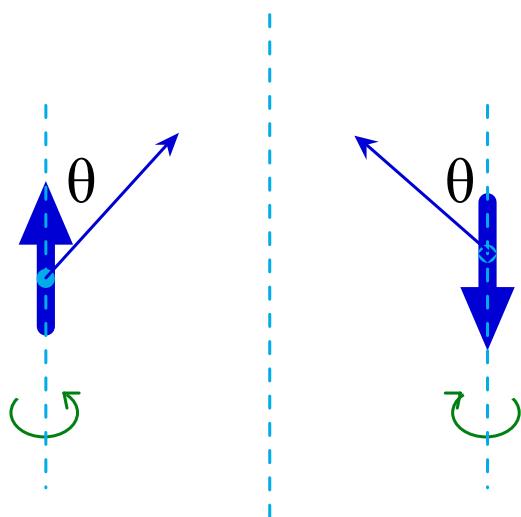
Parity violation



Gravitation +
electromagnetism:

same physical laws
in mirror system,
e.g. planetary motion

parity conserved.



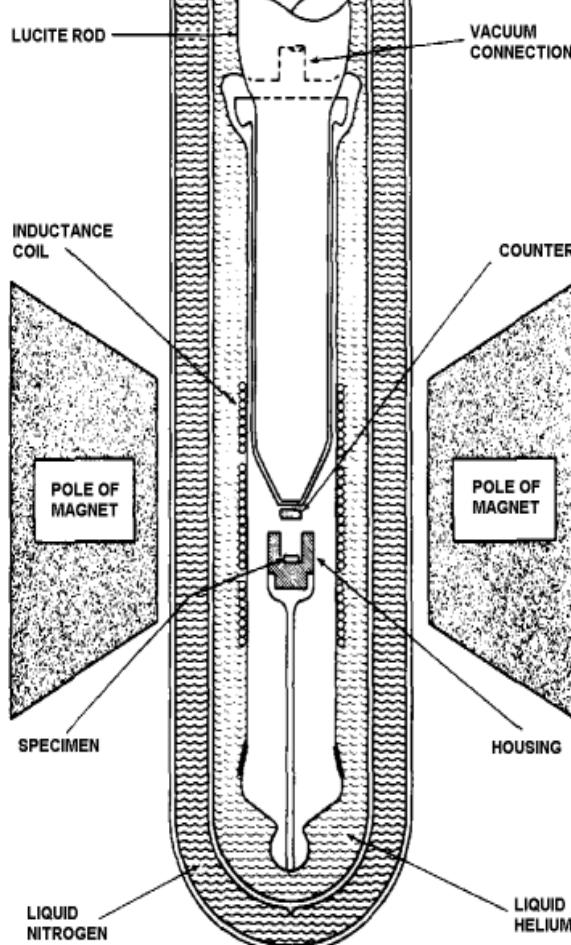
Weak interaction:

β -decay
of polarized ^{60}Co
violates parity !

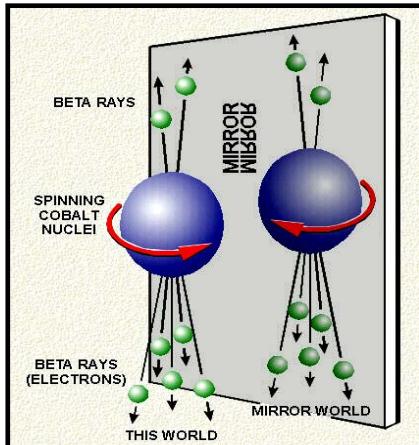
discovered 1956

Parity violation

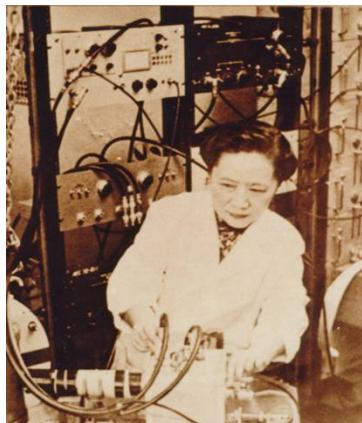
cool to 10 mK



polarize nuclear spins
along magnetic field



C.S. Wu et al.,



Columbia Univ., USA 1957

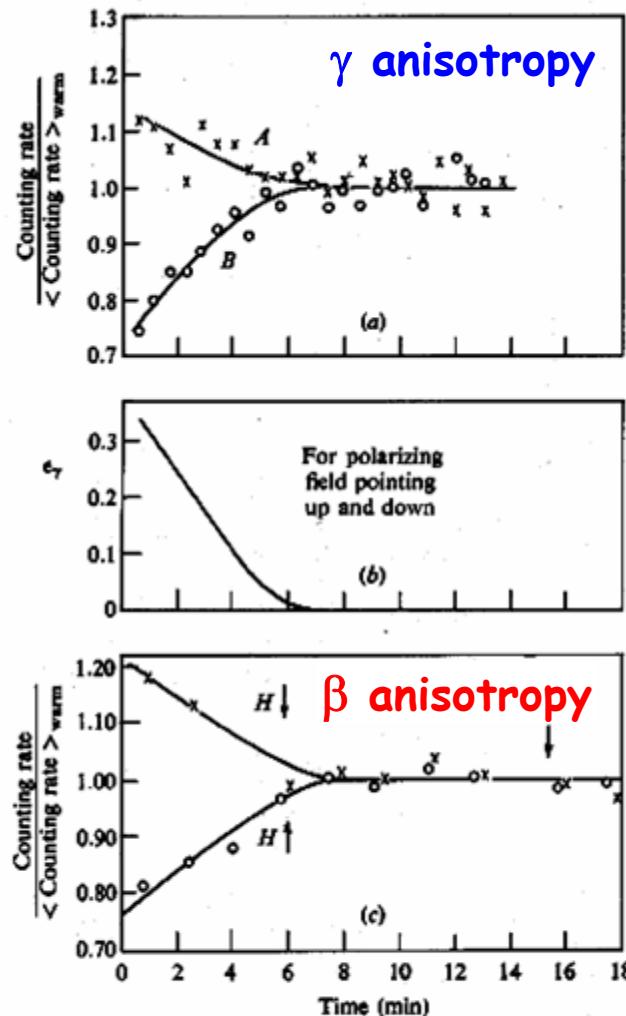


FIG. 8.37 The results obtained by Wu, et al. by using the apparatus shown in Fig. 8.36. [From Wu, et al., Phys. Rev., 105, 1413, (1957).] (a) gamma-anisotropy obtained by equatorial counter (A) and polar counter (B). (b) gamma-anisotropy calculated from (a). (c) beta-asymmetry.

Parity Violation

nuclear β decay:

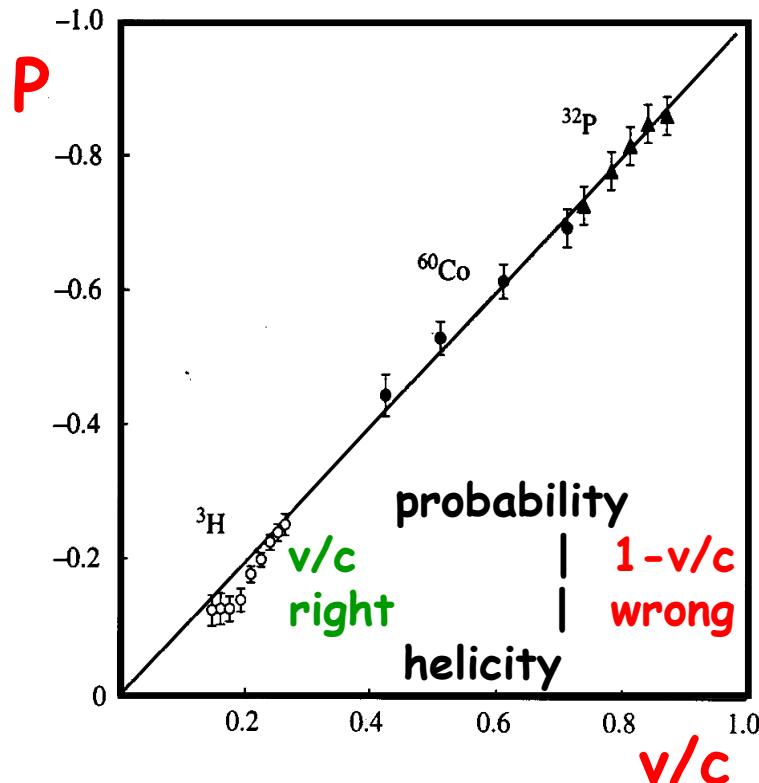
$$\text{helicity } H = (\vec{\sigma} \cdot \vec{p}) / |\vec{p}|$$

in a system with $c > v' > v$
momentum p and helicity H
are inverted:

probability of 'wrong' helicity $1 - v/c$

$$\text{polarization } P = (N^+ - N^-) / (N^+ + N^-)$$
$$= \sigma \cdot p c / E = H v/c$$

longitudinal polarization P
vs. electron velocity v/c :



(V-A) Theory

→
 σ ... unit vector in spin direction
 helicity operator H :

$$H = \frac{\vec{\sigma} \cdot \vec{p}}{|\vec{p}|} \quad H |\Psi\rangle = \pm |\Psi\rangle$$

projection operator L / R left / right:

$$L = (1 - H) / 2 = 0,1 \quad \text{for} \quad H = \pm 1 \quad \text{or} \quad (1 \pm \gamma_5) / 2$$

$$R = (1 + H) / 2 = 1,0$$

chirality = handiness

$$L = (1 - H) / 2 \quad \text{particle: left handed}$$

$$L = (S - P) / 2 \quad \text{space structure: (Scalar-Pseudoscalar)}$$

all interactions via vector bosons V :

$$L = V (S - P) / 2$$

$$L = (V - A) / 2$$

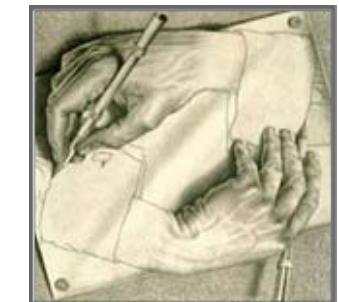
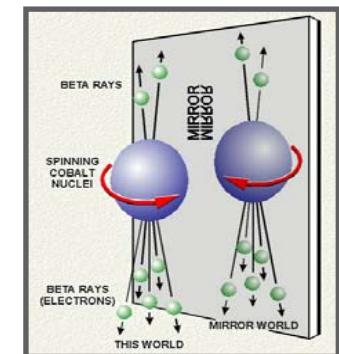
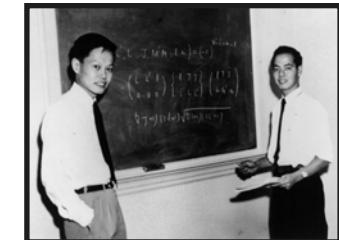
$$L = \gamma^\mu (1 - \gamma^5) / 2$$

space structure of weak force :

(V-A)

why ?

maximum parity violation !



Helicity + Chirality

v_L : FERMIONS are LEFT handed

\bar{v}_R : ANTI-FERMIONS RIGHT handed

$$P \mid v_L \rangle = \mid v_R \rangle$$

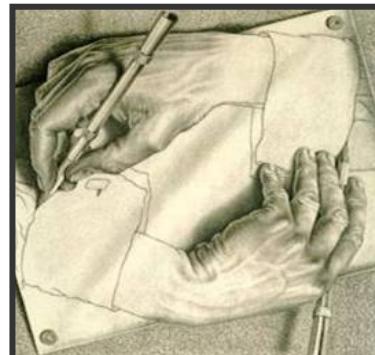
P violated

$$C \mid v_L \rangle = \mid \bar{v}_L \rangle$$

C violated

$$CP \mid v_L \rangle = \mid \bar{v}_R \rangle$$

CP conserved



Manche meinen,
lechts und rechts
kann man nicht
verwechseln.

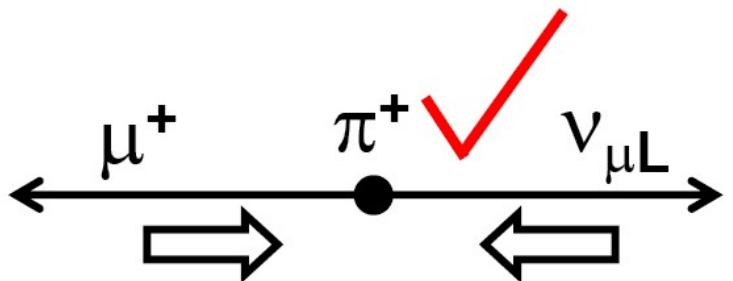
Werch ein Illtum !

E.Jandl

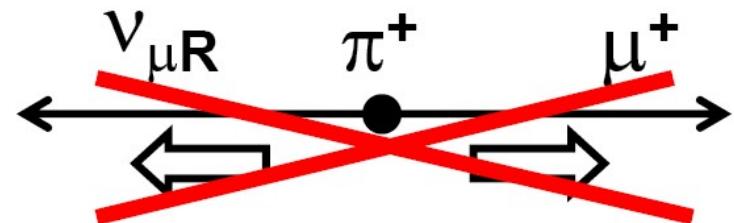
C,P violation - CP conservation

only left-handed neutrinos exist

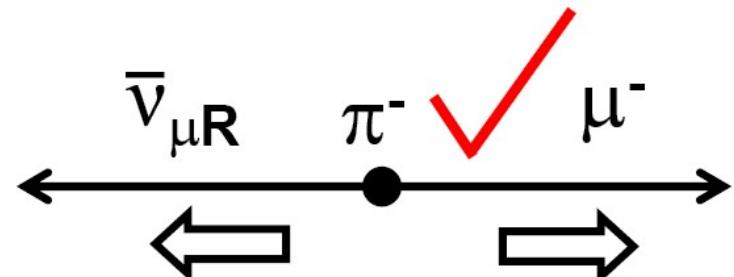
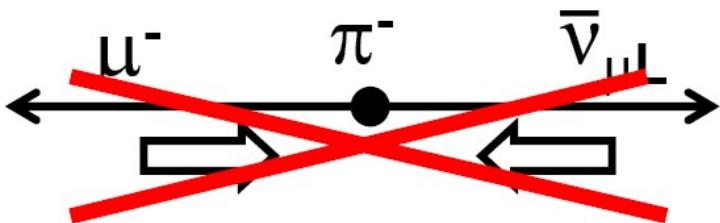
weak interaction **violates C+P separately**:



P:



C:



Only combined CP operation
gives right-handed antineutrino, which exists:
CP conserved !

4. Weak Interaction

Neutrinos, they are very small.
They have no charge and have no mass
And do not interact at all.

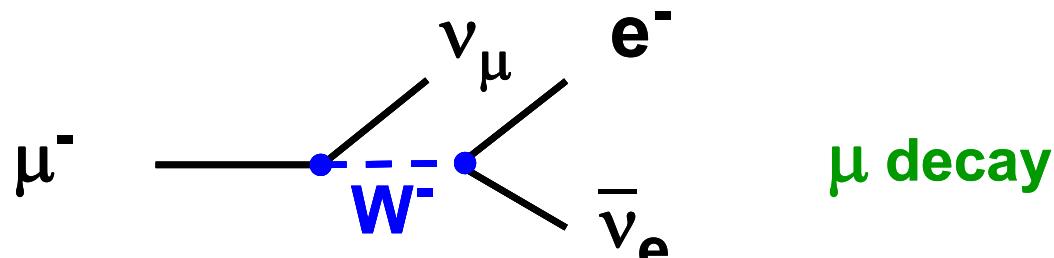
The earth is just a silly ball
To them, through which they simply pass.

John Updike

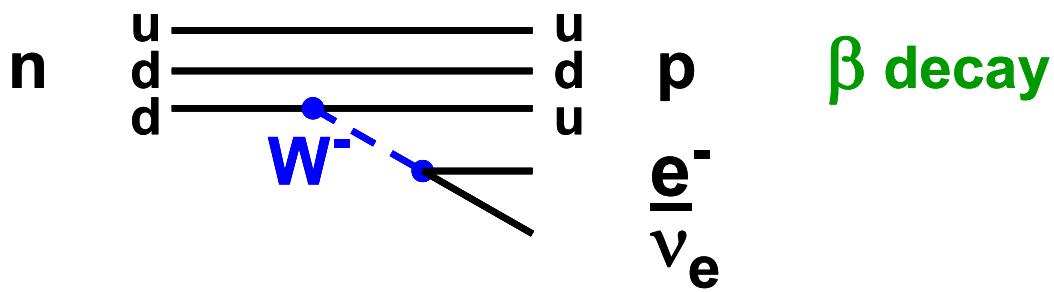
Weak Decays

DECAYS

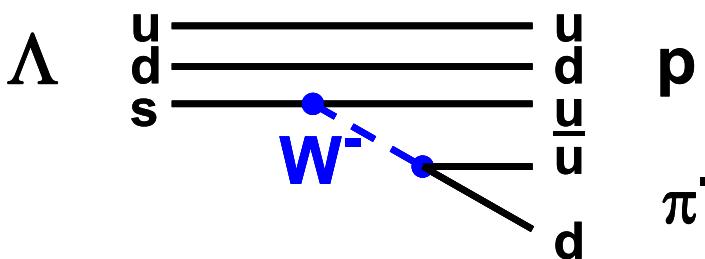
leptonic



semi-leptonic



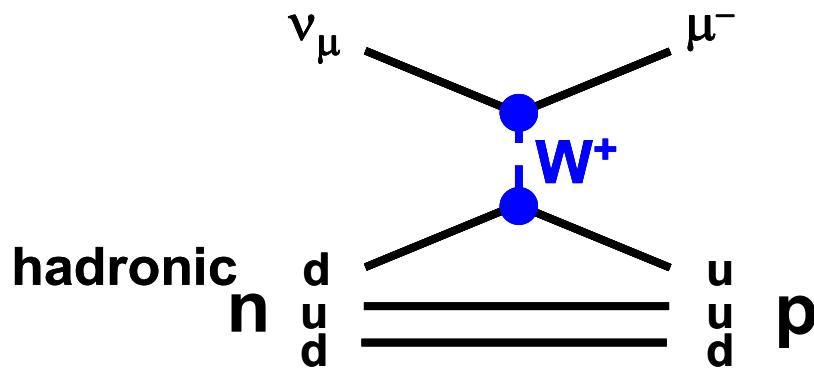
non-leptonic



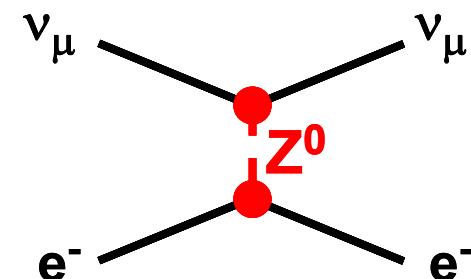
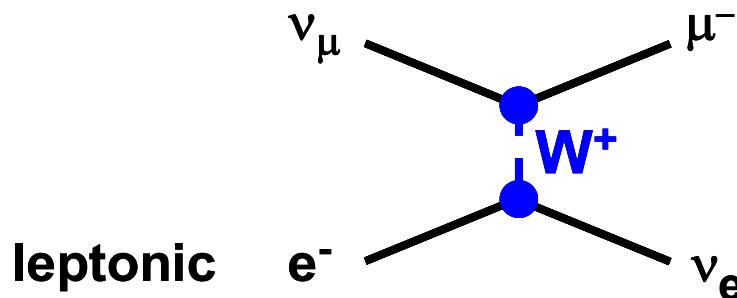
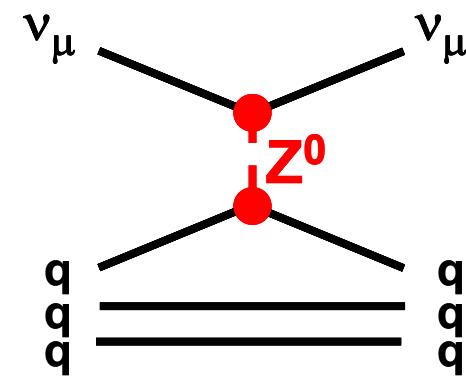
Weak Reactions

current

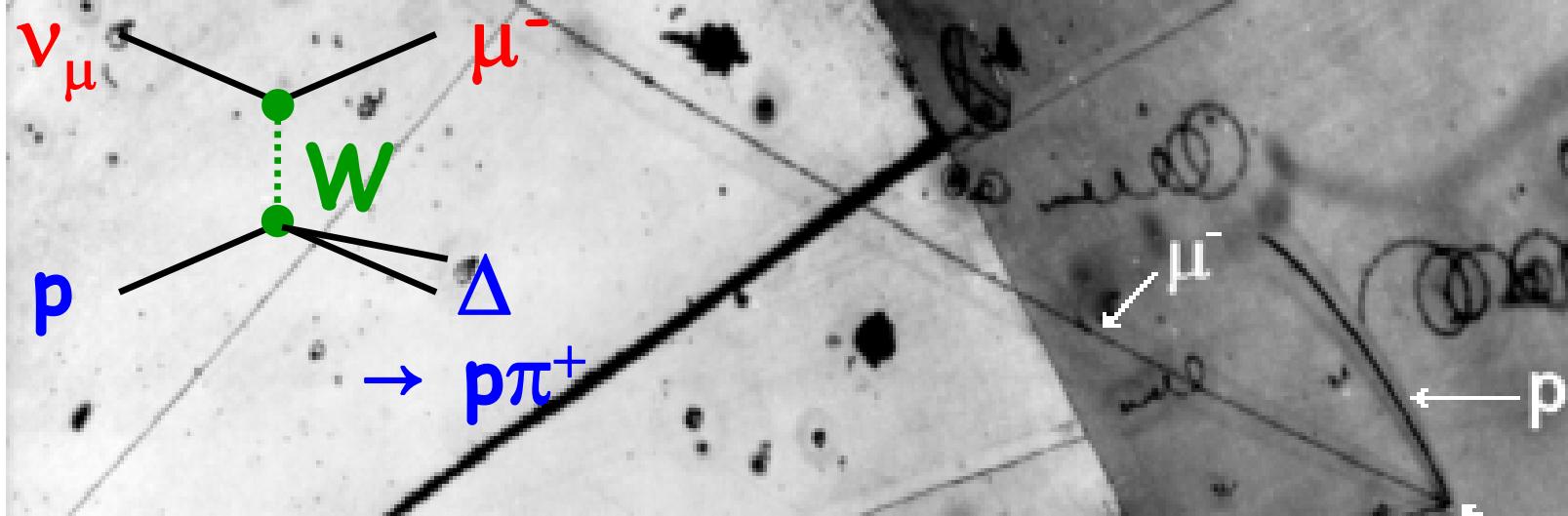
charged



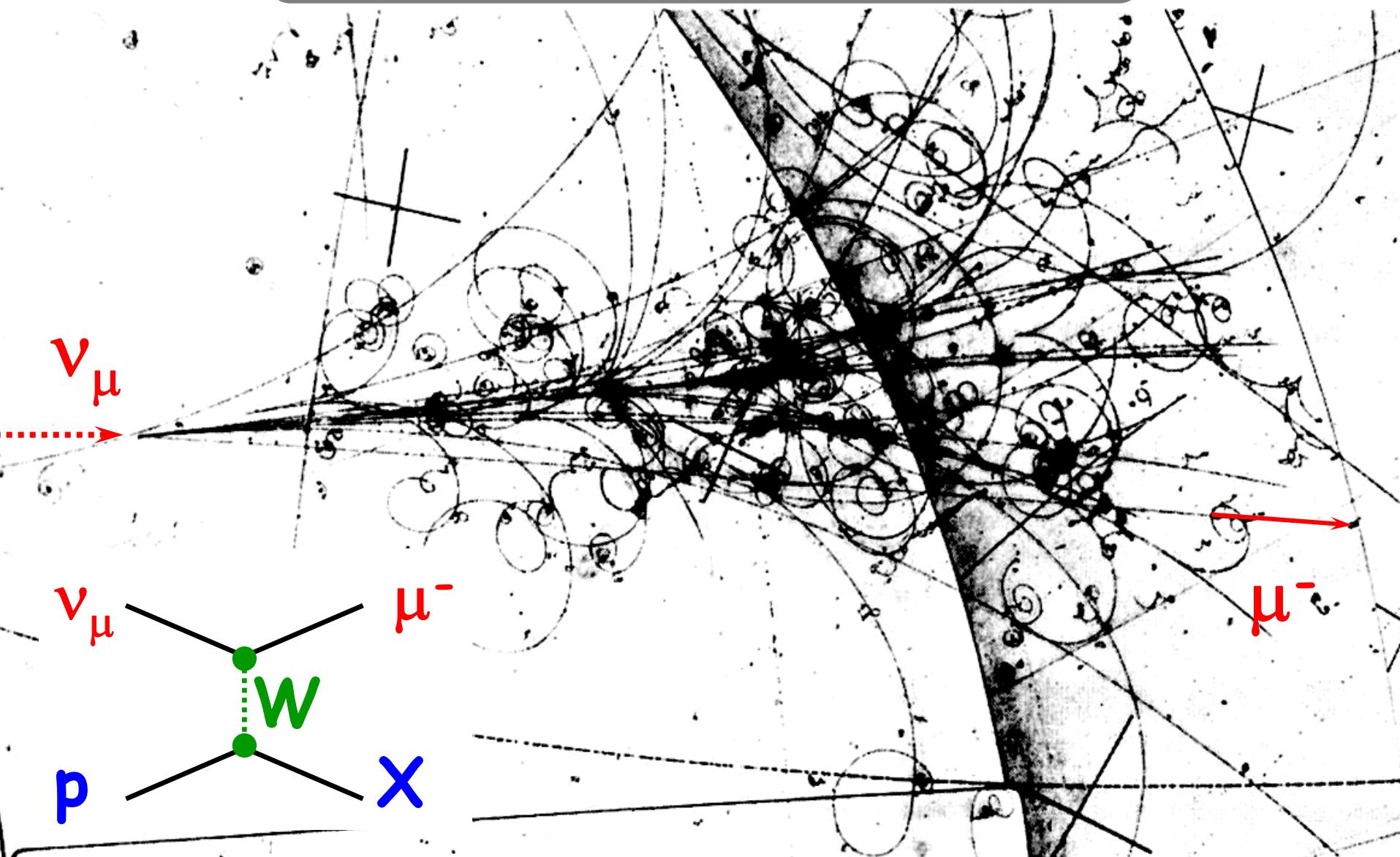
neutral



Charged Weak Current



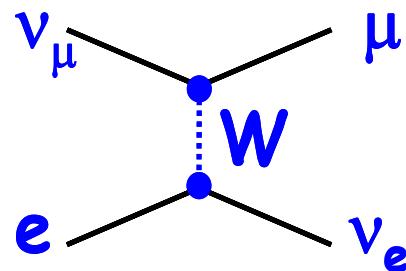
Charged Weak Current



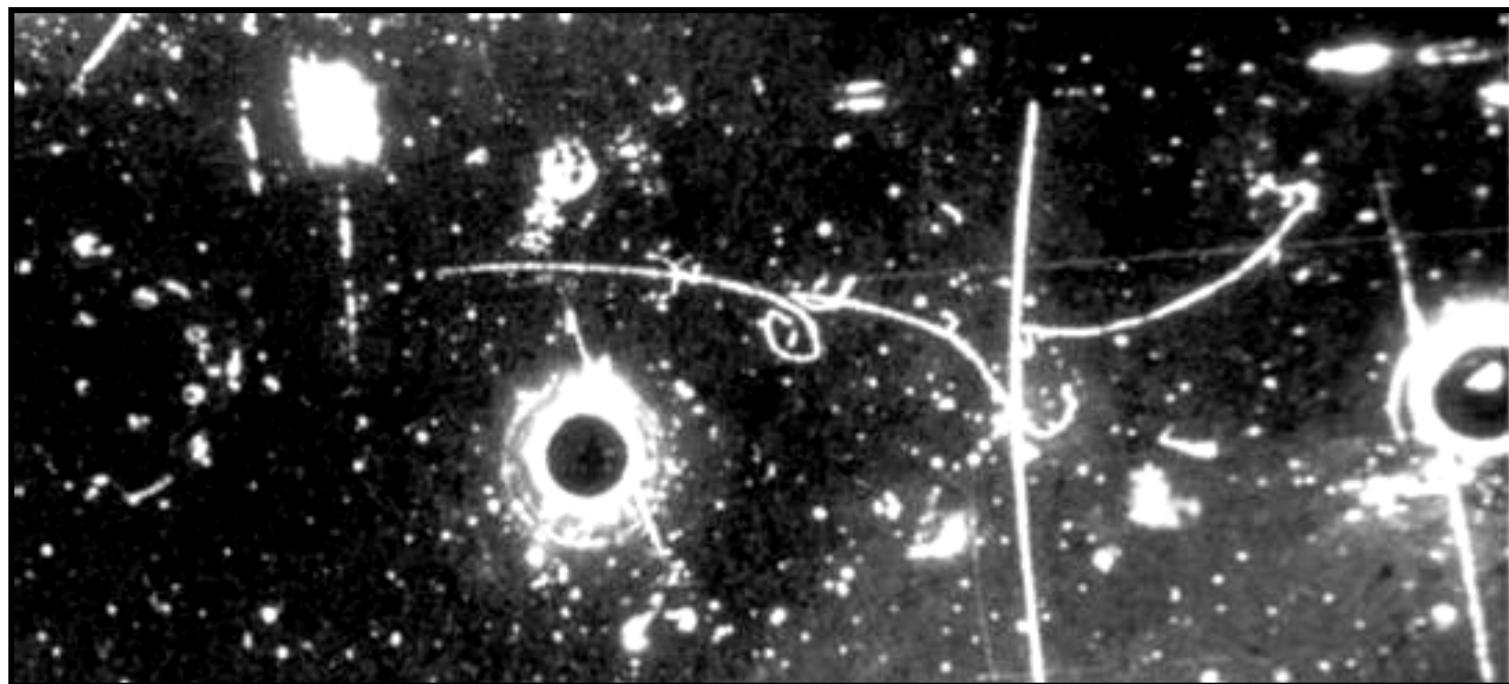
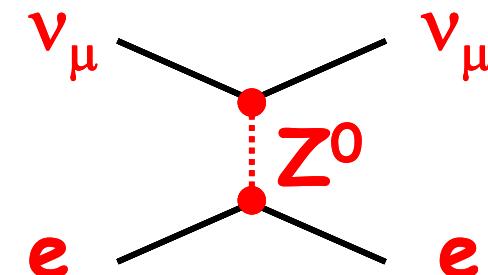
CERN 1984: Big European Bubble Chamber BEBC

Neutral Weak Current

charged current

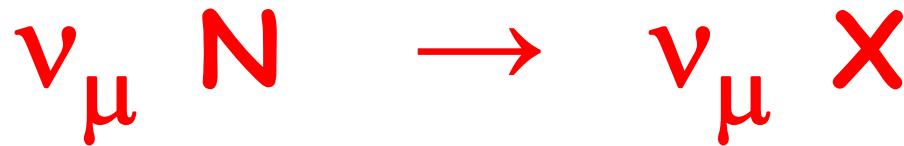


neutral current

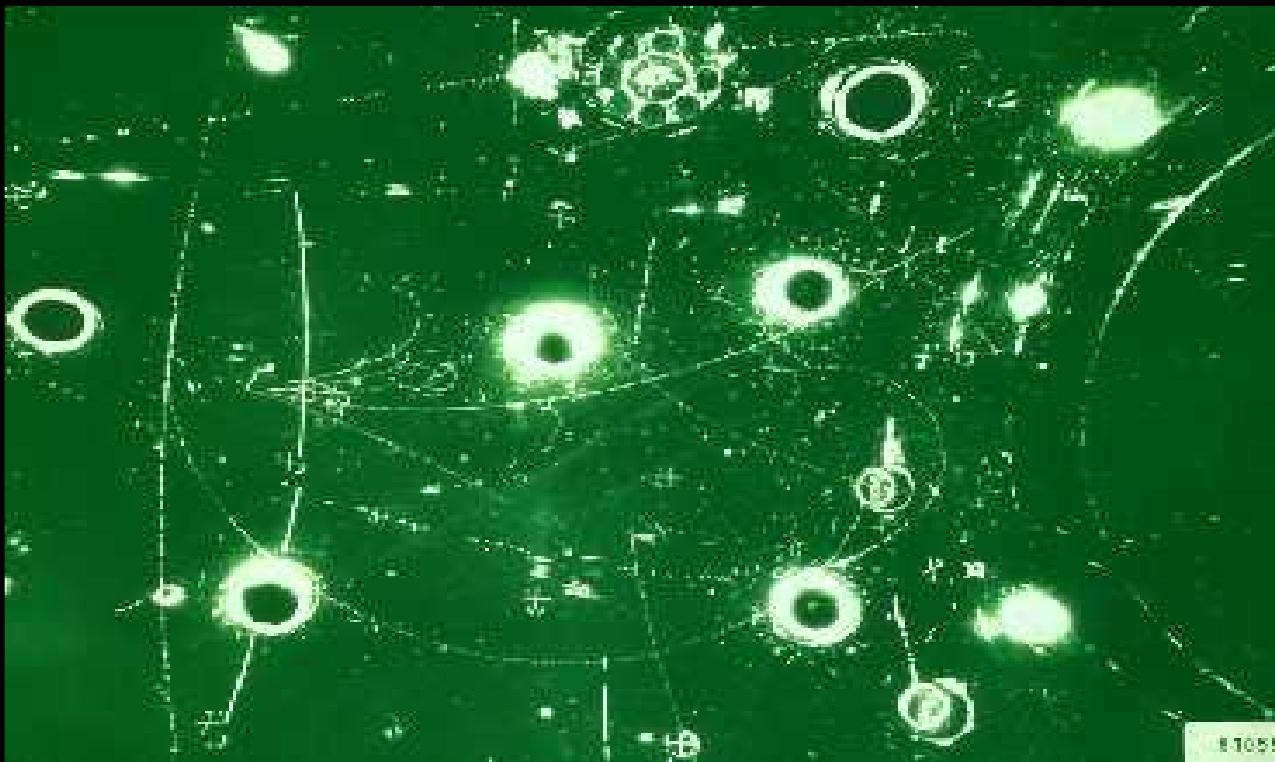


CERN 1973 : Gargamelle Heavy Liquid Bubble Chamber

Neutral Weak Current



CERN 1973 : Gargamelle Heavy Liquid Bubble Chamber



Neutrinos

Neutrinos, they are very small
They have no charge and have no mass
And do not interact at all.
The earth is just a silly ball
To them, through which they simply pass,
Like dustmaids down a drafty hall
Or photons through a sheet of glass.

They snub the most exquisite gas,
Ignore the most substantial wall,
Cold-shoulder steel and sounding brass,
Insult the stallion in his stall,
And, scorning barriers of class,
Infiltrate you and me! Like tall
And painless guillotines, they fall
Down through our heads into the grass.

At night, they enter at Nepal
And pierce the lover and his lass
From underneath the bed. You call
It wonderful; I call it crass.

John Updike

Neutrinos

- pure sources + pure probes of **weak interaction**

- extremely **penetrating** !

- **omnipresent**:

- solar neutrino flux:	$6 \cdot 10^{10} \text{ m}^{-2} \text{ s}^{-1}$
- atmospheric neutrino flux:	$10^4 \text{ m}^{-2} \text{ s}^{-1}$
- supernova SN 1987A flux:	10^{15} m^{-2}
- cosmic relic neutrinos:	335 cm^{-3}

- charged leptons have **mass** :

511 keV

$\begin{bmatrix} e \\ \nu_e \end{bmatrix}$

< 2 eV

$H \rightarrow He e \bar{\nu}_e$

106 MeV

$\begin{bmatrix} \mu \\ \nu_\mu \end{bmatrix}$

< 190 keV

$\pi \rightarrow \mu \nu_\mu$

1.8 GeV

$\begin{bmatrix} \tau \\ \nu_\tau \end{bmatrix}$

< 18 MeV

$\tau \rightarrow 5\pi \nu_\tau$

$\begin{bmatrix} u \\ d \end{bmatrix}$

strong
electromagnetic
weak

$\begin{bmatrix} e \\ \nu_e \end{bmatrix}$

electromagnetic
weak
weak

NEUTRINO
MASS

?

Pauli's Neutrino hypothesis

$n \rightarrow p e^-$? continuous β spectrum
E conservation violated ?

Offener Brief an die Gruppe der Radikaktiviten bei der
Gauvereins-Tagung zu Tübingen.

Abschrift

Physikalisches Institut
der Eidg. Technischen Hochschule
Zürich

Zürich, 4. Des. 1930
Cleriastrasse

Liebe Radioaktive Damen und Herren,

Wie der Ueberbringer dieser Zeilen, den ich halbwollst
anzuhören bitte, Ihnen des näheren auszusinndersetzen wird, bin ich
angesichts der "falschen" Statistik der H- und Li-6 Kerne, sowie
des kontinuierlichen beta-Spektrums auf einen verzweifelten Anweg
verfallen um den "Wechselsatz" (1) der Statistik und dem Energiesatz
zu retten. Mömlich die Möglichkeit, es könnten elektrisch neutrale
Teilchen, die ich Neutronen nennen will, in den Kernen existieren,
welche den Spin 1/2 haben und das Ausschließungsprinzip befolgen und
sich von Lichtquanten außerdem noch dadurch unterscheiden, dass sie
nicht mit Lichtgeschwindigkeit laufen. Die Masse der Neutronen
sollte von derselben Grossenordnung wie die Elektronenmasse sein und
jedemfalls nicht grösser als 0,01 Protonenmassen. Das kontinuierliche
beta-Spektrum wäre dann verständlich unter der Annahme, dass beim
beta-Zerfall mit dem Elektron jeweils noch ein Neutron emittiert
wird, derart, dass die Summe der Energien von Neutron und Elektron
konstant ist.



© CERN, Geneva

$n \rightarrow p e^- \bar{\nu}_e$

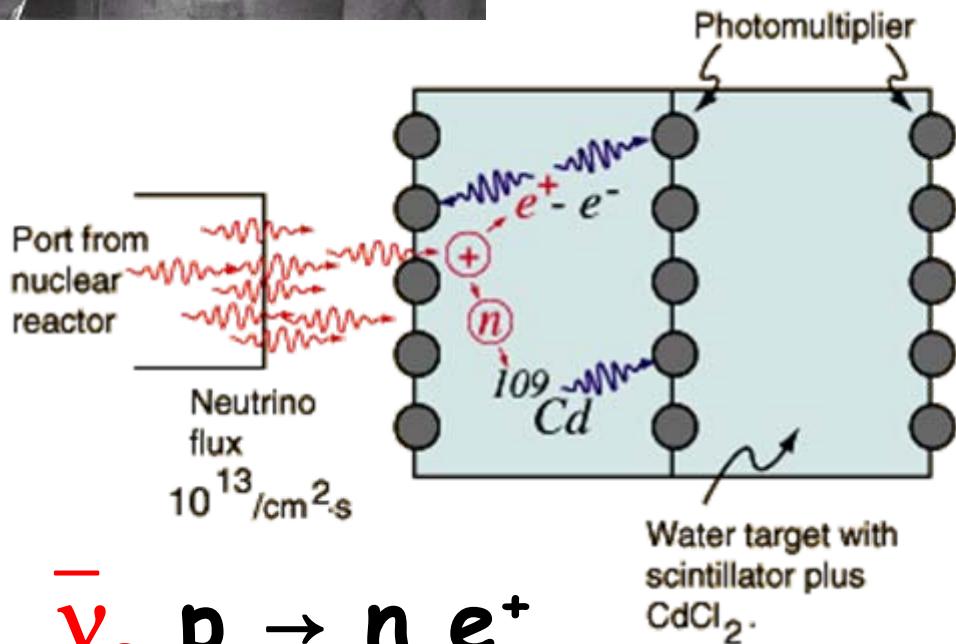
renamed
Neutrino
1933 by E. Fermi
after neutron
discovery



Neutrino Discovery

25 years later: Project Poltergeist:

C.Cowan and F.Reines, 1956 :
Savannah river reactor, USA



Delayed coincident
detection of γ from ^{109}Cd
with pair of γ 's from
 $e^+ - e^-$ annihilation.

$$\begin{aligned} \bar{\nu}_e p &\rightarrow n e^+ \\ e^+ e^- &\rightarrow \gamma \gamma \quad E_\gamma = m_e \\ n \text{ Cd} &\rightarrow \text{Cd}^* \gamma \\ \gamma_n \gamma_e\text{-coincidences} & \quad \delta t \sim \mu\text{s} \end{aligned}$$

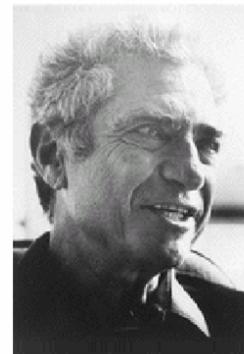
Reines:
Nobel prize



1995



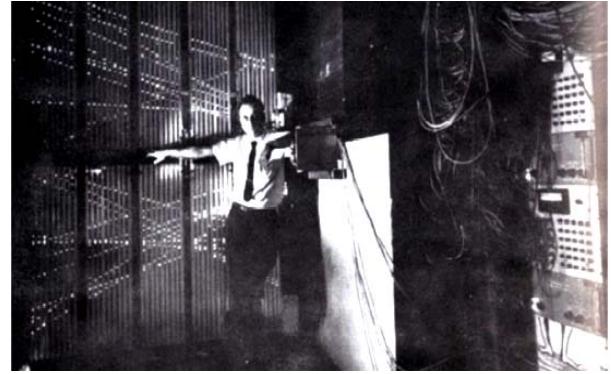
Discovery of muon neutrino



Nobel



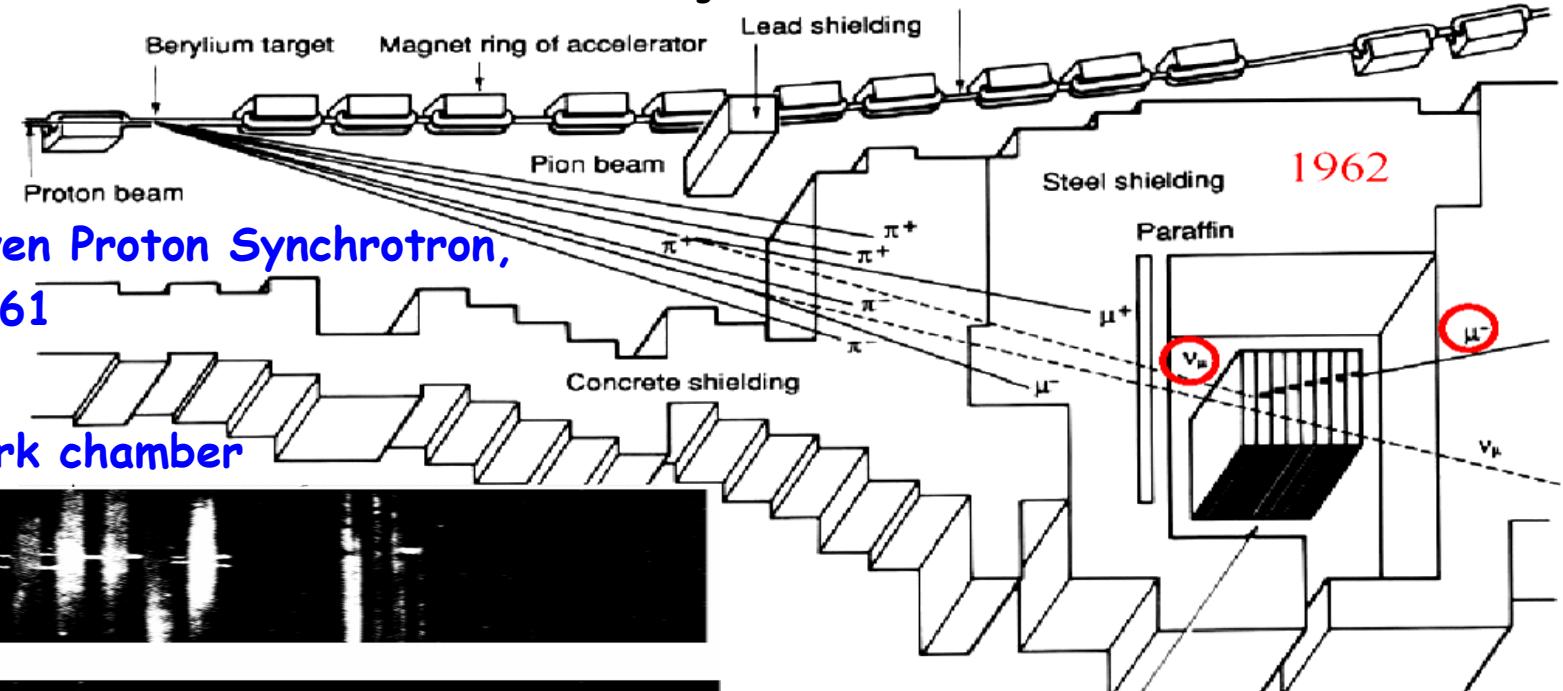
prize
1988



L.Lederman

M.Schwartz

J.Steinberger



Brookhaven Proton Synchrotron,
USA, 1961

10 t spark chamber

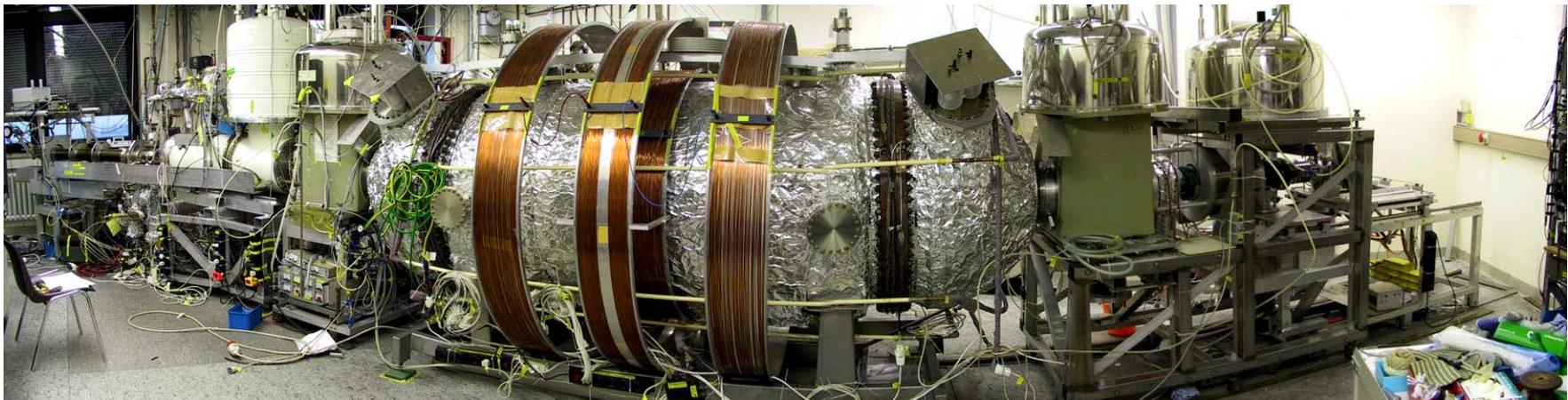
μ

Spark chamber

$\nu_\mu \neq \nu_e$

ν_μ

Neutrino mass



Mainz experiment:

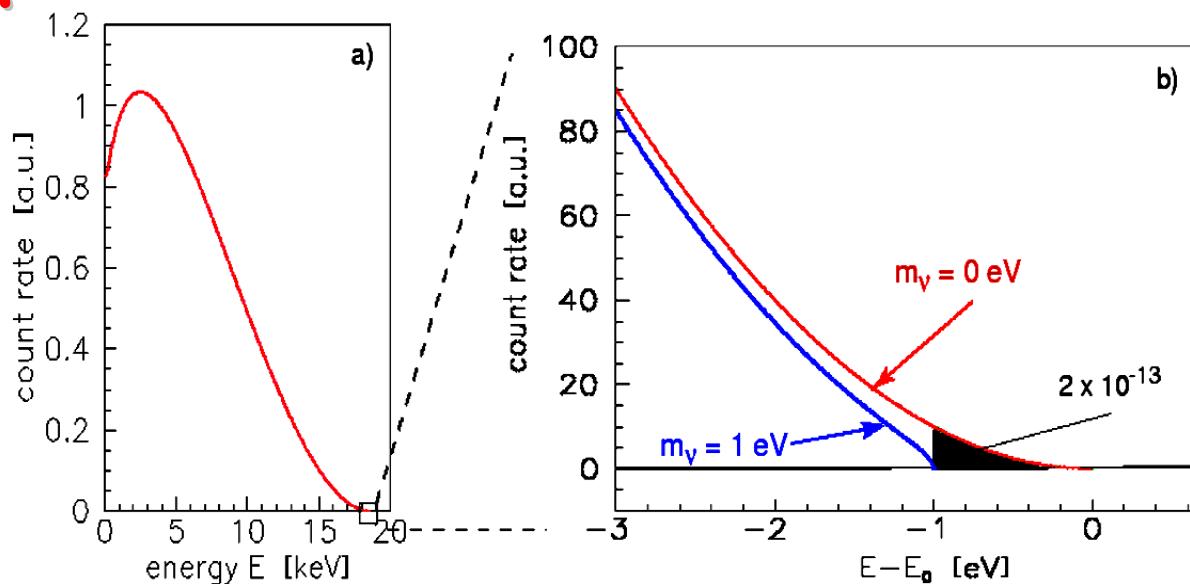
Kurie plot: end point of
Tritium β spectrum



$$m(\nu_e) < 2.3 \text{ eV}$$

(95% CL)

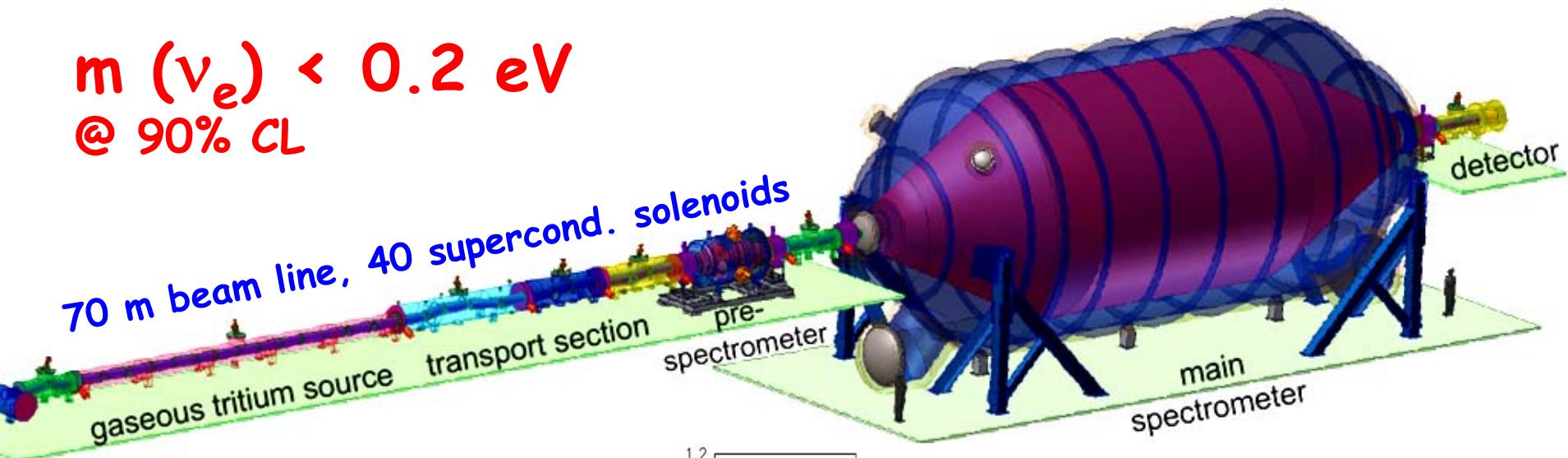
problem: nuclear models



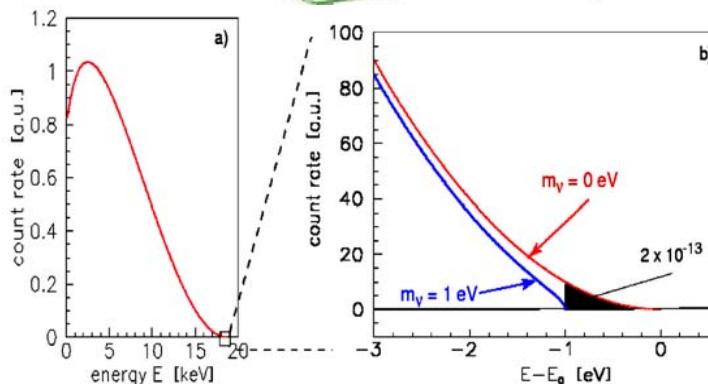
Neutrino Mass

KATRIN: KArlsruhe TRItium Neutrino Experiment

$m(\nu_e) < 0.2 \text{ eV}$
@ 90% CL



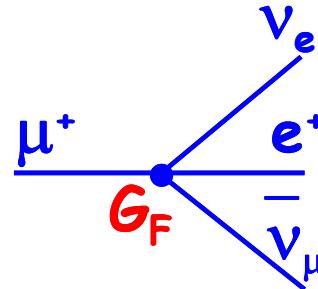
> 2008:
measure end point
of tritium β spectrum:



Fermi Constant

4 Fermion theory of weak interaction
with coupling constant

$$G_F = (1.16637 \pm 0.00001) \times 10^{-5} \text{ GeV}^{-2}$$

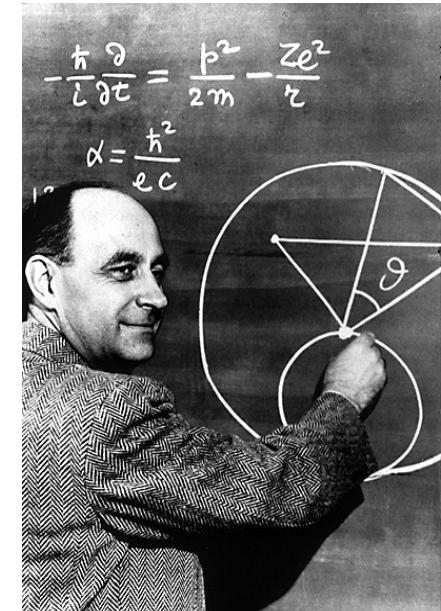
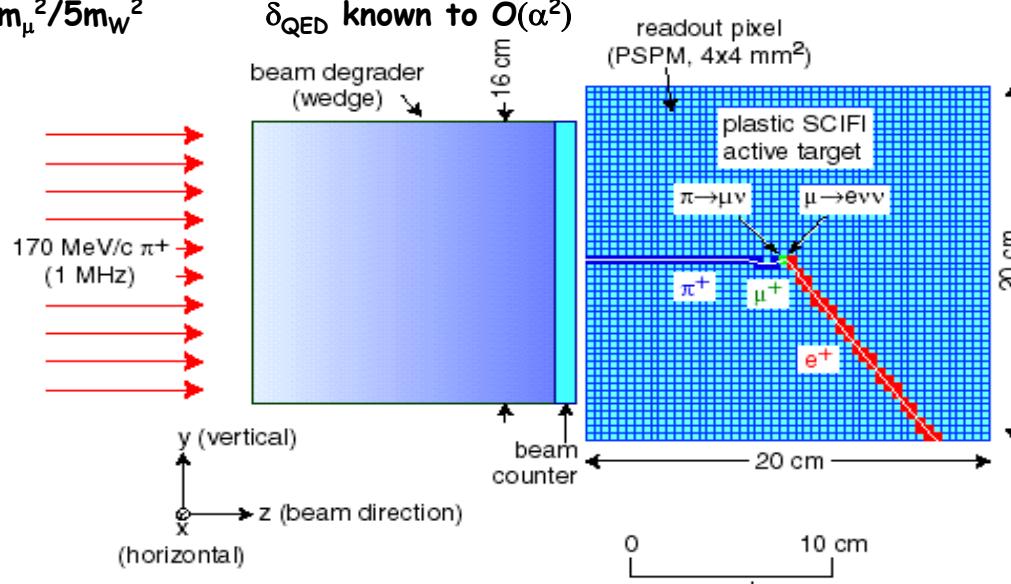


determined from muon lifetime

$$\tau_\mu = 192\pi^3 / (G_F^2 m_\mu^5) (1 + \delta_{\text{kin}})(1 + \delta_{\text{weak}})(1 + \delta_{\text{QED}})$$

$$\delta_{\text{kin}} = -8x + 8x^3 - x^4 - 12x^2 \ln x \quad \text{with } x = m_e^2/m_\mu^2 = 2 \cdot 10^{-5}$$

$$\delta_{\text{weak}} = 3m_\mu^2/5m_W^2 \quad \delta_{\text{QED}} \text{ known to } O(\alpha^2)$$



E. Fermi

Nobel



prize

1938

FAST at CERN

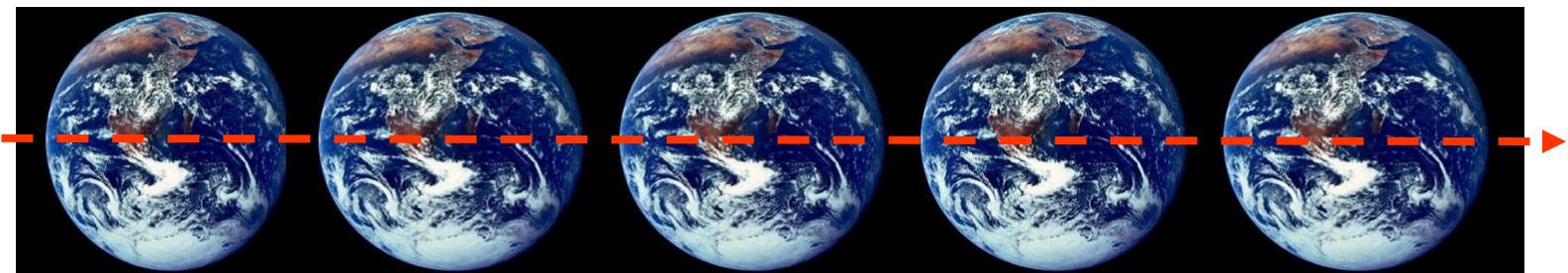
get G_F to 1 ppm
with $>10^{12} \mu$ decays

Neutrino interactions

only 1 solar neutrino/person/human life reacts !
can a solar neutrino penetrate the Earth?

$$L_{\text{reaction}} = \frac{\text{atomic weight}}{\text{density} \times \text{cross section} \times N_{\text{Avogadro}}} = \frac{40 \text{ g/mol}}{5.5 \text{ g/cm}^3 \times 0.7 \times 10^{-38} (\text{E/GeV}) \text{ cm}^2 \times 6.023 \times 10^{23} / \text{mol}}$$

mean free path 2×10^{13} km = 2 light years :



1 billion earths or 10 million suns

Supernova 1987A. 23.
Februar 1987



Neutrinos

Supernovae

SN 1987A:
 10^{15} m^{-2}

Sun

$6 \cdot 10^{10} \text{ m}^{-2} \text{ s}^{-1}$

Reactors

$2 \cdot 10^{20} \text{ GW s}$

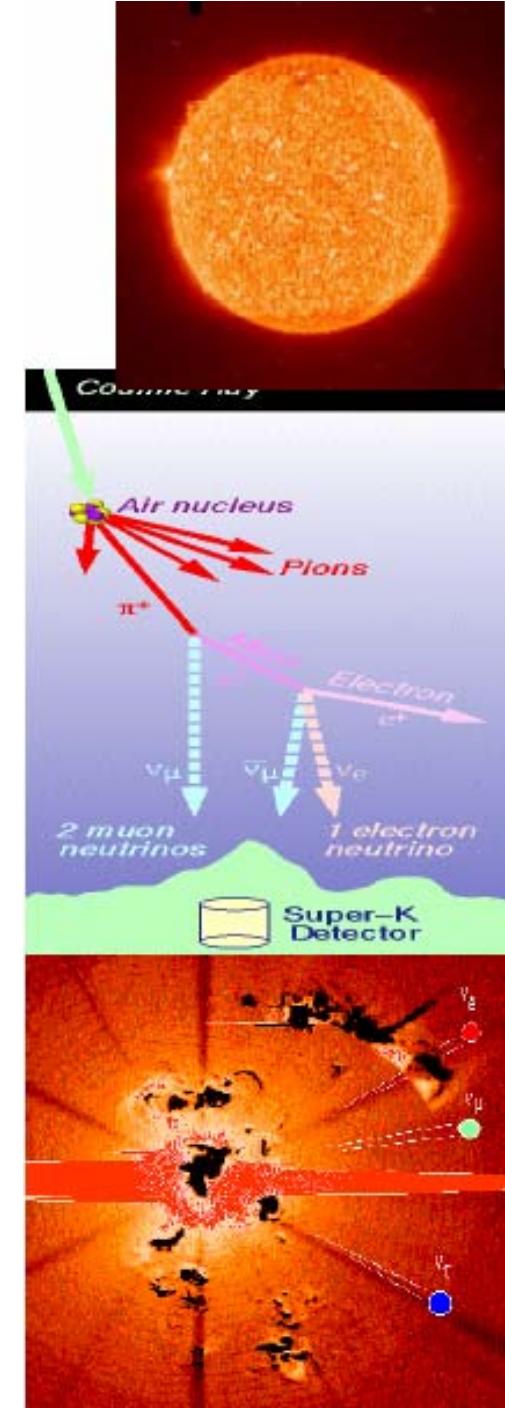
Cosmic radiation - Atmosphere

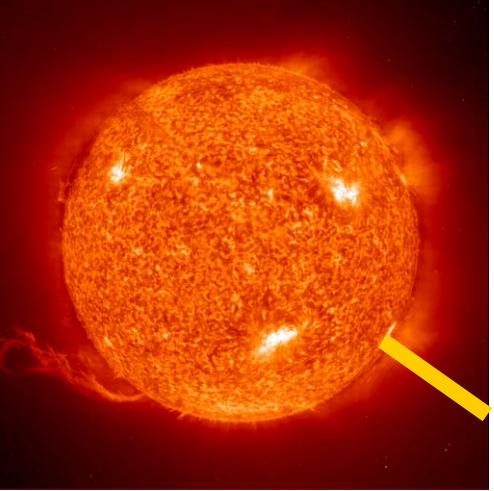
$10^4 \text{ m}^{-2} \text{ s}^{-1}$

Accelerators

Big Bang

335 cm^{-3}



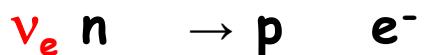
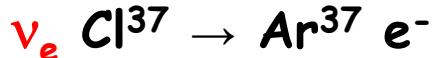


Solar Neutrinos

nuclear fusion in the Sun :



micro-chemistry: detect single atoms

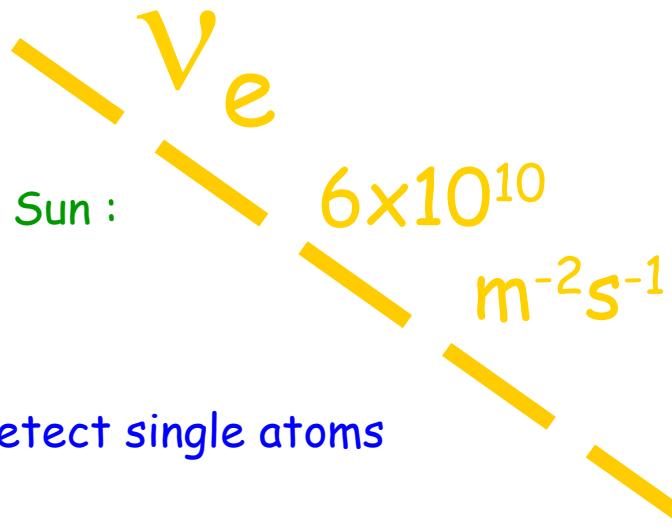


1967-94: ~ 2/3 neutrinos missing !

nuclear fission in reactors produces $\bar{\nu}_e$



$$\nu \neq \bar{\nu}$$



Homestake Gold mine



$600 + C_2Cl_4$

Neutrino oscillations

Flavor eigenstates

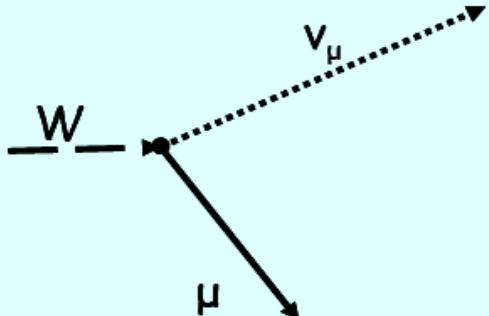
$$v_\mu, v_\tau$$

$$\begin{pmatrix} v_\mu \\ v_\tau \end{pmatrix} = \begin{pmatrix} \cos\theta_{23} & \sin\theta_{23} \\ -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} v_2 \\ v_3 \end{pmatrix}$$

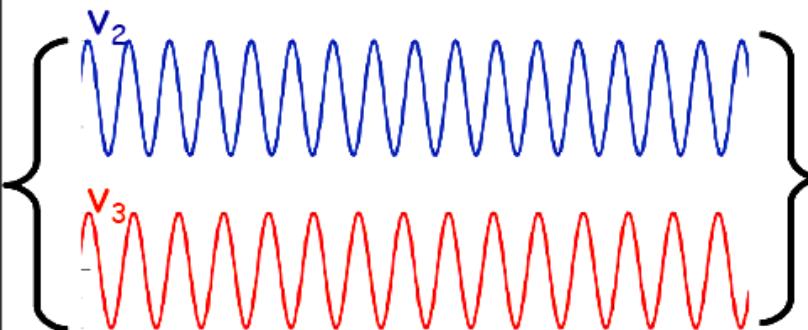
Mass eigenstates

$$v_2, v_3 \text{ with } m_2, m_3$$

source creates flavor-eigenstates



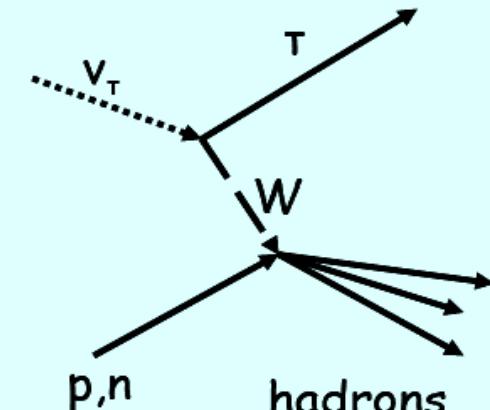
propagation determined by mass-eigenstates



$$\omega_{2,3} = E_{2,3} = \sqrt{p^2 + m_{2,3}^2}$$

slightly different frequencies
 \rightarrow phase difference changes

detector sees flavor-eigenstates



$$\theta=45^\circ$$

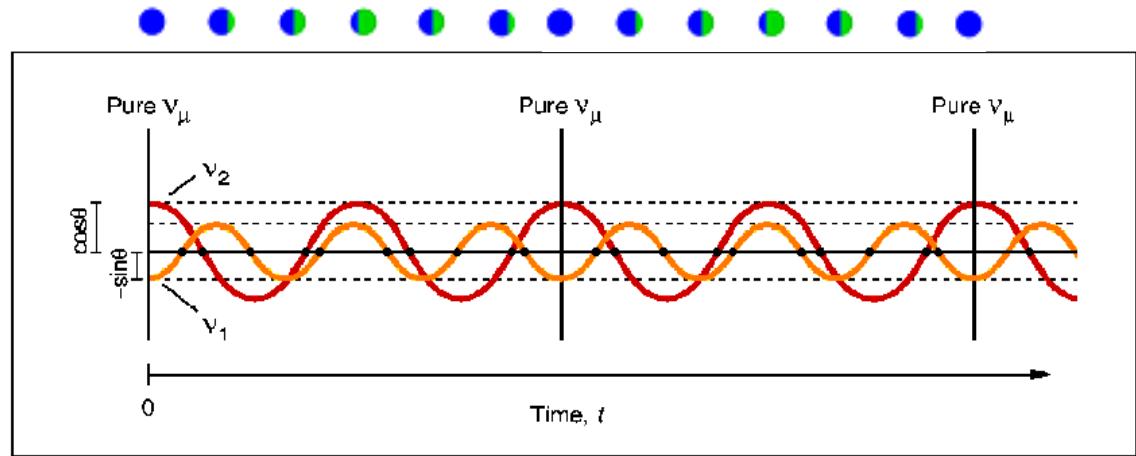
$$\text{max. mixing: } v_e = v_1 - v_2$$

$$v_\mu = v_1 + v_2$$

Neutrino oscillations

2 Neutrinos: ν_e, ν_μ

$$\begin{aligned} |\nu_e(0)\rangle &= \cos\theta |\nu_1\rangle + \sin\theta |\nu_2\rangle \\ |\nu_\mu(0)\rangle &= -\sin\theta |\nu_1\rangle + \cos\theta |\nu_2\rangle \end{aligned}$$



$$|\nu_\mu(t)\rangle = -\sin\theta \exp[-\frac{iE_1t}{\hbar}] |\nu_1\rangle + \cos\theta \exp[-\frac{iE_2t}{\hbar}] |\nu_2\rangle$$

$$E_i = \sqrt{p_i^2 + m_i^2} \xrightarrow{p_i=p \gg m_i} \simeq p + \frac{m_i^2}{2p} \simeq p + \frac{m_i^2}{2E}$$

$$L = c \cdot t \quad \Delta m^2 = m_2^2 - m_1^2 \Rightarrow E_2 - E_1 = \frac{\Delta m^2}{2E}$$

2ν-transition-
probability:

$$P(\nu_\mu \rightarrow \nu_e) = |\langle \nu_\mu(t) | \nu_e(0) \rangle|^2 = \sin^2 2\theta \cdot \sin^2 \left(\frac{\Delta m^2 L}{4E} \right)$$

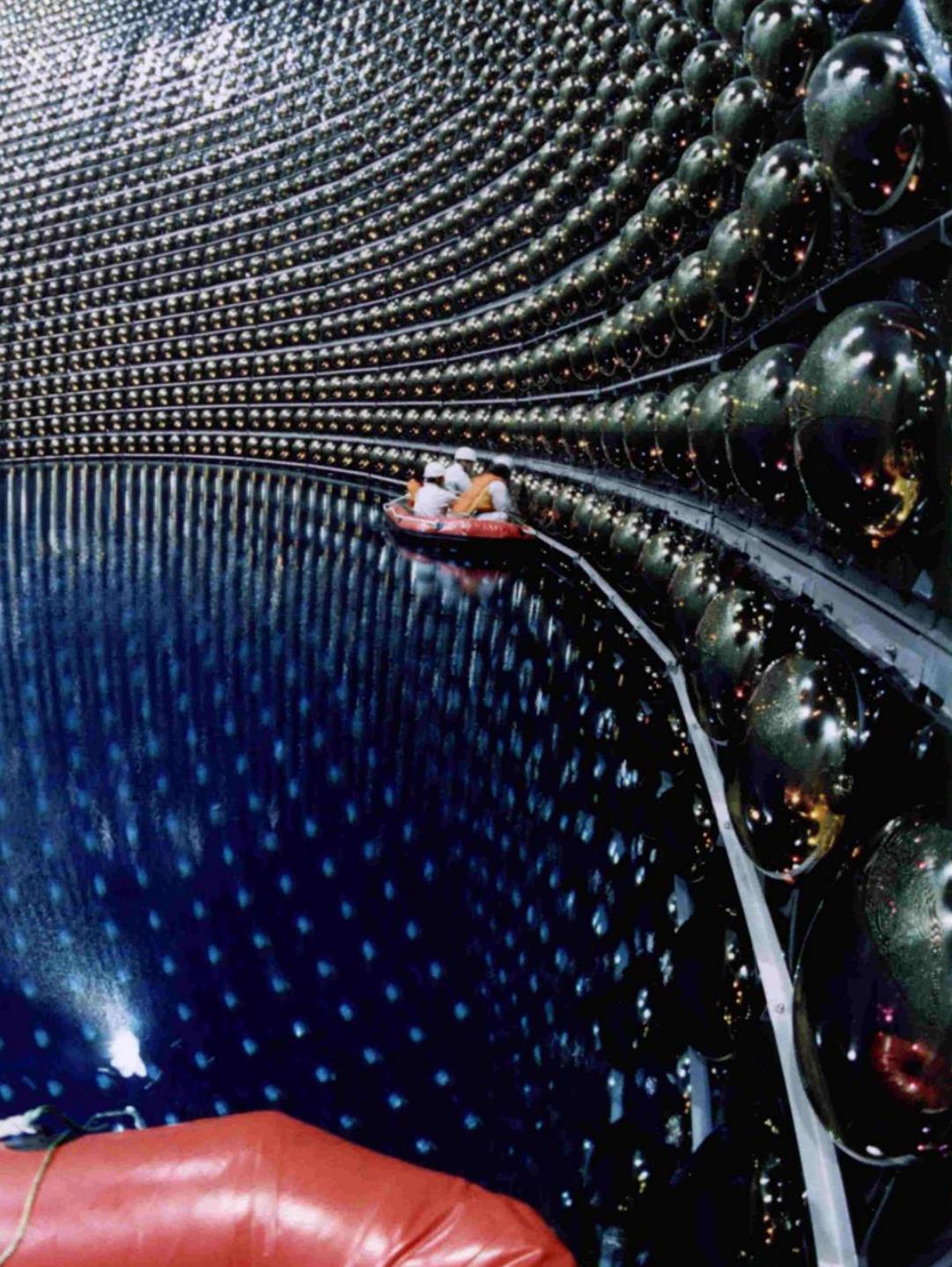
$$= \sin^2 2\theta \cdot \sin^2 (1.27 [\Delta m^2/\text{eV}^2] [L/\text{km}] / [\text{E/GeV}])$$

Super-Kamiokande

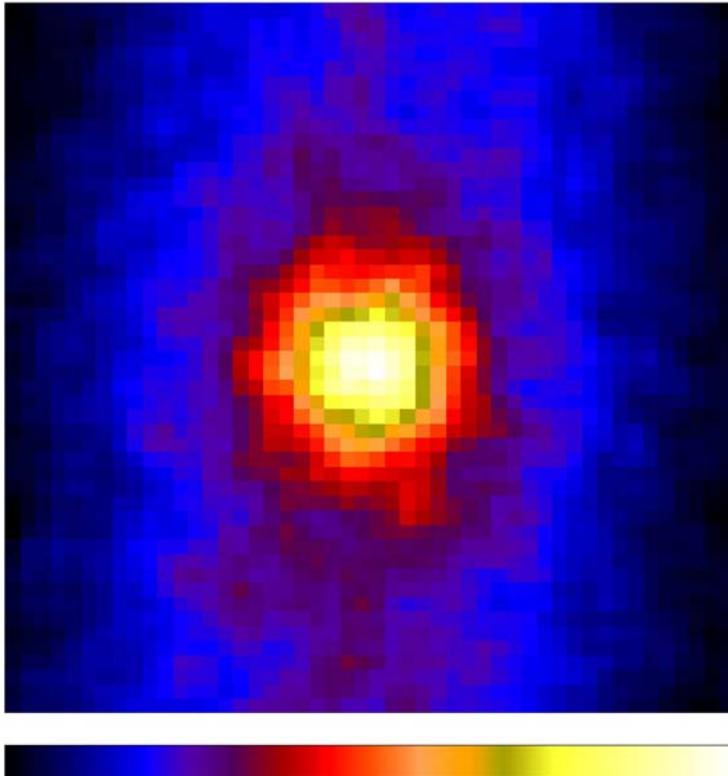
Kamioka mine
Japan :

detect Cerenkov light
in 50.000 t water
12.000 photomultipliers
1000 m underground

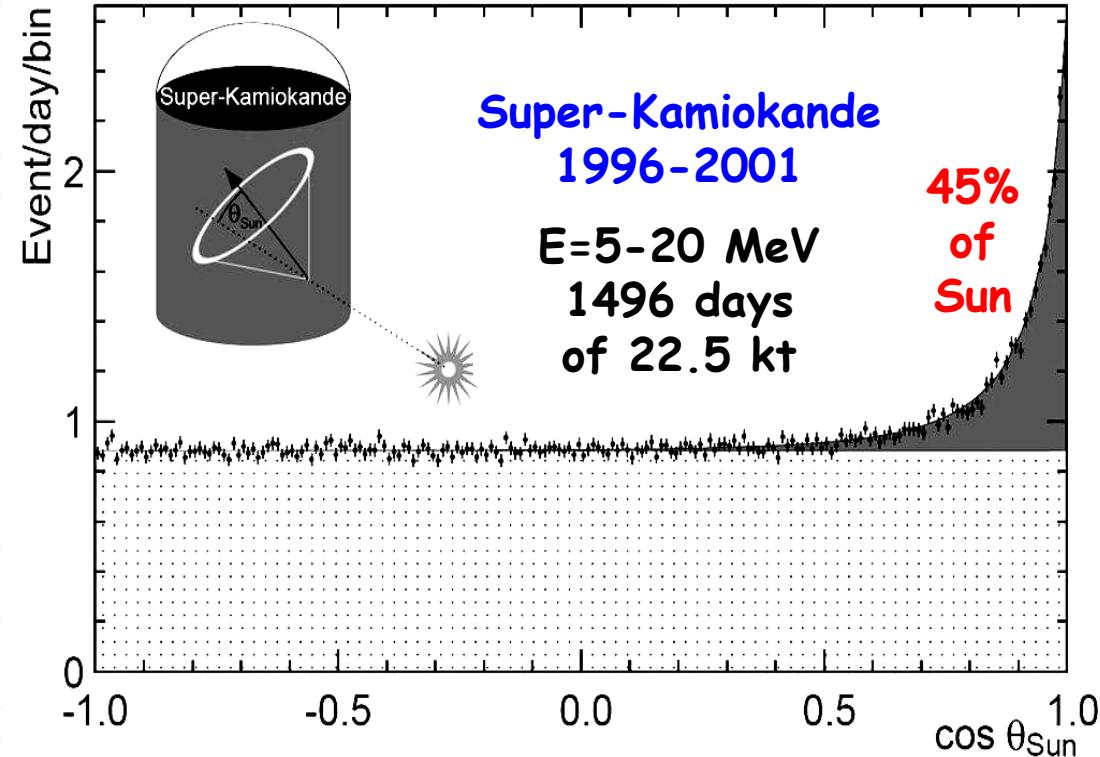
built for p decay
1989:
 $\tau(p \rightarrow \pi^0 e^+) > 10^{33} \text{ a}$



Solar Neutrinos



First Sun neutrino-graphy.
 $90^\circ \times 90^\circ$.
SuperKamiokande 1998.



$45 \pm 2\%$ of Bahcall solar model

where are the
solar neutrinos?

Atmospheric Neutrinos

Super-Kamiokande:

cosmic protons
react in atmosphere:



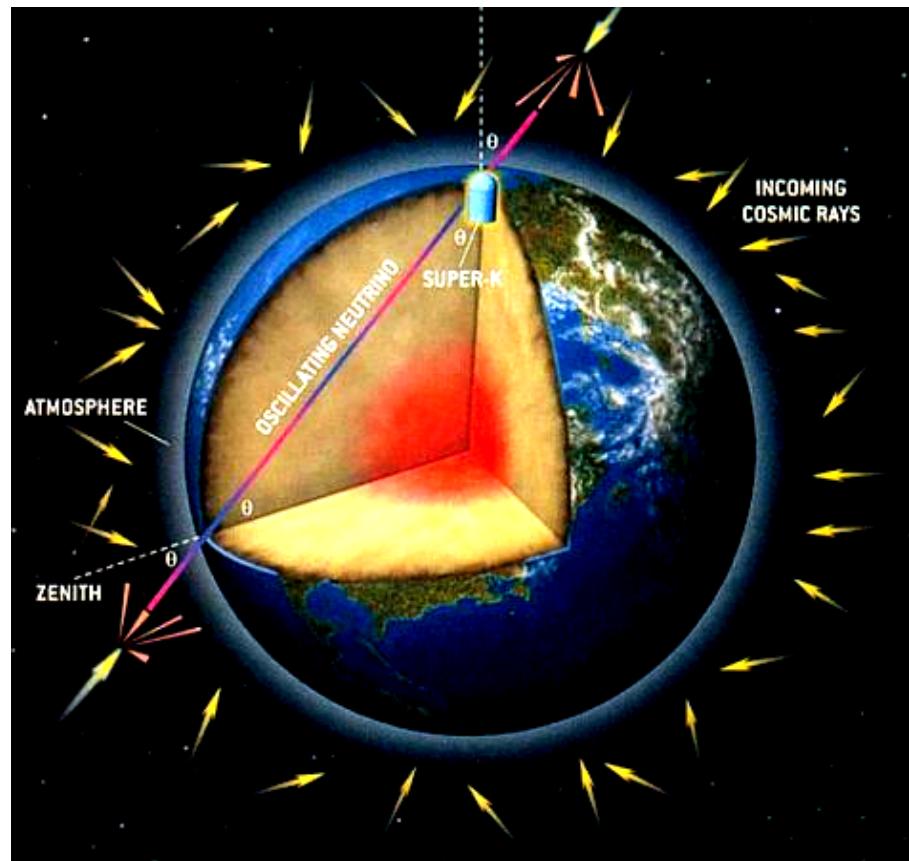
expect $\nu_\mu / \nu_e \sim 2/1$

observe $\sim 1/1$

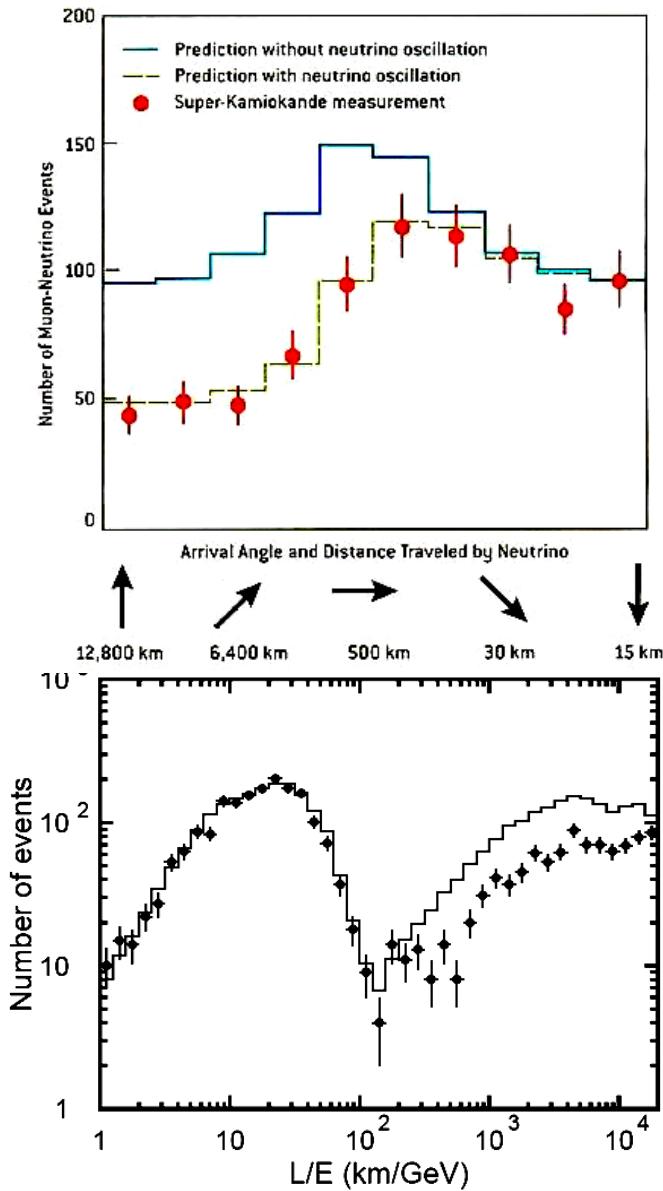
$\nu_\mu \rightarrow \nu_e$ disappear ?!

detect Cerenkov rings

from $\nu_e \rightarrow e$ and $\nu_\mu \rightarrow \mu$



Atmospheric Neutrinos



Super Kamiokande 2004:

rate too small as $f(\theta, E)$!

ν_μ disappear !

ν oscillation with **mass** difference
 $\Delta m_{23}^2 = 2.4 \pm 0.4 \cdot 10^{-3} \text{ eV}^2$

and max. mixing angle:

$\sin^2 2\theta_{23} > 0.92$ @ 90% CL

$\theta_{23} \sim 45^\circ$ max. mixing !

Neutrino Mixing

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ -s_{13}e^{i\delta} & 0 & c_{13} \\ 0 & c_{12} & s_{12} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

($\theta_{13} < 13^\circ$, δ ?)

$\Theta_{23}: 34^\circ - 58^\circ$

$\Theta_{12}: 29^\circ - 39^\circ$

Atmospheric + accelerator
 - MINOS (precision)
 - OPERA (ν_T appearance)

- Is θ_{23} exactly maximal?
- sign of Δm^2_{23} ?

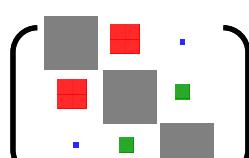
Reactor (CHOOZ)
 - Super+Beta-Beams
 - Neutrino factory

- how small is θ_{13} ?
- CP-violation
Dirac CP-phase?

Solar + reactor

- θ_{12} how far from max.?
- test MSW effect

Quarks: $V_{CKM} = \begin{pmatrix} 1.0 & 0.2 & 0.001 \\ 0.2 & 1.0 & 0.01 \\ 0.001 & 0.01 & 1.0 \end{pmatrix}$



Neutrinos: $V_{MNS} = \begin{pmatrix} 0.8 & 0.5 & 0.2 \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix}$

Why so different?

Solar neutrinos: SNO

Sudbury Neutrino Observatory, Canada

lend ~1 reactor load ~ 1000 t heavy water ~ 100 M\$

CC: $\nu_e d \rightarrow p p e$ get ν_e flux + see difference

NC: $\nu d \rightarrow p n \nu$ get ν_{total} flux + verify solar models

CC: e in water Cerenkov, detect by photomultipliers

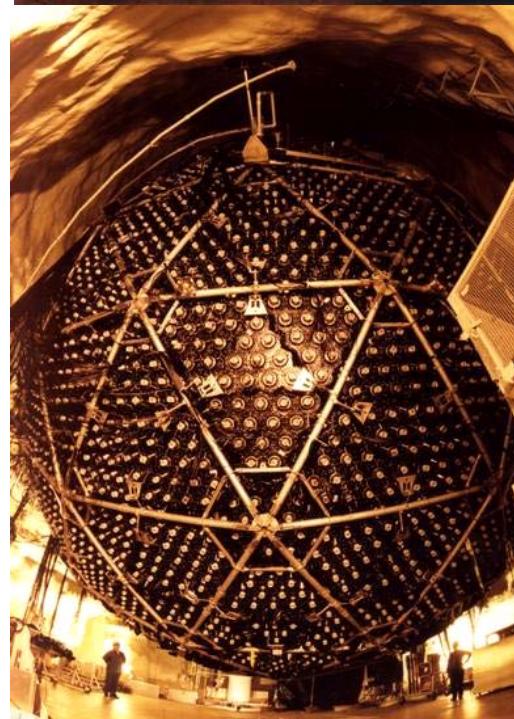
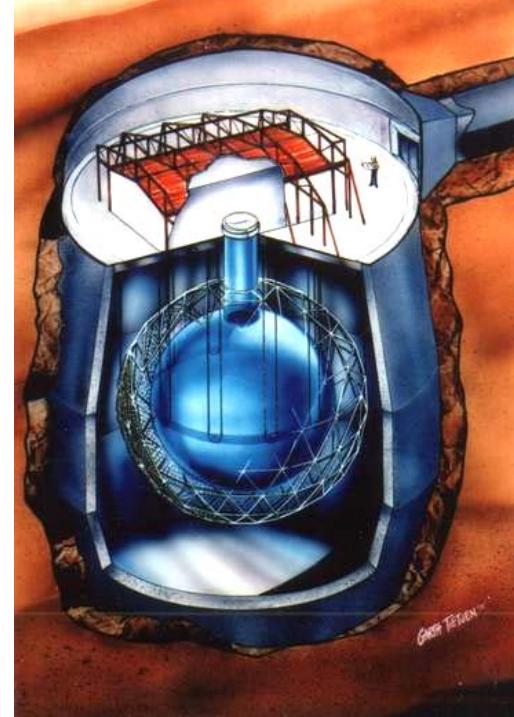
NC: n capture in salt: $n^{35}\text{Cl} \rightarrow ^{36}\text{Cl} \rightarrow \text{Cl} + \gamma's$ (8MeV)

~10 evts/s radioactive background

~30 evts/day solar neutrinos

2002-4: Sun ok , but

$$\nu_e \sim \nu_{\text{all}} / 3$$



Neutrino oscillation results

solar + reactor neutrinos (low E):

$$\Delta m_{12}^2 = 8.0 \pm 0.5 \cdot 10^{-5} \text{ eV}^2$$

$m(\nu_\mu) = 9 \text{ meV} \sim 10^{-7} \text{ m}(\mu)$ in simplest mass hierarchy

large mixing angle $\tan^2 \theta_{12} = 0.4 \pm 0.1$

atmospheric + accelerator neutrinos (high E): ν_μ disappear

$$\Delta m_{23}^2 = 2.4 \pm 0.4 \cdot 10^{-3} \text{ eV}^2$$

max. mixing angle $\sin^2 2\theta_{23} > 0.92$ @ 90% CL

min. mixing angle $\sin^2 2\theta_{13} < 0.09$ @ 90% CL fitted

quark mixing: CKM matrix - lepton mixing matrix

more expts:

FermiLab, USA: MiniBoonE, MINOS

KEK, Japan: K2K, KamLand (reactor)

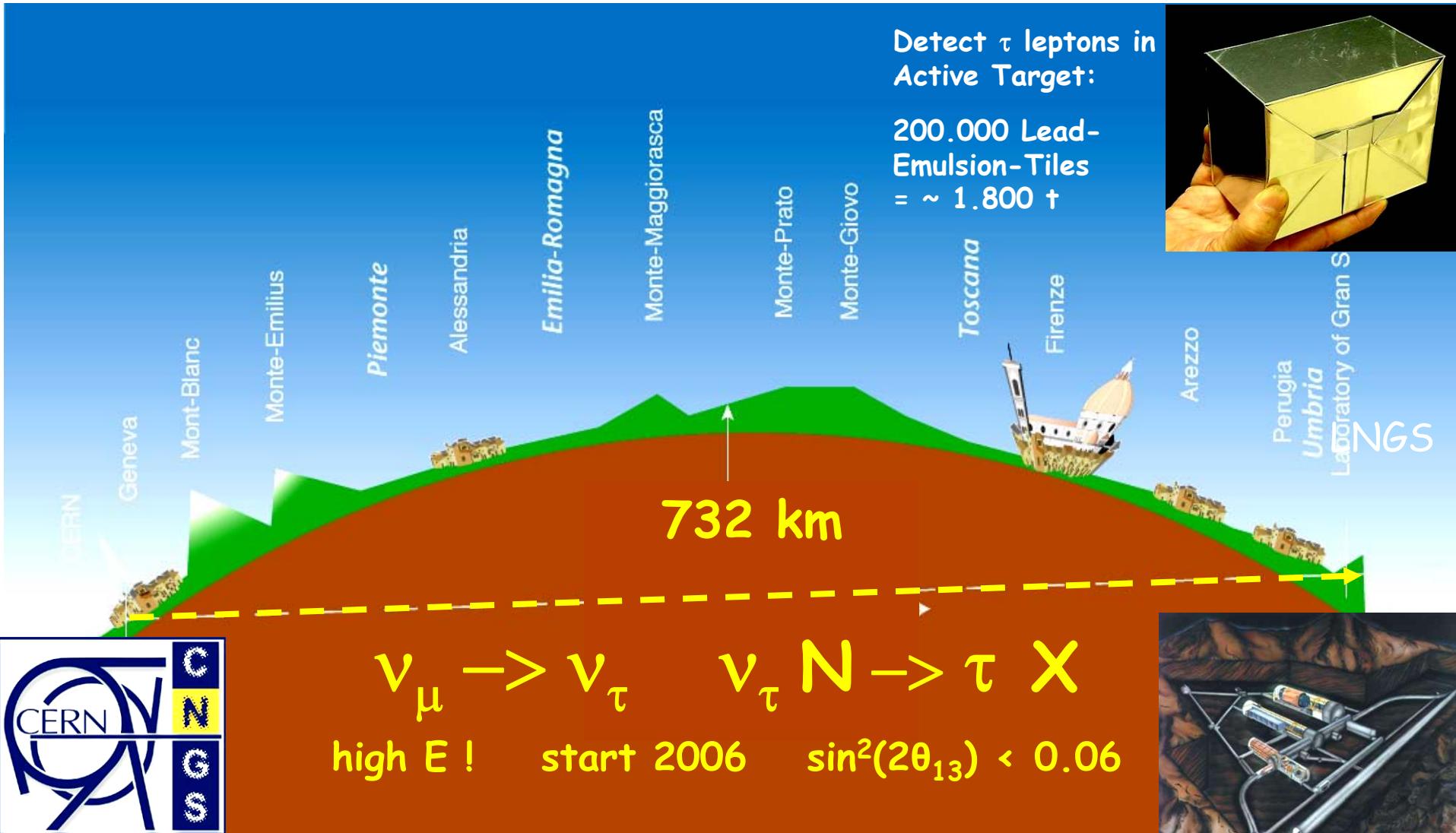
France: Double-CHOOZ (reactor) θ_{13}

CERN, Geneva: CERN-Gran Sasso OPERA $\nu_\mu - \nu_\tau$ appearance



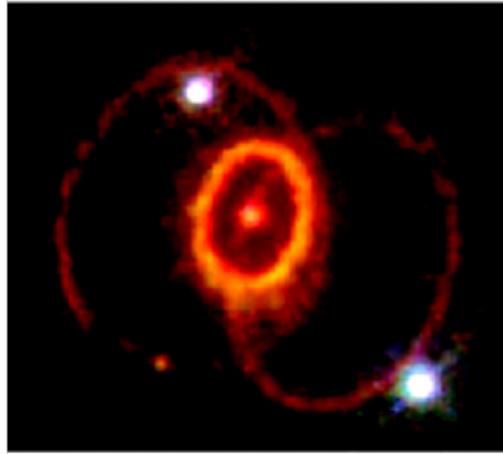
ν_τ appearance: CNGS+OPERA

CERN Neutrino ν_μ beam to Gran Sasso Underground Lab near Rome

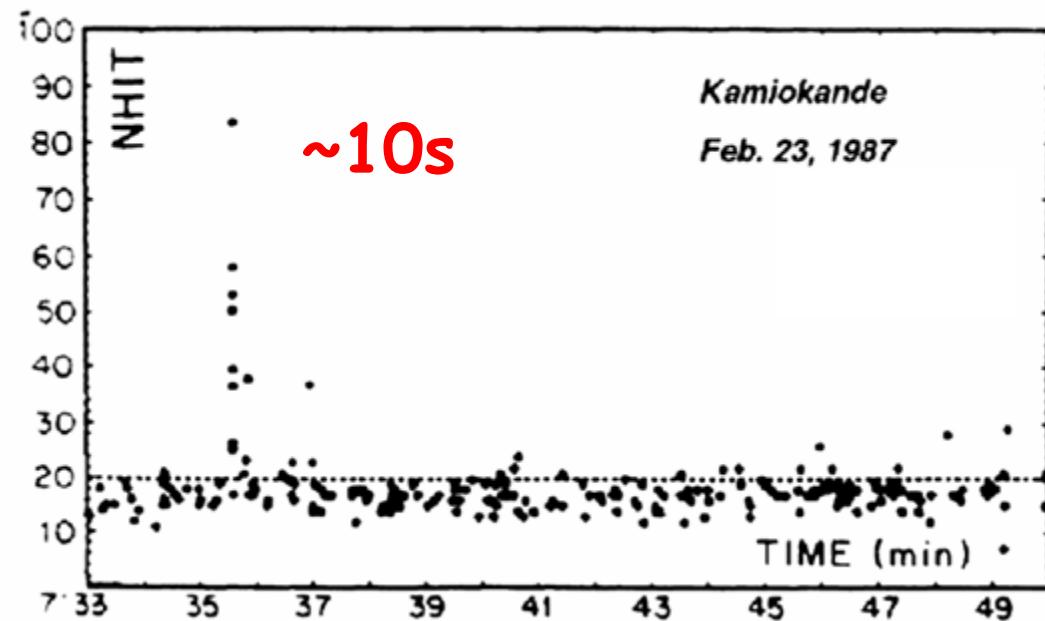


Cosmic Neutrinos

Supernova SN 1987A: Magellan cloud, 150 million ly
 10^{57} ν total, 10^{15} ν/m²

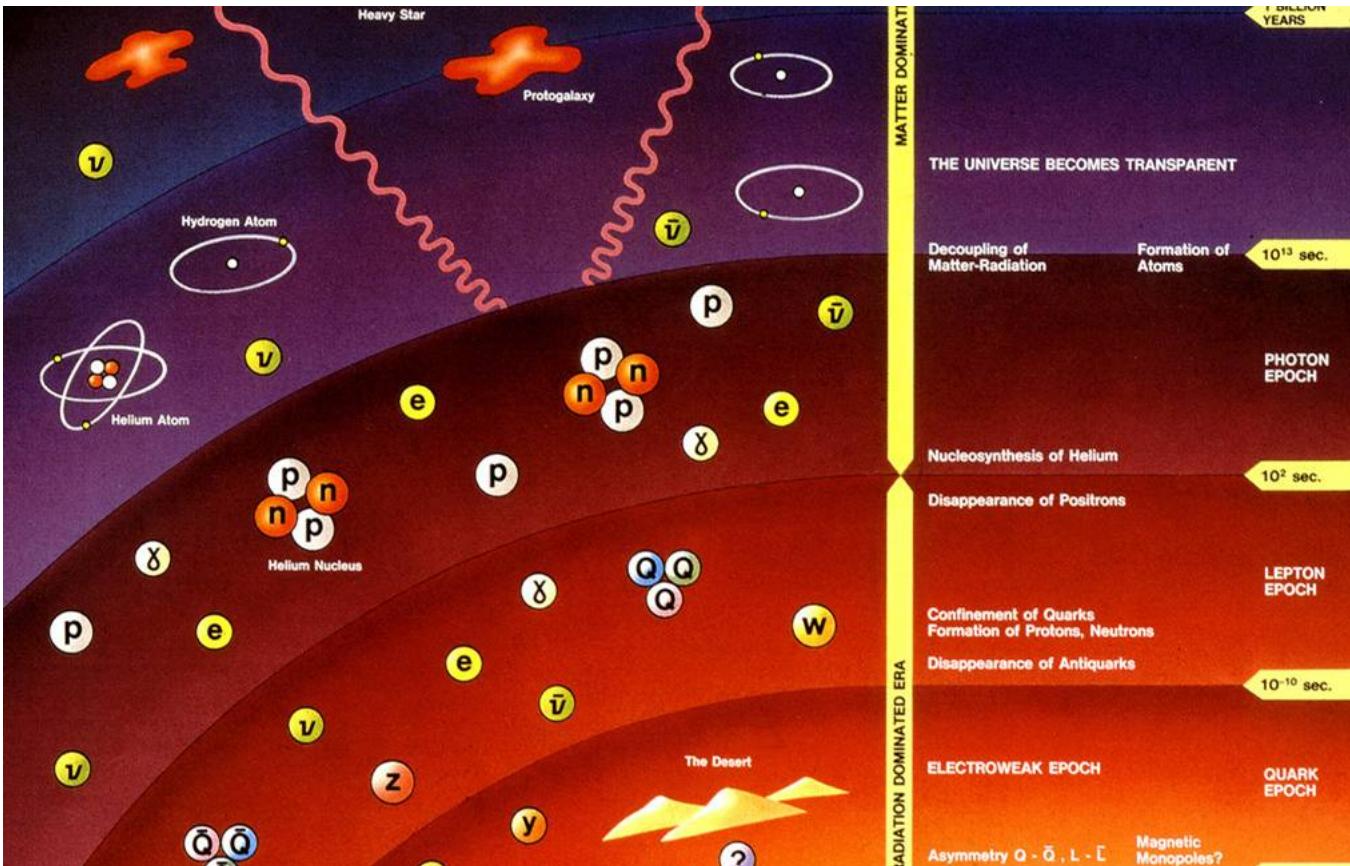


2002



First Supernova visible by naked eye since Kepler 1604 !

Big Bang Neutrinos



now: γ : 2.7 K background radiation $410/\text{cm}^3$

ν : 1.9 K background radiation $335/\text{cm}^3$

Penzias, Wilson 1964

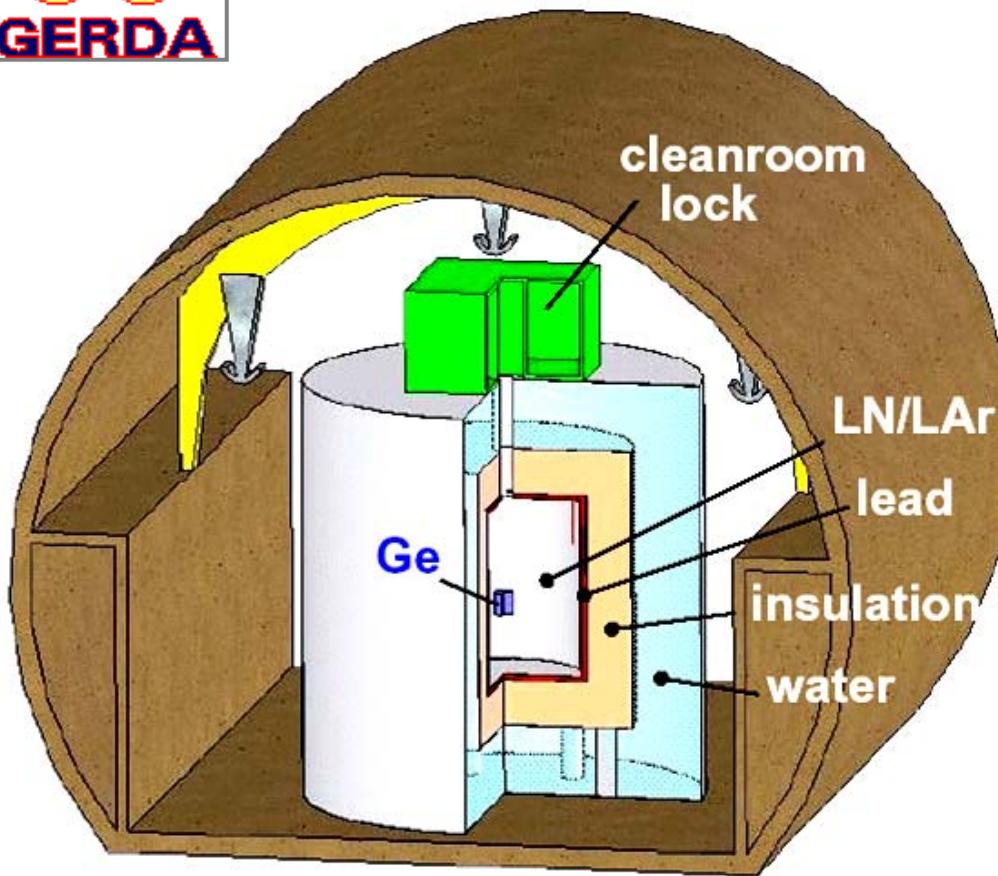
Nobel



1978



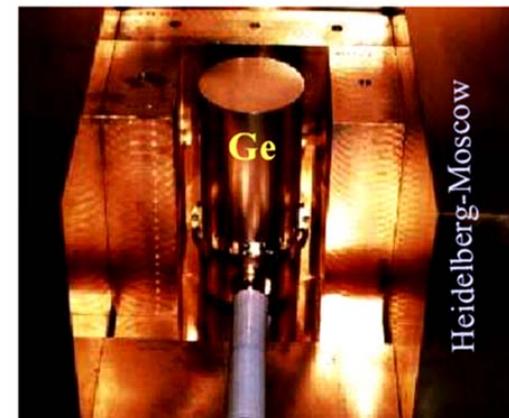
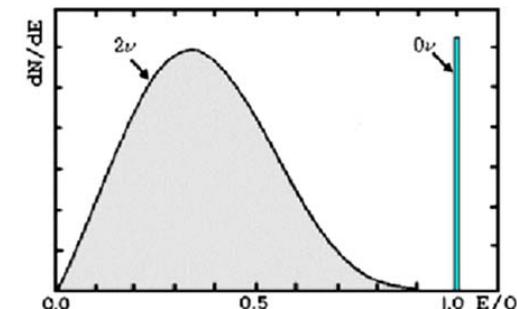
GERmanium Detector Array



phase 1: 40 kg
phase 2: 500 kg

high purity
86% enriched ^{76}Ge diodes
in cryogenic fluid shield

$$Q_{\beta\beta} = 2039 \text{ keV:}$$



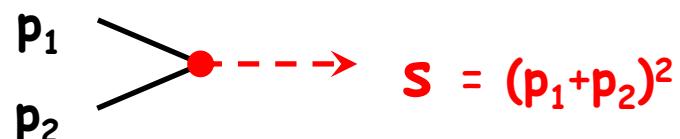
Heidelberg-Moscow expt.

Weak Interactions

Kinematics

- four-momentum : $\mathbf{p} = (E, \mathbf{p})$ $c=1$
- four-momentum² : relativistic invariant
- effective mass² : $m^2 := p^2 = E^2 - \mathbf{p}^2$
- ultra-relativist.: $m \ll E \quad E = p$
- classic: $p \ll m \quad E = m$
- $s = \text{invariant reaction energy}^2$

omit $p_x = p_y = 0$



- Center-of-Mass System : $e^- \xrightarrow{\hspace{2cm}} \xleftarrow{\hspace{2cm}} e^+$

$$\mathbf{p}_1 + \mathbf{p}_2 = (2E, 0)$$

$$s = 4E^2$$

- Lab system :

$$m_{1,2} \ll E_1$$

$$1 \xrightarrow{\hspace{2cm}} \bullet 2$$

$$(E_1, \mathbf{p}_1) \quad (m_2, 0)$$

$$\mathbf{p}_1 + \mathbf{p}_2 = (E_1 + m_2, E_1)$$

$$s = 2m_2E_1$$

target mass effect: $m_p = 2000 m_e$

Weak Interaction

$$\sigma(v_\mu e^- \rightarrow v_e \mu^-) = G_F^2 s / \pi$$

$$G_F = 1.2 \times 10^{-5} \text{ GeV}^{-2}$$

$$s = 2 m_e E_\nu = 1 \text{ MeV } E_\nu = 10^{-3} \text{ GeV}^2 [E_\nu/\text{GeV}]$$

$$\sigma = 4 \times 10^{-14} \text{ GeV}^{-2} [E_\nu/\text{GeV}]$$

$$\hbar c = 1 = 0.2 \text{ GeV fm}$$

$$(\hbar c)^2 = 1 = 0.04 \text{ GeV}^2 \text{ fm}^2 = 0.4 \text{ mb GeV}^2$$

$$\sigma = 1.6 \times 10^{-14} \text{ mb } [E_\nu/\text{GeV}] = 10^{-16} \sigma (\text{strong: } \pi p \rightarrow \Delta)$$

$$\sigma = 1.6 \times 10^{-15} \text{ fm}^2 [E_\nu/\text{GeV}]$$

$$\sigma = 1.6 \times 10^{-45} \text{ m}^2 [E_\nu/\text{GeV}] = \text{extremely weak !}$$

$m_N / m_e = 2.000 \Rightarrow \text{target mass effect !}$

mean free path $\sim 10^{13} \text{ km} \sim \text{light years:}$



1 billion earths or 10 million suns

Weak Interaction

$$[G_F] = \text{GeV}^{-2}$$

Which energy scale hidden in G_F ?

$$G_F^2 s / \pi = \sigma = \pi R^2$$

$R \sqrt{s} = \hbar c = 1$ Uncertainty relation

$$G_F^2 s / \pi = \sigma = \pi / s \quad 1/s : \text{point like interaction}$$

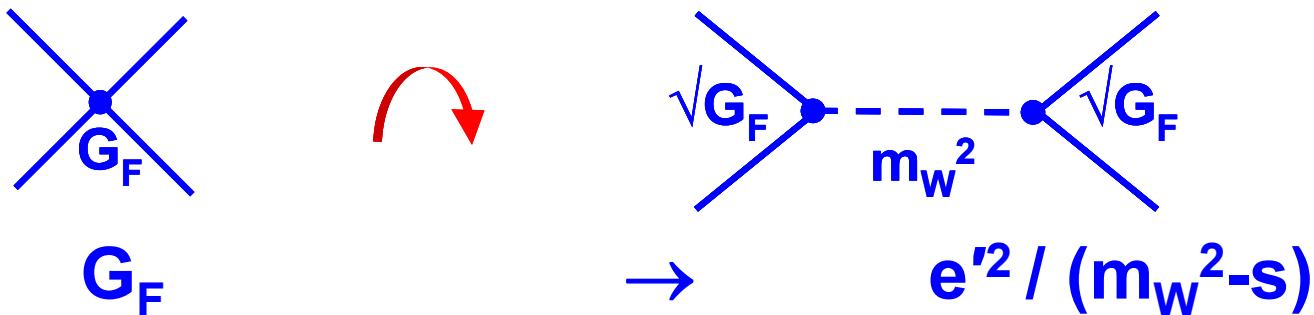
$$s^2 = \pi^2 / G_F^2$$

$$\sqrt{G_F^{-1}} \sim \sqrt{10^5} \text{ GeV} \sim 300 \text{ GeV}$$

collapse of Fermi theory

transition to
massive exchange boson W

Weak Interaction

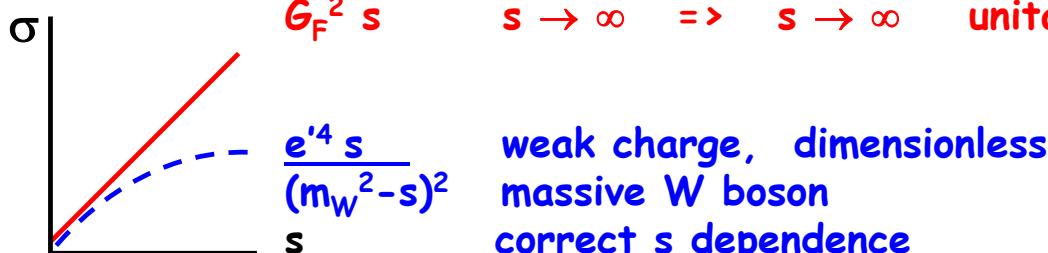


e' ... weak charge, dimensionless
massive W boson, correct s dependence.

- $s \rightarrow 0$: $e'^2 = m_W^2 G_F = e^2 = 4\pi\alpha \sim 0.1$
weak \sim electric force !

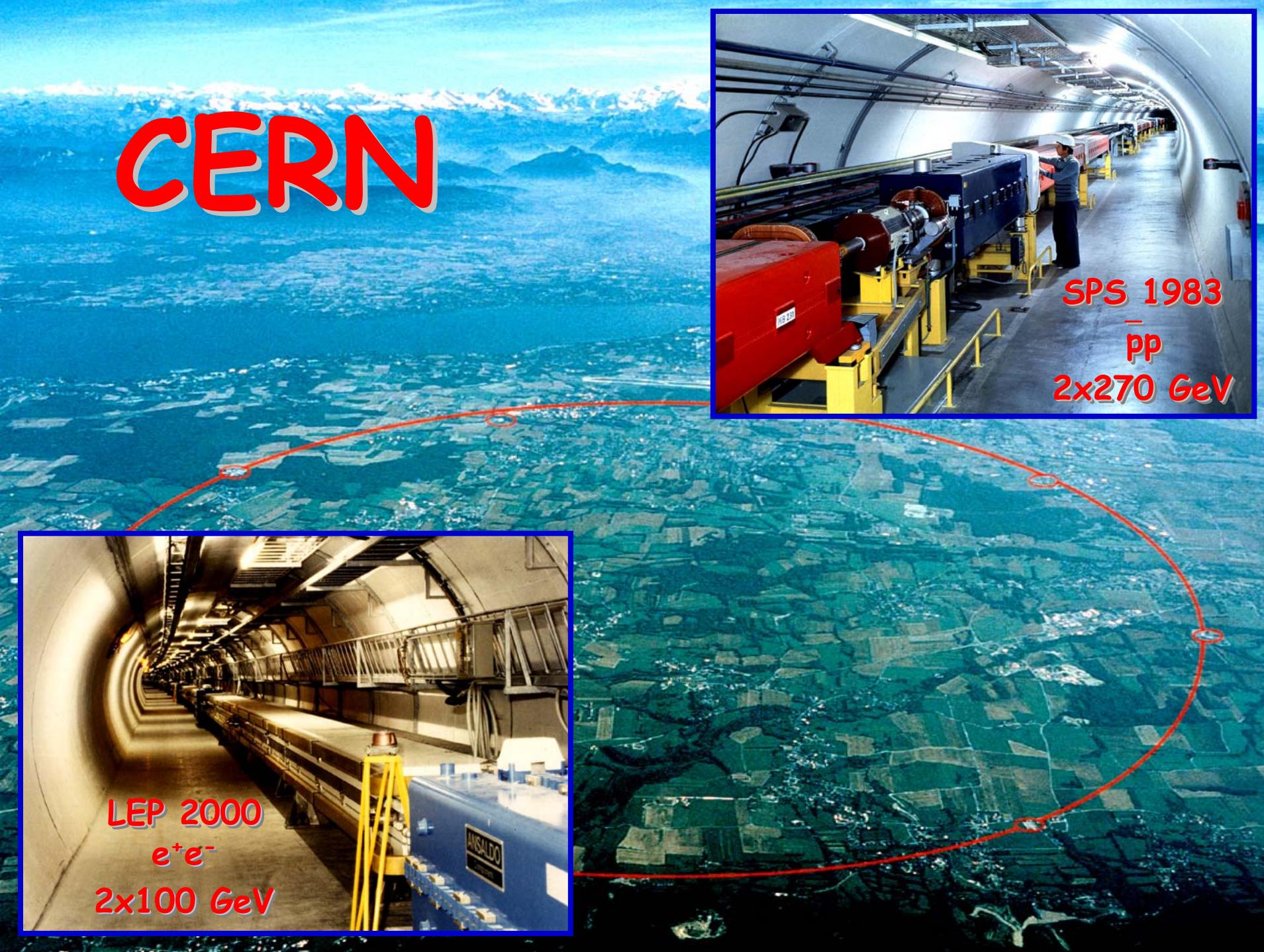
$$m_W^2 = 4\pi\alpha / G_F \quad m_W \sim 90 \text{ GeV}$$

- $s \gg m_W^2$: $\sigma \sim e'^4 / s$ scatter on pointlike charge

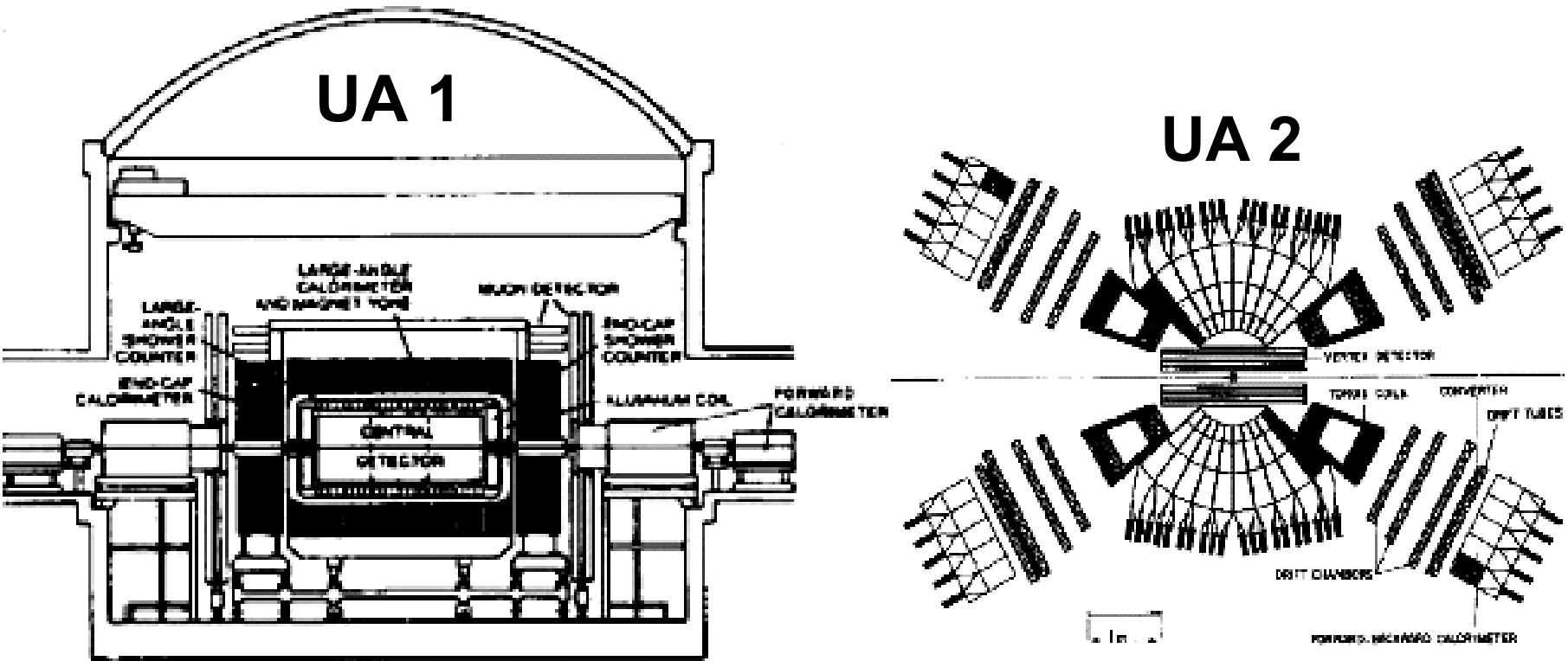


~ 1970 :
find the
 $W^\pm + Z^0$ bosons !

CERN



W+Z Discovery



Expts. UA1+UA2 1983
C. Rubbia



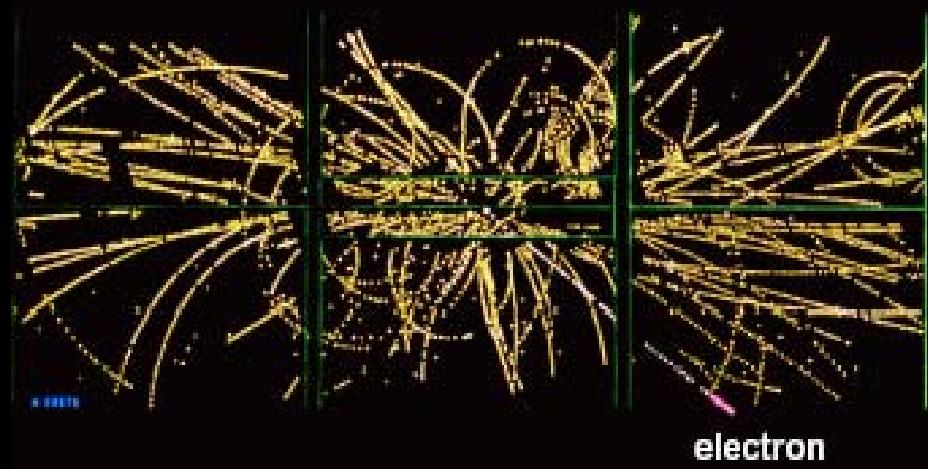
Nobel prize 1984

SPS p \bar{p} collider at CERN
S. van der Meer

W+Z Discovery

EVENT 2004, 179.

W Event in UA1:



Z Event in UA1:



$$u \bar{d} \rightarrow W^+ \rightarrow e^+ \nu_e$$

$$q \bar{q} \rightarrow Z^0 \rightarrow e^+ e^-$$

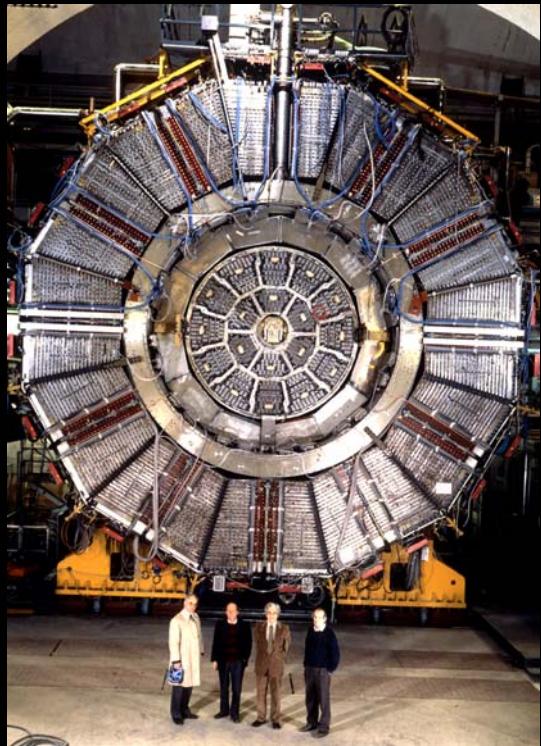
Expts. UA1+UA2 1983

C. Rubbia

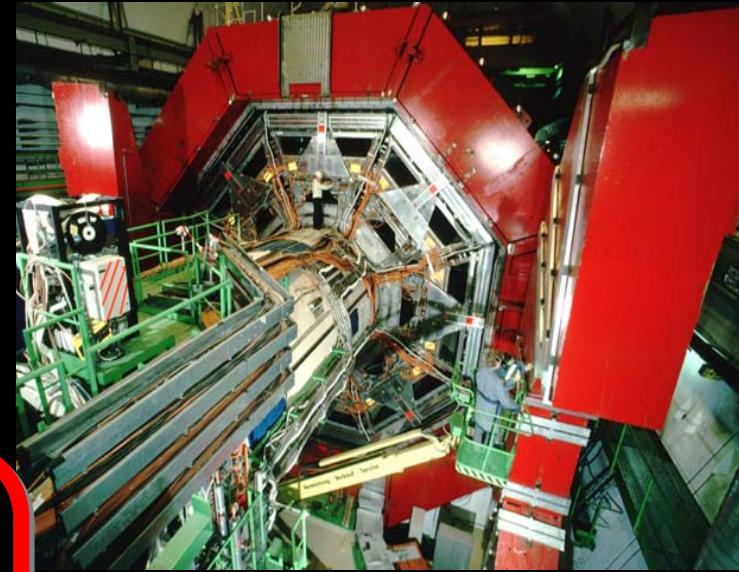


SPS p \bar{p} collider at CERN
S. van der Meer

Nobel prize 1984



ALEPH

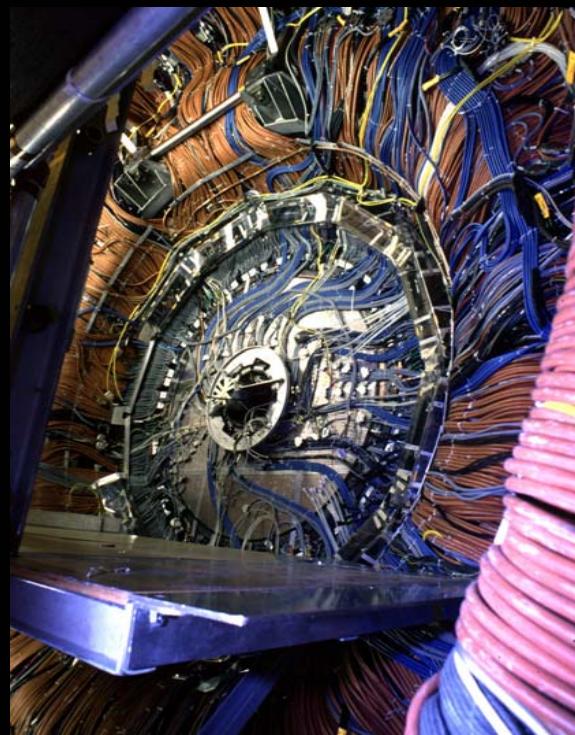


L3

LEP
detect
ors



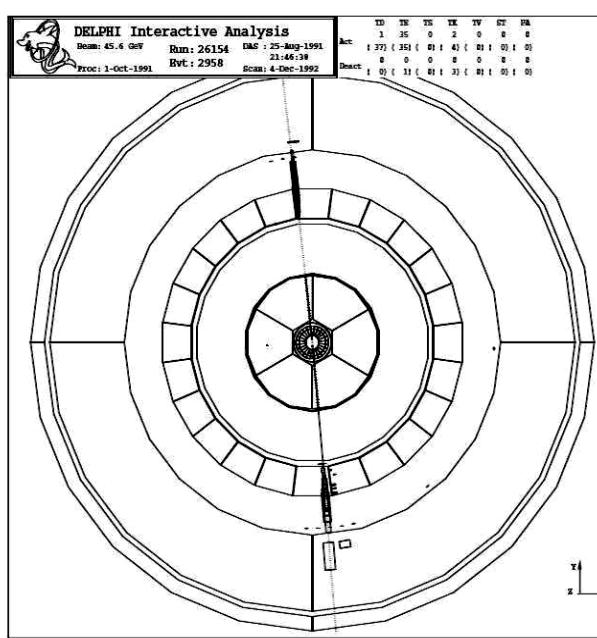
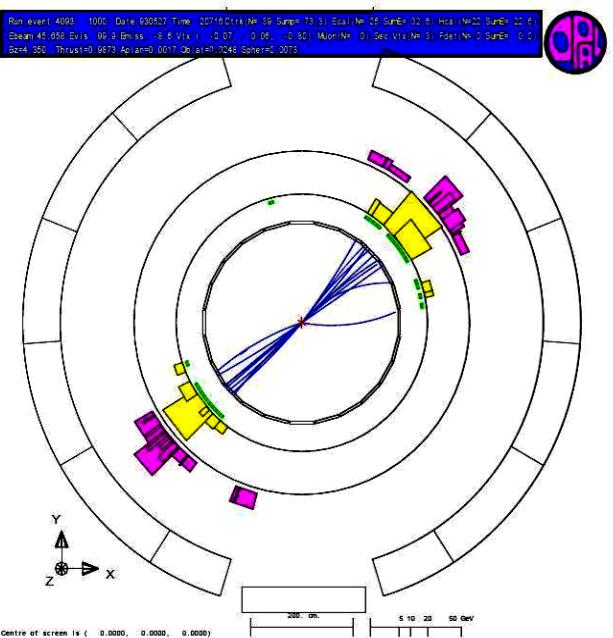
OPAL



DELPHI

$q\bar{q}$

OPAL



e^+e^-

DELPHI

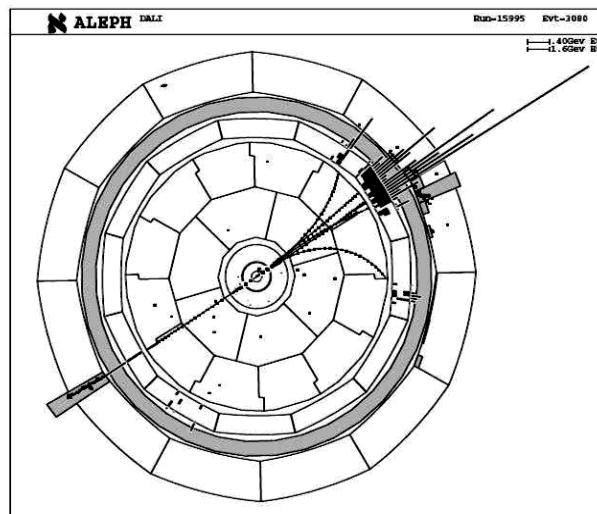
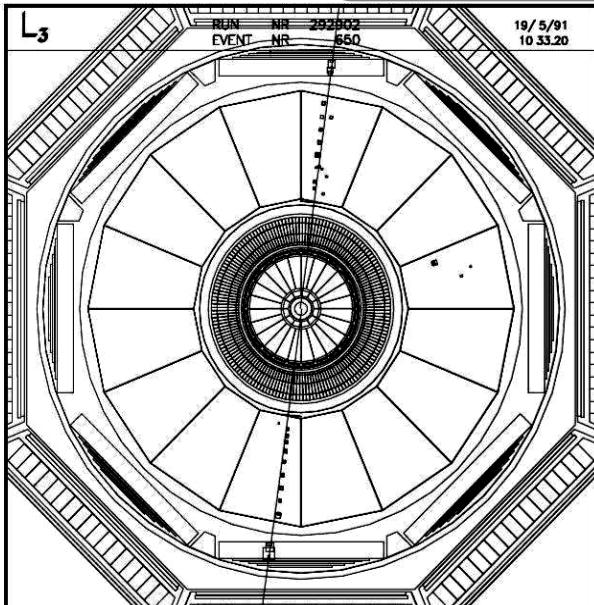
LEP

Z decays

CERN

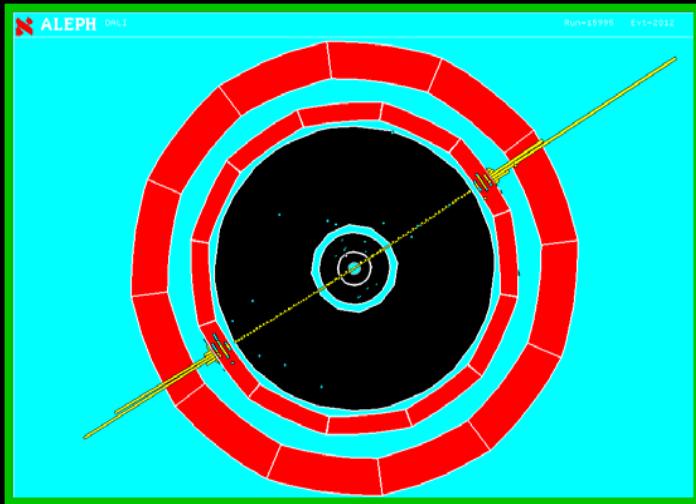
L3

$\mu^+\mu^-$

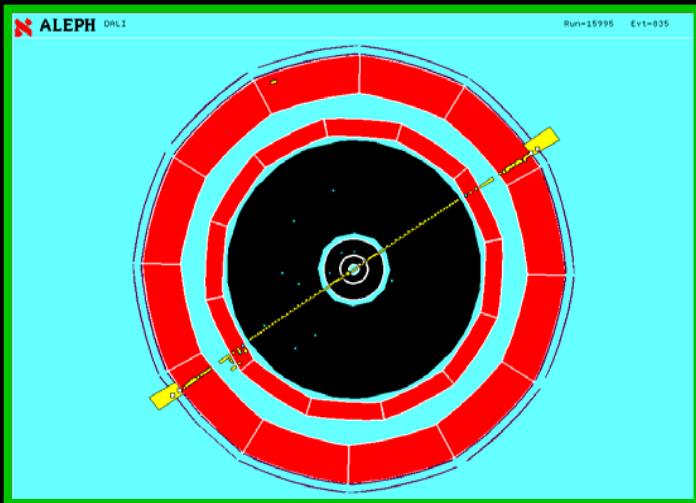


ALEPH

$\tau^+\tau^-$



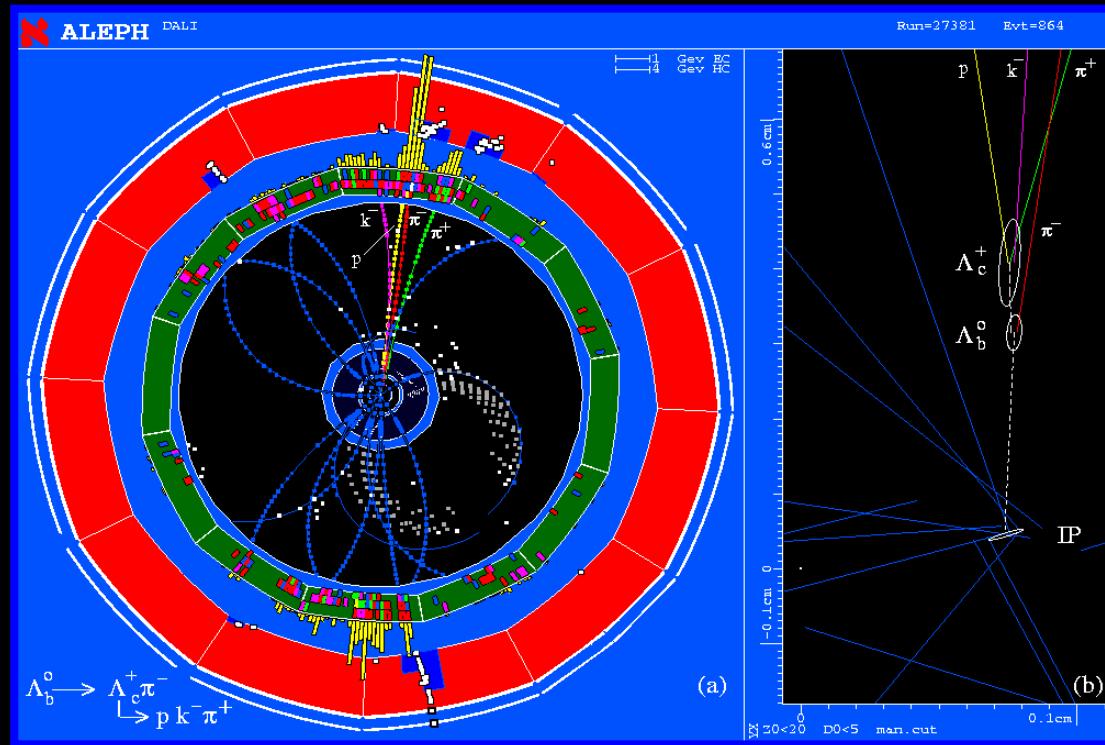
$Z^0 \rightarrow e^+ e^-$
leptonic
 $Z^0 \rightarrow \mu^+ \mu^-$



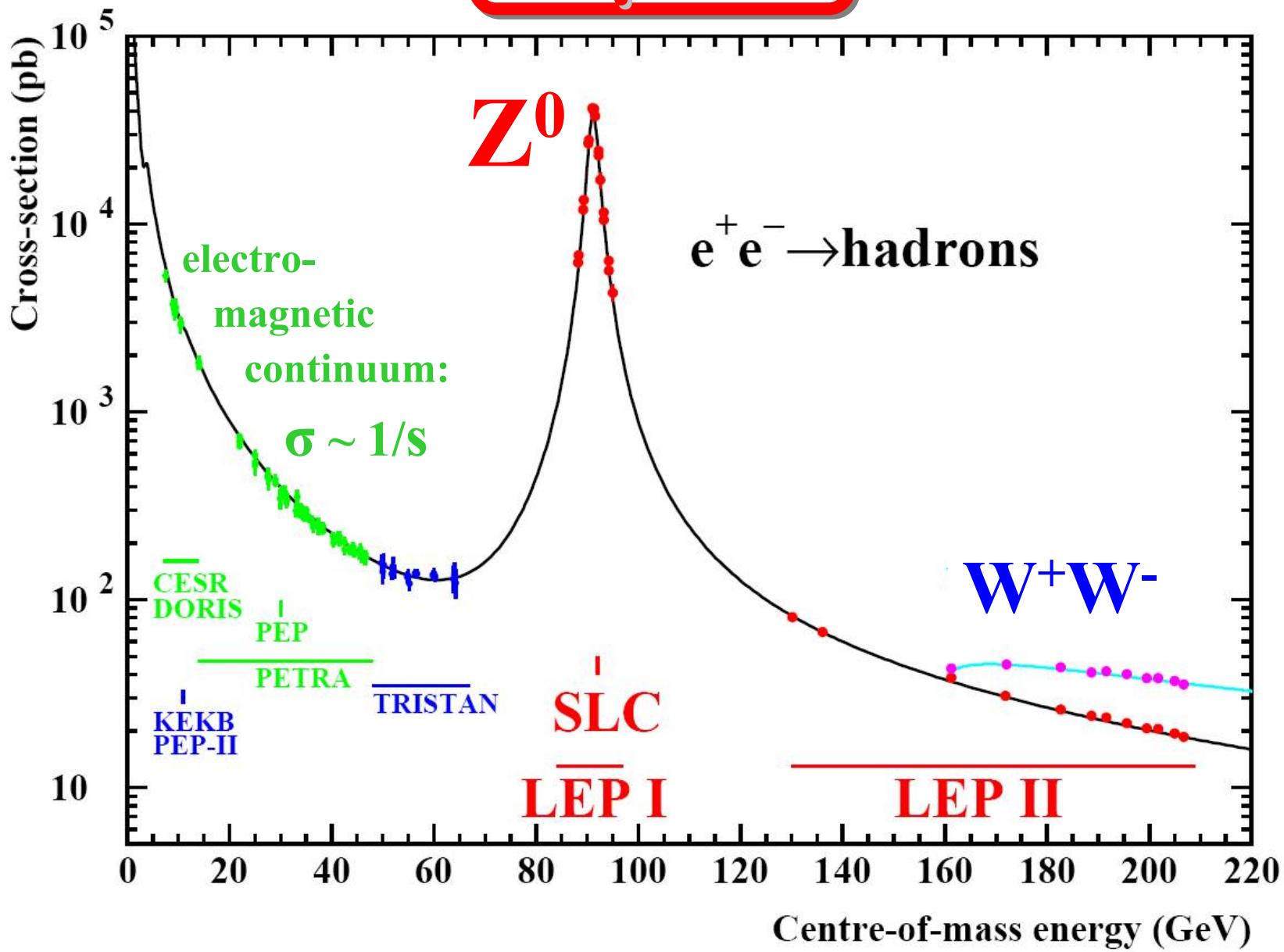
Z decays

ALEPH @ LEP

hadronic
 $Z^0 \rightarrow \bar{b}b$



Z peak



5. Gauge Theories

for an overview, see e.g.:

A. Pich, The Standard Model of Electroweak Interactions, <http://arxiv.org/pdf/hep-ph/0502010>

The LEP Electroweak Working Group, <http://arxiv.org/pdf/hep-ex/0511027> , <http://arxiv.org/pdf/hep-ex/0509008>

Higgs Mechanism

- Fermion mass term gauge invariant :

$$m \psi'^* \psi' = m \psi^* \psi$$

- Boson mass term gauge violating :

$$\begin{aligned} m^2 A'_\mu A'^\mu &= m^2 (A_\mu + \partial_\mu \alpha)(A^\mu + \partial^\mu \alpha) = \\ m^2 (A_\mu A^\mu + \dots) \end{aligned}$$

- introduce auxiliary scalar boson field:

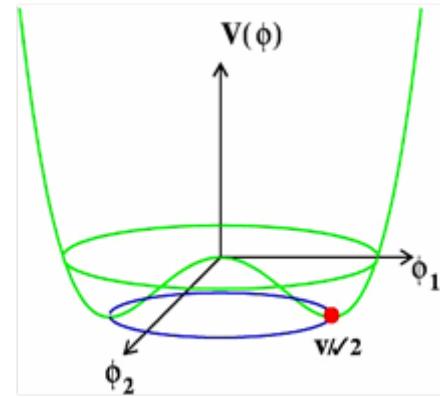
$$\Phi = v \exp(i \xi(x_\mu))$$

- specify gauge transformation:

$$A'_\mu = A_\mu + 1/gv \partial_\mu \xi(x_\mu)$$

now :

- **cancel** gauge viol. terms of vector boson mass term (long. polar.) with ξ degrees of freedom of Higgs field :
- **vector boson mass term gauge invariant !**
- **Higgs boson mass term: witness of trick !**



- Higgs produces **BOSON** masses:
- mass term for 2 W bosons in Lagrangian:

Higgs in $SU(2)$

$$g^2 v^2 / 2 A_\mu A^\mu = (2m_W)^2 / 2 A_\mu A^\mu$$

$$v = 2m_W/g = (\sqrt{2} G_F)^{-1/2} = 246 \text{ GeV}$$

v ... 1 parameter of Higgs potential = scale of electroweak unification !

$U(1)$ toy model $\rightarrow SU(2)$ electroweak Higgs doublet (Φ^+, Φ^0)

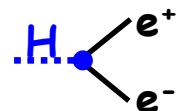
$$\mathcal{L}_H = G_e \left\{ (\bar{\nu}_e, e^+)_L \begin{pmatrix} \Phi^+ \\ \Phi^0 \end{pmatrix} e^-_R + e^+_R (\Phi^-, \Phi^0) \begin{pmatrix} \nu_e \\ e^- \end{pmatrix} \right\} \quad (\text{P violation, no } \nu_R)$$

spontaneous symmetry breaking - only 1 real excitation remains:

$$\Phi = \begin{pmatrix} 0 \\ v + \eta(x) \end{pmatrix} e^{i\xi(x)} \quad (\text{no neutrino mass})$$

$$\mathcal{L}_H = G_e v \ e^- e^+ + G_e \ e^- \eta \ e^+$$

$$G_e v := m_e$$

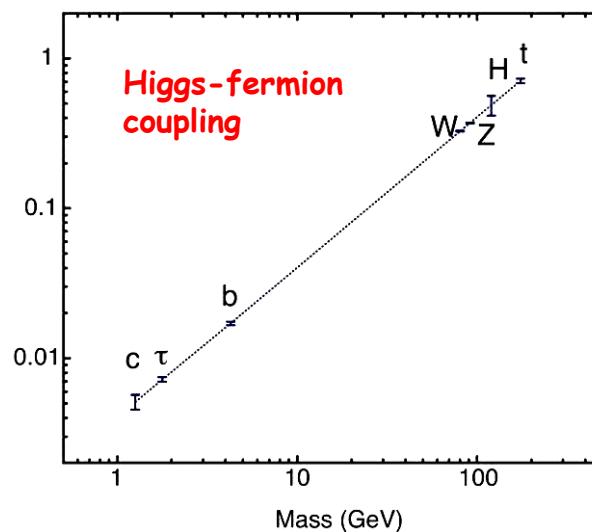


$$= m_e \ e^- e^+ + m_e/v \ e^- \eta \ e^+$$

- Higgs produces **FERMION** masses !
- Higgs-fermion coupling \sim fermion masses:

$$m_e/v = 511 \text{ keV} / 246 \text{ GeV} \sim 10^{-6}$$

$$m_t/v = 172 \text{ GeV} / 246 \text{ GeV} \sim 1$$



Higgs boson

The discovery of a Higgs boson will be of profound importance.
It is one of the most important experimental results of all time.

The Higgs boson is a new kind of matter:

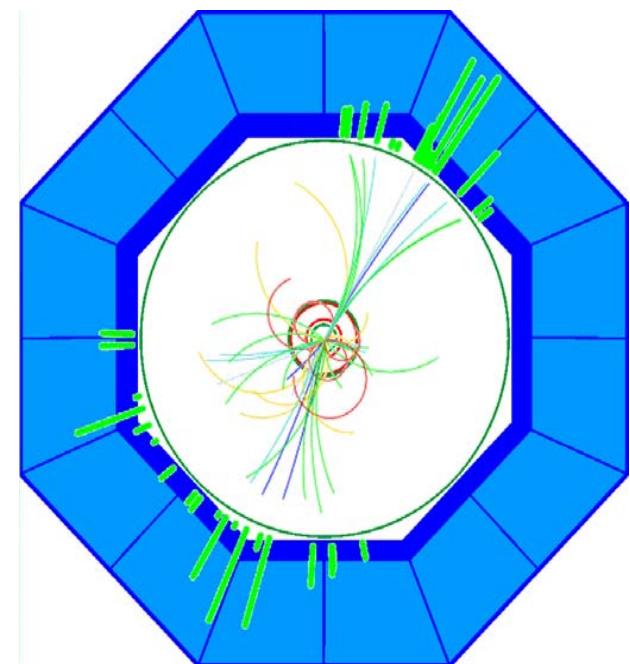
New **quarks and leptons** have been found, but they are all particles like the electron that carry charges.

New **force-mediating bosons** have been found but they are all quanta like the photon.

The Higgs boson is the quantum of a new kind of field,

for which the **energy density of the universe** is lower when the field has a non-zero value, and who

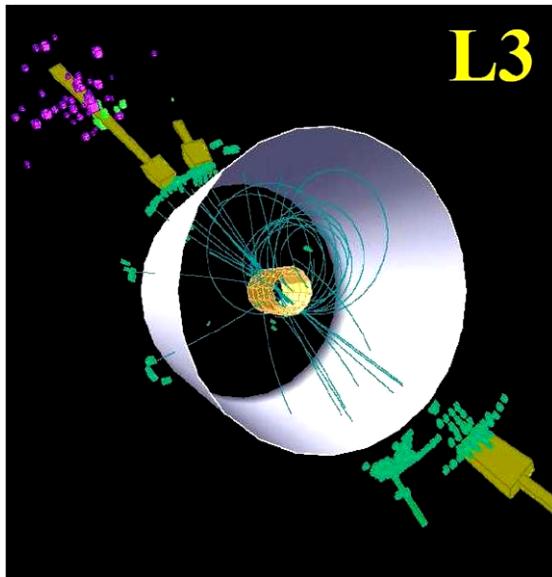
gives mass to bosons and fermions through its interactions.



simulated Higgs decays

LEP Higgs candidates

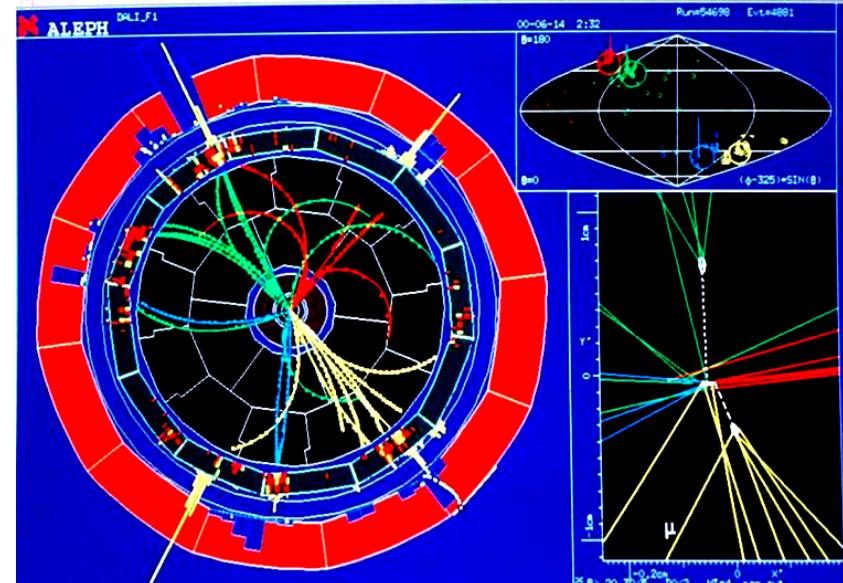
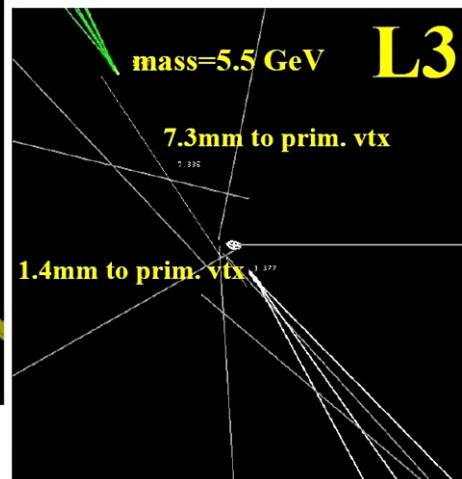
$e^+e^- \rightarrow H Z$



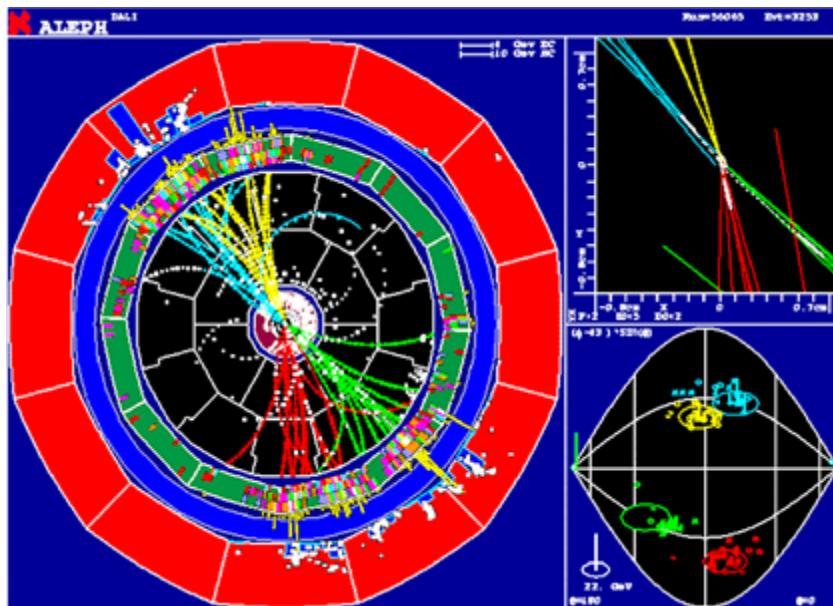
measured H mass=115 GeV
H mass resolution ~3 GeV

$H \rightarrow bb$
 $Z \rightarrow vv$

Secondary vtx's view

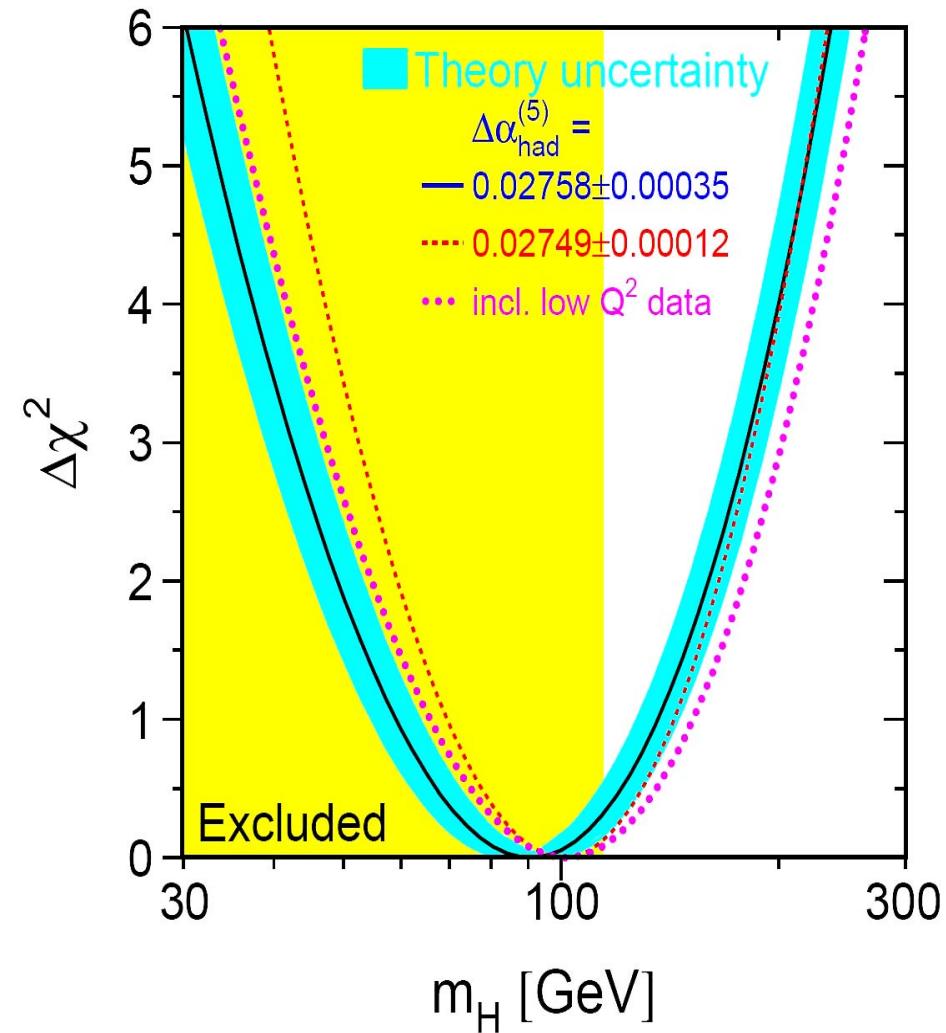


$H \rightarrow 4 \text{ jets}, 2 \text{ b tags}, m_H = 114.3 \text{ GeV}$



$H \rightarrow 4 \text{ jets}, 2+2 \text{ jets back-to-back}$

Higgs Search @ LEP



LEP direct search:

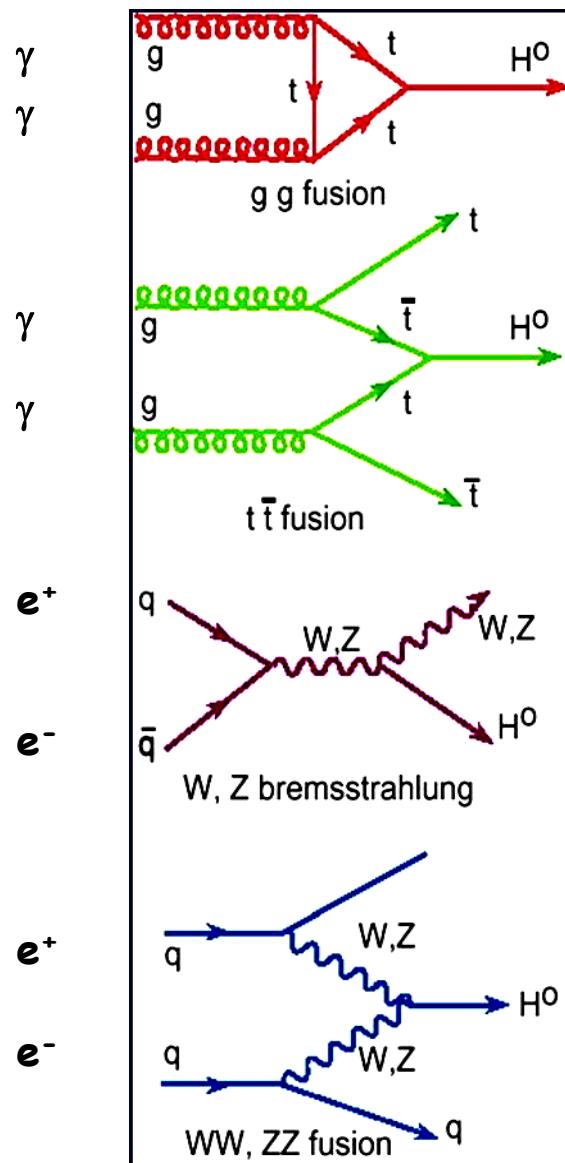
$$m_H > 114.4 \text{ GeV}$$

Standard Model

fits 2006:

$$m_H < 194 \text{ GeV} \quad @ \text{95\% CL}$$

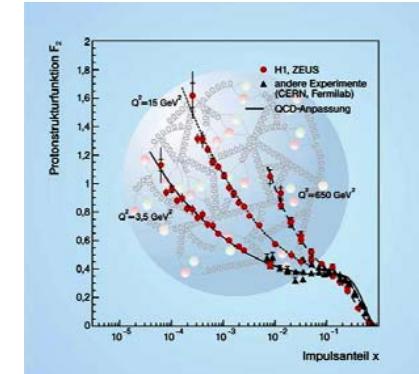
$$m_H = 89^{+38}_{-28} \text{ GeV}$$

ILC $e^+ e^-$
2x0.5-1**LHC** $p p$
2x7 TeV

Higgs production

HERA**p structure function:**

gluon + sea quark
distributions very soft
 $x_{g,sea}^2 \ll 1$

**weak:**

$$\sim G_F \sim g^2/m_W^2$$

superweak:

$$\sim G_F^2 \sim g^4/m_W^4$$

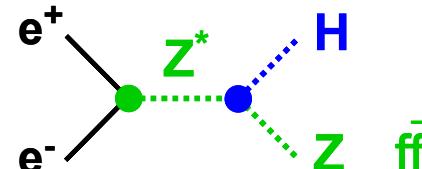
but valence quarks: $ud \rightarrow W^+W^-$ or $qq \rightarrow ZZ$

Higgs Production

Higgs couples to heaviest possible particle !

Higgs coupling $h = m_f / (v=246 \text{ GeV})$ $h_Z = 91/246 \sim 0.4$ $h_t = 172/246 \sim 0.7$

1. e^+e^- : Higgs-strahlung:



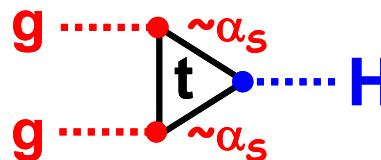
detect Higgs in missing mass: $m_H < \sqrt{s} - m_Z$

LEP: $m_H < \sqrt{s} - m_Z = 2 \times 102 - 91 \text{ GeV} = 114 \text{ GeV}$ 1997-2000

ILC: $m_H < \sqrt{s} - m_Z = 2 \times 250 - 91 \text{ GeV} = 400 \text{ GeV}$ 2015-

- Higgs coupling
- strong coupling
- weak coupling

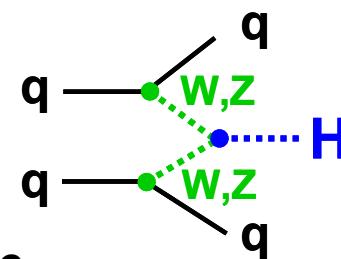
2. pp: boson fusion:



Tevatron, FNAL, USA:

$2 \times 1 \text{ TeV}$, 1996-2008

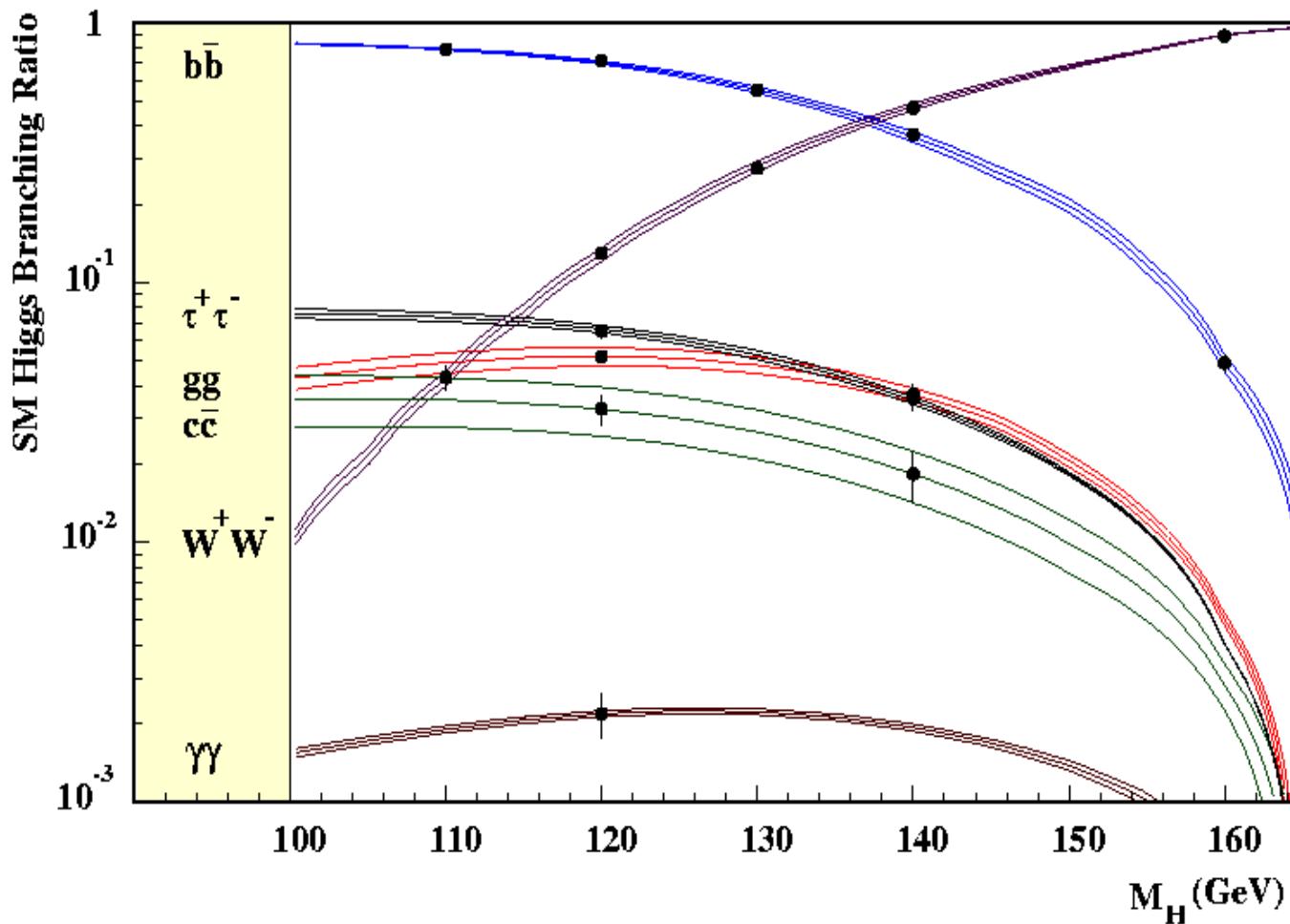
Large Hadron Collider LHC, CERN: $2 \times 7 \text{ TeV}$, 2008-



gluon carries only tiny fraction of p momentum

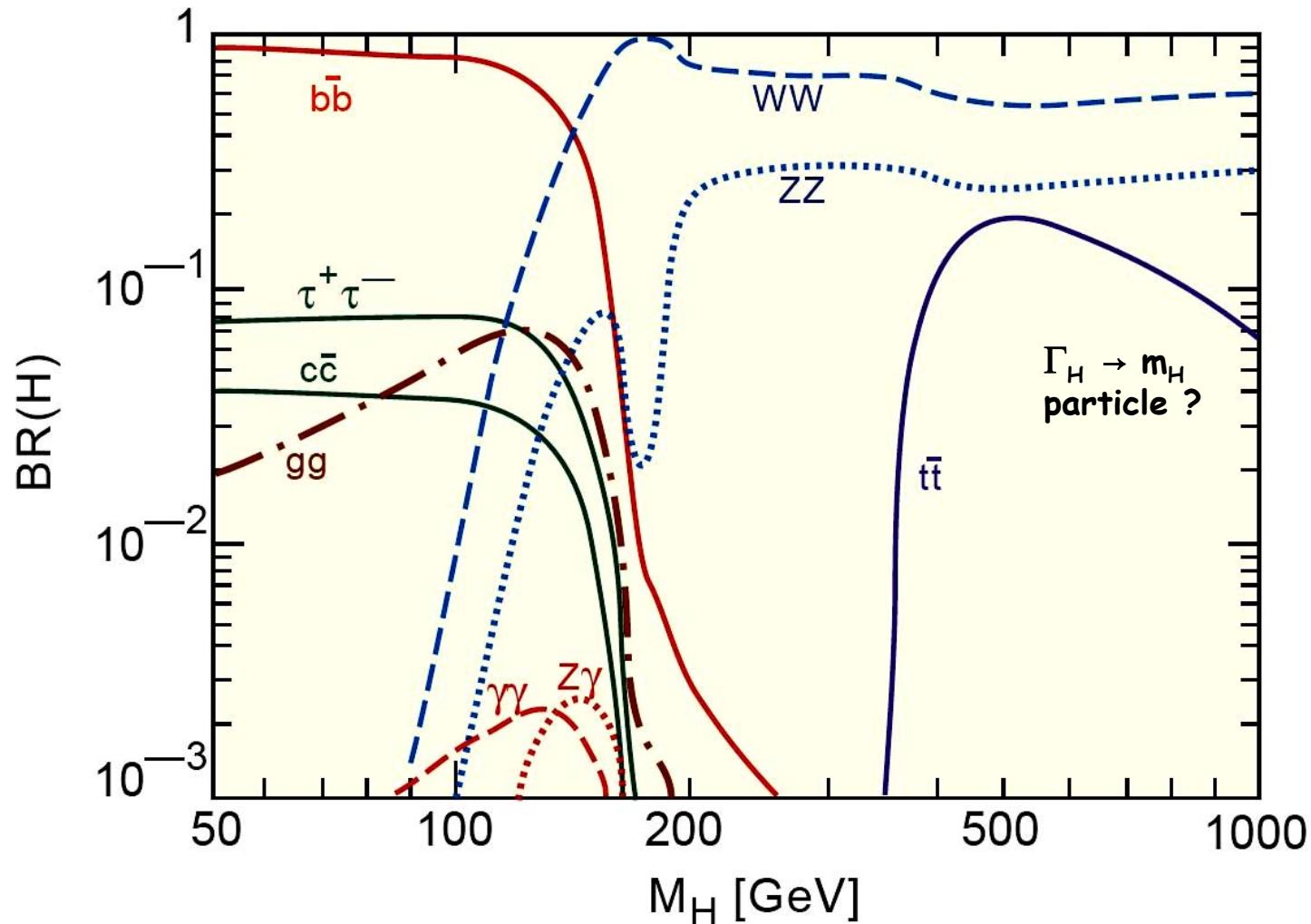
quark carries ~1/6 of p momentum: ~1 TeV / 7 TeV @ LHC

Higgs Decay



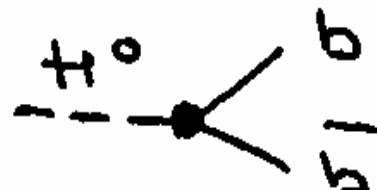
- Higgs = origin of mass in physics
- Higgs coupling = Yukawa coupling \sim mass of reaction partner

Higgs Decay



- Higgs = origin of mass in physics
- Higgs coupling = Yukawa coupling \sim mass of reaction partner

Higgs Decay

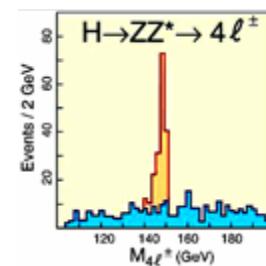


$m_b \sim 4.5 \text{ GeV}$

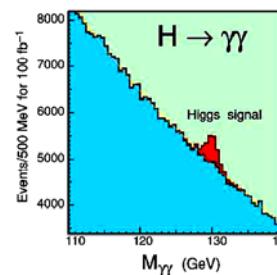
dominant for light Higgs $m_H < 2m_{W,Z}$
 identify 2 b jets !

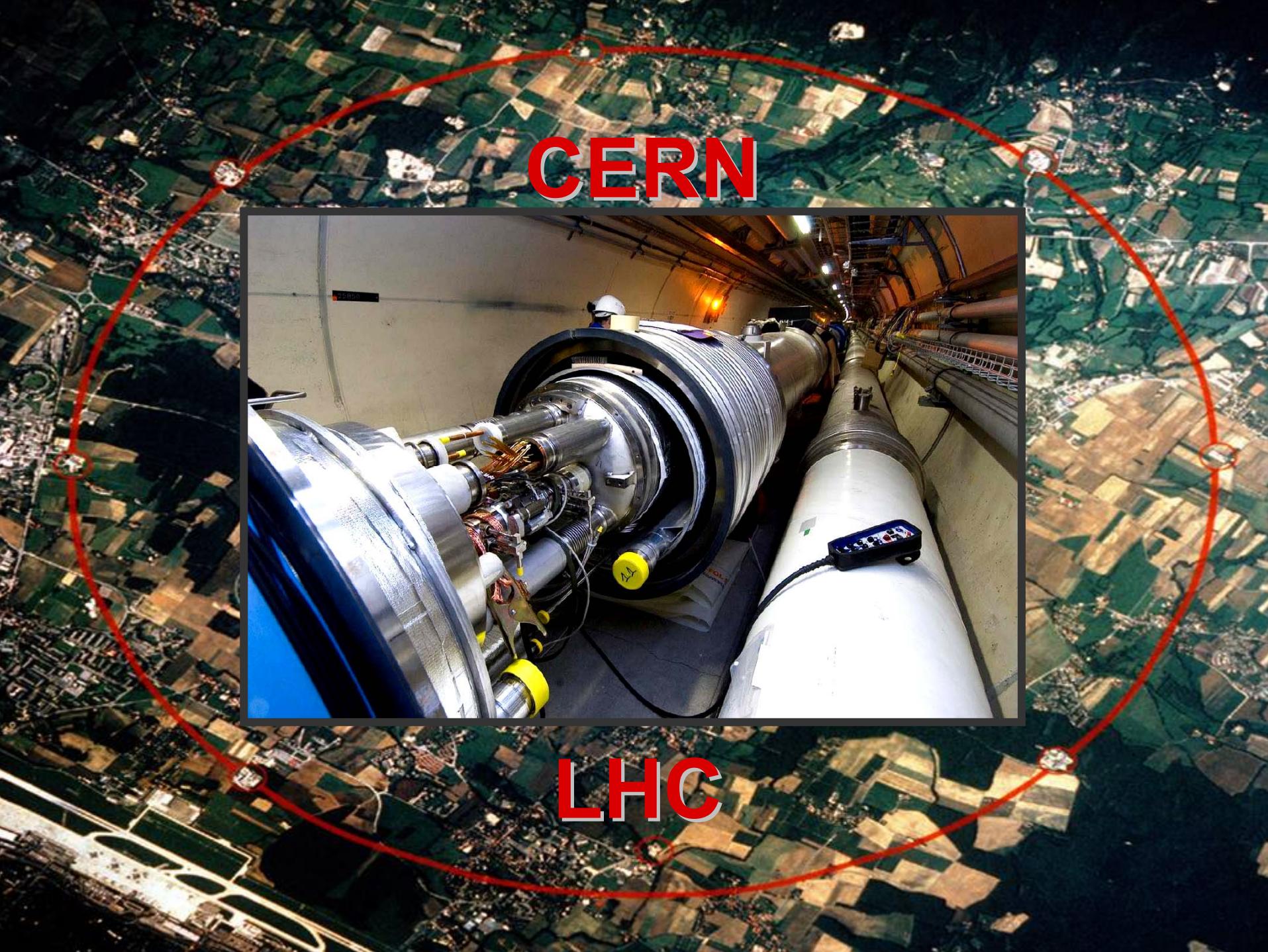


good to detect: find 4 leptons
 rare: $(\Gamma_1/\Gamma_Z)^2 \sim (3\%)^2 \sim 10^{-3}$

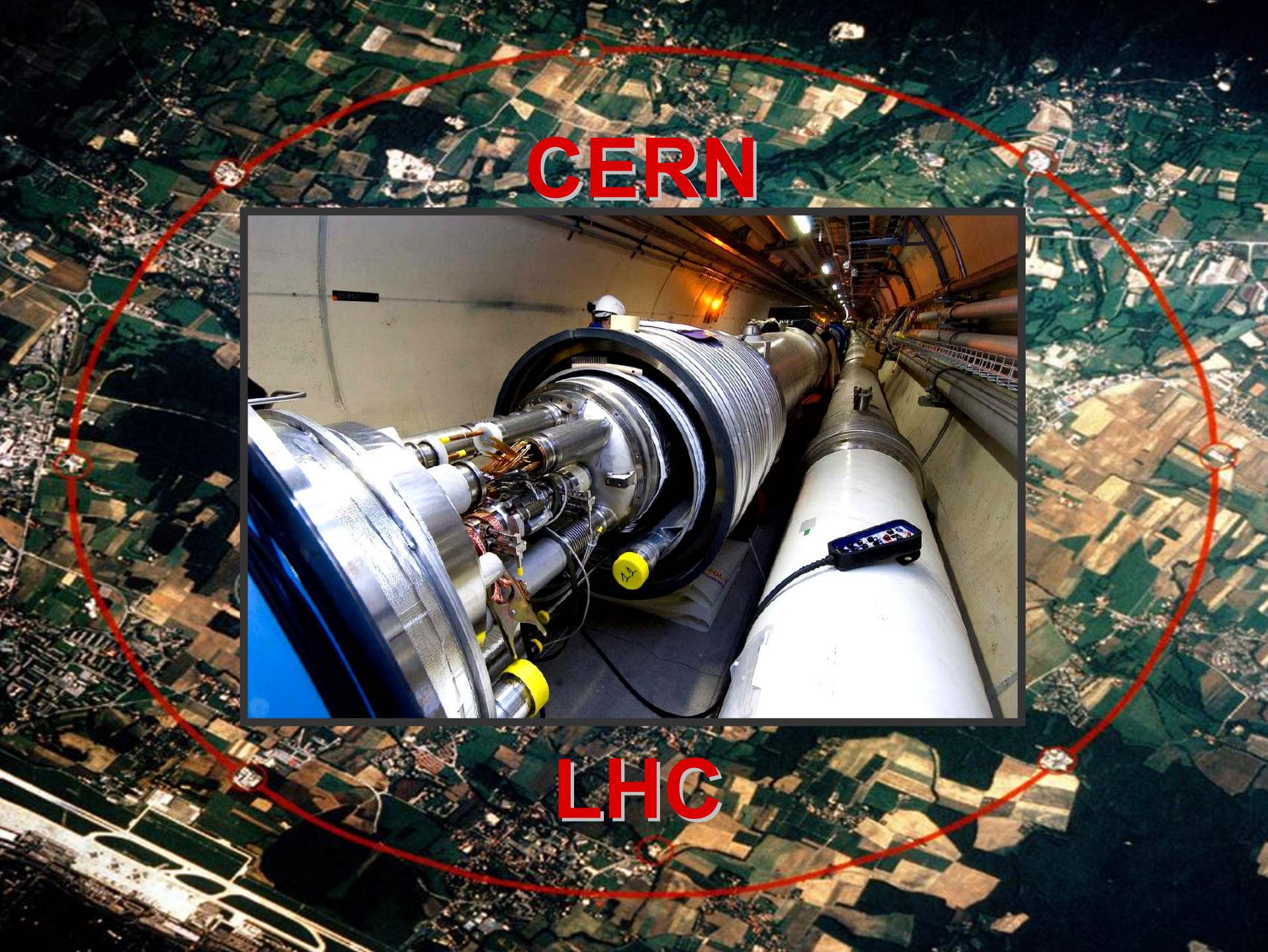
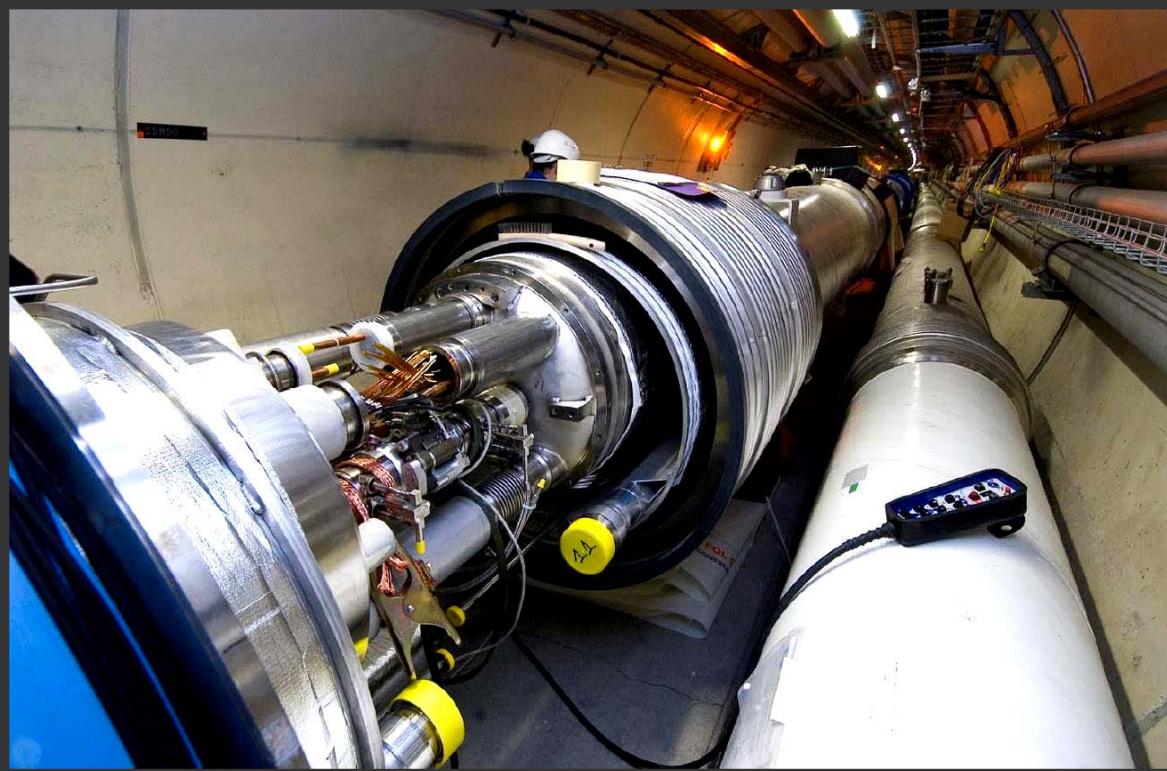


good to detect: find 2 photons
 rare: $\sim \alpha^2 < 10^{-4}$





CERN



LHC



LHC @ CERN

**Large Hadron Collider
at CERN in Geneva :**

27 km long
8 T superconducting magnets
7x7 TeV proton-proton collisions



> 2007 search for :

Higgs particle
Super-Symmetry

...

LHC parameters

	LHC pp	HERA ep
beam energy	7 TeV \times 7 TeV	27 GeV \times 920 GeV
luminosity	$10^{32-34} \text{ cm}^{-2}\text{s}^{-1}$	$10^{31} \text{ cm}^{-2}\text{s}^{-1}$
nr of bunches	2800	180
protons / bunch	10^{10}	10^{11}
bunch crossing rate	25 ns / 40 MHz	96 ns / 10 MHz
beam current	0.5 A	$40 \times 100 \text{ mA}^2$

LHC beam + magnets

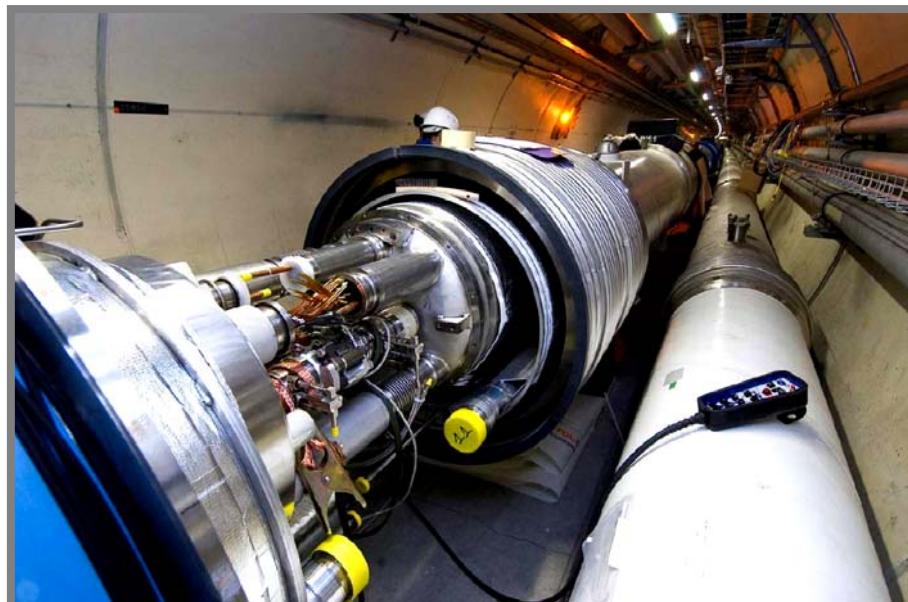
CMS magnet: world's largest SC magnet!

- $6 \times 12.5 \text{ m} @ 4 \text{ T} = 2.6 \text{ GJ}$
= enough to melt 18 t of gold!



LHC magnets: world's largest superfluid cryosystem !

- 27 km or 96 t He @ 8.4 T and 1.8 K
- energy stored 10 GJ
= airbus A380 @ 700 km/h
- = enough to melt 50 t of copper!

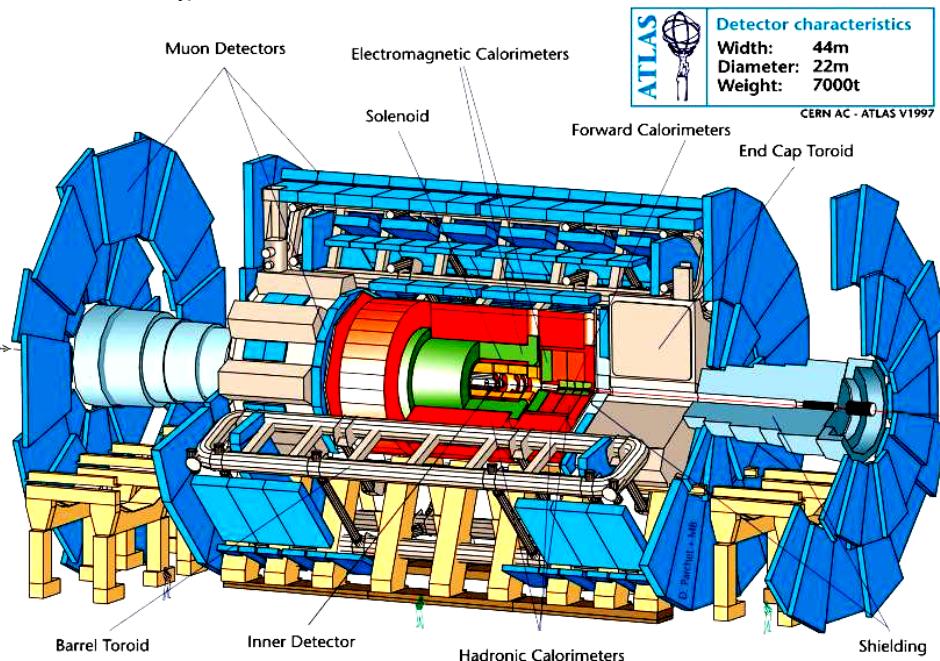


LHC beam:

- $0.5 \text{ A} @ 7 \text{ TeV}: E = 362 \text{ MJ}$
= 120 kg TNT
- = 200 t train locomotive @ 200 km/h
- = melt 25 m deep in Cu beam dump

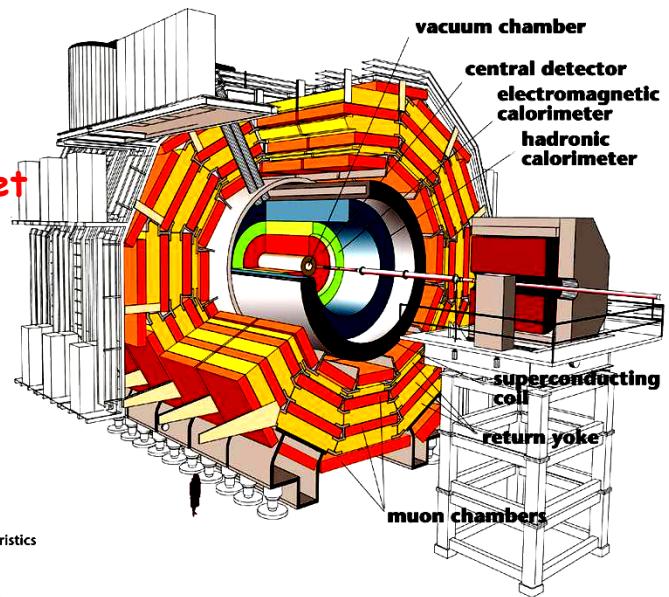
Liquid Argon Calorimeter

83 m³



SC magnet

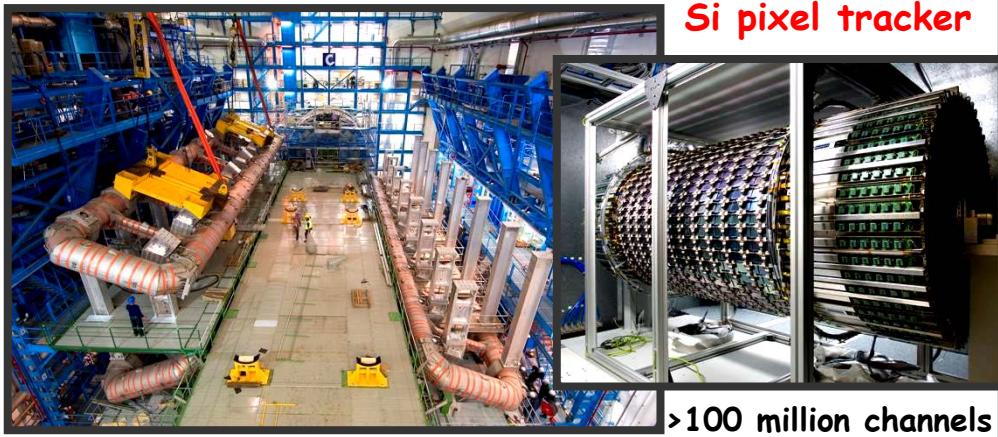
6x12 m
20 kA
4 T
2.6 GJ
world record



Detector characteristics

Width: 22m
Diameter: 15m
Weight: 14'500t

Si pixel tracker

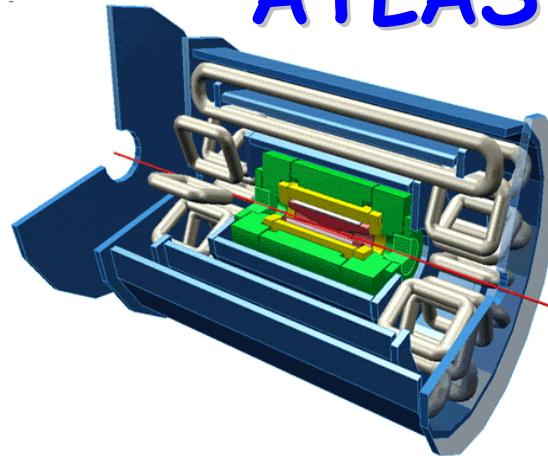


>100 million channels



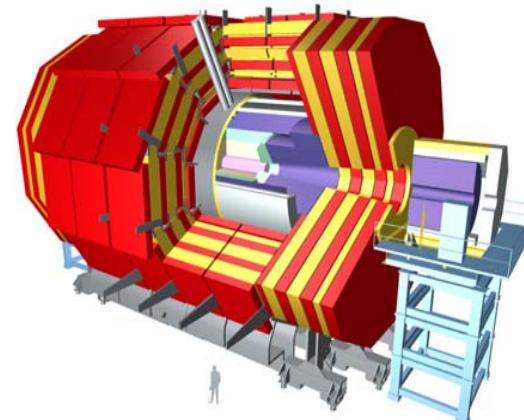
LHC experiments

ATLAS

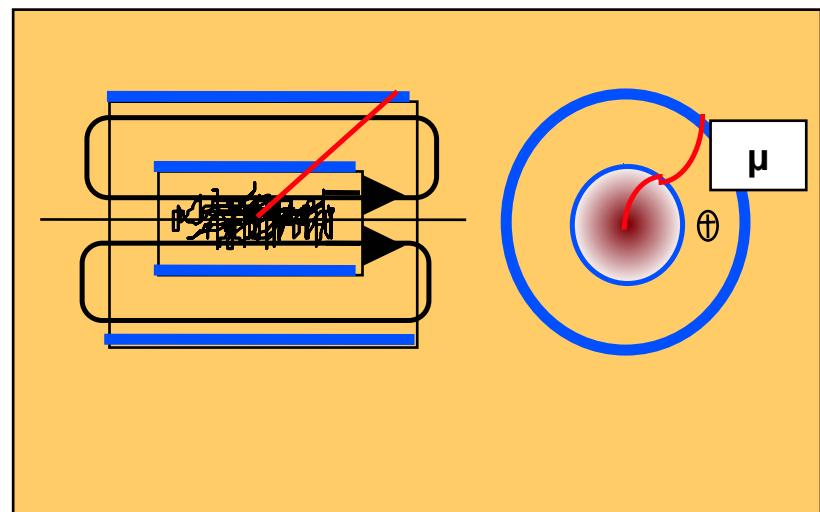
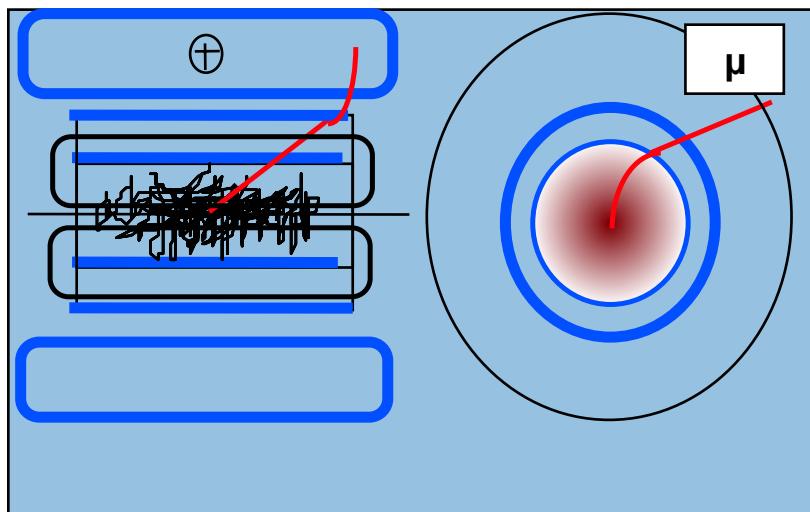


A Toroidal LHC Apparatus

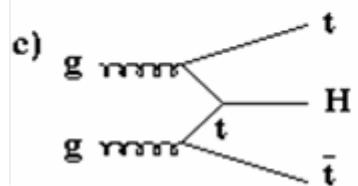
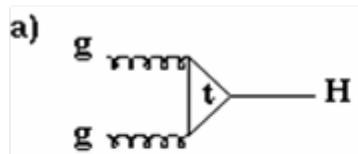
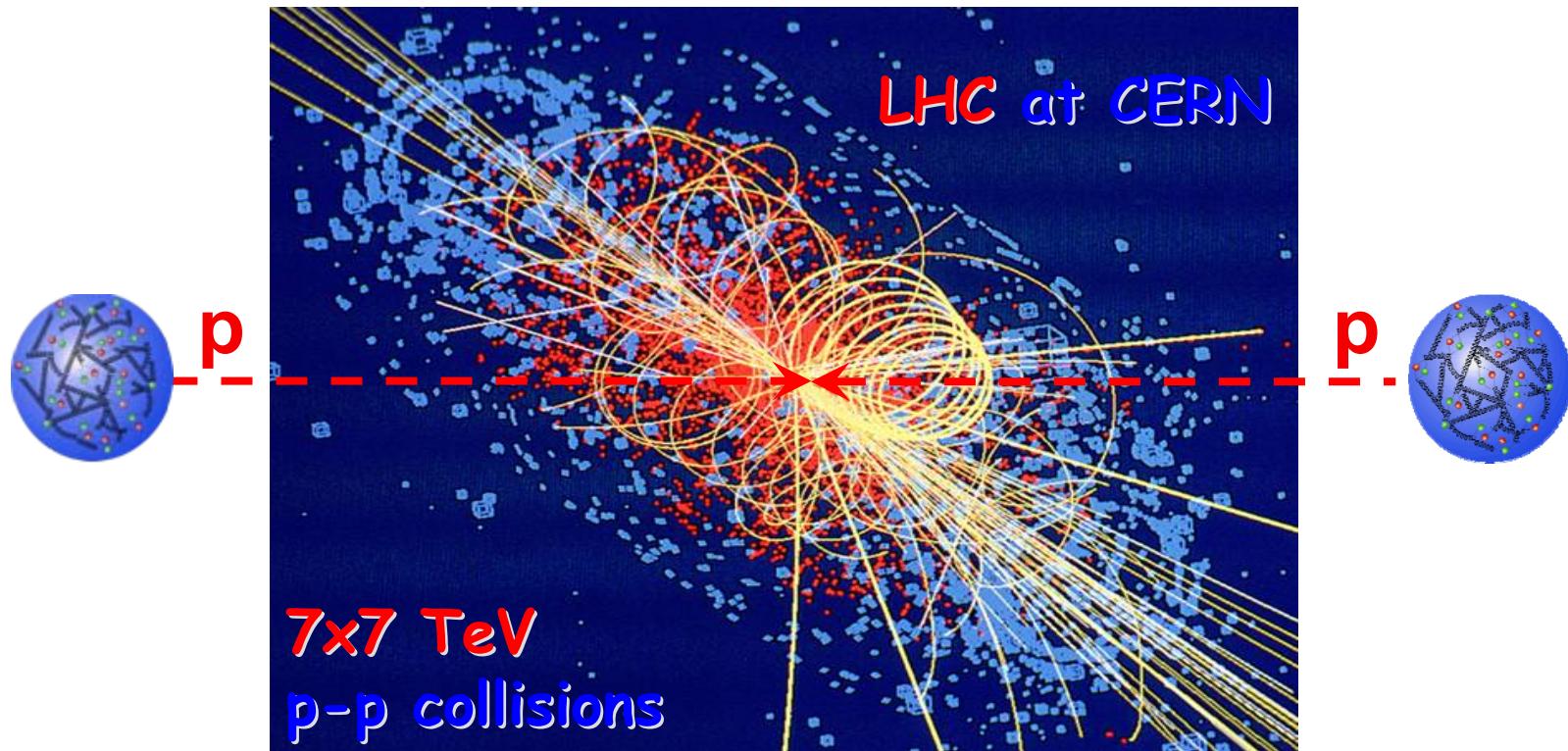
CMS



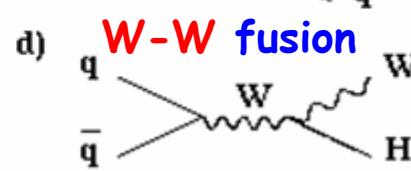
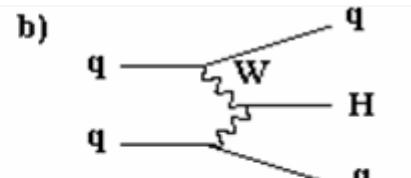
Compact Muon Solenoid



LHC Higgs production

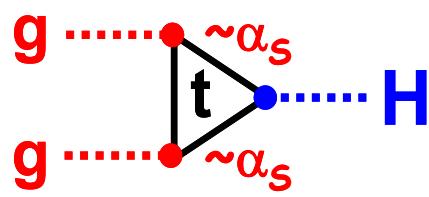
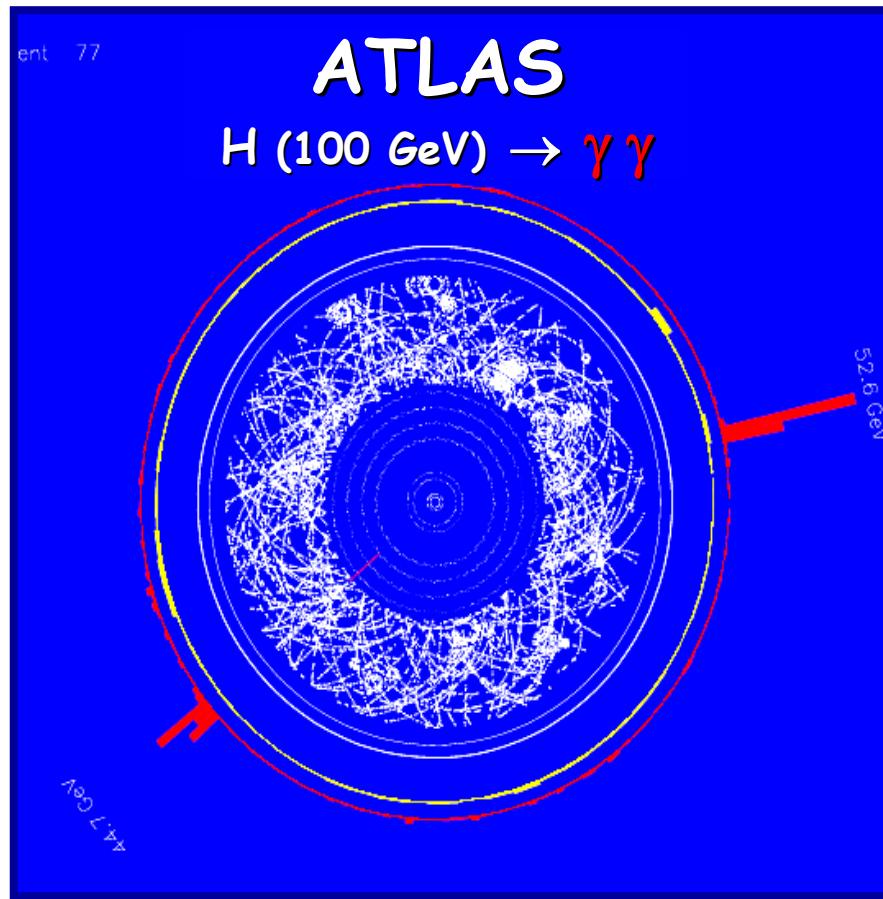


gluon-gluon fusion

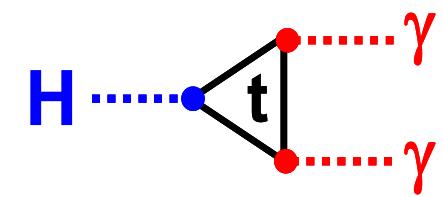


Higgs radiation

Higgs Search

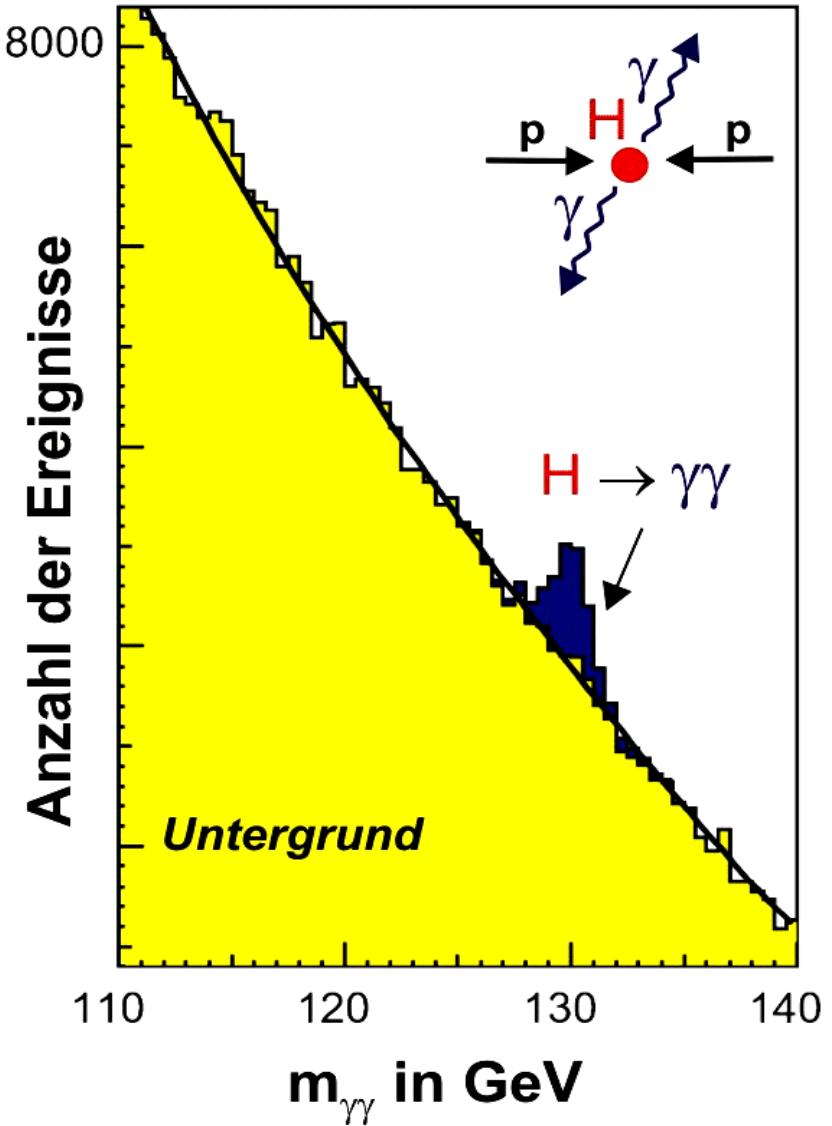


production

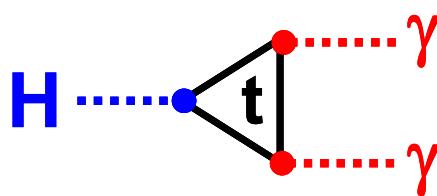
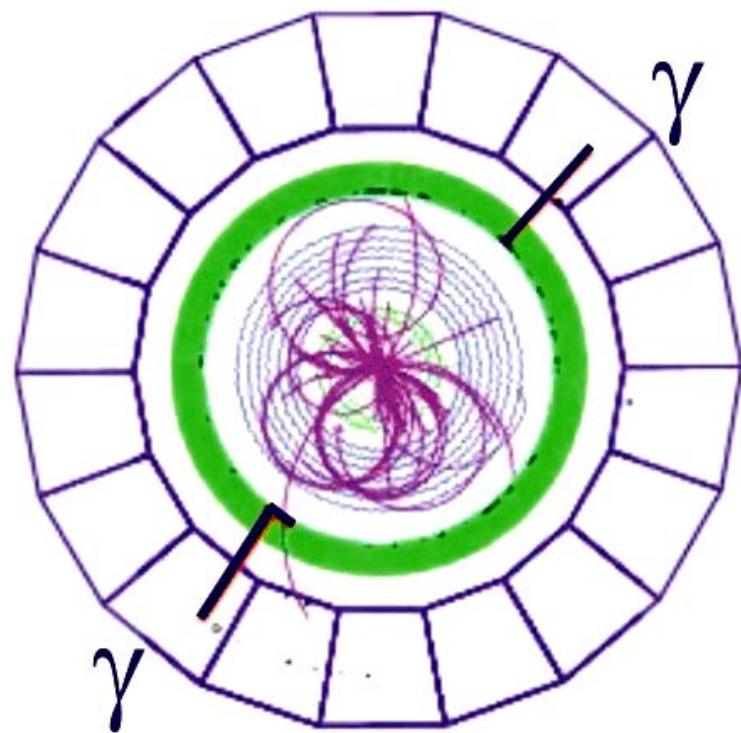


decay

simulation in the ATLAS detector at the LHC at CERN

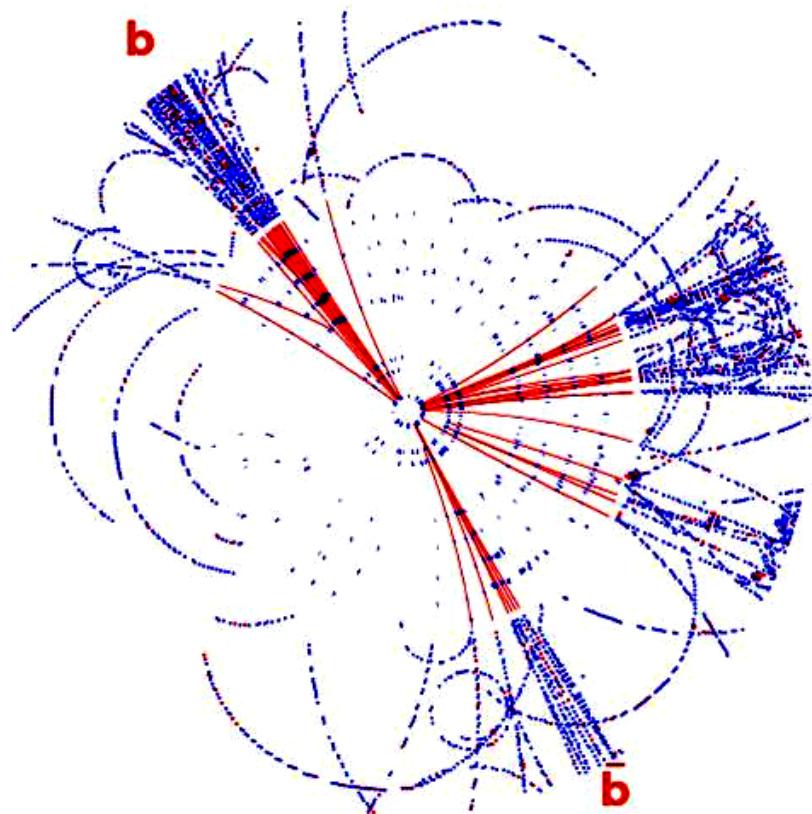
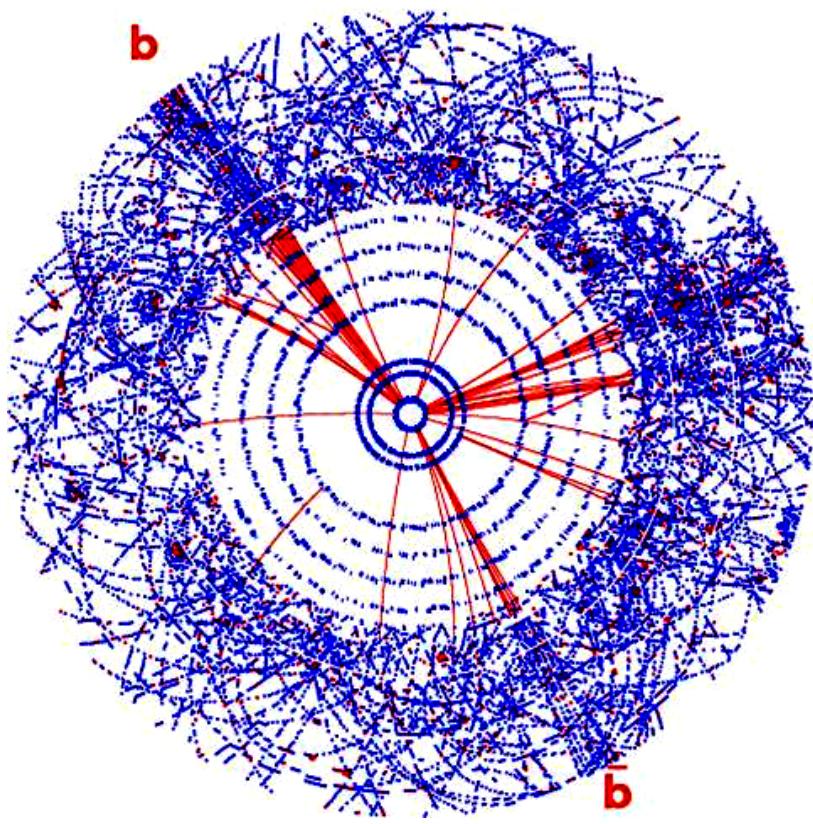


Higgs → $\gamma\gamma$



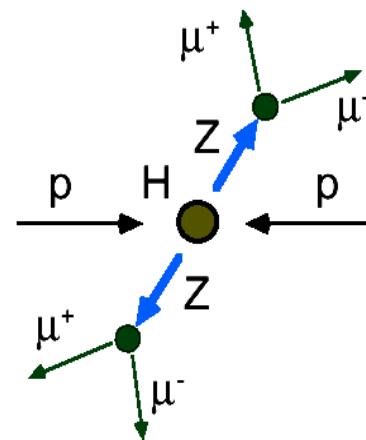
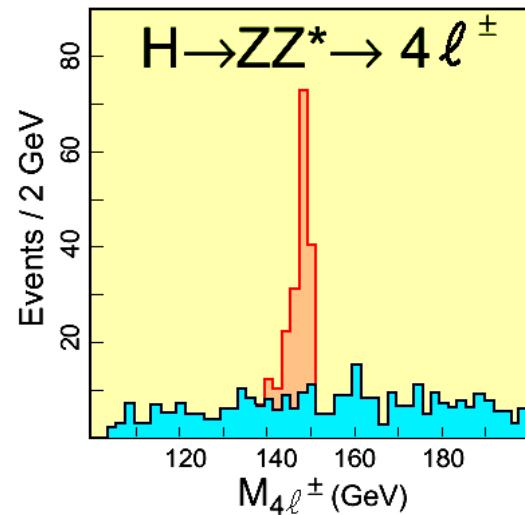
Higgs Search

$p\bar{p} \rightarrow H X \rightarrow b\bar{b} X$



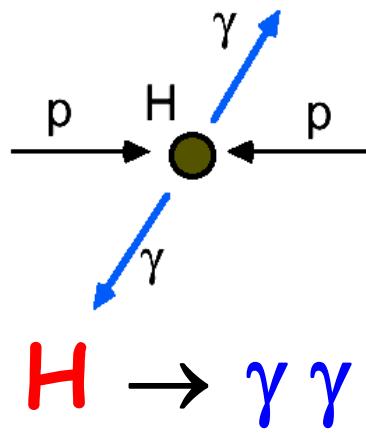
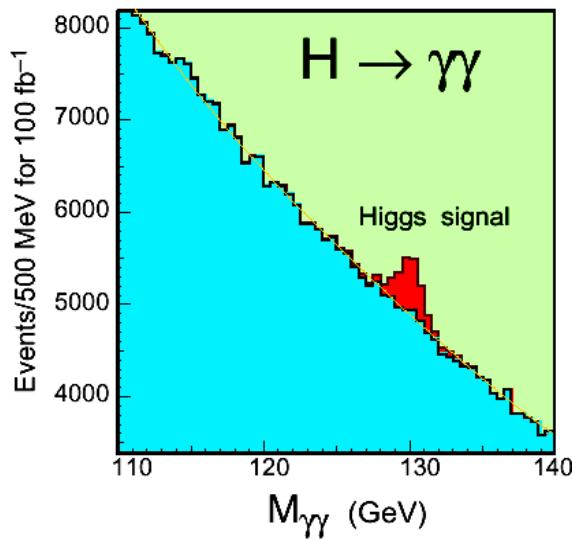
simulation in the ATLAS barrel inner detector

Higgs Search

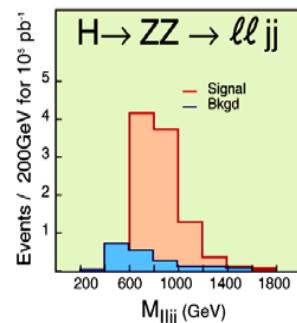
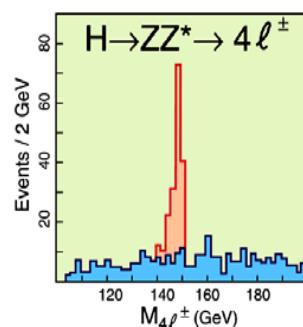
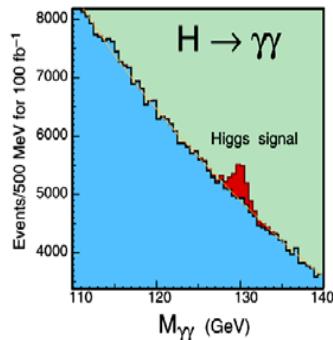
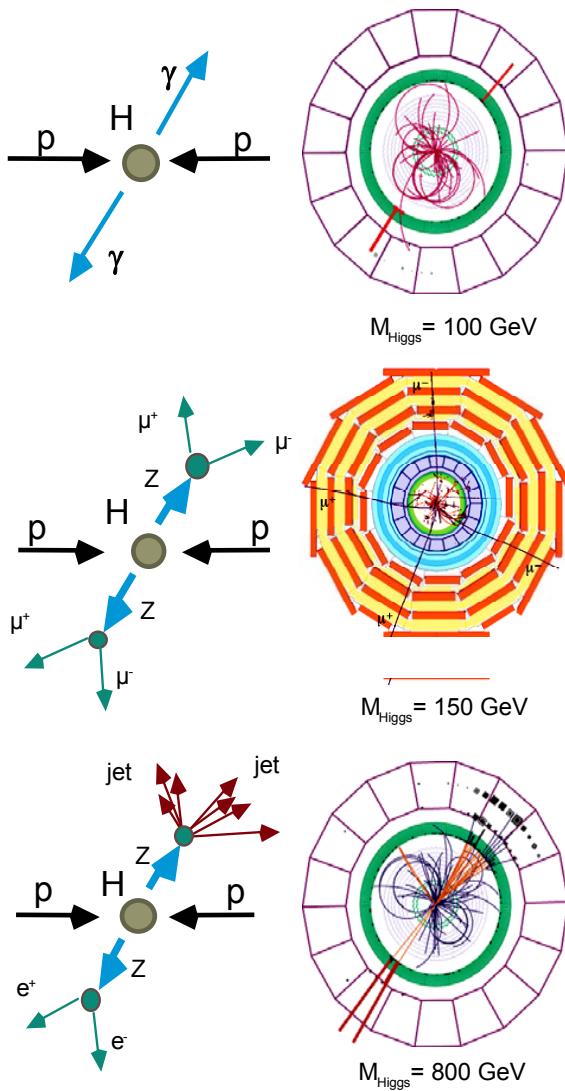


golden decay:
little background

$$H \rightarrow ZZ \rightarrow \mu^+\mu^- \mu^+\mu^-$$



LHC physics: Higgs



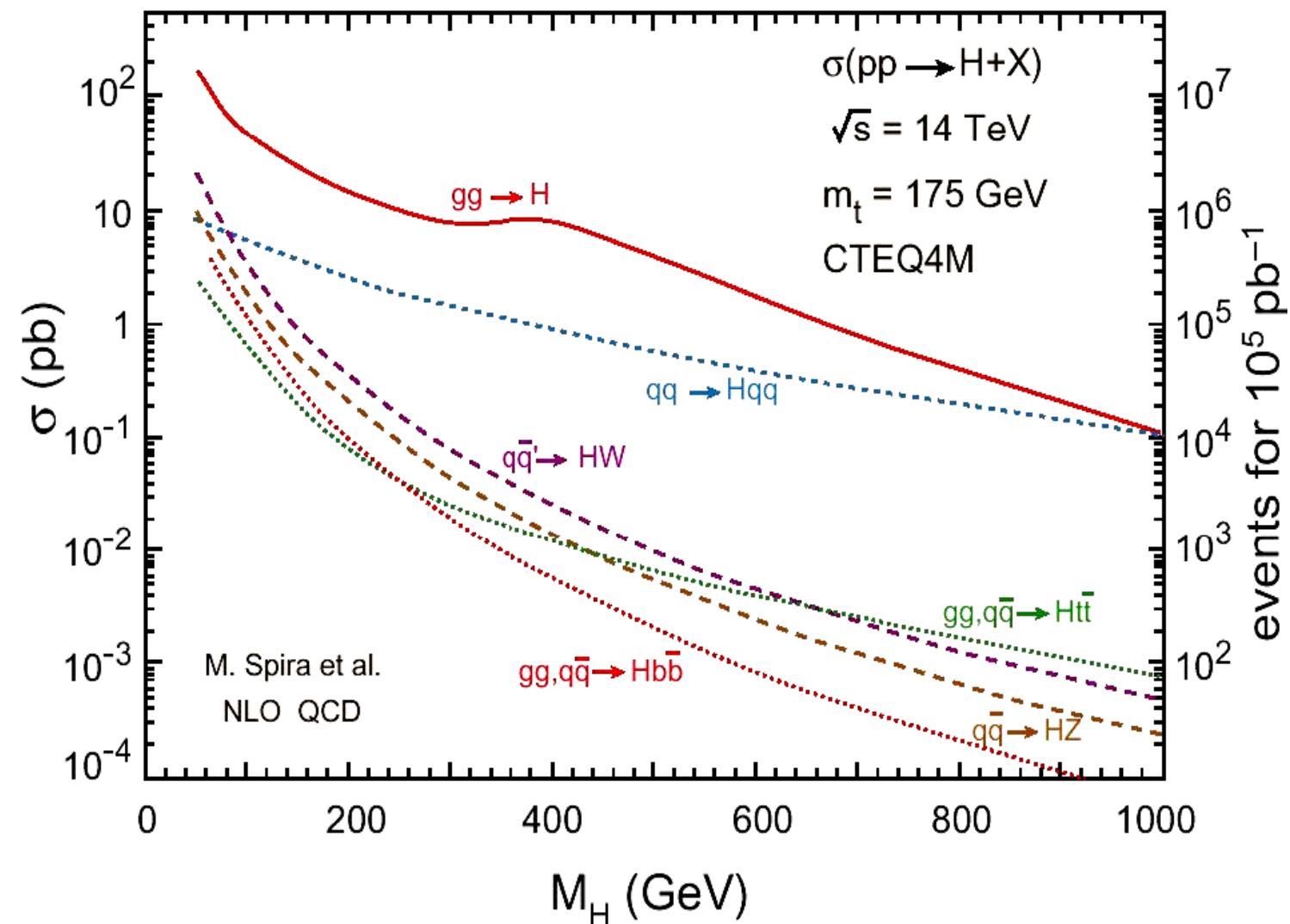
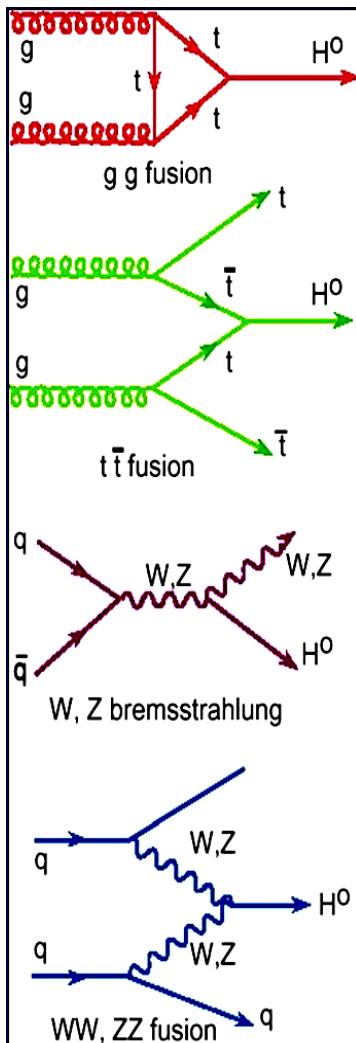
Higgs to 2 photons ($M_H < 140$ GeV). $H^0 \rightarrow \gamma\gamma$ is the most promising channel if M_H is in the range 80 – 140 GeV. The high performance PbWO₄ crystal electromagnetic calorimeter in CMS has been optimized for this search. The $\gamma\gamma$ mass resolution at $M_{\gamma\gamma} \sim 100$ GeV is better than 1%, resulting in a S/B of -1/20

Higgs to 4 leptons ($140 < M_H < 700$ GeV). In the M_H range 130 - 700 GeV the most promising channel is $H^0 \rightarrow ZZ^* \rightarrow 2, + 2, -$ or $H^0 \rightarrow ZZ \rightarrow 2, + 2, -$. The detection relies on the excellent performance of the muon chambers, the tracker and the electromagnetic calorimeter. For $M_H \gtrsim 170$ GeV a mass resolution of ~ 1 GeV should be achieved with the combination of the 4 Tesla magnetic field and the high resolution of the crystal calorimeter

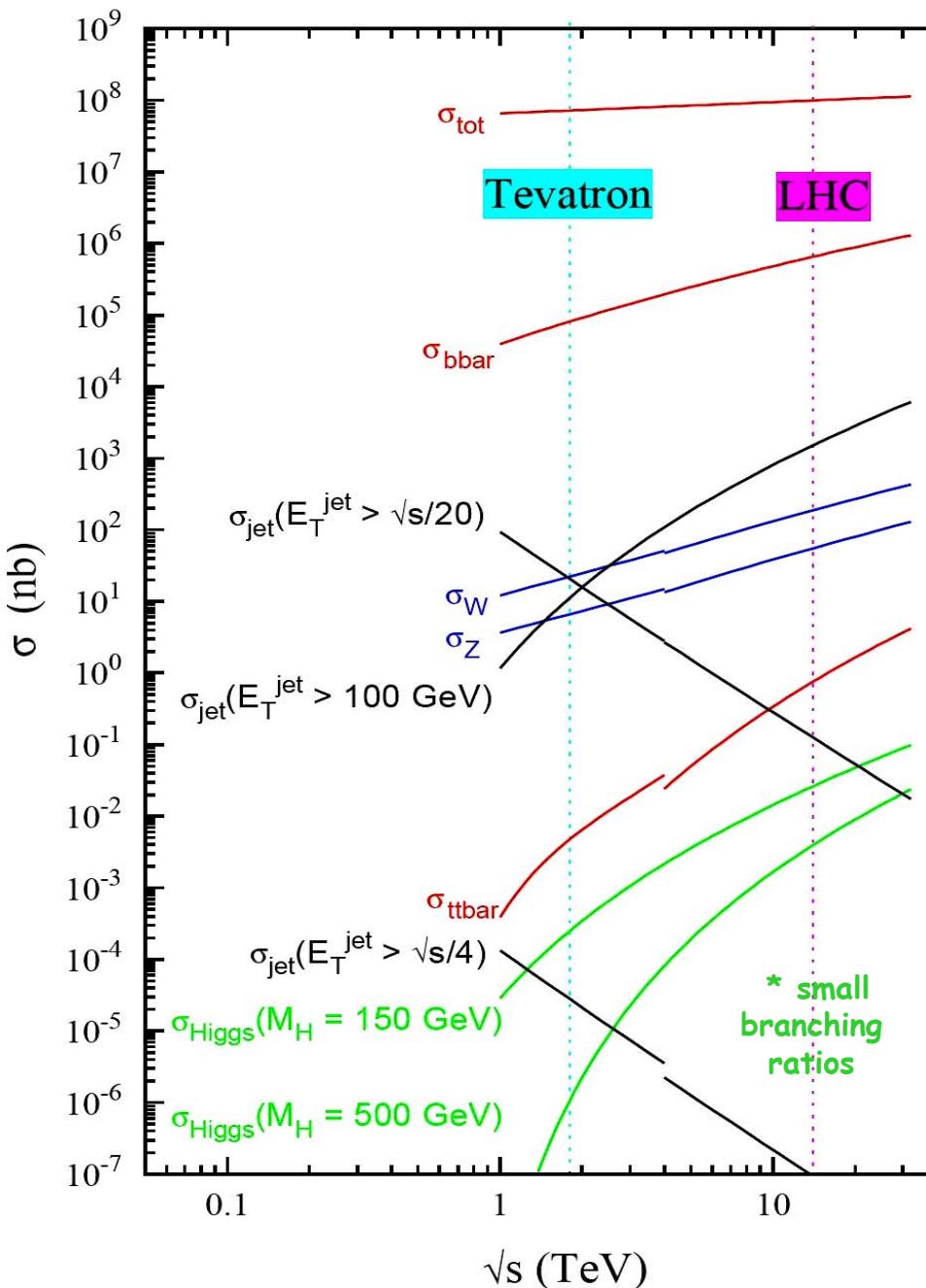
Higgs to 2 leptons+2 jets ($M_H > 500$ GeV). For the highest M_H , in the range 0.5 - 1 TeV, the promising channels for one year at high luminosity are $H^0 \rightarrow ZZ \rightarrow , + , - vv$, $H^0 \rightarrow ZZ \rightarrow , + , - jj$ and $H^0 \rightarrow W^+ W^- \rightarrow , \pm v jj$. Detection relies on leptons, jets and missing transverse energy (E_t^{miss}), for which the hadronic calorimeter (HCAL) performance is very important

Higgs cross sections

dominant



No color exchange:
2 clear forward jets



pp cross sections

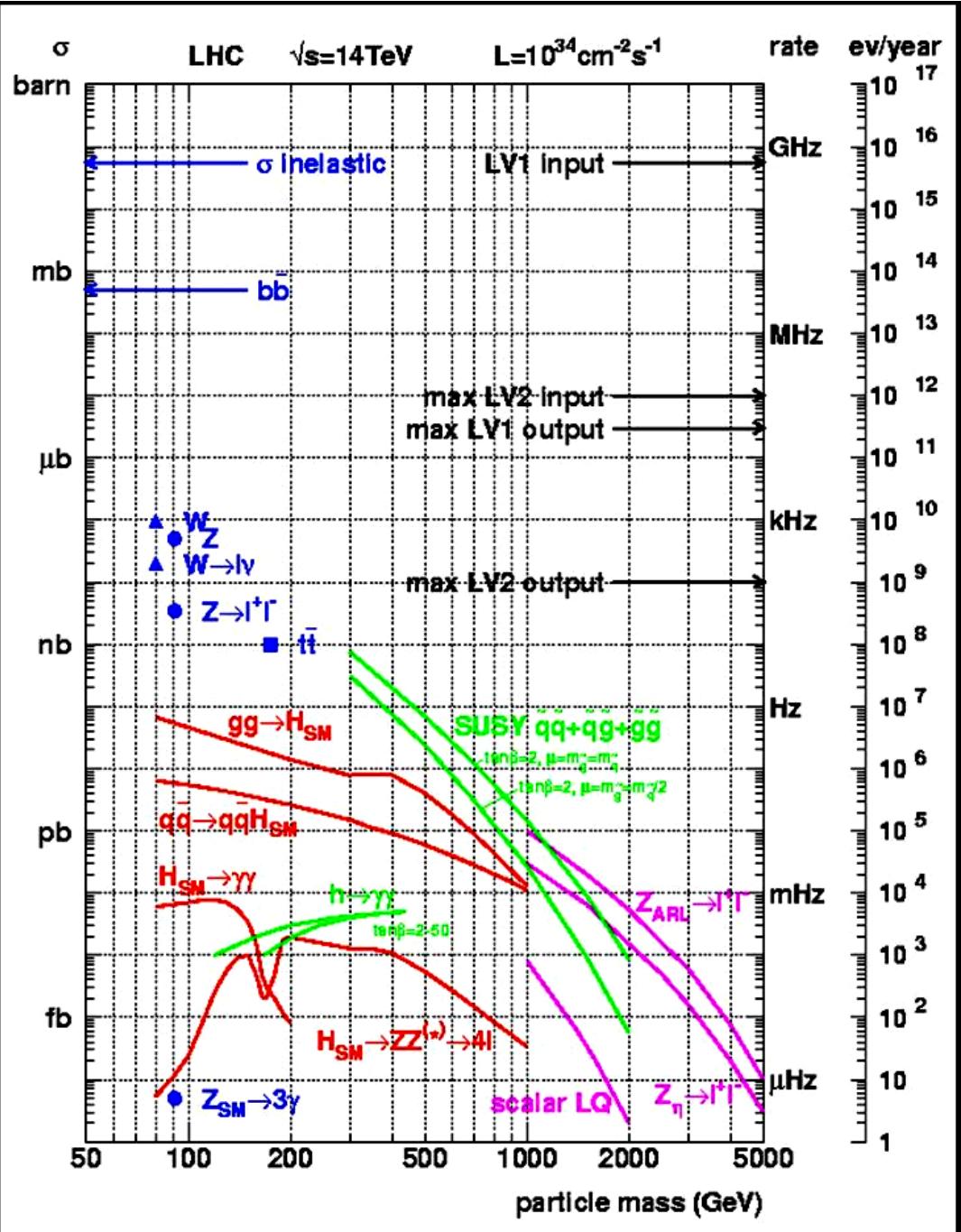
$\sigma_{\text{tot}} \sim 100 \text{ mb}$:
gigantic soft hadronic background

$E_T^{\text{jet}} > 100 \text{ GeV}$:
huge hard QCD backgr.

hadronic H decays
swamped by backgr.
need leptonic decays:

$H \rightarrow \gamma\gamma$,
 $H \rightarrow ZZ \rightarrow 4l$

look for weak signal
in a strong int. machine !



design luminosity
 $10^9 \text{ evts/s} * 10 \text{ MB} = 10 \text{ PB/s}$

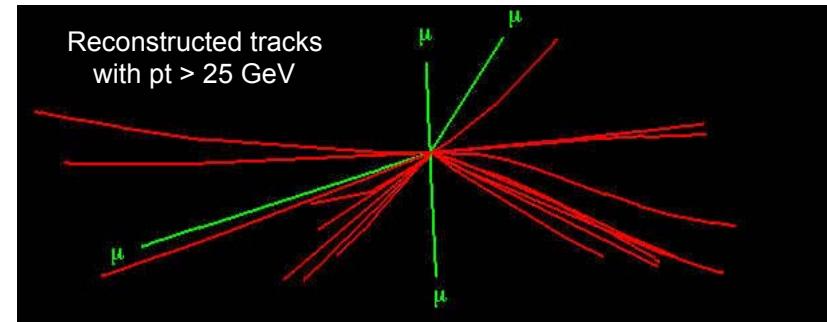
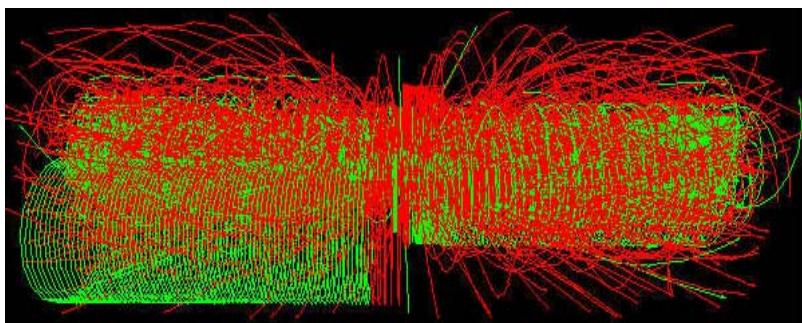
LHC event rates

$10^9 \text{ evts/a} * 2 \text{ MB} = 2 \text{ PB/a}$

$10^3 \text{ evts/a} * \text{efficiency} * \text{acceptance}$

LHC rates

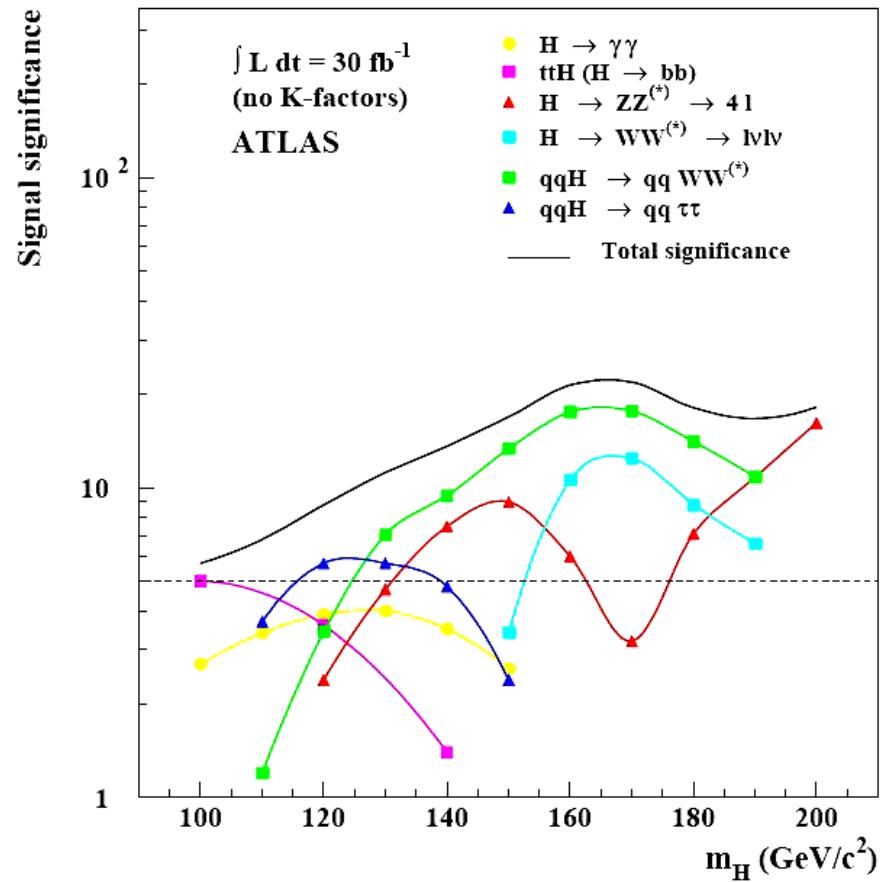
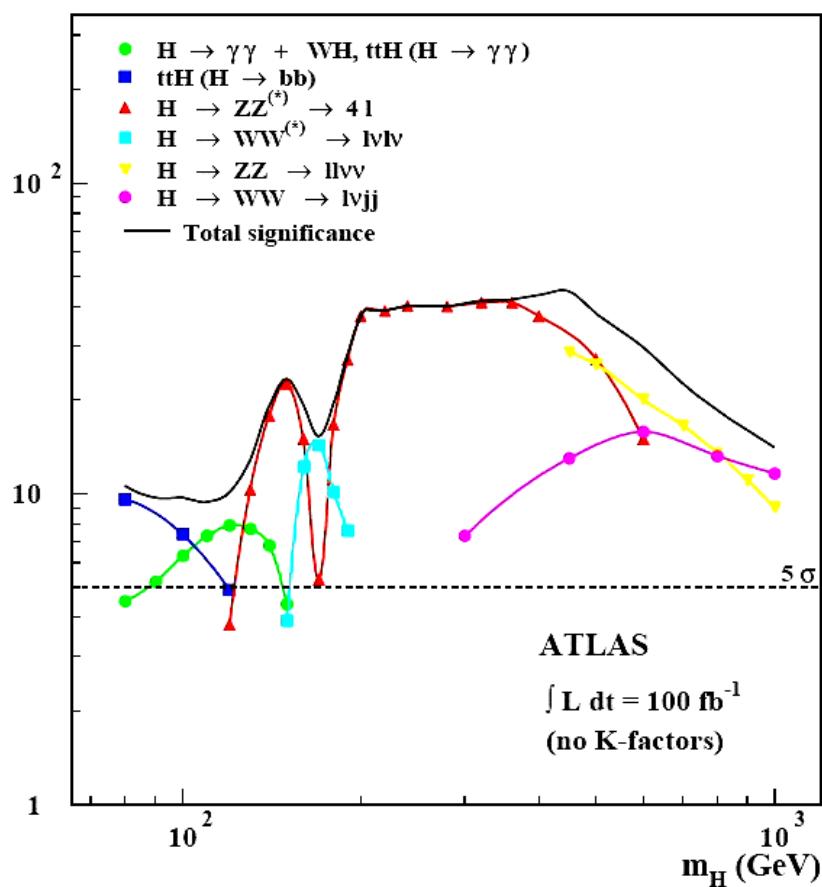
Process	evts/s @ $10^{34} \text{ cm}^{-2}\text{s}^{-1}$	evts/10 fb $^{-1}$ $\sim 1/2 \text{ y}$ @ $10^{33} \text{ cm}^{-2}\text{s}^{-1}$	before LHC
• Inelastic p-p reactions	10^9		
• q-q/g interactions, $p_t > 20 \text{ GeV}$	10^8		
• $b\bar{b}$ pairs	$5 \cdot 10^6$	10^{12-13}	10^9 BaBar/Belle
• $t\bar{t}$ pairs	8	10^7	10^4 FNAL
• $W \rightarrow e \nu$	150	10^8	10^4 LEP, 10^6 FNAL
• $Z \rightarrow e e$	15	10^7	10^6 LEP
• Higgs (150 GeV)	0.2	10^5	
• SUSY: Gluino, Squarks (1 TeV)	0.03	10^4	
• Black Holes $m_D=3 \text{ TeV}$ $n=4$	0.001	10^3	



Find a needle in a haystack:

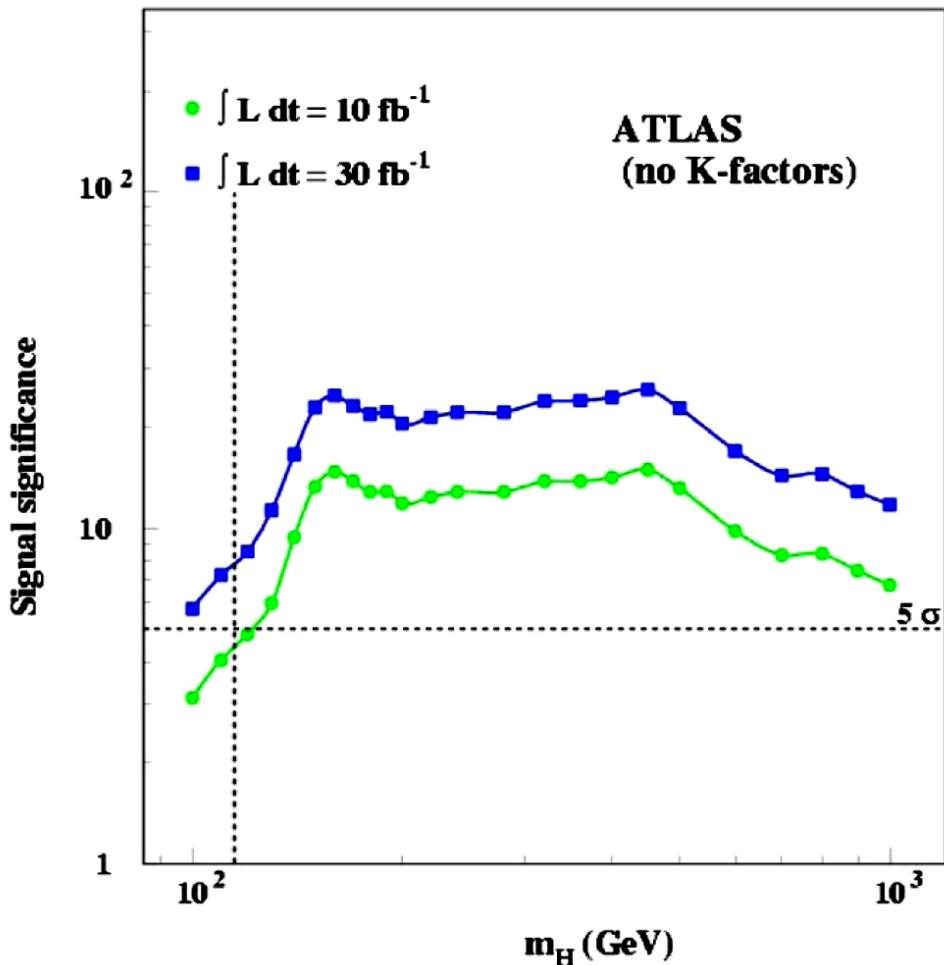
- high luminosity $L \sim 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$, 40 MHz event rate, 100 GB/s data
- excellent detector, excellent trigger

Higgs Significance



- $m_H < 150 \text{ GeV: } H \rightarrow \gamma\gamma$
- $m_H > 150 \text{ GeV: } H \rightarrow ZZ \rightarrow 4l$

Higgs Significance



1 γ @ $10^{33} \text{ cm}^{-2}\text{s}^{-1}$:
10 fb^{-1} (≥ 2008)

1 γ @ $10^{34} \text{ cm}^{-2}\text{s}^{-1}$:
100 fb^{-1} (≥ 2010)

discovery vs precision

SPS $2 \times 270 \text{ GeV}$

1983 W,Z discovery
1984 Nobel prize



LHC $2 \times 7 \text{ TeV}$

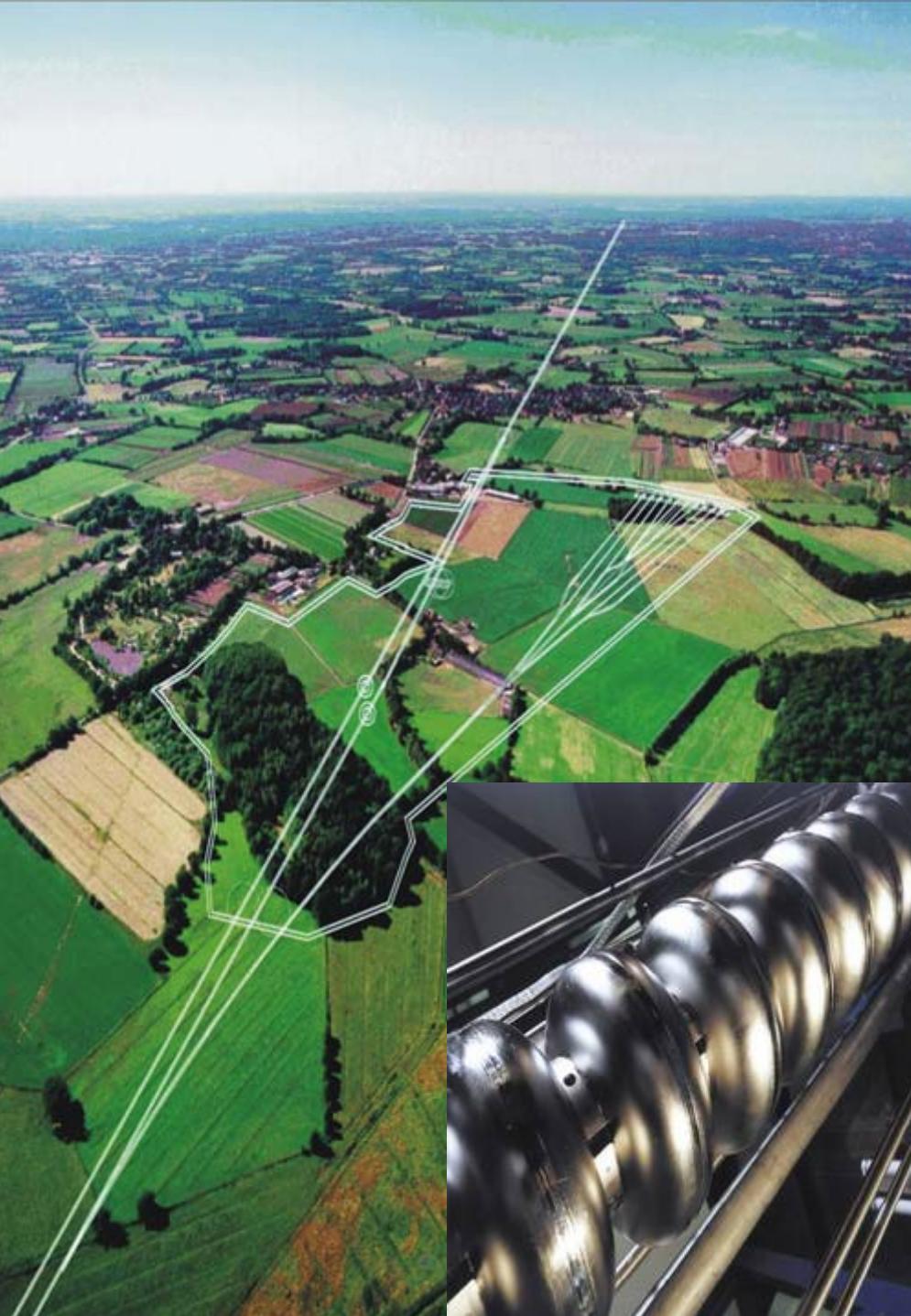
>2007 Higgs, SUSY discovery ?

LEP $2 \times 50\text{-}100 \text{ GeV}$

1992- establish Standard
2000 Model to 10^{-3}

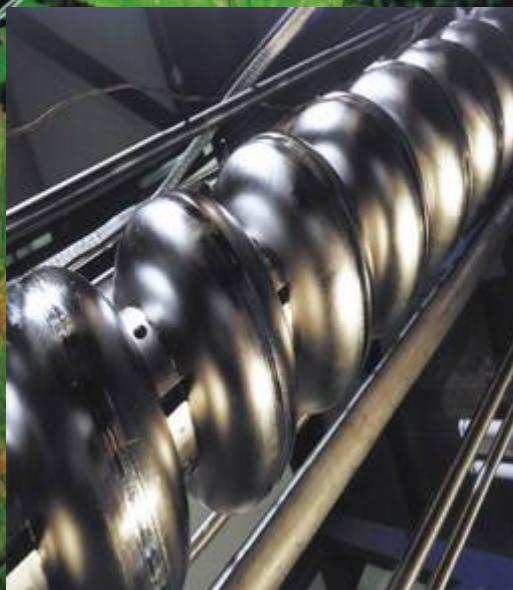
ILC $2 \times 500 \text{ GeV} - 1 \text{ TeV}$

>2015 establish
Higgs + SUSY ?



International Electron-Positron- Linear Collider

Superconducting
500-800 GeV
30 km long



Aim:

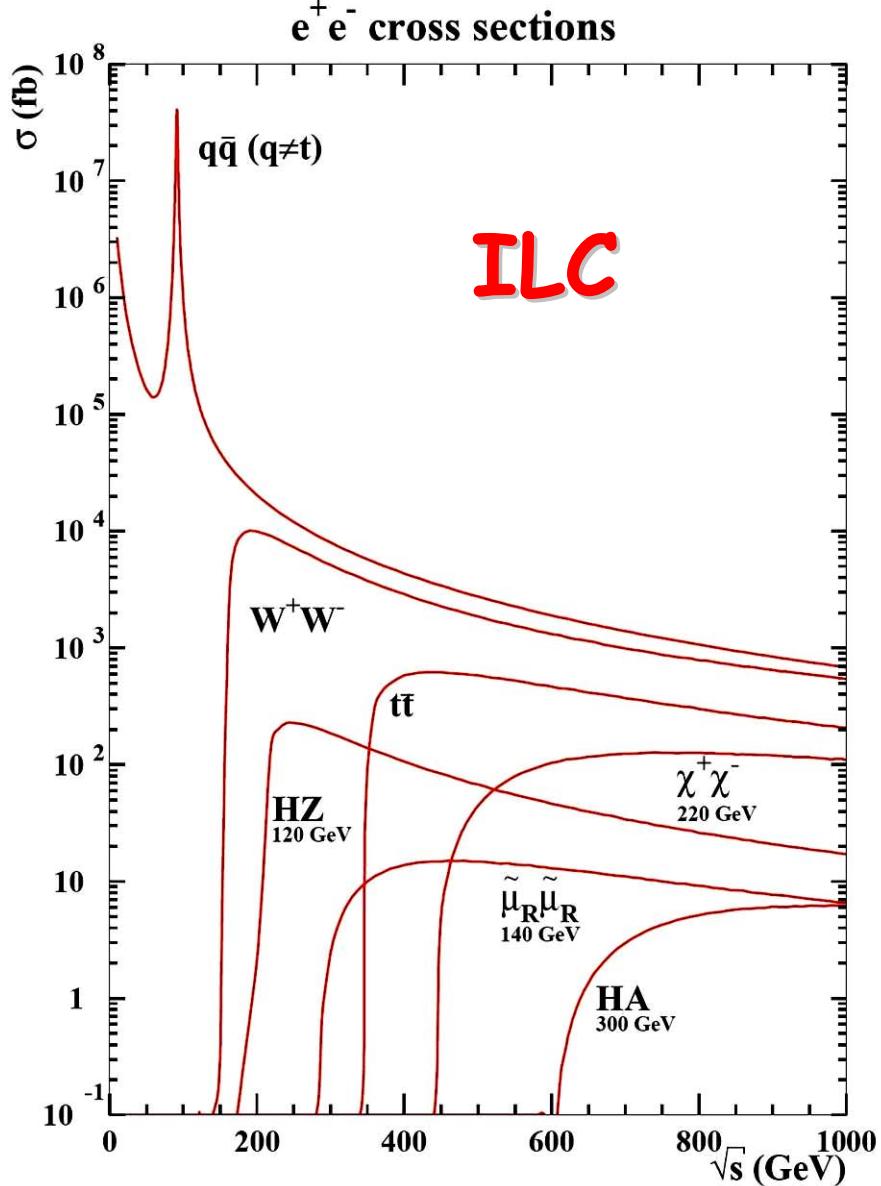
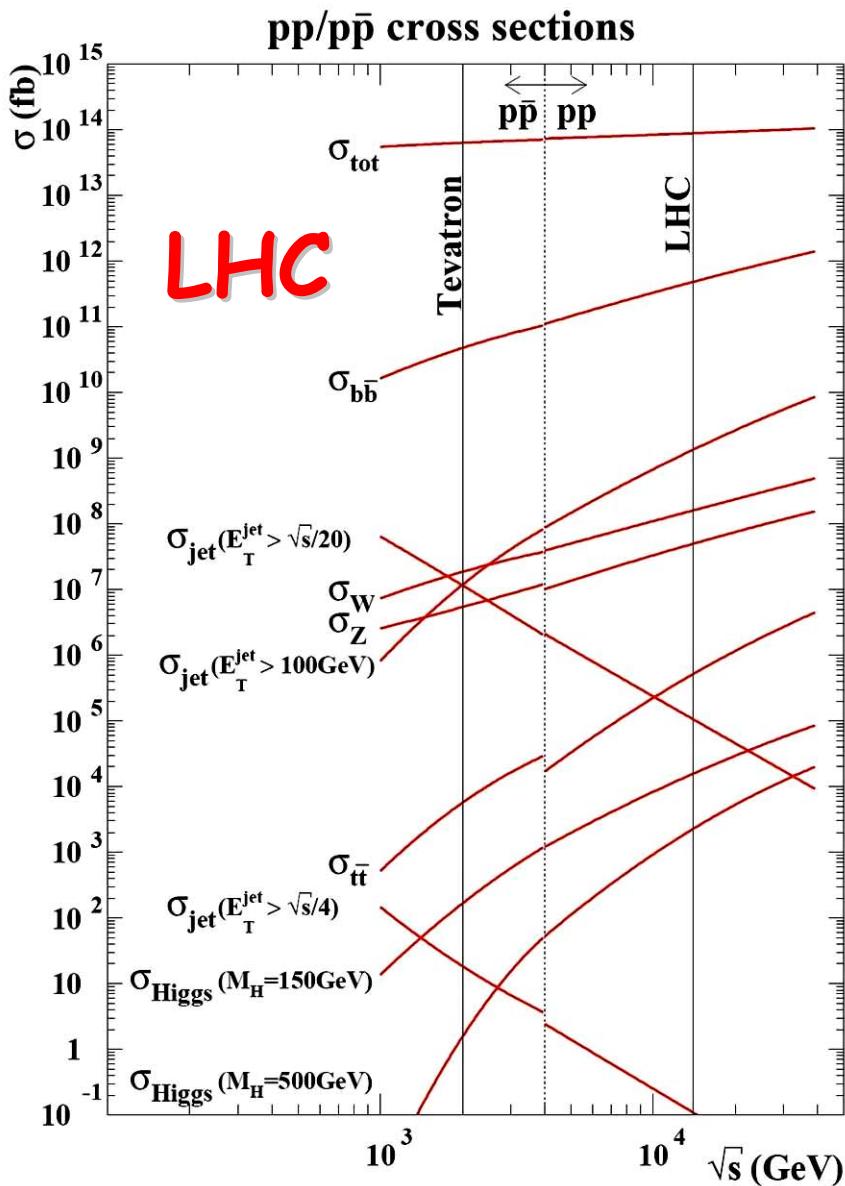
Higgs- and SUSY-
precision physics

> 2015

- signal $\sim 10^4$ fb
- background $\sim 10^{10} \times$ signal

LHC vs ILC

- signal $\sim 10^2$ fb
- background $\sim 10^2 \times$ signal



6. QCD

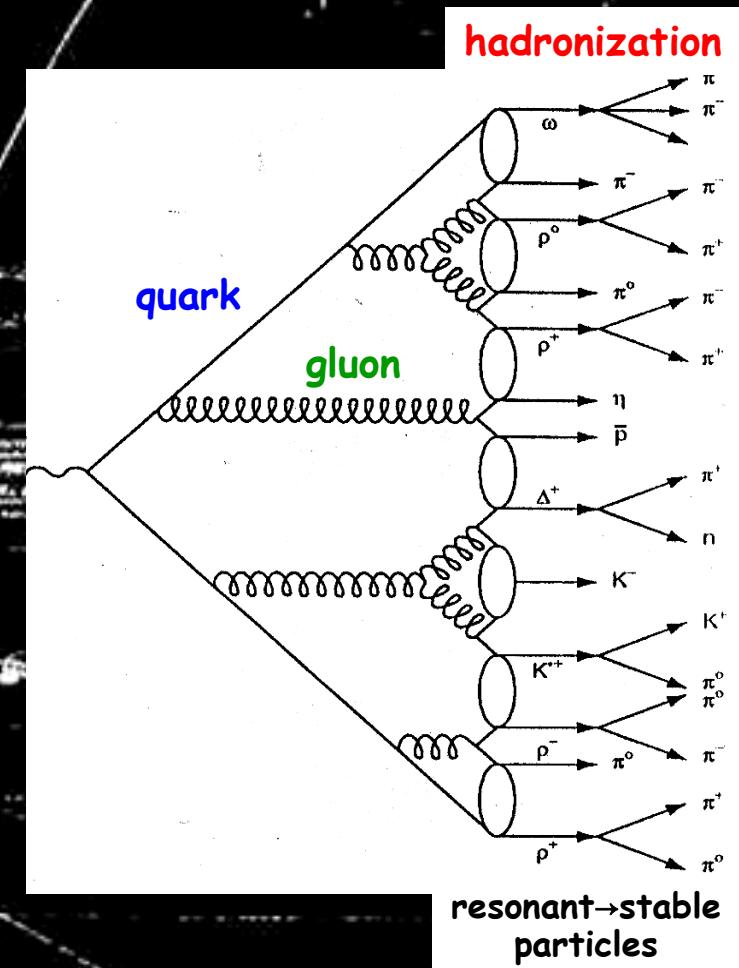
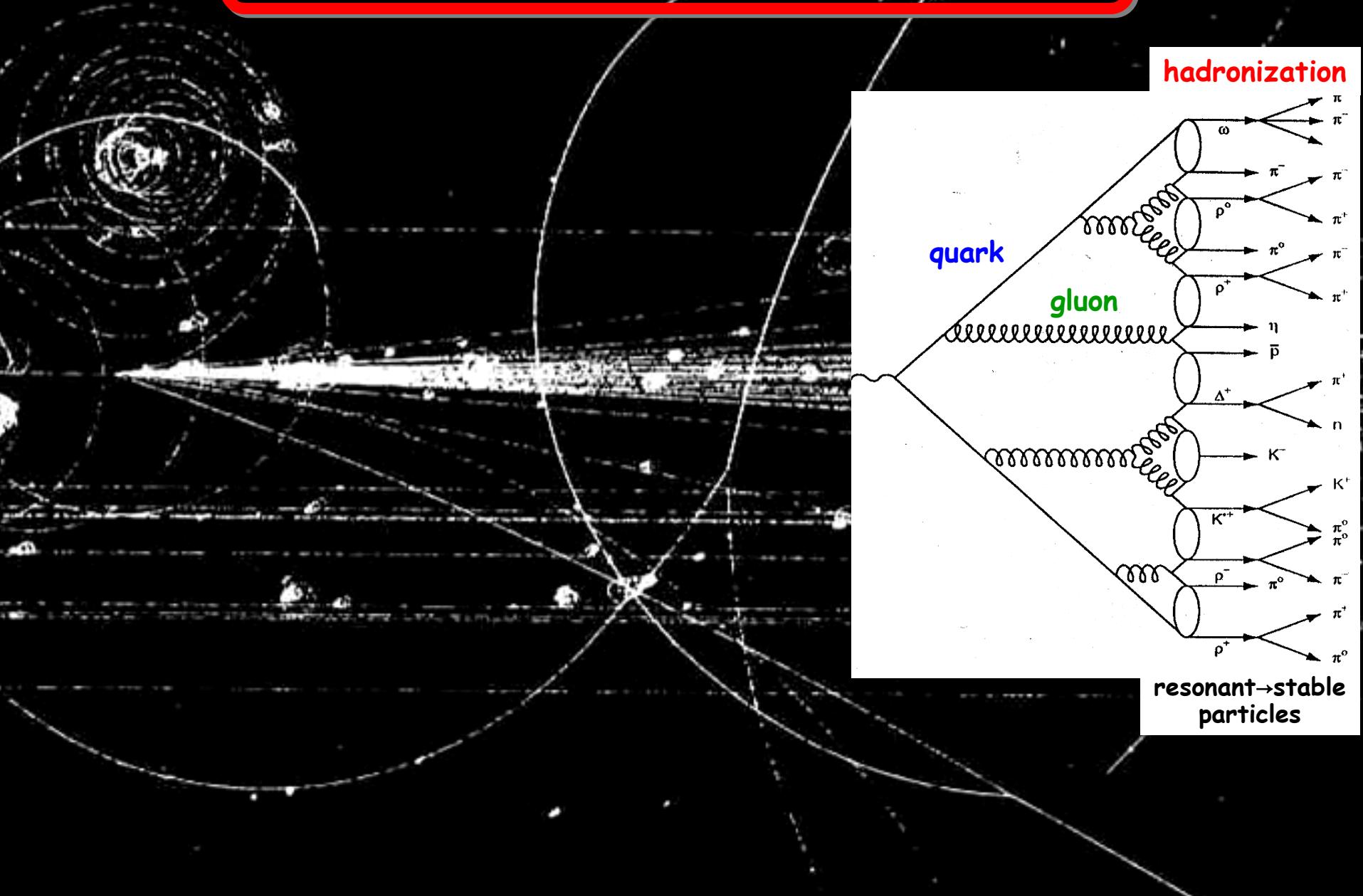
Quantum Chromo-Dynamics

Gauge field theory of strong interactions

in $SU(3)_{\text{COLOR}}$

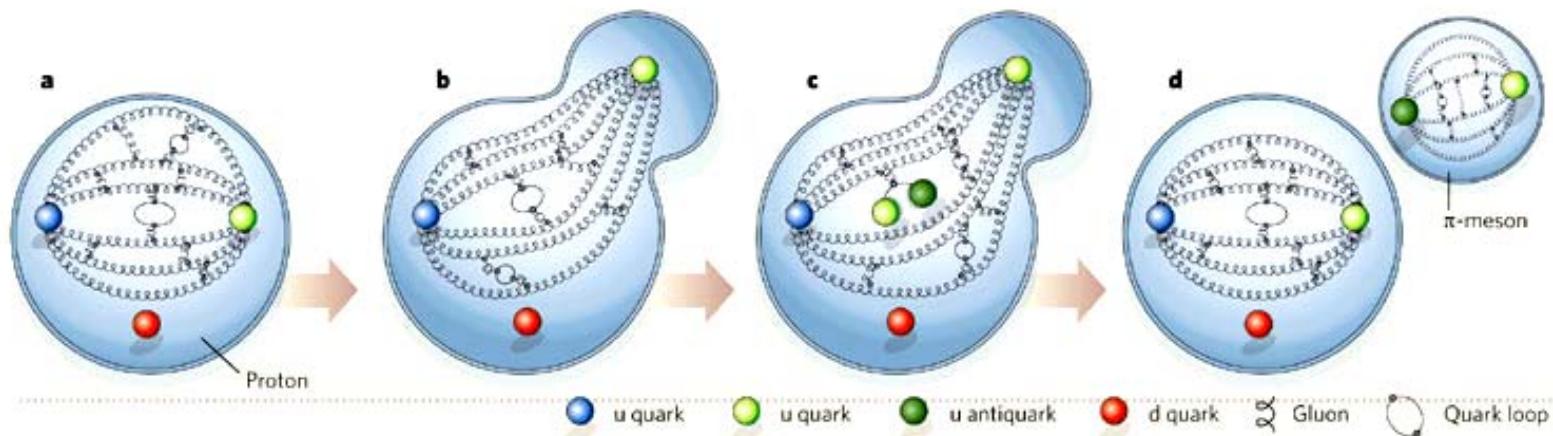
1. Color
2. Gluons
3. Quarks
4. Confinement + asymptotic freedom
5. Running coupling constant

Hadronic interactions

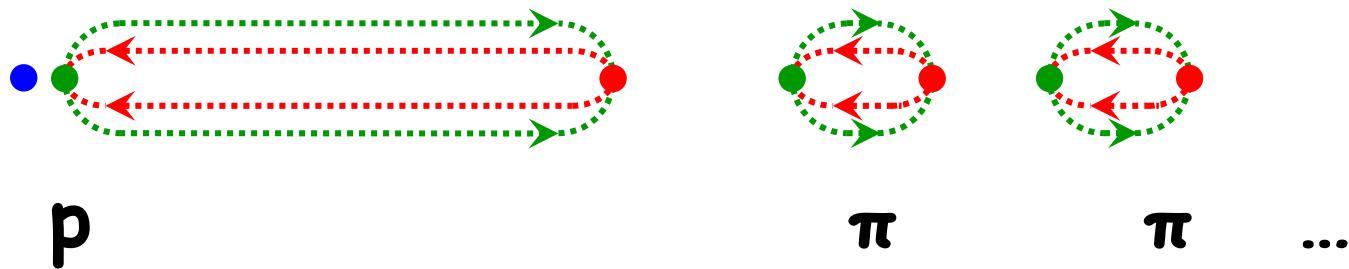


Confinement

Why no free quarks, no fractional charge observed ?!



color string: tension 200 MeV/fm $\sim m_\pi$



1. Confinement:

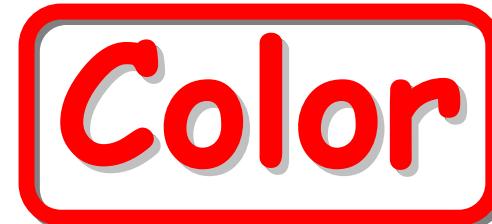
no free quarks, no $1/3$ charge observed !

which law forbids that + enforces trinity (qqq) ?

Color: hadrons have to be color **singlets**

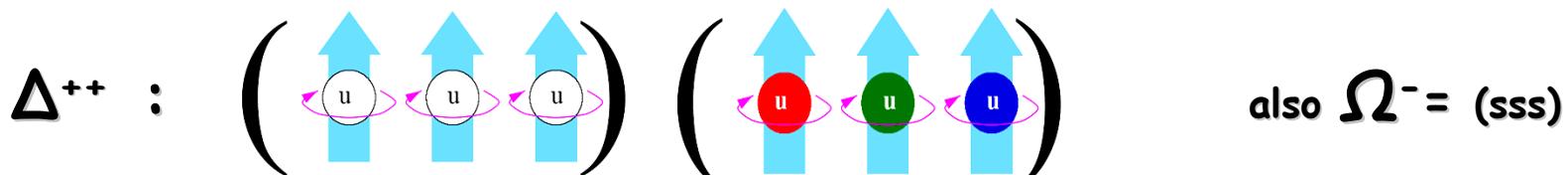
Mesons: $\{N\} \otimes \{\bar{N}\} = \{1\} \oplus \{N^2-1\}$

Baryons: $\{3\} \otimes \{3\} \otimes \{3\} = \{1\} \oplus \dots$



2. Spin-Statistics problem (Nambu 1964)

baryon decuplet $J^P = 3/2^+$



$$\Psi(qqq) = \Psi(\text{space}) \Psi(\text{spin}) \Psi(\text{flavor}) \Psi(\text{color})$$

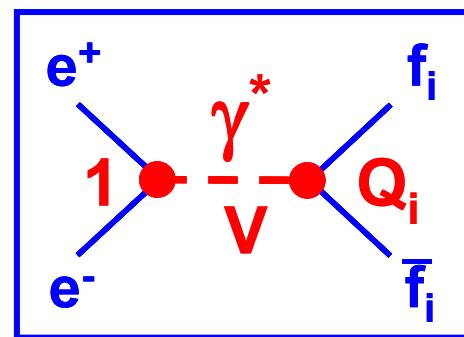
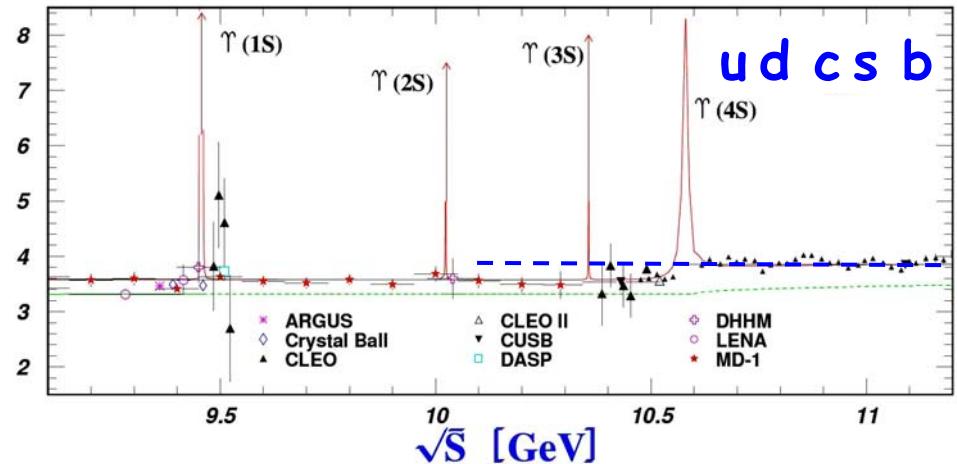
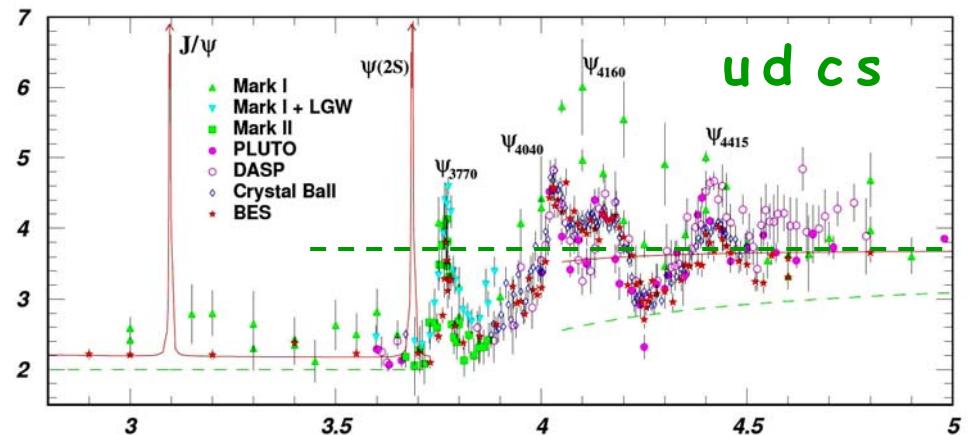
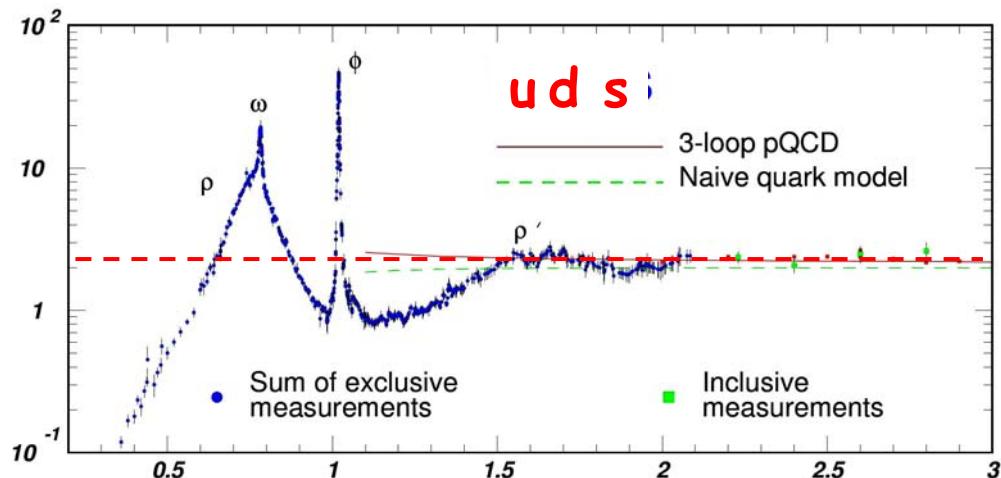
$$\text{asy} = \text{sym} \times \text{sym} \times \text{sym} \times \text{asy}$$

$$\Psi_c(q\bar{q}) = (\bar{r}r + \bar{g}g + \bar{b}b) / \sqrt{3}$$

$$\Psi_c(qqq) = (rgb - grb + rbg - gbr + brg - bgr) / \sqrt{6} \quad \text{asym. color wave fct.!}$$

3. $R_{QCD} = \sigma(e^+e^- \rightarrow q\bar{q}) / \sigma(e^+e^- \rightarrow \mu^+\mu^-) \sim N_c :$

Color counting



$$\begin{aligned}
 R &= \frac{\sigma(e^+e^- \rightarrow q\bar{q} \rightarrow \text{hadrons})}{\sigma(e^+e^- \rightarrow \mu^+\mu^-)} \\
 &= \sum_{\text{flavors}} Q_i^2 =
 \end{aligned}$$

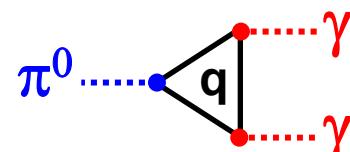
N_F	Q_i^2	N_C
3	$\left(\frac{2}{3}\right)^2 + \left(\frac{1}{3}\right)^2 + \left(\frac{1}{3}\right)^2$	$= \frac{6}{9} \cdot 3 = \frac{6}{3}$
4	$\left(\frac{2}{3}\right)^2 + \left(\frac{1}{3}\right)^2 + \left(\frac{2}{3}\right)^2 + \left(\frac{1}{3}\right)^2$	$= \frac{10}{9} \cdot 3 = \frac{10}{3}$
5	$\left(\frac{2}{3}\right)^2 + \left(\frac{1}{3}\right)^2 + \left(\frac{2}{3}\right)^2 + \left(\frac{1}{3}\right)^2 + \left(\frac{1}{3}\right)^2$	$= \frac{11}{9} \cdot 3 = \frac{11}{3}$
	$(u)(d)(s)$	
	$(u)(d)(c)$	
	$(u)(d)(s)(b)$	

4. OZI rule:

Why are $\Xi(ss)$, $\Psi(cc)$, ... so narrow? Why is  suppressed?

5. Triangle anomaly (Adler, Bell, Jackiw 1969)

$$\Gamma(\pi^0 \rightarrow \gamma\gamma) = (\alpha/2\pi)^2 (Q_u^2 - Q_d^2) N_c^2 m_\pi^2 / (8\pi f_\pi)$$



$$\Gamma = 0.86 \text{ eV} \quad \text{THEORY} \quad N_c=1$$

$$\Gamma = 7.8 \pm 0.5 \text{ eV} \quad \text{EXPT.}$$

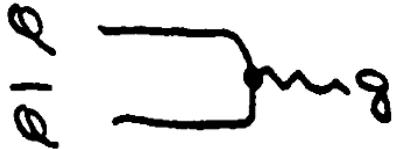
$$\Gamma = 7.75 \text{ eV} \quad \text{THEORY} \quad N_c=3$$

6. τ^- branching ratios

$$\frac{\Gamma(\tau^- \rightarrow e^- \bar{v}_e v_\tau)}{\Gamma_{\text{tot}}} = \frac{\Gamma_e}{\Gamma_e + \Gamma_\mu + \Gamma_{\text{had}} \cdot N_c} = \frac{1}{1+1+3} = \frac{1}{5} = (17.8 \pm 0.1) \%$$

Vector Mesons

$J^{PC} = 1^{--}$

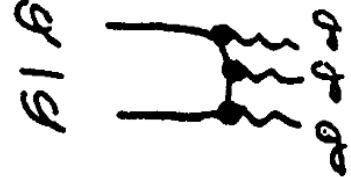


OZI rule

α_s^1 forbidden: free color !

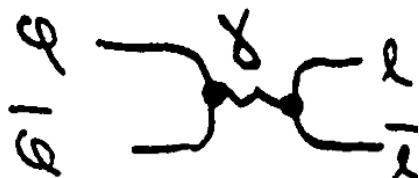


α_s^2 forbidden: $1^{--} \neq (...)^2$



α_s^3 ok: measure α_s

$V \rightarrow ggg \rightarrow 3$ gluon jets



$$\frac{\Gamma(V \rightarrow 3g \rightarrow \text{hadrons})}{\Gamma(V \rightarrow \gamma \rightarrow l^+l^-)} = \frac{\alpha_s^3(m_V)}{\alpha^2 q_i^2} \frac{10(\pi^2 - 9)}{81\pi} (1 + \dots)$$

$\alpha^2 q_i^2$ measure quark charges q_i

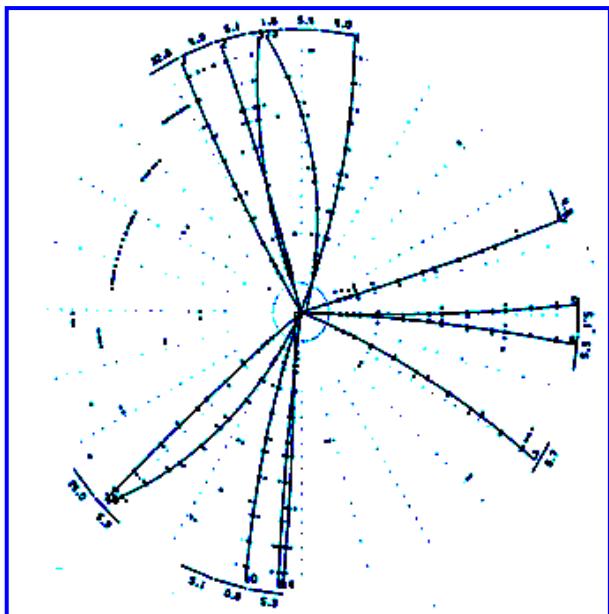


Gluon



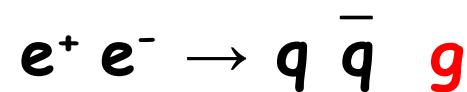
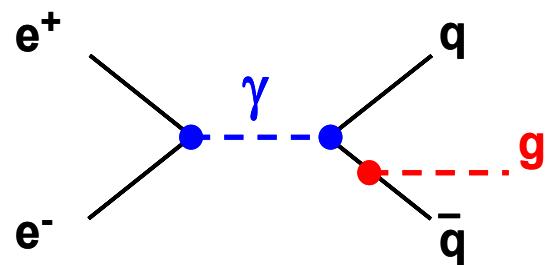
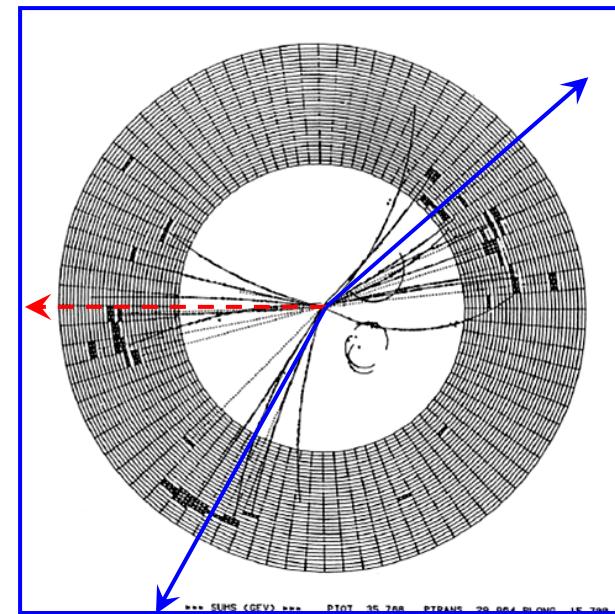
discovered 1979 at PETRA at DESY in

TASSO



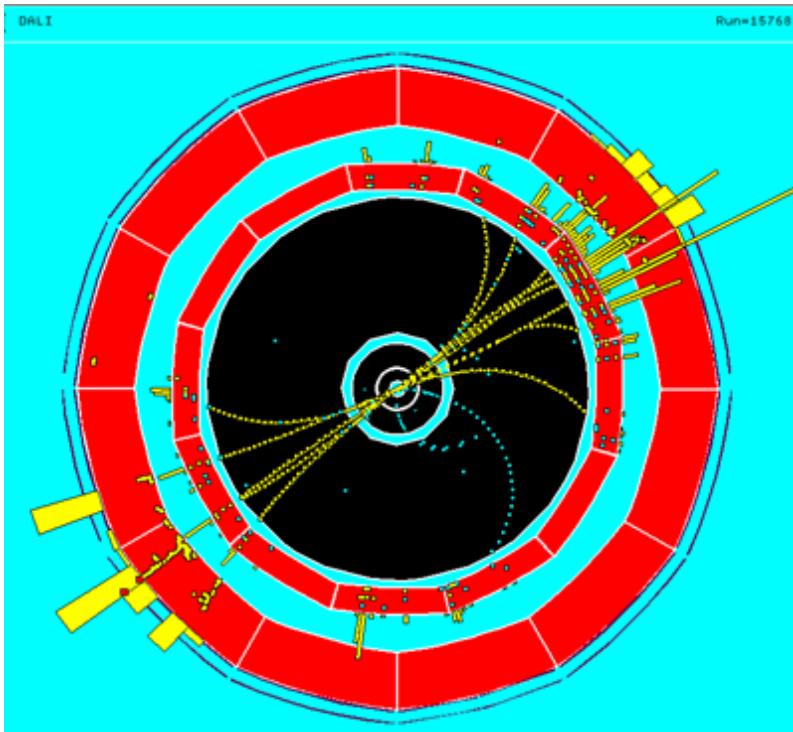
q

JADE

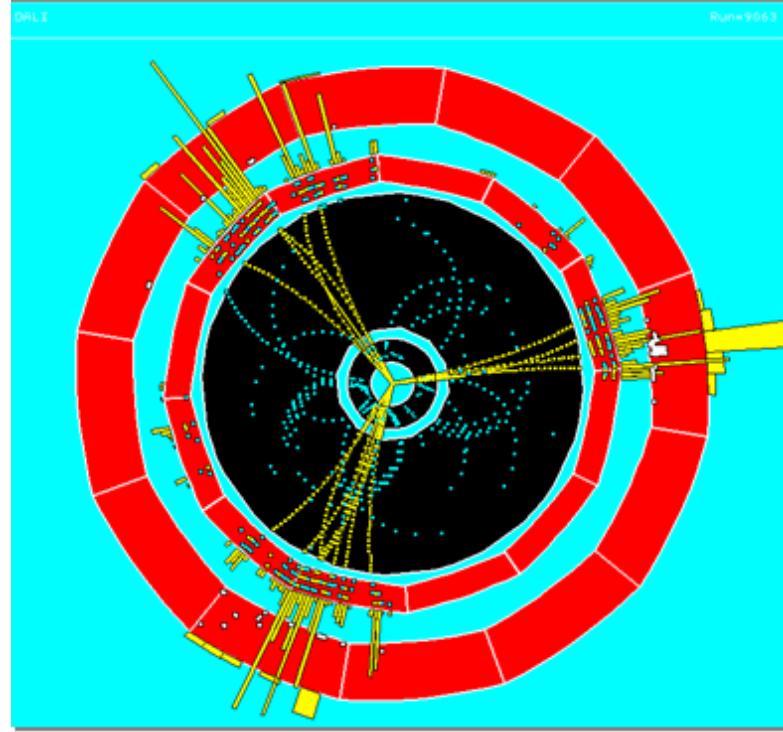


Europhysics Prize 1995

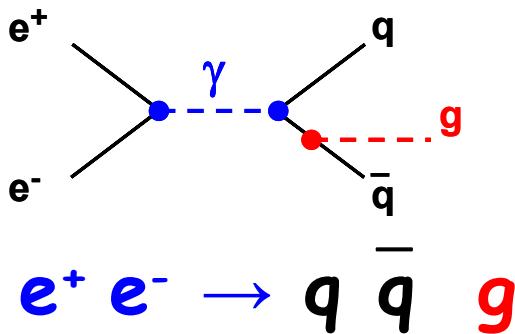
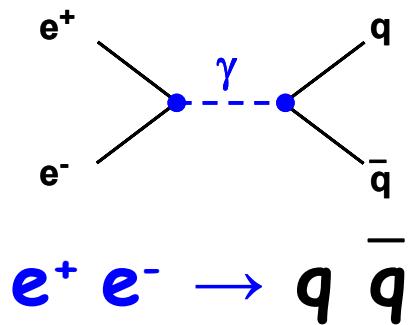
Gluon



A
L
E
P
H



LEP



Gluons

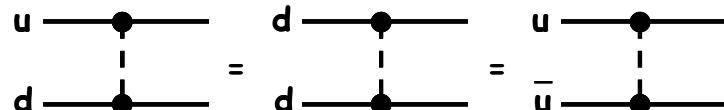
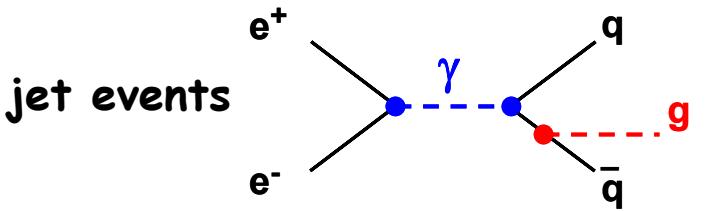
1. Existence: DESY PETRA 1978: 3 jet events

2. Couplings

SU(3): N=3 color charges
group has $N^2-1 = 8$ generators

$$\{3\} \otimes \{\bar{3}\} = \{1\} \oplus \{8\}$$

strong int. conserves isospin,
not flavor dependent

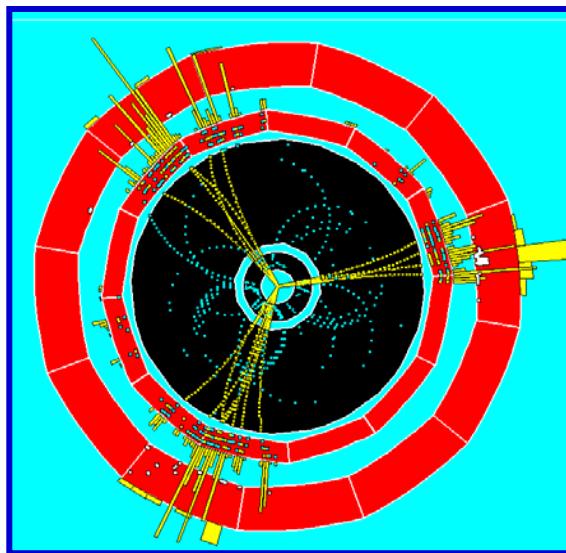


$$\{1\} = (\bar{r}r + \bar{g}g + \bar{b}b) / \sqrt{3} \quad \text{color singlet, meson, color blind}$$

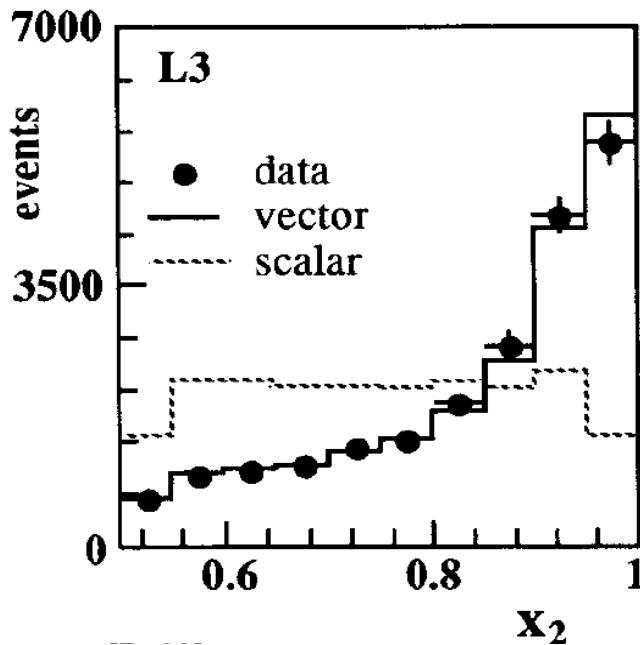
$$\{8\} = \bar{r}g + \bar{r}b + \bar{b}g + \bar{b}r + \bar{g}b + \bar{g}r + (\bar{r}r - \bar{g}g) + (\bar{r}r + \bar{g}g - 2\bar{b}b)$$

8 colored gluons

3. Spin: jet-jet angle $\Rightarrow J^P = 1^-$



collinearity of 2nd energetic jet:



Gluon spin

LEP

$e^+ e^- \rightarrow 3$ jets:
order the scaled
jet energies x_i :

$$x_i = 2E_i / \sqrt{s}$$

$$x_1 > x_2 > x_3$$

infrared collinear singularity
of gluon bremsstrahlung

Running coupling constant

A constant is not constant:

- running coupling constant
- asymptotic freedom
- confinement

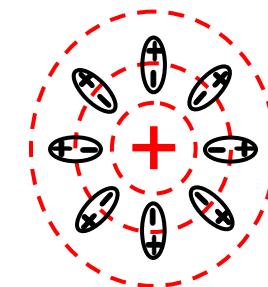


Vacuum polarization in e-e scattering:

$$\begin{array}{c}
 k \quad \quad \quad k' \\
 \gamma \quad \quad \quad Q^2 \\
 | \quad \quad \quad | \\
 \text{---} \quad \quad \quad \text{---} \\
 e \quad \quad \quad e_0
 \end{array}
 = - \quad - \quad + \dots = \left[\frac{1}{1 + \frac{1}{e_0}} \right]$$

physical naked vacuum $Q^2 = -(k^2 - k')^2$
 charge charge polarization

dielectric
screening:



- coupling \sim charge 2 : $F = \alpha/r^2$ $\alpha = e^2/4\pi$
- infrared stable: $\alpha = 1/137$
ultraviolet divergent - naked charge infinite !?
- Cutoff at arbitrary scale: renormalization !
energy scale: $Q^2 = -(k-k')^2$
- consider only evolution from energy scale Q to scale μ
UV divergences cancel

classical electron
radius
 $r = \alpha/m_e \sim 3 \text{ fm}$

Compton
wavelength
 $\lambda_c = 1/m_e$

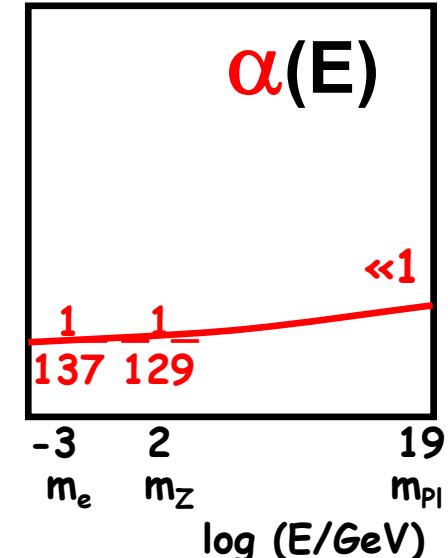
Running coupling

$\alpha = e^2/4\pi$ charge ~ coupling = fine structure constant

constant **not** constant:

$$2\beta = \partial \alpha(\mu) / \partial \ln \mu = \frac{2}{3\pi} \alpha^2 + \dots$$

$$\alpha(Q^2) = \frac{\alpha(\mu^2)}{1 - \frac{\alpha(\mu^2)}{3\pi} \log \frac{Q^2}{\mu^2}}$$



$\alpha(E)$ running (or crawling):

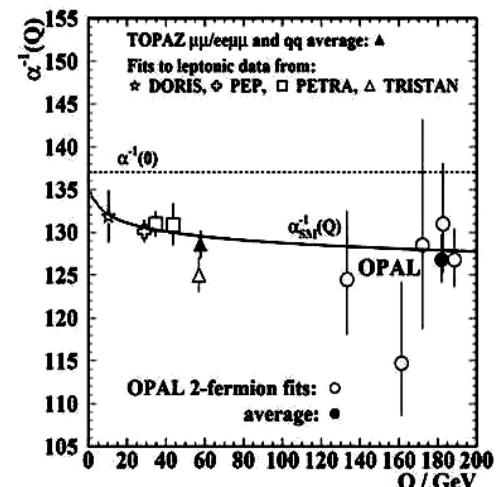
$$\delta \alpha(Q^2) / \delta Q^2 > 0$$

$$\alpha(m_e = 0.5 \text{ MeV}) = 1/137$$

$$\alpha(m_Z = 91 \text{ GeV}) = 1/128.9 \quad \text{CERN LEP}$$

$$\alpha(m_{\text{Pl}} = 10^{19} \text{ GeV}) \ll 1 \quad \text{more Fermion-Loops}$$

$$\alpha(\infty) \quad \text{undefined} \quad \text{no electric point interaction ?!}$$



QCD: the Lagrangian

$$\mathcal{L} = \frac{1}{4g^2} G_{\mu\nu}^a G_{\mu\nu}^a + \sum_j \bar{q}_j (\not{\partial}^\mu D_\mu + m_j) q_j$$

where $G_{\mu\nu}^a = \partial_\mu A_\nu^a - \partial_\nu A_\mu^a + i f_{bc}^a A_\mu^b A_\nu^c$

and $D_\mu = \partial_\mu + i t^a A_\mu^a$

That's it!

j ... quark flavors

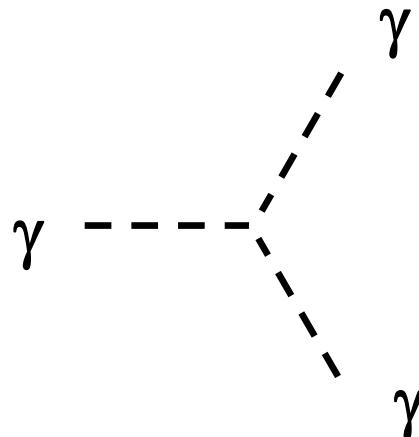
a,b,c ... 3 colors

μ, ν ... space-time

Quantum Chromo-Dynamics

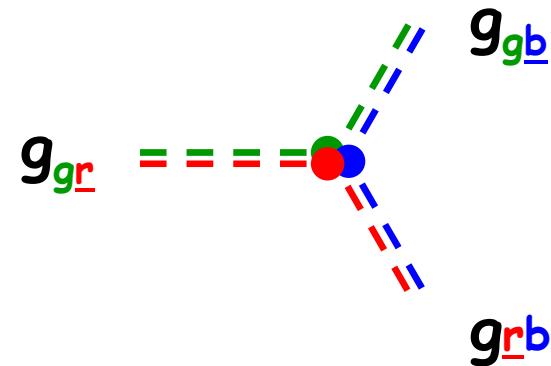
QED

- $U(1)$, abelian
- 1 charge type
- 1 photon:
- electric **neutral**
- **no** photon-photon coupling:
 - light does not clump !



QCD

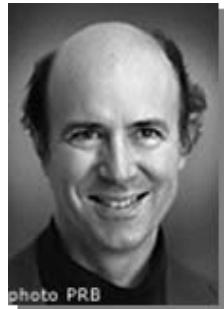
- $SU(3)_{\text{COLOR}}$, non-abelian
- 3 charge types: **r,g,b**
- $\{3\} \otimes \{\bar{3}\} = \{1\} \oplus \{8\}$: 8 gluons:
- carry color charges
- gluon-gluon self-coupling
 - gluonium, glue balls



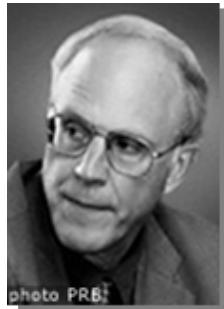
Nobel prize 2004



for the discovery of asymptotic freedom
in Quantum Chromo-Dynamics



D. Gross



D. Politzer



F. Wilczek

The most dramatic of these [consequences],
that protons viewed at ever higher resolution
would appear more and more as field energy (soft glue),
was only clearly verified at



HERA

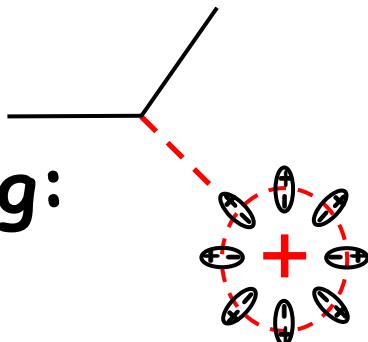


twenty years later.

F. Wilczek, 2001

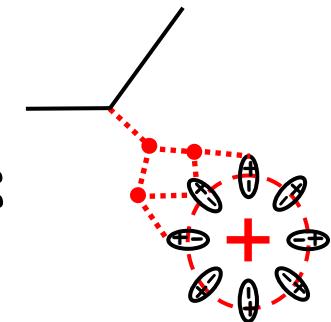
QED and QCD

Screening:



$$\gamma = \dots - \text{dipole} + \dots$$

Anti-Screening:



$$g = \dots - \text{dipole} + \text{gluon} + \dots$$

$$-2N_F + 11N_c$$

$$\alpha(Q^2) = \frac{\alpha(\mu^2)}{1 + b_0 / 4\pi \alpha(\mu^2) \log(Q^2 / \mu^2)}$$

$$b_0 = -4/3$$

$$\delta \alpha(Q^2) / \delta Q^2 > 0$$

$$b_0 = (-2N_F + 11N_c)/3$$

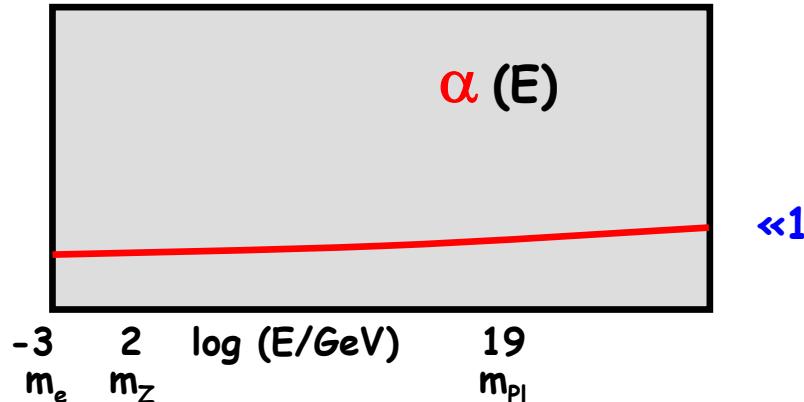
$$\delta \alpha(Q^2) / \delta Q^2 < 0$$

gluon massless ! $SU(2)_W$: $m_W > 10^5 m_e$

QED and QCD

QED: $\delta \alpha(\mu^2) / \delta \mu^2 > 0$ **screening**

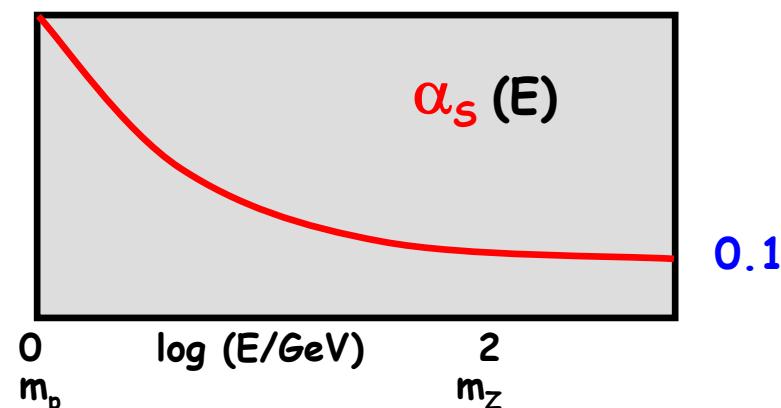
IR:
 $\alpha = 1/137$



UV:
 Landau
 Singularity

QCD: $\delta \alpha(\mu^2) / \delta \mu^2 < 0$ **anti-screening**

IR:
 $\alpha_s \rightarrow \infty$
 collapse of
 perturbation theory
 infrared slavery
 confinement



UV:
 $\alpha_s \rightarrow 0$
 asymptotic
 freedom

Confinement

instead of $\alpha_s(\mu^2)$ define

$$\Lambda = \mu \exp [-2\pi/(b_0 \alpha_s(\mu^2))]$$

$$\alpha_s(Q^2) = \frac{4\pi}{9 \ln(Q^2/\Lambda^2)} + \dots \quad (N_F=3)$$

$$\alpha_s(Q^2 \rightarrow \Lambda^2) \rightarrow \infty$$

collapse of perturbation theory
nuclear force confines - infrared slavery:

no free quarks !

$\hbar c \approx 200 \text{ MeV} \cdot \text{fm}$

QCD scale • **proton radius**

proton = 'QCD black hole'

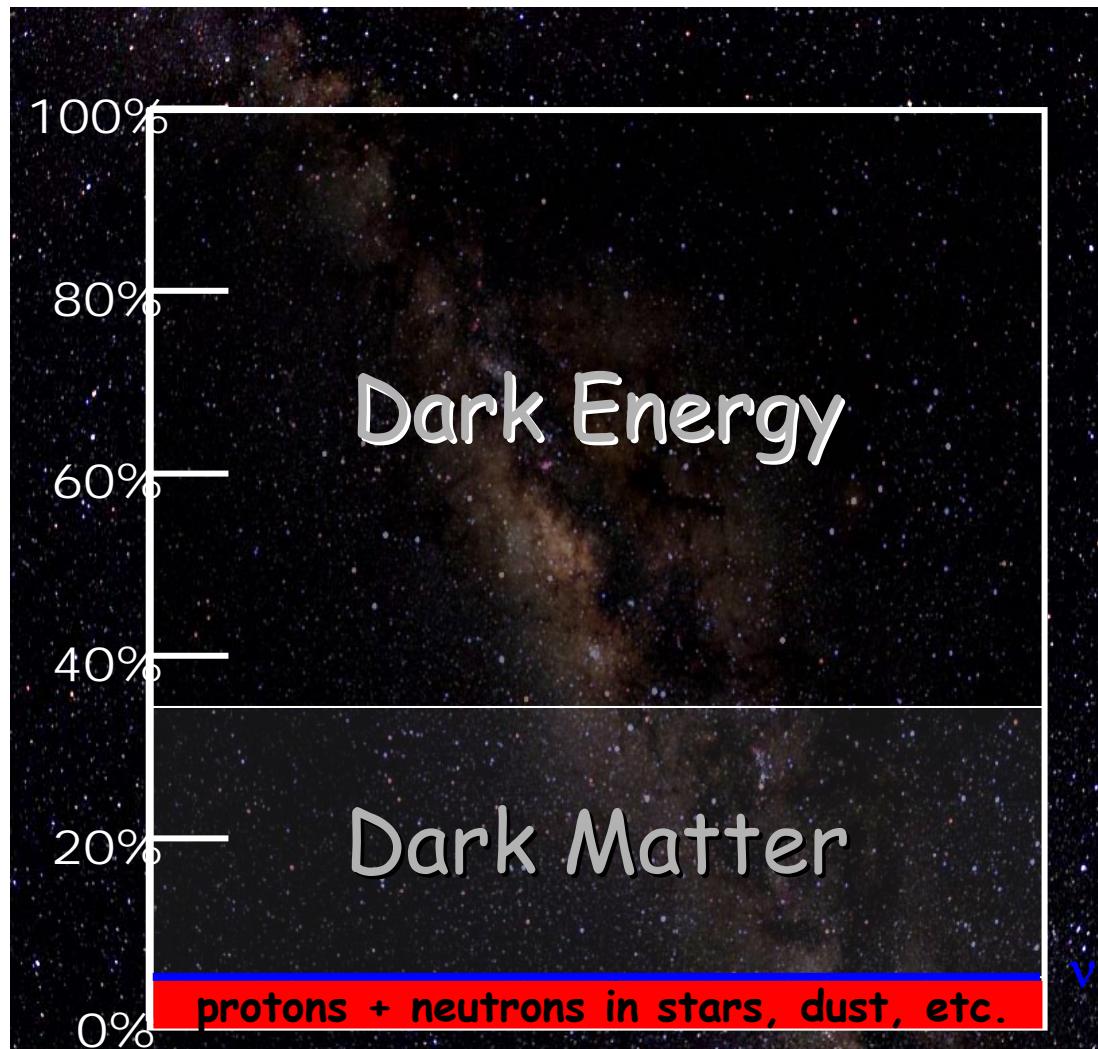
From asymptotic freedom to infrared slavery

Quarks are
born free,
but everywhere
they are
in chains.

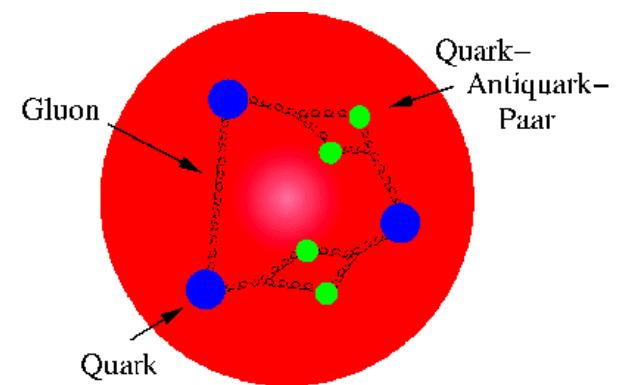
F.Wilczek, Nobel talk, 2004.

J.J. Rousseau, Du Contrat Social, 1762:
«L'homme est né libre et partout il est dans les fers.»

Mass vs Energy



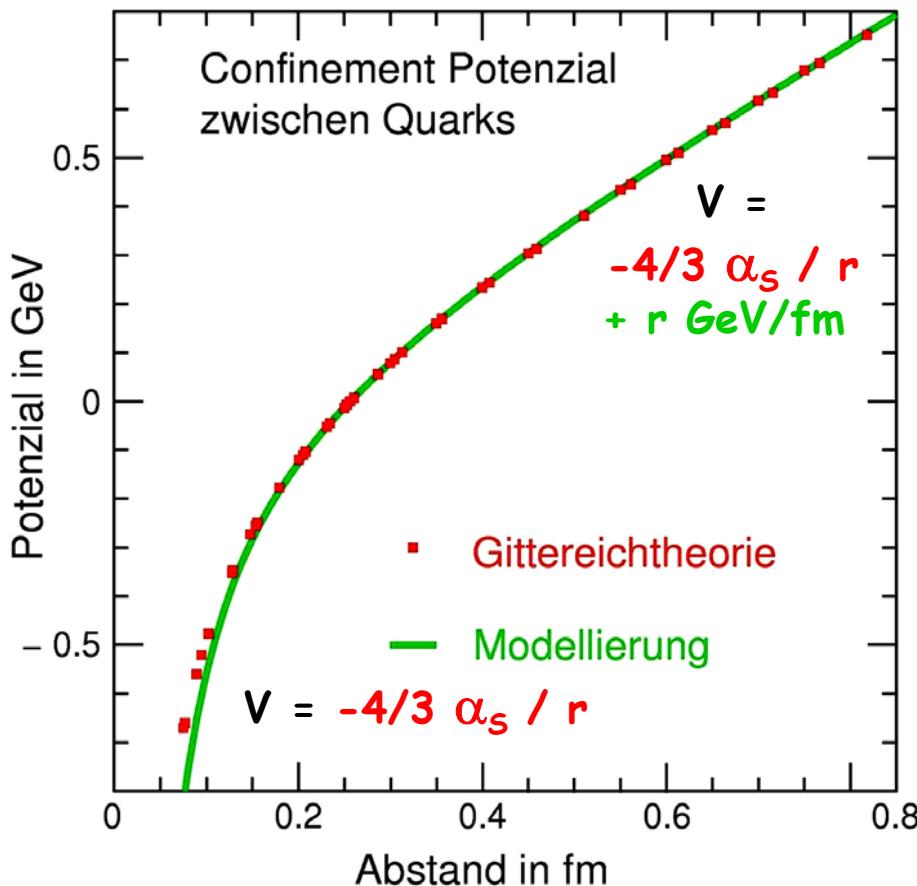
Nucleon mass



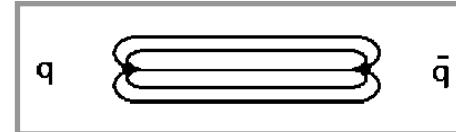
~ % due to
quark masses

rest:
binding energy
of partons

Confinement



hadron radius:
confinement



color string: constant force =
energy/length: $k = 1 \text{ GeV} / \text{fm}$

describes spectroscopy of
heavy quark bound states:

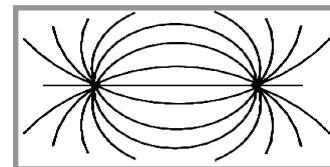
$$\Psi, \Psi', \Psi'', \dots = (c \bar{c})$$

$$Y, Y', Y'', \dots = (b \bar{b})$$

like positronium $= (e \bar{e})$

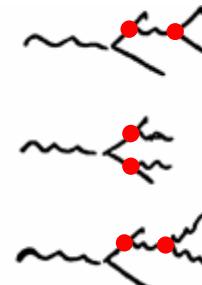
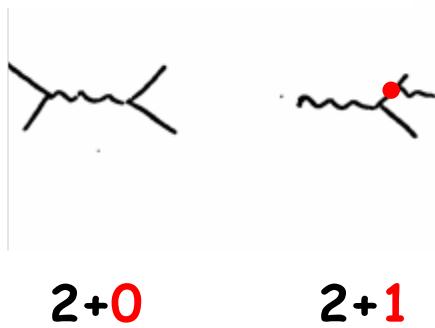
asymptotic freedom

at short distances = high energies:
Coulomb law



α_s measurement

e^+e^- interactions :

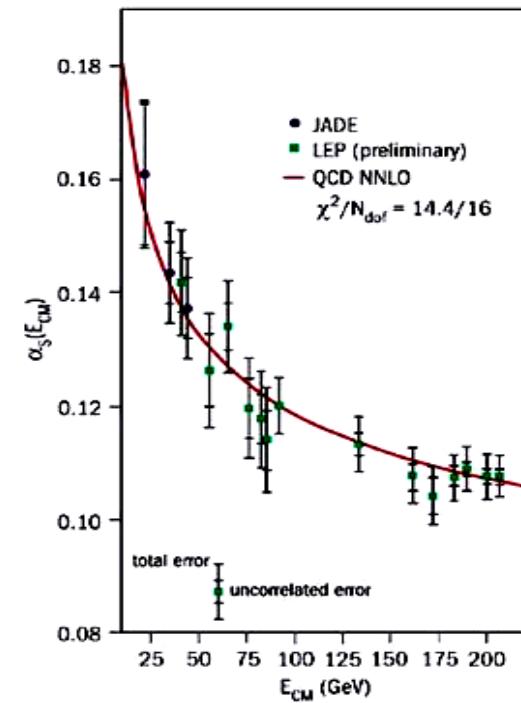


2+2 jets

$$R = \Gamma_h / \Gamma_l = R_0 [1^0 + (\alpha_s / \pi)^1 + 1.4(\alpha_s / \pi)^2 + \dots]$$

$\delta R/R \sim 10^{-3}$ syst. errors cancel: luminosity, ...

$$\delta \alpha_s / \alpha_s = \pi / \alpha_s \delta R/R \sim 25 \delta R/R$$



- CLEO, PEP, PETRA, TRISTAN: $\alpha_s(34 \text{ GeV}) = 0.15 \pm 0.03$
- LEP, SLC @ Z pole: $\alpha_s(91 \text{ GeV}) = 0.120 \pm 0.003 \text{ (exp)} \pm 0.002 \text{ (th)}$
- LEP2: $\alpha_s(172 \text{ GeV}) = 0.102 \pm 0.006$

α_s from jets

$e^+e^- \rightarrow 2+n$ jets



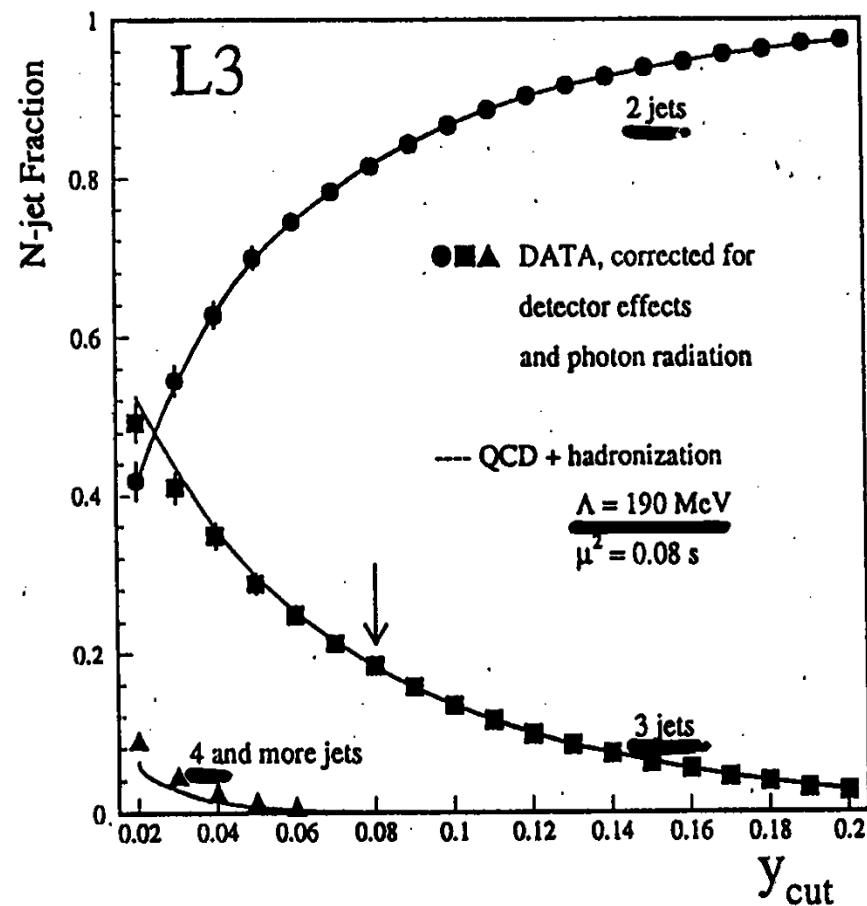
2+0



2+1

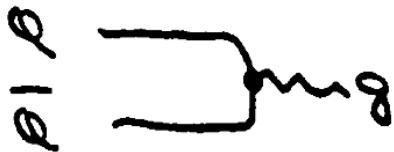


2+2 jets



Heavy Vector Mesons

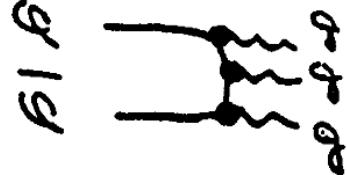
α_s in Quarkonium



α_s^1 forbidden: free color !

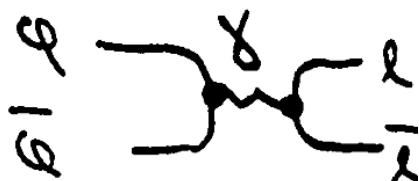


α_s^2 forbidden: $1^{--} \neq (...)^2$



α_s^3 ok: measure α_s

$$\gamma \rightarrow ggg \rightarrow 3 \text{ gluon jets}: R = \frac{\Gamma(V \rightarrow 3g \rightarrow \text{hadrons})}{\Gamma(V \rightarrow \gamma \rightarrow l^+l^-)} = \frac{\alpha_s^3(m_V)}{\alpha^2 q_i^2} \frac{10(\pi^2 - 9)}{81\pi} (1 + \alpha_s/\pi [\dots])$$



$\alpha^2 q_i^2$ measure quark charges q_i

$$\alpha_s(m_b=4.25 \text{ GeV}) = 0.18 \pm 0.01$$

$$\alpha_s(m_c=1.25 \text{ GeV}) \sim 0.4$$

Running coupling

$$\alpha_s(Q^2) = \frac{4\pi}{9 \ln(Q^2/\Lambda^2)} + \dots$$

$$\alpha_s(M_Z) = 0.1182 \pm 0.0027$$

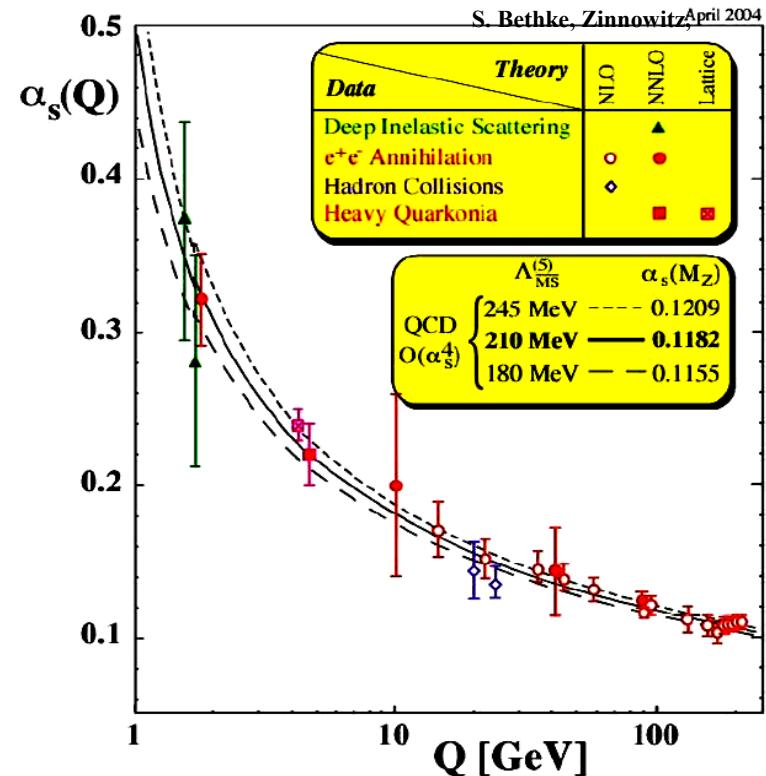
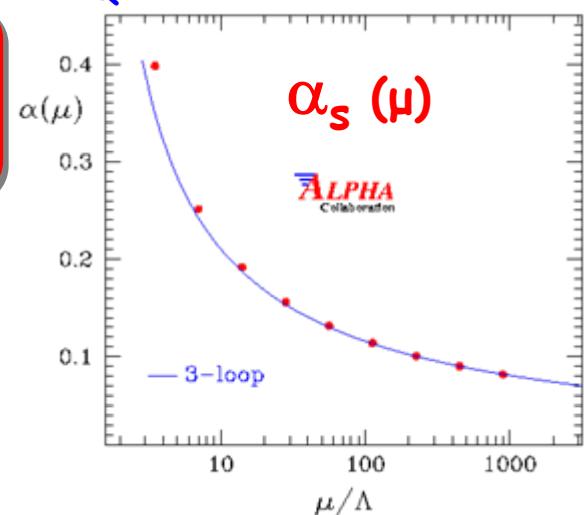
QCD scale:

$$\Lambda = 210 \pm 30 \text{ MeV} \quad (\text{MS}, N_F=5)$$

$$\hbar c \approx 200 \text{ MeV} \cdot \text{fm}$$

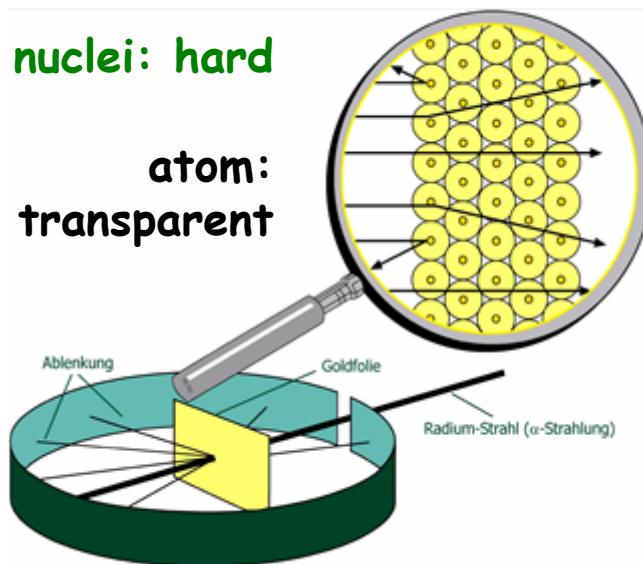
QCD scale • proton radius

QCD on a 3D discrete lattice:

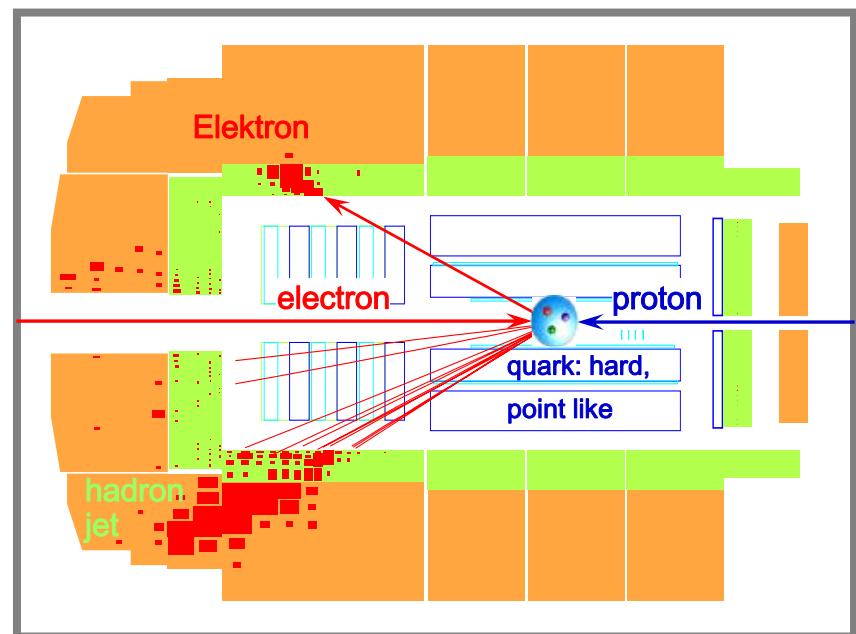


Structure of Matter

Rutherford 1910
atomic nucleus



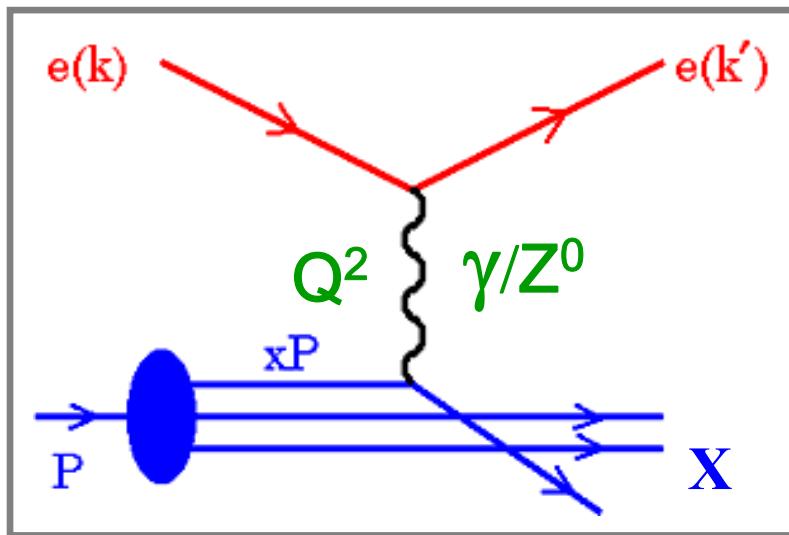
HERA 2000
proton structure



DIS Kinematics

$$s = (k+P)^2 = 4 E_e E_p \quad \text{center-of-mass energy}^2 \quad (m^2 \ll s)$$

$$s = 4 \cdot 27.6 \text{ GeV} \cdot 920 \text{ GeV} = (319 \text{ GeV})^2 \quad @ \text{HERA}$$



$$x = Q^2 / (2P \cdot q) \quad \text{Bjorken-}x$$

parton momentum fraction in proton
(quark, gluon)

$$Q^2 = -q^2 = -(k-k')^2$$

four-momentum transfer²

Rutherford 1911:
photon propagator

$$d\sigma / dQ^2 = 4\pi\alpha^2 / Q^4$$

$$d\sigma / d\cos(\theta) = \pi\alpha^2 / 2E^2 \sin^4(\theta/2)$$

Mott 1929, electron spin:

$$\sigma_{\text{Mott}} = \cos^2(\theta/2) \sigma_{\text{Ruth}}$$

$$W^2 = (P+Q)^2$$

mass² of the hadronic system X

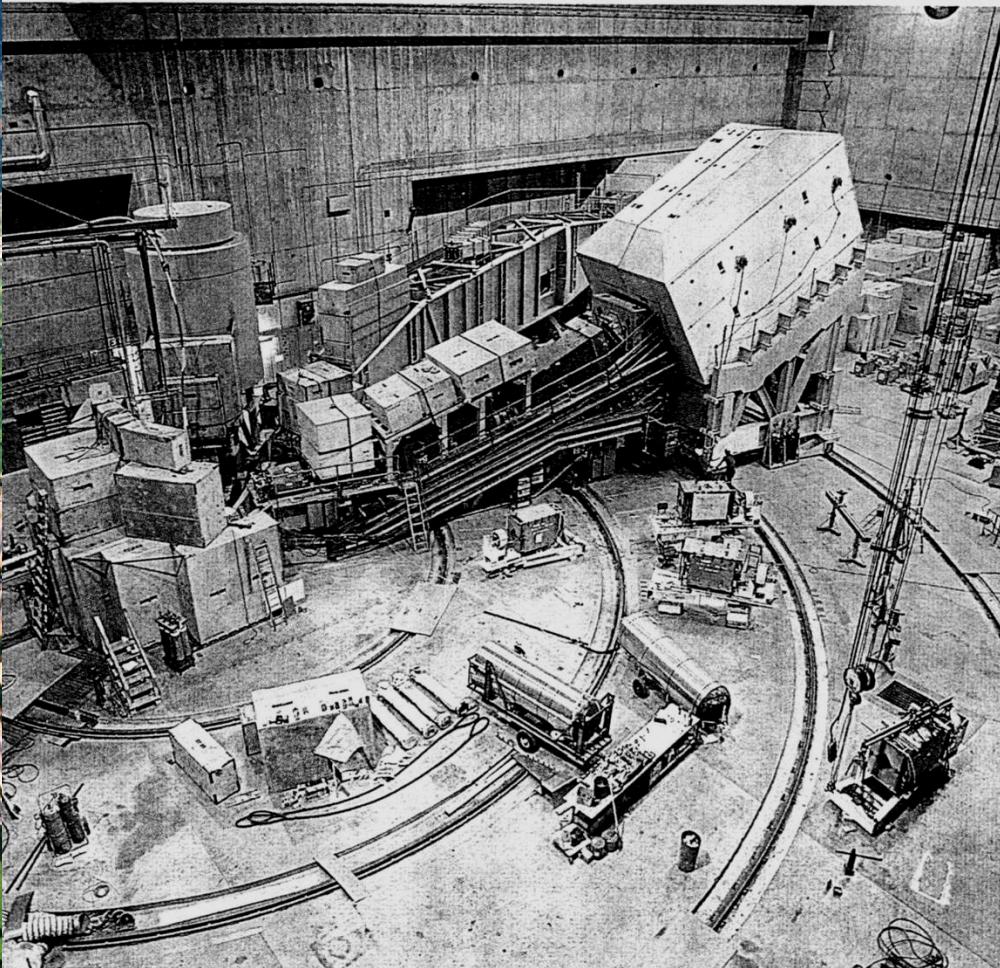
$$s = x \cdot Q^2$$

SLAC - Stanford Linear Accelerator



SLAC-MIT expt.

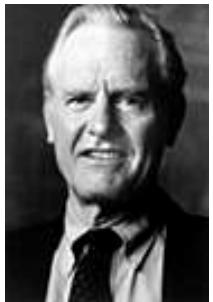
$E = 1.5\text{-}20 \text{ GeV}$ $\theta = 6^\circ\text{-}26^\circ$
 $Q^2 = 1\text{-}7 \text{ GeV}^2$



Quark Discovery



Friedman



Kendall



Taylor

Nobel prize



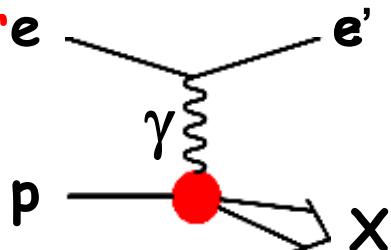
1990

Stanford Linear Acceleratore

USA, 1968-71:

electron-proton

deep inelastic scattering



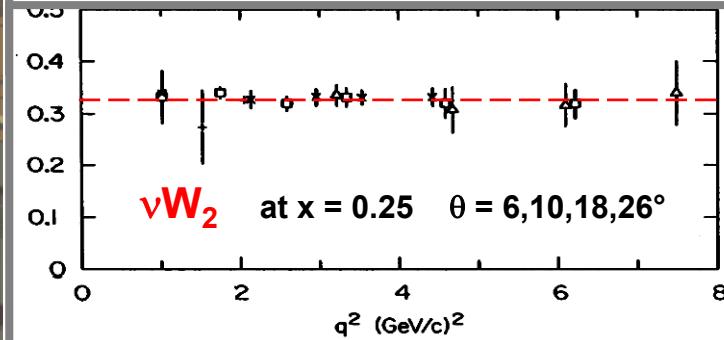
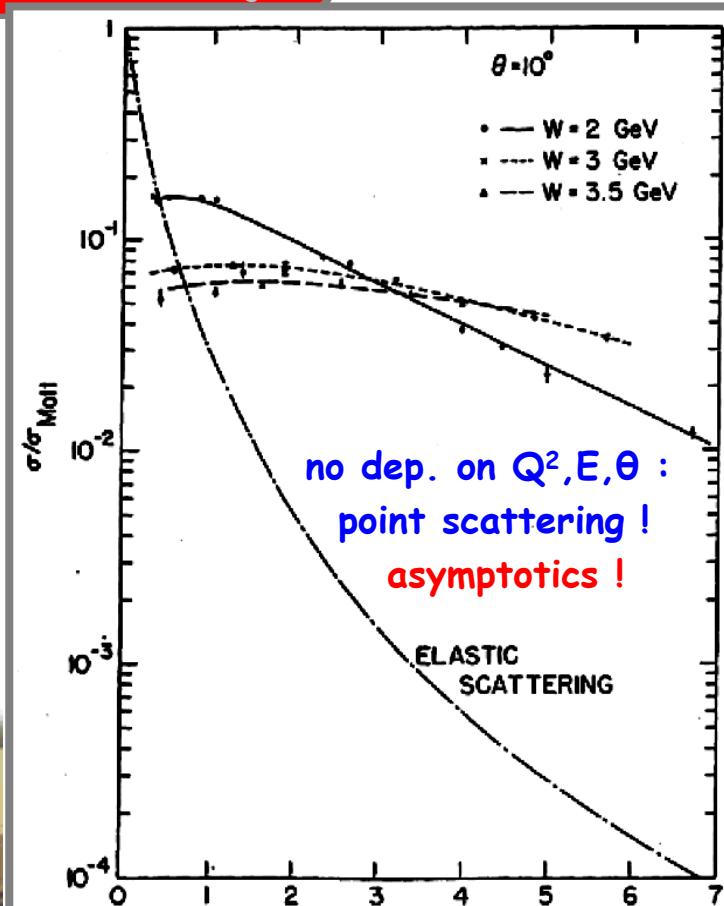
$E = 1.5 - 20 \text{ GeV}$

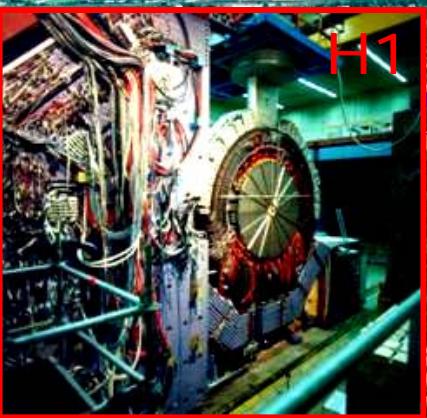
$\theta = 6^\circ - 26^\circ$

$Q^2 = 1 - 7 \text{ GeV}^2$



1 person





H1

Protons 920 GeV

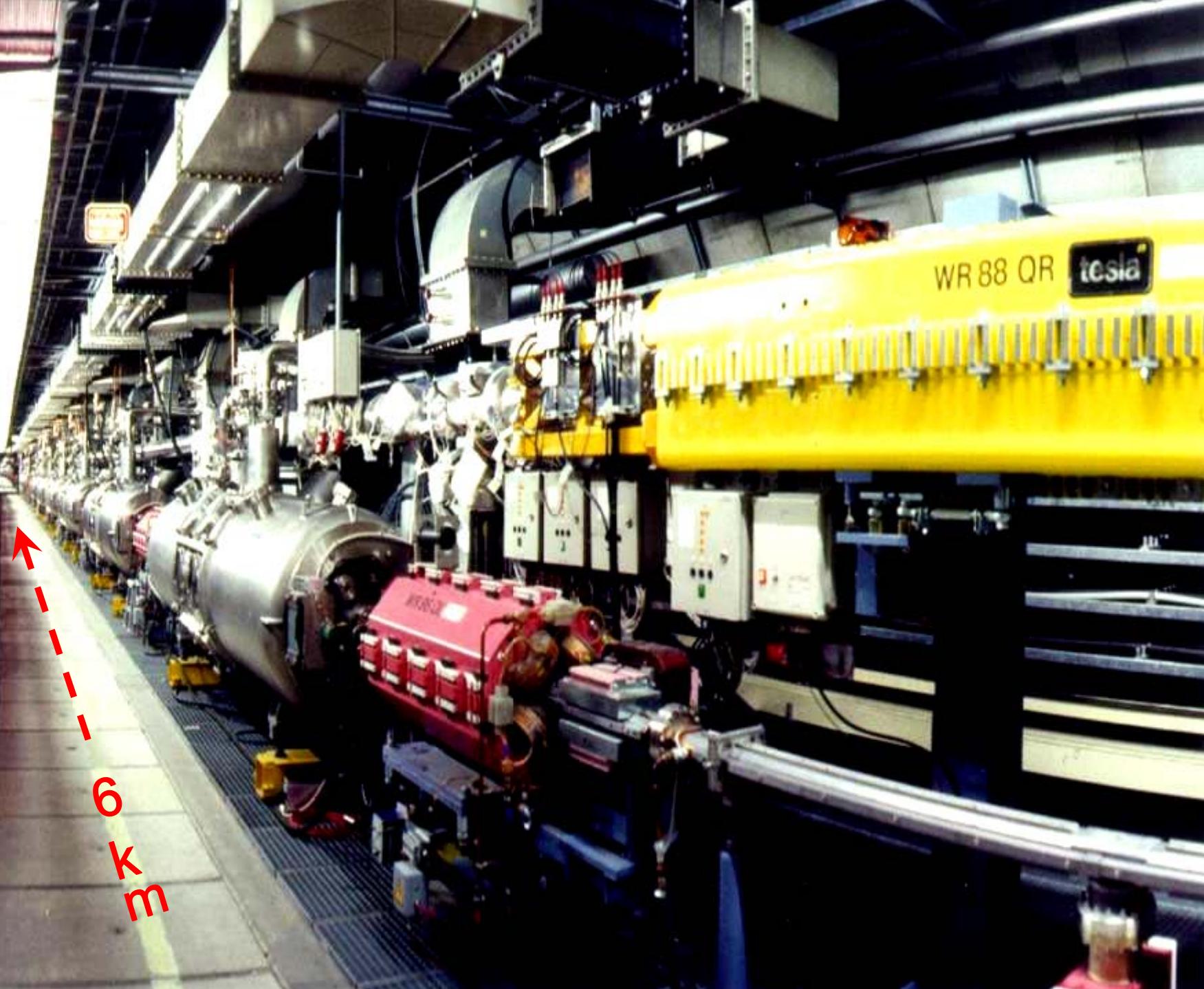


ZEUS



Electrons 27.6 GeV

PETRA



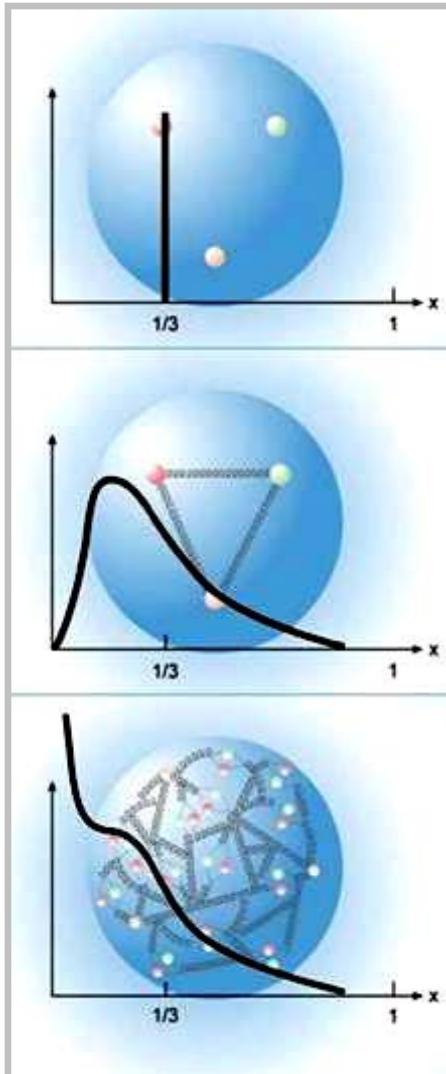
6
km



ZEUS

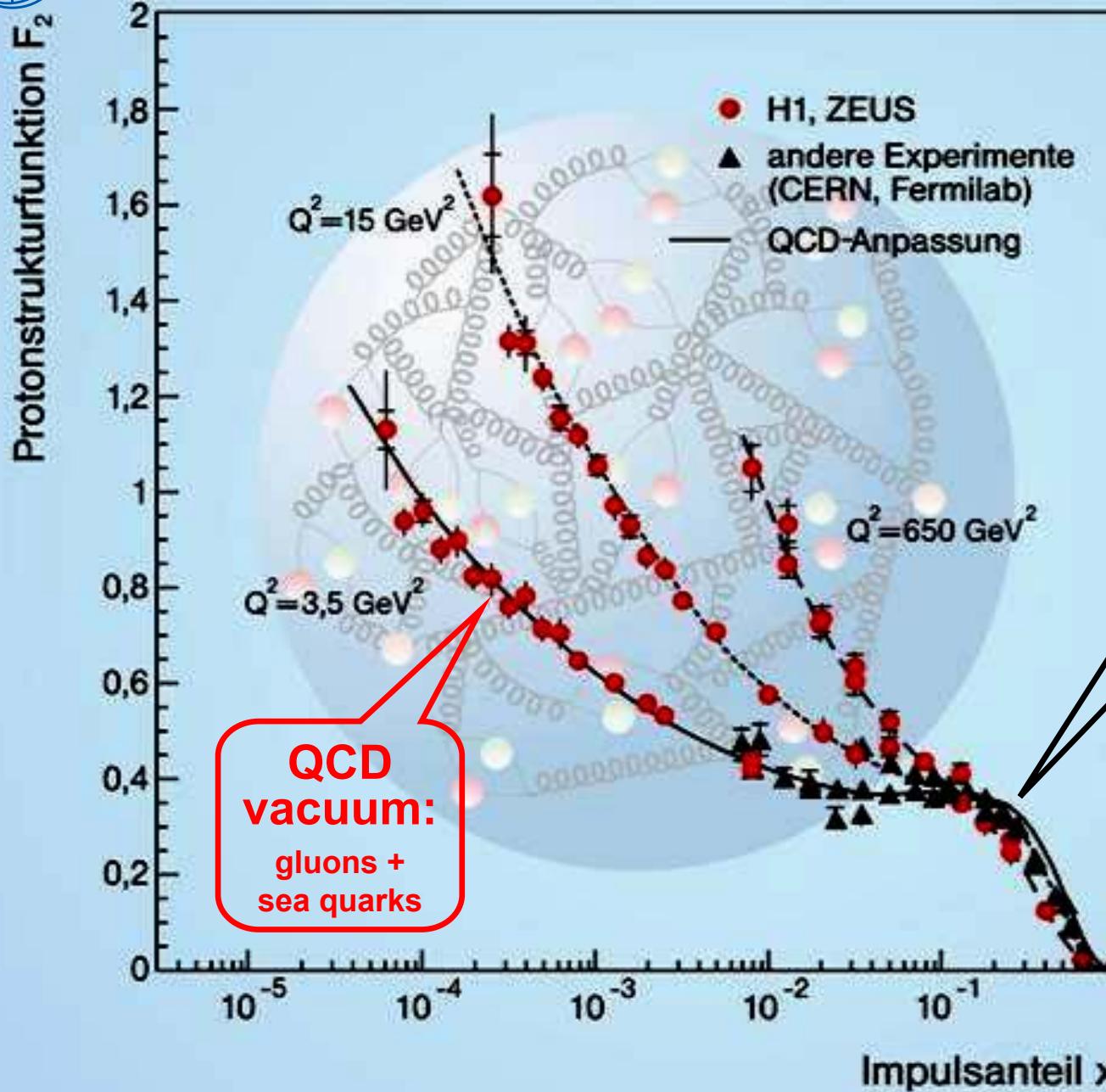
Proton Structure

QCD
Order



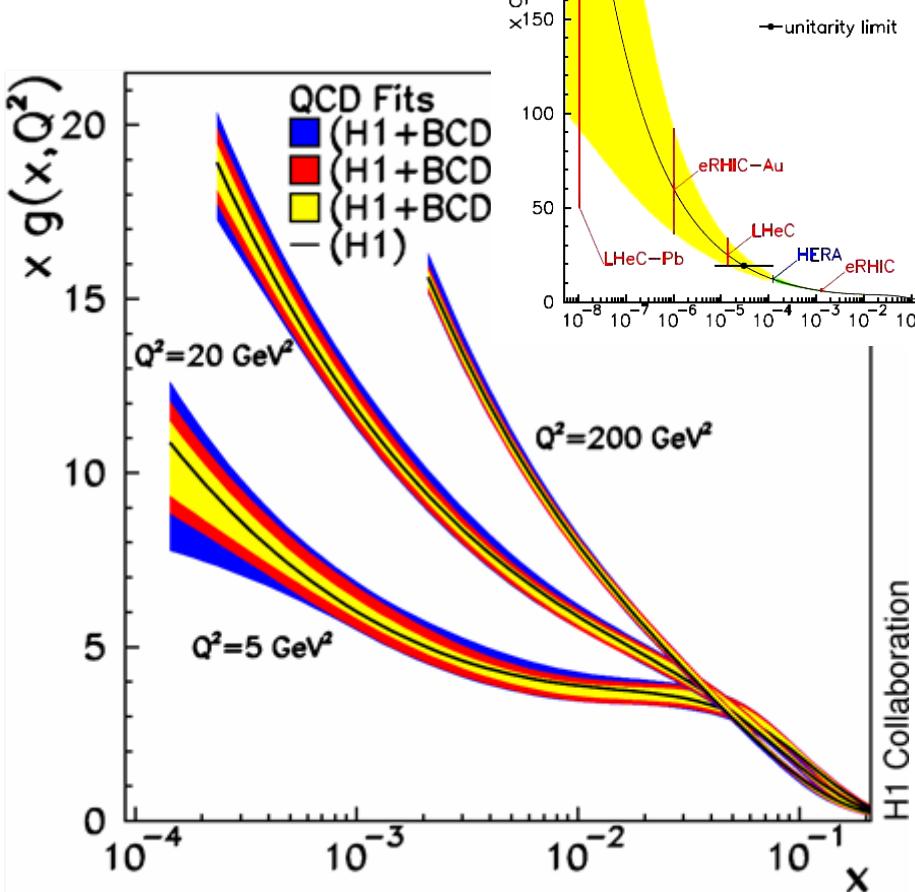
- **SLAC 1968:**
3 valence quarks:
 $x = 1/3$
- 3 bound quarks:
~ 50% of p in 3 quarks +
~ 50% of p in N gluons
quark momentum fraction in hadron:
 $x \sim 1/6$
- **HERA 1994:**
 $x > 0.1$: valence quarks
 $x \ll 0.1$: sea quarks + gluons :
pure QCD !

Proton Structure





gluon density



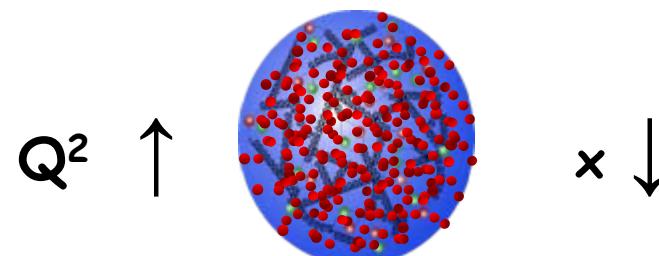
gluon self coupling

drives proton structure + $\alpha_s(Q^2)$

$$\delta \sigma(x, Q^2) / \delta \ln Q^2 \sim \alpha_s(Q^2) x g(x, Q^2)$$

test non-abelian QCD :

boil up QCD vacuum:

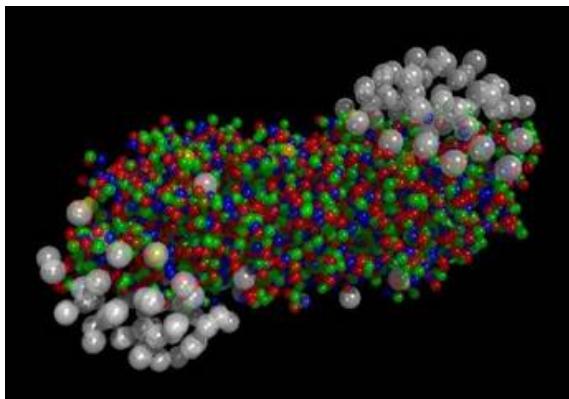
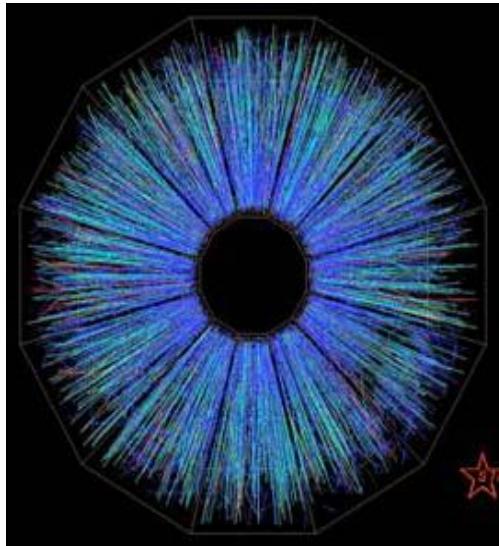
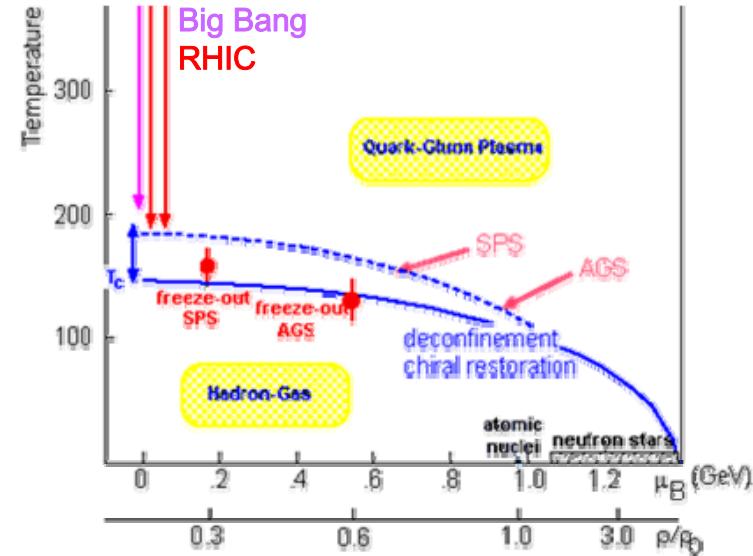


$\alpha_s \rightarrow 1$: end of perturbation theory?

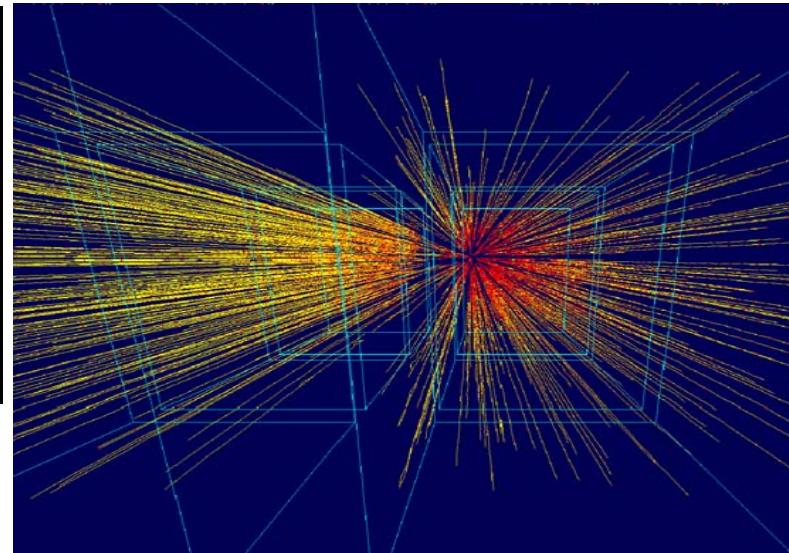
- gluons = bosons !
- gluon saturation ? shadowing ?
- gluonium, glue-balls ?
- quark-gluon plasma ?

Quark-Gluon Plasma

Big Bang: quarks are born free
after μs : QCD confinement:
phase transition quagma-hadron gas
nucleons freeze out at $T_c = 173 \pm 15$ MeV (lattice)



>1000 quarks:
thermalisation



RHIC at BNL Brookhaven:

- Au-Au collisions at 100 GeV/nucleon

ALICE at CERN LHC >2008:

- Pb-Pb at 3 TeV/nucleon

Quarks + Hadrons

HADRONS consist of : **QUARKS** are

- point like partons
 - with spin 1/2
 - with charges +2/3, -1/3
 - 3 valence quarks
 - + sea quarks carry $\sim p/2$
 - + gluons carry $\sim p/2$
 - colored
 - confined in the hadrons
 - asymptotically ($Q^2 \gg \Lambda^2$) free
- QCD** violates scaling:
- $\partial F_2(x, Q^2) / \partial \log Q^2$
 - $F_2 \neq 2x F_1$

7. Unification

PARTICLES

FORCES

			Electro-Magnet.	Weak	Nuclear	Gravitation
	Charge	Electric	Weak	Color	Mass	
	Symmetry	U(1)	SU(2)	SU(3)		
Matter Particles		Fermions J=1/2				
Quarks	Up	u	c	+2/3	I_W, Y_W	$r \ g \ b$
	Down	d	s	-1/3		
Leptons	Electrons	e	μ	-1	I_W, Y_W	
	Neutrinos	ν_e	ν_μ	0		
Force Particles			Bosons J=1			
Photon		γ				
Weak Bosons		W^+, Z^0, W^-				
Gluons		$8 g_{ij}$				
Graviton	(J=2)	G				

The Standard Model

- predictive power: c, b, t, ... ok
- Higgs, SUSY to be discovered !
- theory + experiment agree to 10^{-3}
with 3rd order radiative corr.: test of theory
- consistency of all parameters
- no new building blocks :
 - quarks
 - leptons
 - bosons: W' , Z'
- no new structure level :
 - composite leptons: e^* , ...
 - quarks: q^*
 - bosons: W^* , Z^*
 - lepto-quarks
- no new couplings:
lepton-quark universality
- no proton decay: baryon nr ok
- neutrino oscillations: lepton nr violated NEW PHYSICS !!!
- no magnetic monopole

The Standard Model

DESCRIBES THE PROPERTIES OF

ELEMENTARY PARTICLES

AND THEIR

- WEAK
- ELEKTRO-MAGNETIC
- STRONG

INTERACTIONS

PRECISELY + COMPLETELY.

HOWEVER ,

MANY QUESTIONS REMAIN OPEN ...

PARTICLES:

- NR OF FAMILIES ?
- LEPTON-QUARK SYMMETRY ?
- SUBSTRUCTURE of Quarks + Leptons ?
- MASS: Higgs Mechanism ?
 Quark + Lepton Masses ?
- NEUTRINO: Dirac or Majorana ?
- MIXING ANGLES of Quarks + Neutrinos ?
- SUSY ?

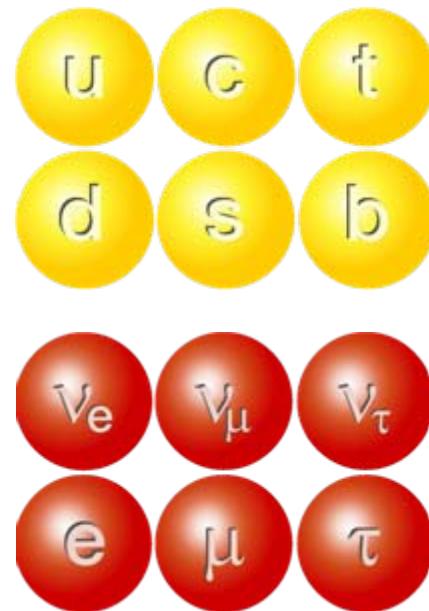
FORCES:

- STRUCTURE: $U(1) \otimes SU(2) \otimes SU(3)$?
- COUPLINGS: Values ?
- GRAND UNIFICATION: Scale + Scheme ?
- GRAVITATION AND SUPER-STRINGS ?
- EXTRA DIMENSIONS ?

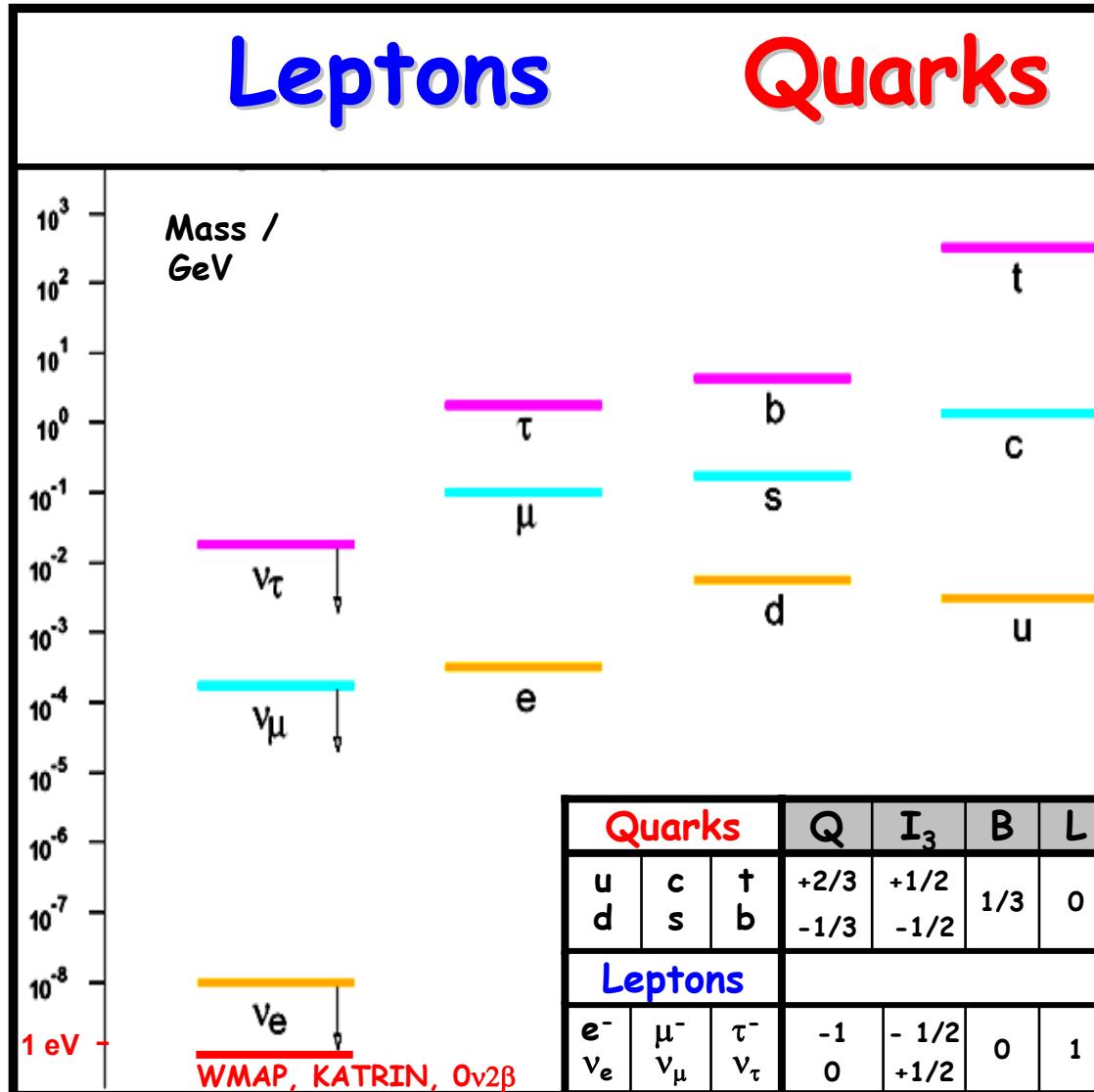
SYMMETRIES:

- P-VIOLATION ?
- CP-VIOLATION ?
- BARYON-NR ? Baryon Asymmetry of Universe ?
- LEPTON-NR ? Neutrino Oscillations !
- MAGNETIC MONOPOLES ?
- SYMMETRY BREAKING: HOW ?
- SUPER-SYMMETRY ?

Quest
ions



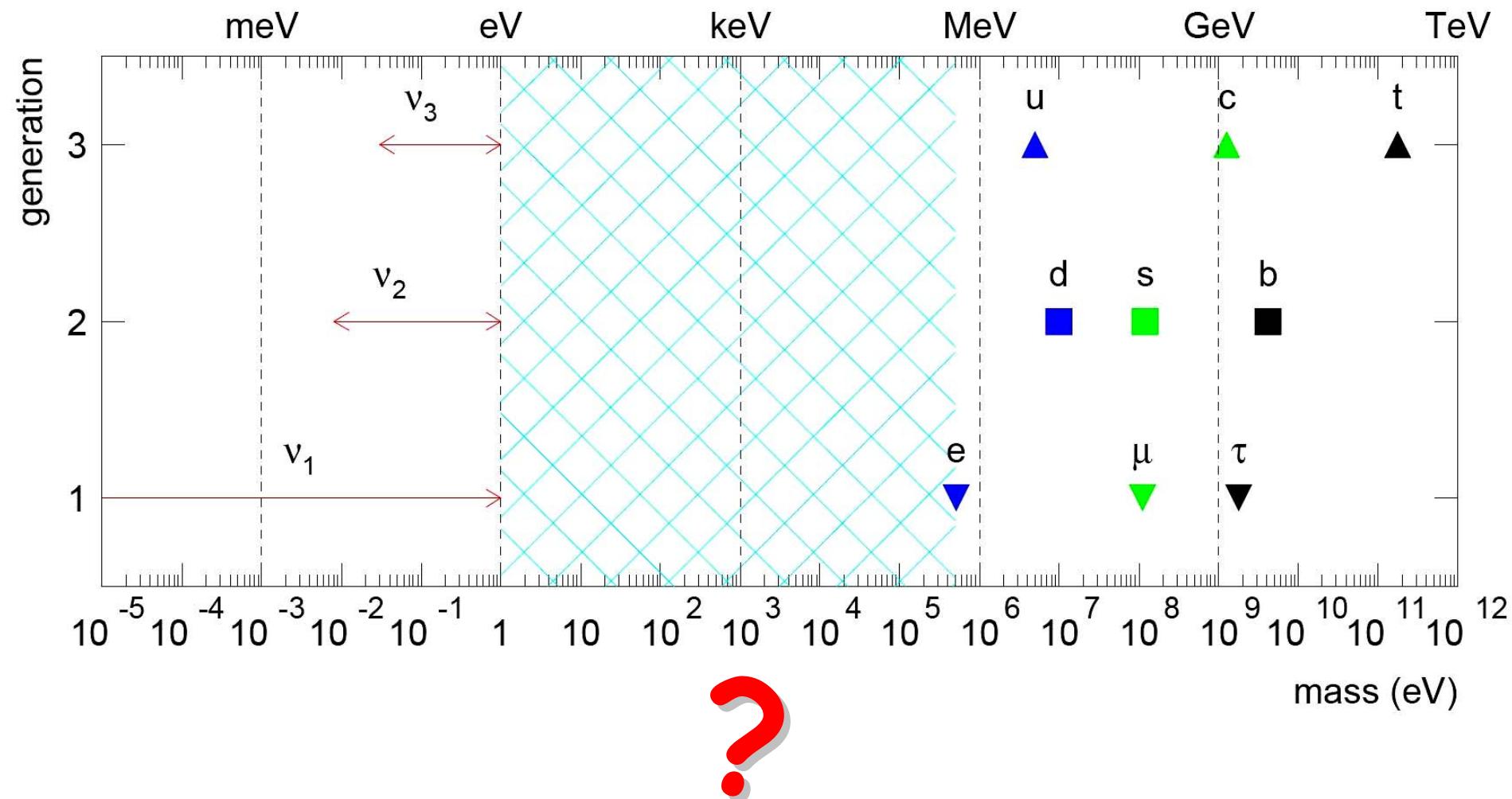
Fermion Masses



explained
by
Higgs
mechanism

?

Neutrino masses



What tells us Nature with this **mass spectrum**

Cosmology

Natural Units

M. Planck, 1899:

besides speed of light c and
Newton's constant G
find a third quantity $b=h$ that allows

„Einheiten für Länge, Masse, Zeit und Temperatur aufzustellen, welche ... ihre Bedeutung für alle Zeiten und für alle, auch außerirdischen und außermenschlichen Culturen nothwendig behalten.“

one year before Planck's law !

Planck

$$\begin{aligned} \text{mass} &= (\hbar c / G_N)^{1/2} = 1.2 \cdot 10^{19} \text{ GeV/c}^2 \\ \text{time} &= (\hbar G_N / c^5)^{1/2} = 5.4 \cdot 10^{-44} \text{ s} \\ \text{length} &= (\hbar G_N / c^3)^{1/2} = 1.6 \cdot 10^{-35} \text{ m} \end{aligned}$$

Die Mittel zur Festsetzung der vier Einheiten für Länge, Masse, Zeit und Temperatur werden gegeben durch die beiden erwähnten Constanten a und b , ferner durch die Grösse der Lichtfortpflanzungsgeschwindigkeit c im Vacuum und durch die der Gravitationskonstante f . Bezogen auf Centimeter, Gramm, Secunde und Celsiusgrad sind die Zahlenwerthe dieser vier Constanten die folgenden:

$$a = 0.4818 \cdot 10^{-10} [\text{sec} \times \text{Celsiusgrad}]$$

$$b = 6.885 \cdot 10^{-27} \left[\frac{\text{cm}^3 \text{gr}}{\text{sec}} \right] = \hbar$$

$$c = 3.00 \cdot 10^{10} \left[\frac{\text{cm}}{\text{sec}} \right]$$

$$f = 6.685 \cdot 10^{-8} \left[\frac{\text{cm}^3}{\text{gr} \cdot \text{sec}^2} \right]^1$$

Wählt man nun die »natürlichen Einheiten« so, dass in dem neuen Maasssystem jede der vorstehenden vier Constanten den Werth 1 annimmt, so erhält man als Einheit der Länge die Grösse:

$$\sqrt{\frac{bf}{c^3}} = 4.13 \cdot 10^{-33} \text{ cm},$$

als Einheit der Masse:

$$\sqrt{\frac{bc}{f}} = 5.56 \cdot 10^{-8} \text{ gr},$$

als Einheit der Zeit:

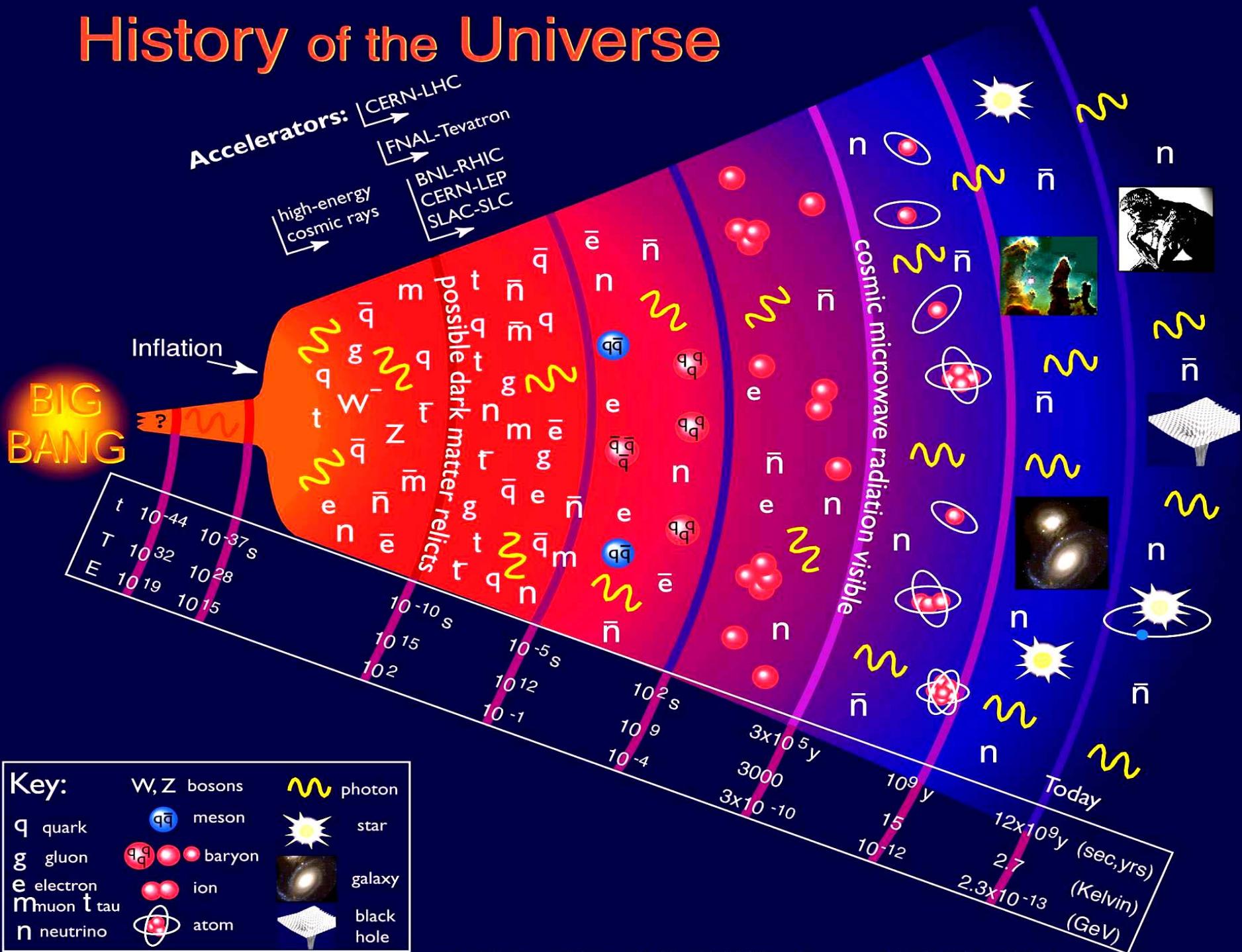
$$\sqrt{\frac{bf}{c^4}} = 1.38 \cdot 10^{-43} \text{ sec},$$

als Einheit der Temperatur:

$$a \sqrt{\frac{c^3}{bf}} = 3.50 \cdot 10^{30} \text{ Cels.}$$

Diese Grössen behalten ihre natürliche Bedeutung so lange bei, als die Gesetze der Gravitation, der Lichtfortpflanzung im Vacuum und die beiden Hauptsätze der Wärmetheorie in Gültigkeit bleiben, sie müssen also, von den verschiedensten Intelligenzen nach den verschiedenen Methoden gemessen, sich immer wieder als die nämlichen ergeben.

History of the Universe



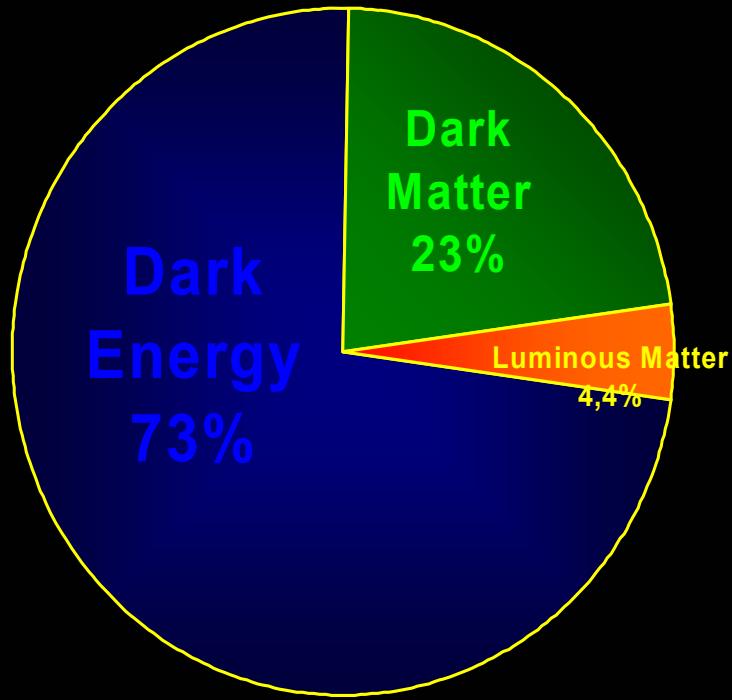
Big Bang Cosmology

	time/s	E/GeV	T/K	R/m
Planck time	10^{-44}	10^{19}	10^{32}	$10^{-5/-35}$
Grand Unification	10^{-37}	10^{16}	10^{29}	10^{-2}
electroweak unification	10^{-10}	10^2	10^{15}	10^{12}
quark confinement	10^{-5}	10^{-1}	10^{12}	10^{15}
ν decouple, e^+e^- annihilate	1	10^{-3}	10^{10}	10^{17}
light nuclei form	10^2	10^{-4}	10^9	10^{18}
from radiation to matter dom.	$10^{11} = 10 \text{ ky}$	10^{-8}	10^4	10^{22}
atoms form, γ decouple	$10^{13} = 379 \text{ ky}$	10^{-9}	3000	10^{23}
now	$10^{17} = 13 \text{ Gy}$	10^{-13}	2.73	10^{26}

$\rho \sim T^4$ radiation, $\sim T^3$ matter dominated universe

inflation !

Microcosm - Macrocosm



**Supernovae Type Ia
accelerated expansion:**

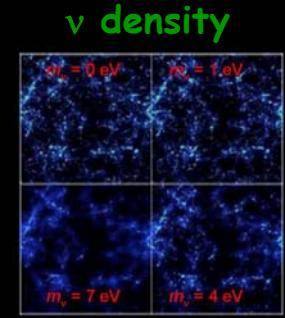
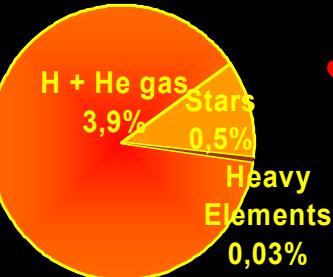
- Newton's law ?
- Einstein's cosmological constant Λ ?
- quintessence ?

HIGGS ?

Motion in galaxies + clusters, gravit. lenses:

- WIMPs = Weakly Interacting Massive Particles =

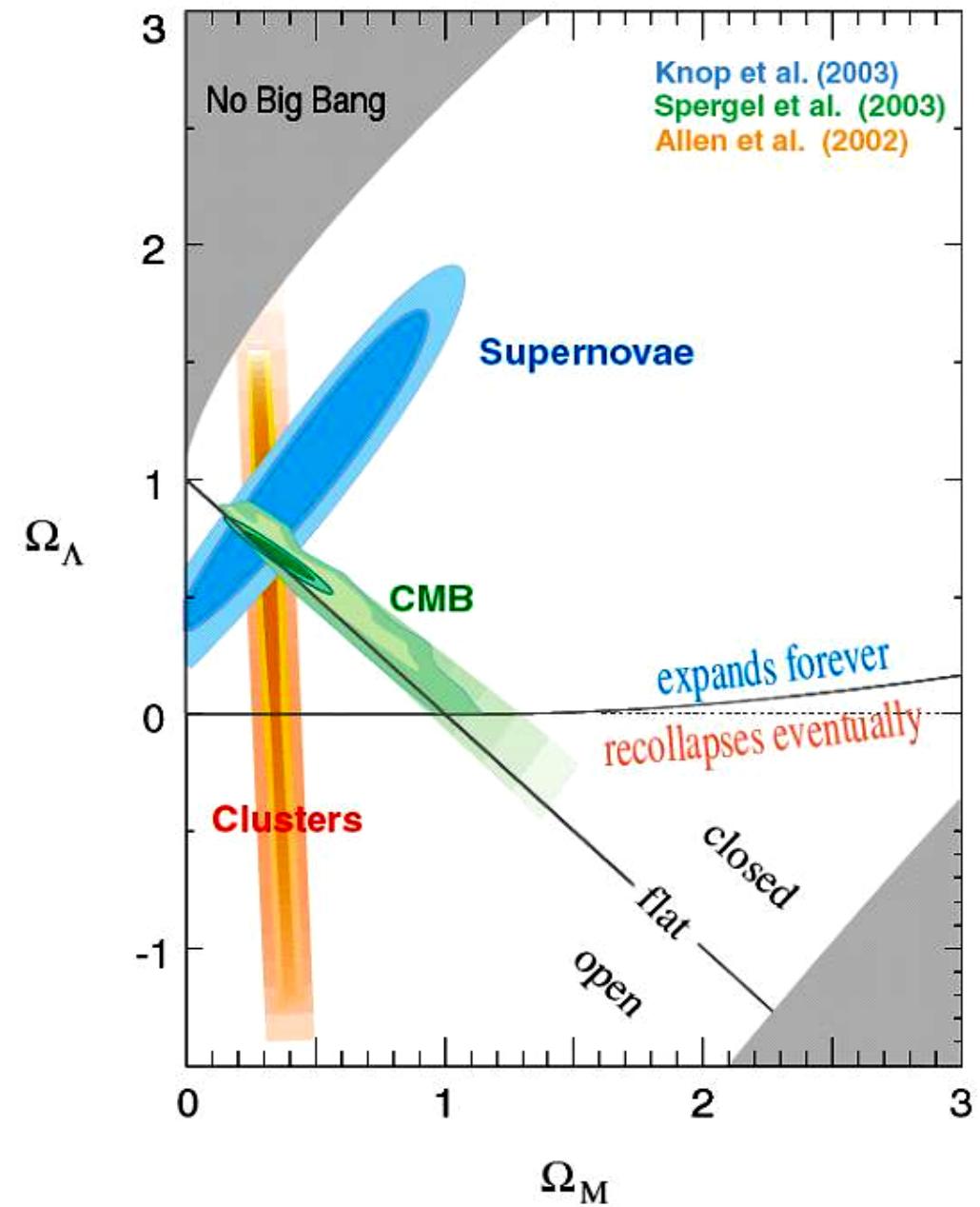
SUSY ?



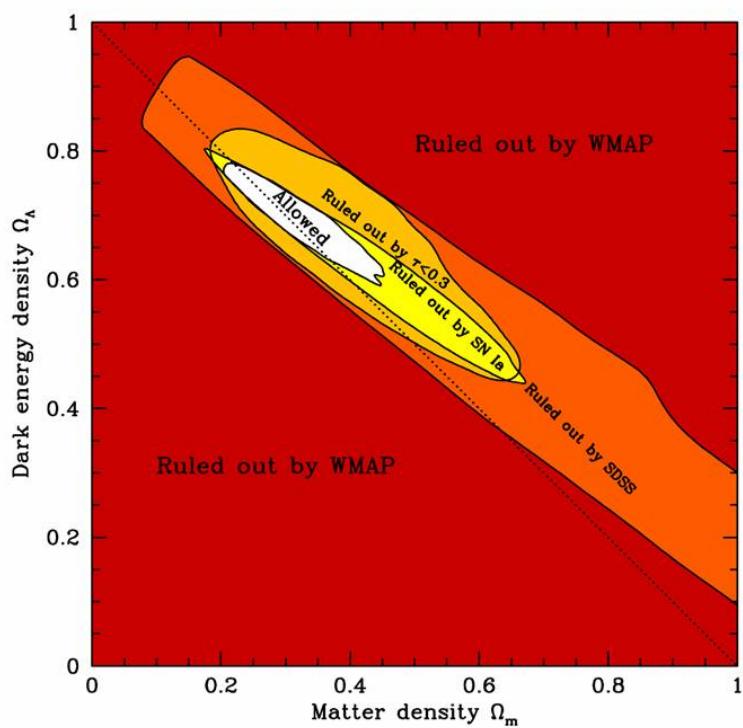
<5% dark matter !

Quantity	Symbol	Value	Comment
Total density	Ω_{tot}	1.02 ± 0.02	why = 1 ?
Dark energy density	Ω_Λ	0.73 ± 0.04	what is dark energy ?
Matter density	Ω_m	0.27 ± 0.04	
Baryon density	Ω_b	0.041 ± 0.002	
Dark matter fraction	$1 - \Omega_b / \Omega_m$	0.83 ± 0.01	what is dark matter ?
Baryon-to-photon ratio	η	$(6.1 \pm 0.3) \cdot 10^{-10}$	why >0 + so small ?
CMB temperature / K	T_{CMB}	2.725 ± 0.002	
CMB photon density / cm ⁻³	n_γ	410.4 ± 0.9	
Age at decoupling / kyr	t_{dec}	379 ± 8	causal horizon 2°
Age of the Universe / Gyr	t_0	13.7 ± 0.02	very precise !

Supernova Cosmology Project



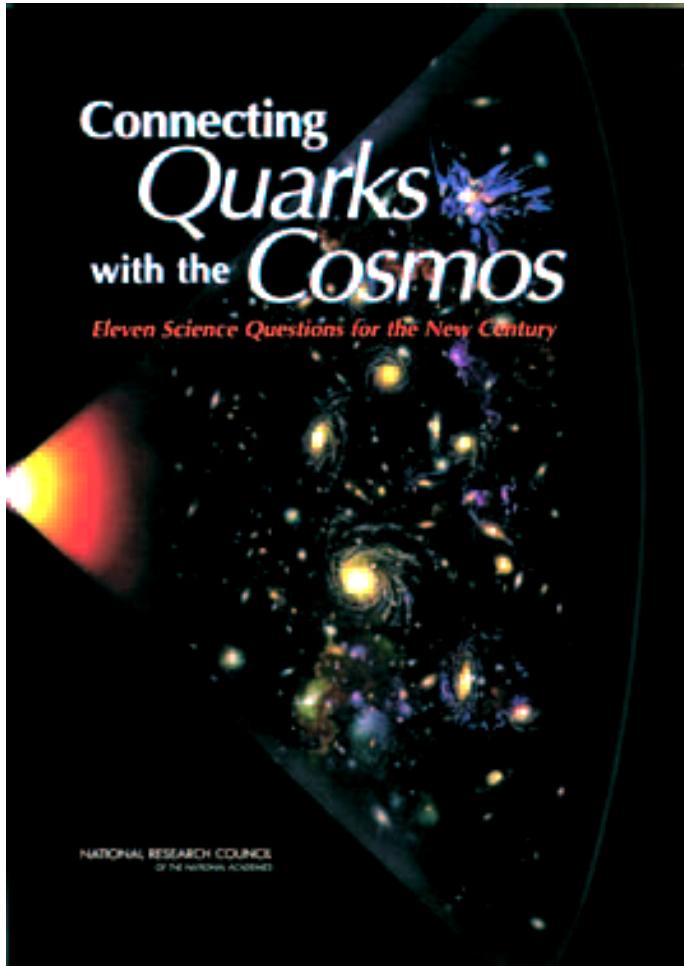
Dark Energy Dark Matter



Connecting Quarks with the Cosmos:

Eleven Science Questions for the New Century

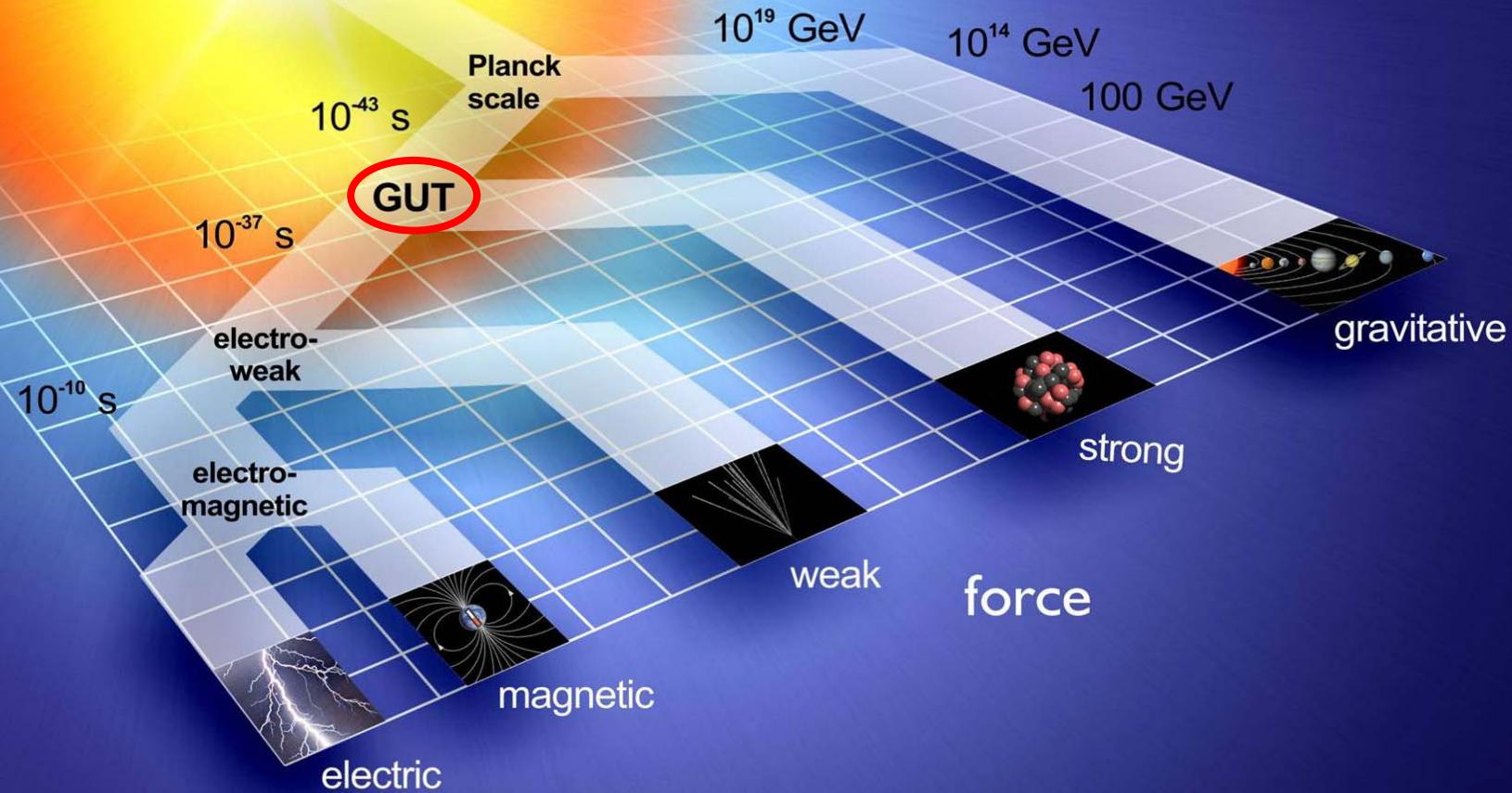
The US National Research Council, 2001.



1. What is **Dark Matter**?
2. What is the Nature of **Dark Energy**?
3. How Did the Universe Begin?
4. Did Einstein Have the Last Word on **Gravity**?
5. What are the **Masses of the Neutrinos** and How Have They Shaped the Evolution of the Universe?
6. How do Cosmic Accelerators Work and What are They Accelerating?
7. Are **Protons Unstable**?
8. What Are the New States of Matter at Exceedingly High Density and Temperature?
9. Are There **Additional Space-Time Dimensions**?
10. How Were the Elements from Iron to Uranium Made?
11. Is a New Theory of Light and Matter Needed at the Highest Energies?

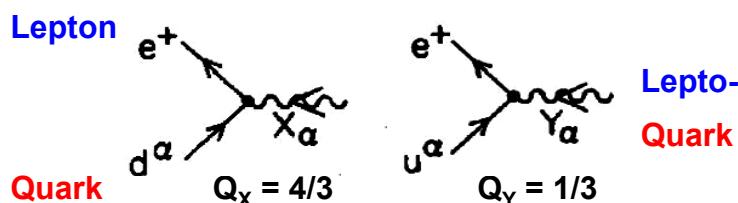
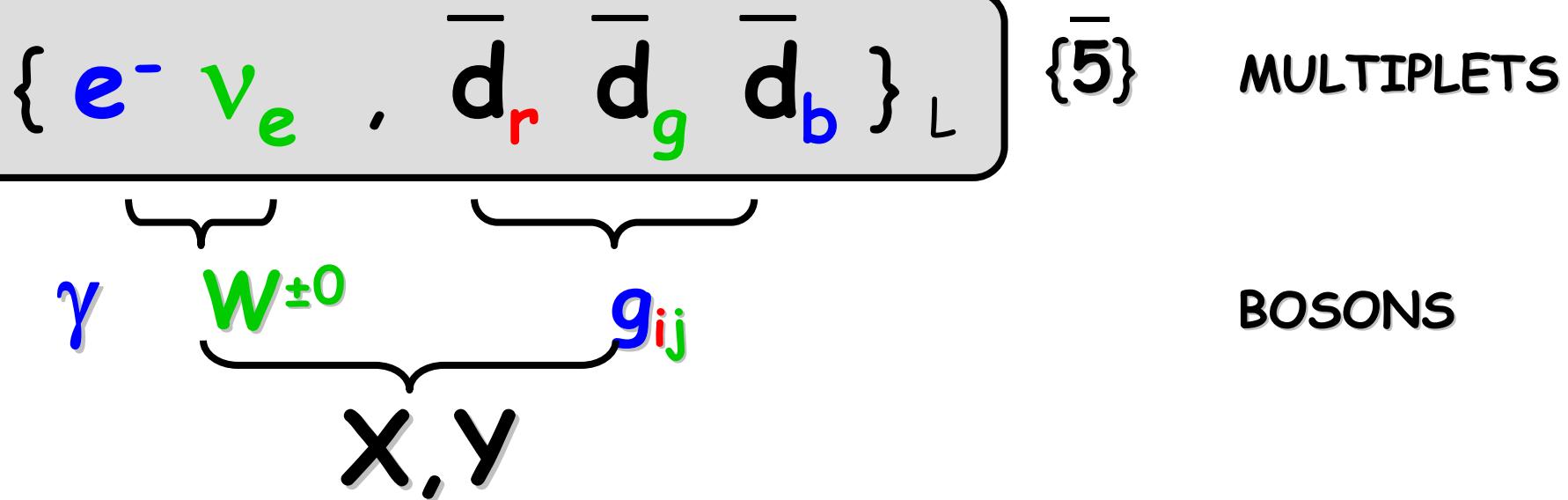
Unification of Forces

Big Bang



Grand Unification

elm. weak strong
 $U(1) \otimes SU(2) \otimes SU(3)$ $\subset SU(5)$ SYMMETRIES



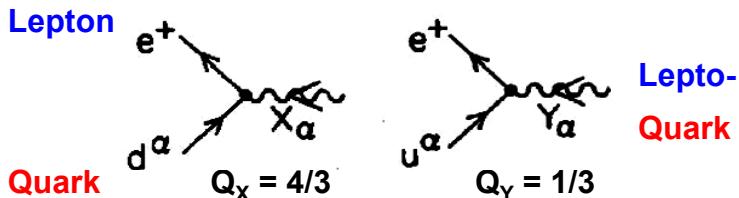
HERA: $e p \rightarrow LQ$ $m > 300 \text{ GeV}$
proton decay: $p \rightarrow e \pi$ $\tau > 5 \cdot 10^{33} \text{ a}$

Grand Unification

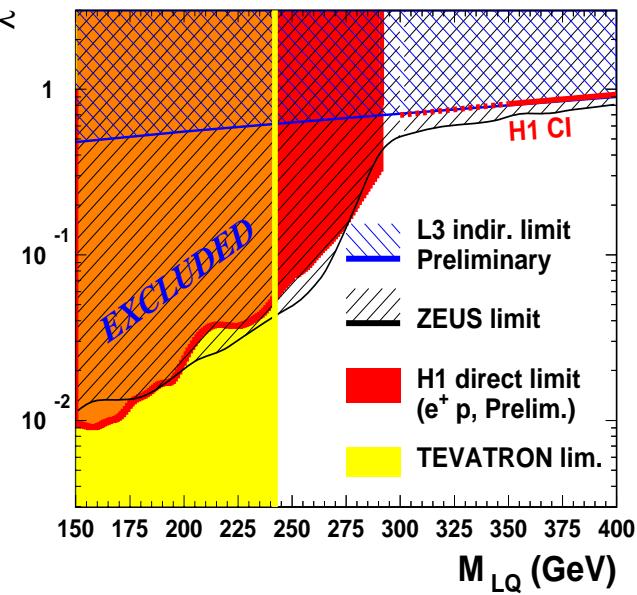
$SU(N)$	$N^2 - 1$	gauge bosons
$SU(5)$	$25 - 1 = 24$	
$SU(3)$	$9 - 1 = -8$	gluons
$SU(2)$	$4 - 1 = -3$	Z, W
$U(1)$	$1 = -1$	photon
	$6 (X+Y) =$	$12 X, Y$ bosons



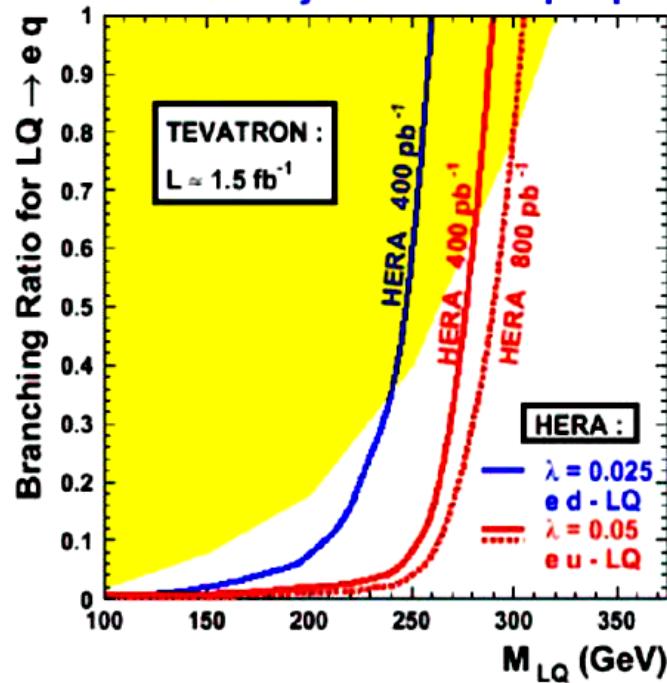
Leptoquarks @ HERA



SCALAR LEPTOQUARKS WITH $F=0$ ($\tilde{S}_{1/2,L}$)



Future Sensitivity on Scalar Leptoquarks



LQ (RPV q)	H1 M_{LQ}	LEP GeV/g
S_0L	710	2150
S_0R	640	1700
\check{S}_0R	330	660
$S_{1/2}L$	850	590
$S_{1/2}R$	370	770
$\check{S}_{1/2}L$	430	-
S_1L	490	1190
V_0L	730	3030
V_0R	580	540
$V_{\sim 0}R$	990	1610
$V_{1/2}L$	420	1000
$V_{1/2}R$	950	750
$V_{\sim 1/2}L$	1020	580
V_1L	1360	2170

HERA competitive with LEP and Tevatron for small branching ratios

Charge quantization + $\sin^2 \theta_W$

- $U(1) \times SU(2)$:

Q not conserved, def. $Q = I_3^W + Y^W/2$

$$\text{Tr}_2(I_3^W) = 0$$

- $SU(5)$:

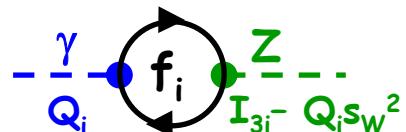
$$\text{Tr}_5(Q) = N_C Q_d - Q_e = 0 \Rightarrow$$

$$Q_d = e/N_C \quad \text{or} \quad Q_p = -Q_e$$

'explain' fractional quark charges from nr. of colors !

- prediction of electroweak mixing angle:

γ, Z orthogonal \Rightarrow no coupling, but:



$$\sum_i Q_i (I_{3i} - Q_i s_W^2) = 0$$

$$s_W^2 = \sum_i Q_i I_{3i} / \sum_i Q_i^2 = "SU(2)/U(1)" = 3/8 \quad @ \text{true unification point } m_{GUT}$$

	I_3	Q
ν_e	+1/2	0
e^-	-1/2	-1
\bar{d}_r	0	1/3
\bar{d}_g	0	1/3
\bar{d}_b	0	1/3

$$\bar{d}_L \dots \{1\}_{SU(2)}$$

Grand Unification

- quark-lepton symmetry:

quarks + leptons in one multiplet

- quantization of electric charge:

$$N_c Q_q - Q_e = 0 = 3 \times 1/3 - 1$$

$$\text{or } Q_p = Q_e$$

- prediction for electro-weak mixing angle:

- $\sin^2 \theta_W (M_X) = g^2 / (g^2 + g'^2) = 3/5 / (1+3/5) = 3/8$
- $\sin^2 \theta_W (M_Z) = 0.20 \quad \text{GUT} \quad (M_X \rightarrow M_Z)$
- $\sin^2 \theta_W (M_Z) = 0.22 \quad \text{expt.}$

- lepton number violation:

neutrino masses + oscillations

- baryon number violation: proton decay: $\tau_p \sim M_X^{-4} / g^4 m_p^5$

baryon asymmetry of universe: $N_p / N_\gamma = 10^{-10}$

- SU(5) GUT: $M_X \sim 10^{15} \text{ GeV}$ $\tau \sim 10^{29 \pm 2} \text{ a}$

Super-K 1998: $\tau(p \rightarrow e^+ \pi^0) > 5 \cdot 10^{33} \text{ a}$ $\tau(p \rightarrow K^+ \nu) > 2 \cdot 10^{33} \text{ a}$

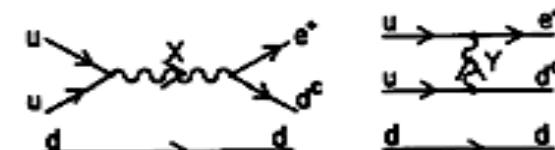
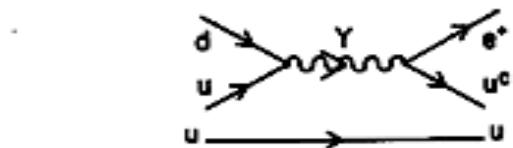
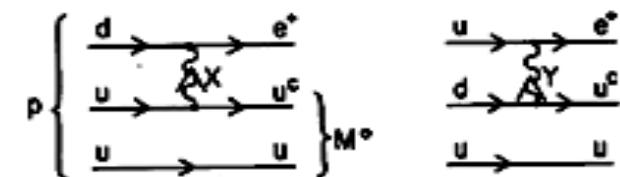
- SUSY GUT: $M_X \sim 10^{16} \text{ GeV} :$

$\tau(p \rightarrow e^+ \pi^0) \sim 10^{35 \pm 1} \text{ a}$ $\tau(p \rightarrow K^+ \nu) \sim 10^{32 \pm 3} \text{ a}$

Hyper-K 201x+10: $\tau(p \rightarrow e^+ \pi^0) > 10^{35} \text{ a}$ $\tau(p \rightarrow K^+ \nu) > 3 \cdot 10^{34} \text{ a}$

- GUT magnetic monopoles

$$\{e^- \nu_e, \bar{d}_r, \bar{d}_g, \bar{d}_b\}_L$$



$$SU(2) \uparrow \begin{pmatrix} \nu_L^0 & d_L^{0c} \\ e_L^{-0} & \end{pmatrix} \quad \begin{pmatrix} e_L^{+0} & u_L^0 & u_L^{0c} \\ d_L^0 & \end{pmatrix} \uparrow SU(2)$$

5*

10

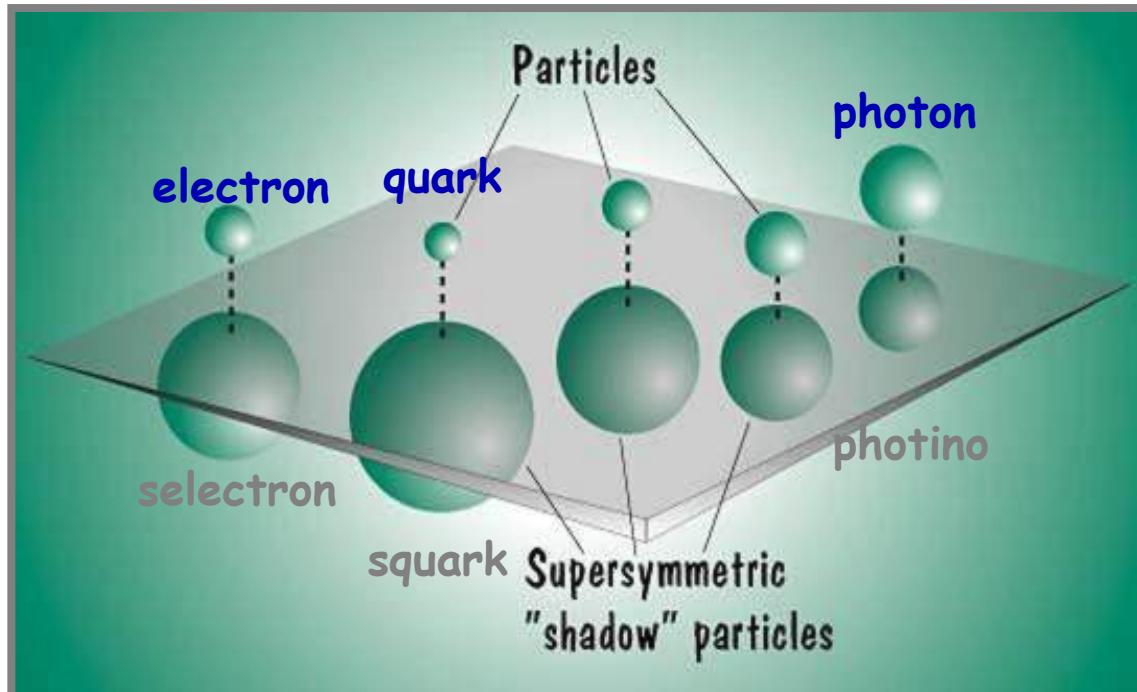
Super-Symmetry

Fermion

Boson

Boson

Fermion



AntiSymmetry - SuperSymmetry

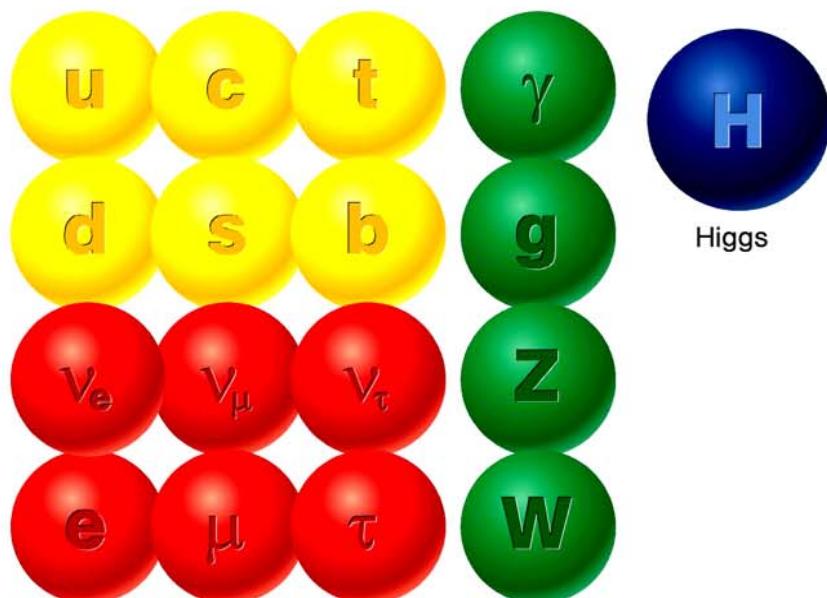
mirror world

unifies

bosons with fermions
force with matter

Super-Symmetry

Standard-Teilchen

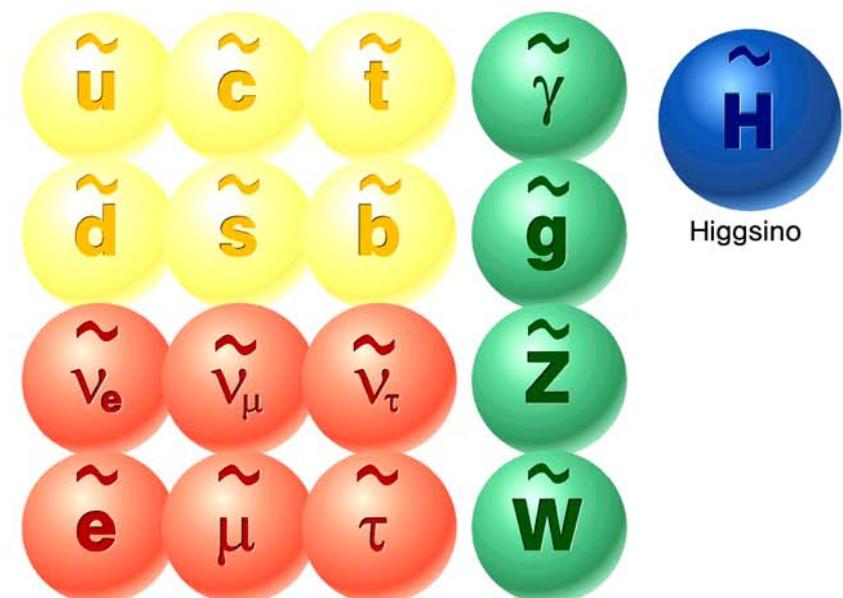


Quarks

Leptonen

Kraftteilchen

SUSY-Teilchen



Squarks

Sleptonen

SUSY-Kraftteilchen

Super-Symmetry

Neveu, Schwarz, Ramond 1971
Wess, Zumino 1974
Fayet, Farrar 1976

SUSY transformation:

turn boson to fermion state + vice versa:

$$Q |Boson\rangle = |Fermion\rangle \quad Q |Fermion\rangle = |Boson\rangle$$

Q ... anticommuting Weyl spinor (spin $\frac{1}{2}$) operator

P^μ ... 4-momentum generator of space-time translations
vector (spin 1) under Lorentz trafo.

$$\{Q, Q^\dagger\} = 2 \sigma^\mu P^\mu \quad \sigma^\mu \dots \text{Pauli matrices}$$

$$\{Q, Q\} = 0$$

$$\{P^\mu, Q\} = 0$$

SUSY: connects space-time with statistics properties

local SUSY needs metric tensor: $g_{\mu\nu}(x) P^\mu$

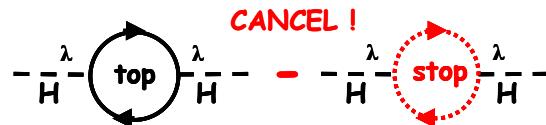
$J=2$ graviton , $J=3/2$ gravitino: gauge fields of local SUSY

general relativity from **local SUSY** =>

SUPERGRAVITATION => **Superstrings** !

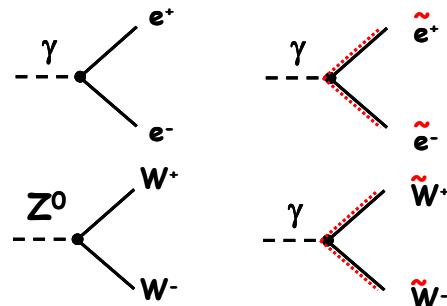
Super-Symmetry

- Higgs mass terms divergent:
- cancel divergences by
a **mirror world**:

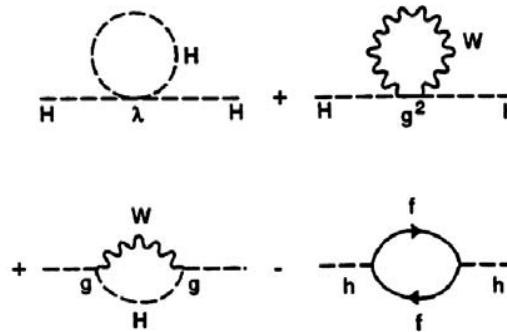


n (bosons) = n (fermions)

- SUSY adds new decay channels:



changes coeff. b_N of renorm. group equ.
changes evolution of coupling constants:
evolution of weak coupling changes sign!
changes Grand Unification scale!



particle	spin	sparticle	spin
bosons		fermions	
photon	γ	photino	$\tilde{\gamma}$
gluon	g	gluino	\tilde{g}
W, Z boson	W, Z	wino, zino	\tilde{W}, \tilde{Z}
graviton	G	gravitino	\tilde{G}
higgs	H	higgsino	\tilde{H}
		neutralino	$\tilde{\gamma}, \tilde{Z}, \tilde{H}$
fermions	1/2	bosons	
lepton	l	slepton	\tilde{l}
quark	q	squark	\tilde{q}

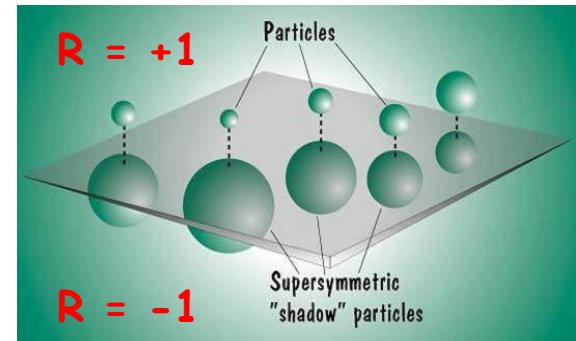
R parity

conserved multiplicative quantum number:

$$R = (-1)^{3B+L+2S}$$

B,L: baryon+lepton nr, S: spin

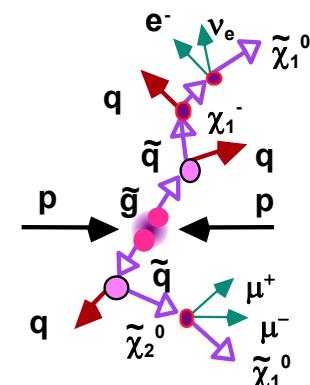
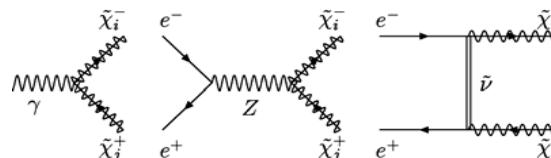
particles: $R=+1$ sparticles: $R=-1$



- sparticles produced in pairs: $pp \rightarrow \tilde{q}\tilde{q} X$, $e^+e^- \rightarrow \tilde{\mu}^+\tilde{\mu}^-$
- heavier sparticles decay to lighter ones: $\tilde{q} \rightarrow q \tilde{g}$, $\tilde{\mu} \rightarrow \mu \tilde{\gamma}$
- the lightest particle is stable - it has to conserve R parity !
not seen in universe => no electric + strong charge
- **neutralino** $\tilde{\chi} = (\tilde{\gamma}, \tilde{Z}, \tilde{H}_{1,2})$!?

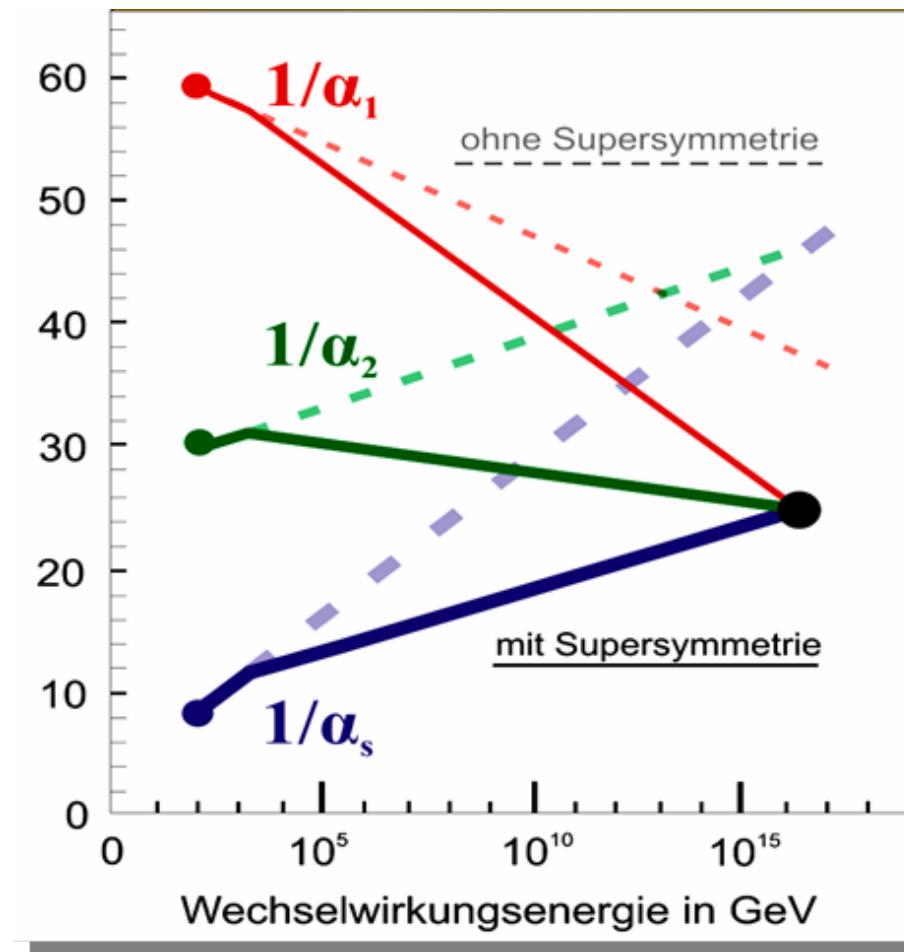
Experiment:

- 2 neutralinos end all SUSY decay chains
- interact only weakly + escape => large missing E_T = SUSY signature !



SUSY + Grand Unification

elm.
weak
strong



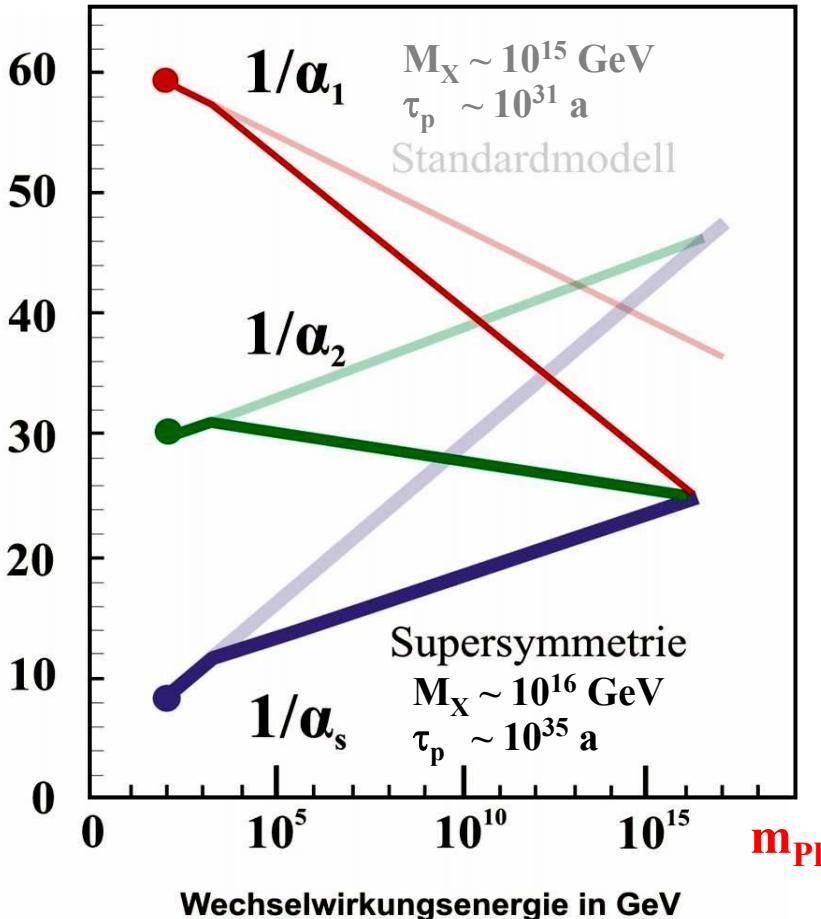
$U(1)$
 \otimes
 $SU(2)$
 \otimes
 $SU(3)$
 \subset
 $SU(5),$
 $SO(10)$

unified at the Planck scale ?

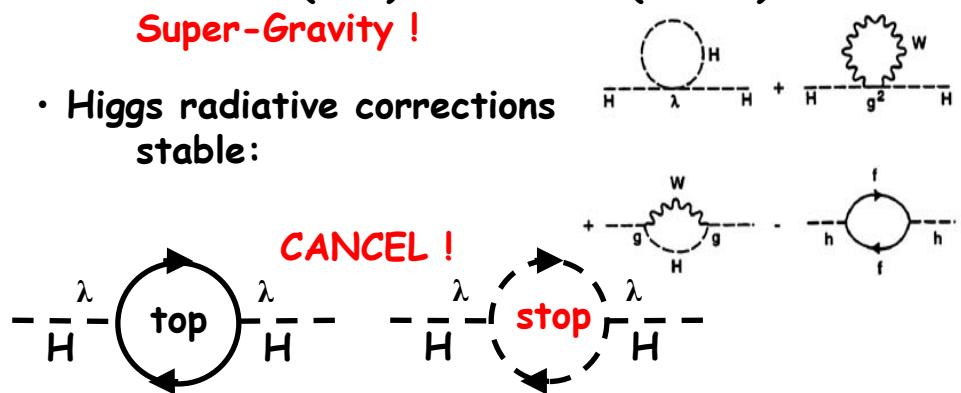
...

Super-Symmetry

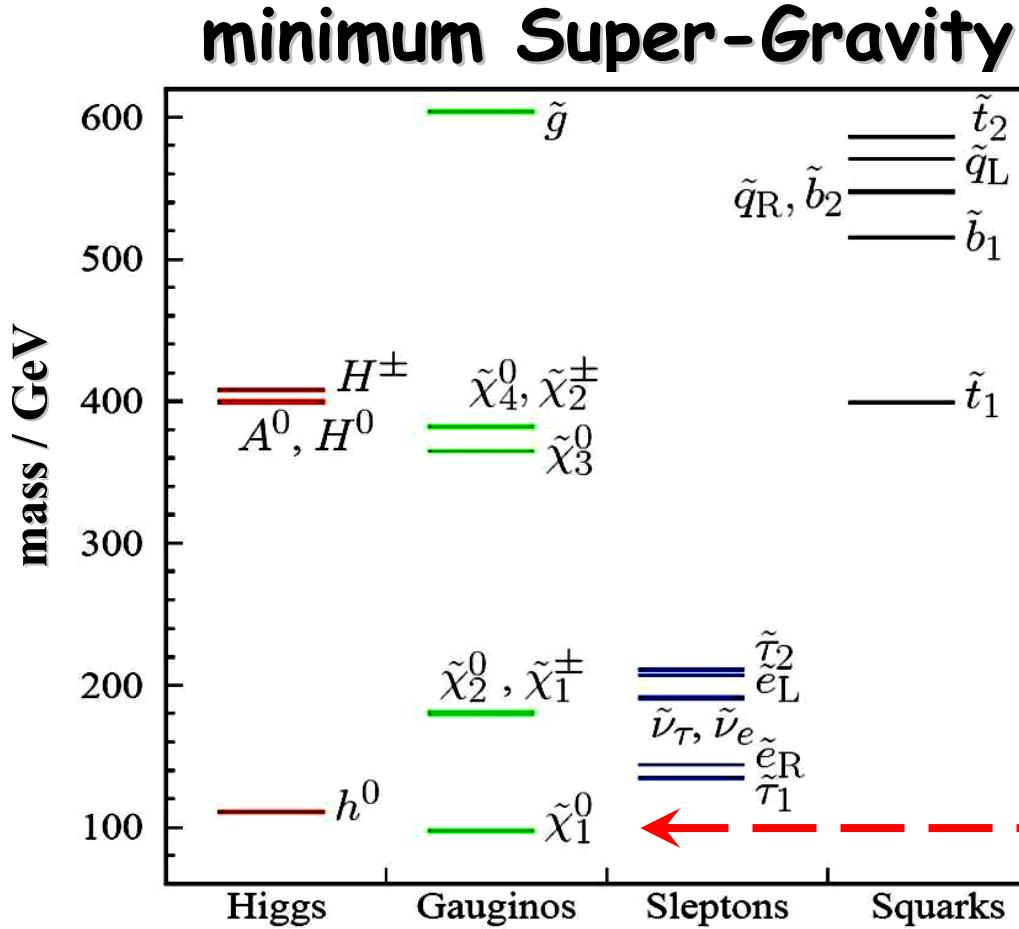
unifies forces + couplings



- changes energy dependence of couplings:
- one unification point at $M_X = 2 \cdot 10^{16} \text{ GeV}$!
- proton lifetime > exptl. limit
- Neutralino ($\tilde{\gamma}, \tilde{Z}, \tilde{H}_{1,2}$): lightest SUSY particle
Dark Matter in the Universe !
- connects continuous space-time symmetry with spin-statistics + symmetries
 $U(1) \otimes SU(2) \otimes SU(3) \supset SU(5), SO(10), E(6), \dots$
- gauge fields of local SUSY:
Graviton ($J=2$) + Gravitino ($J=3/2$)
Super-Gravity !
- Higgs radiative corrections stable:



SUSY mass spectrum

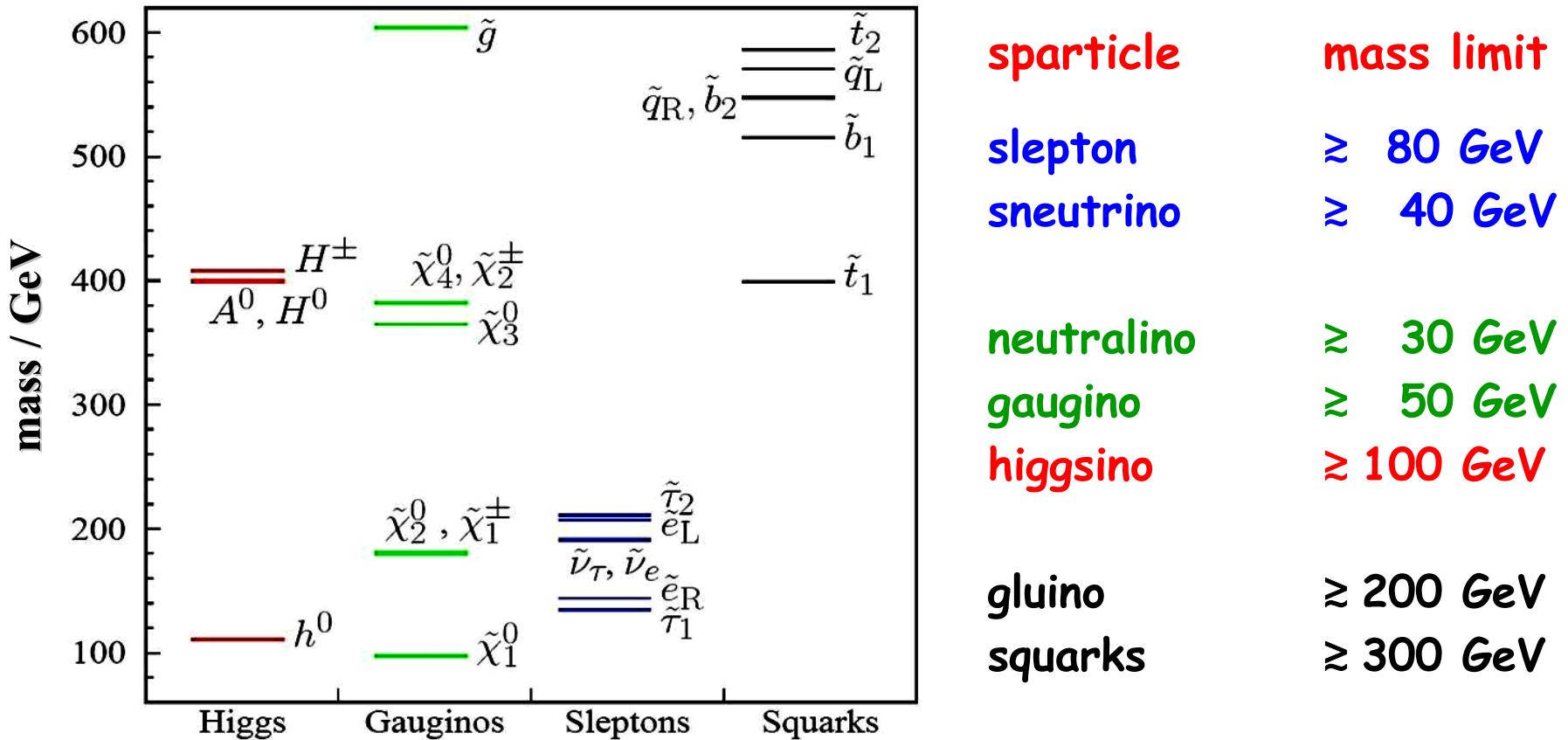


**SUSY
WIMP:**
dark matter candidate !

lightest neutral SUSY
particles mix -
Neutralino
=

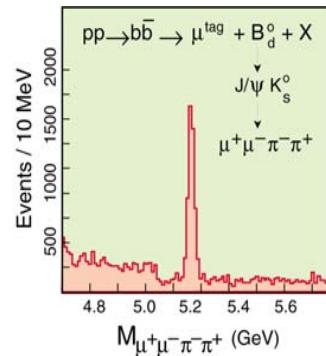
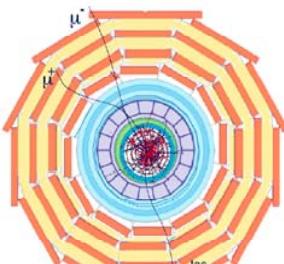
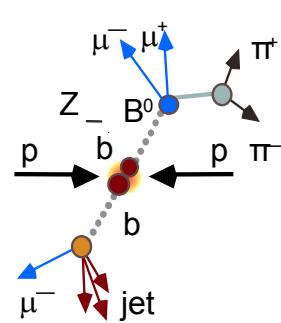
Higgsino+Photino+Zino

SUSY mass limits

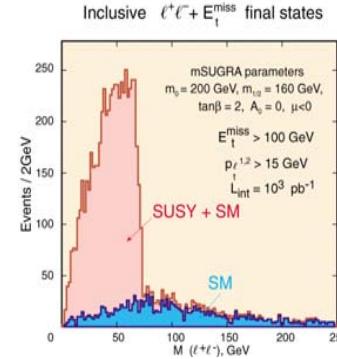
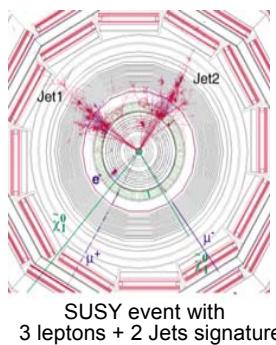
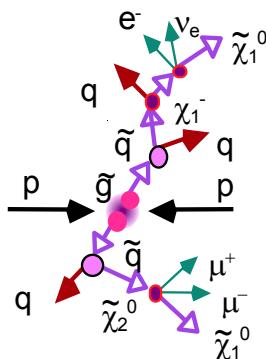
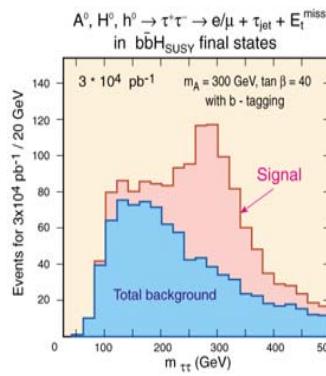
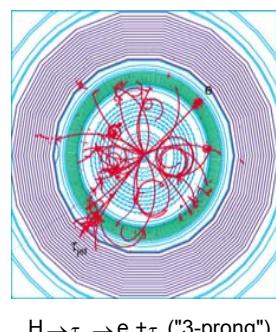
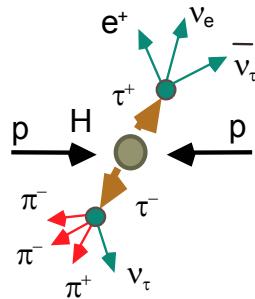


lowest limit + prediction for neutralino = WIMP/dark matter candidate

LHC physics: Supersymmetry



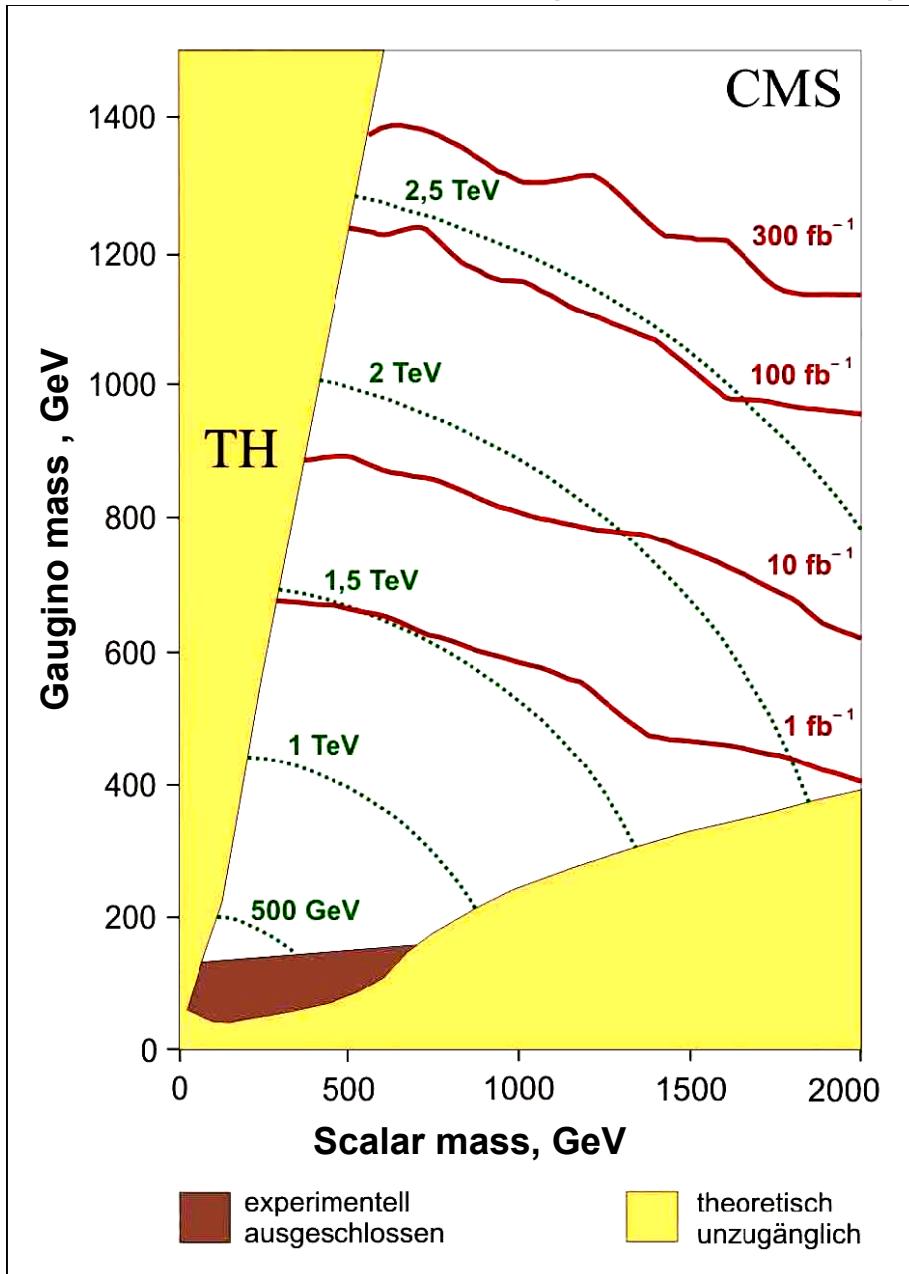
The decay B^0 or $B^0 \rightarrow J/\psi K_s^0$ presents a very clean experimental signature. The particle content (B^0 or B^0 meson) that gave the decay can be determined from a muon from the second b-flavored hadron in the event. An asymmetry in the two rates (B^0 vs B^0) would signal CP violation. This would be the first time that CP violation is observed outside the neutral kaon system



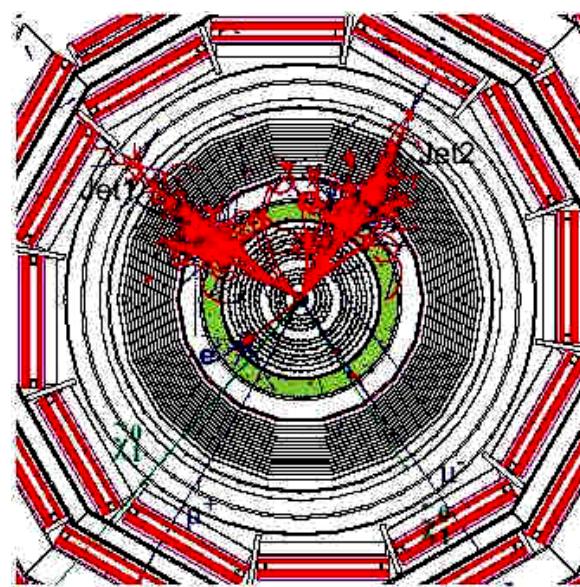
SUSY Higgs bosons. In the MSSM there are 5 Higgs bosons: h^0 , H^0 , A^0 and H^\pm decaying through a variety of decay modes to γ , e^\pm , μ^\pm , τ^\pm and jets in final states. Above: an example of a SUSY Higgs decay to $\tau\tau$ in CMS. On the right is the reconstructed $\tau\tau$ mass spectrum

Sparticles. Production of sparticles may reveal itself through some spectacular kinematical spectra, with a pronounced "edge" in the l^+, l^- mass spectrum reflecting $\chi_2^0 \rightarrow l^+, l^- \chi_1^0$ production and decay. An example of such a spectrum in inclusive $l^+, l^- + E_t^{\text{miss}}$ and of a 3 l^\pm production event are shown below

LHC: minimum Super-Gravity



SUSY @ LHC



SUSY event with 3 leptons + 2 jets

production:

- squarks+gluino strongly prod.

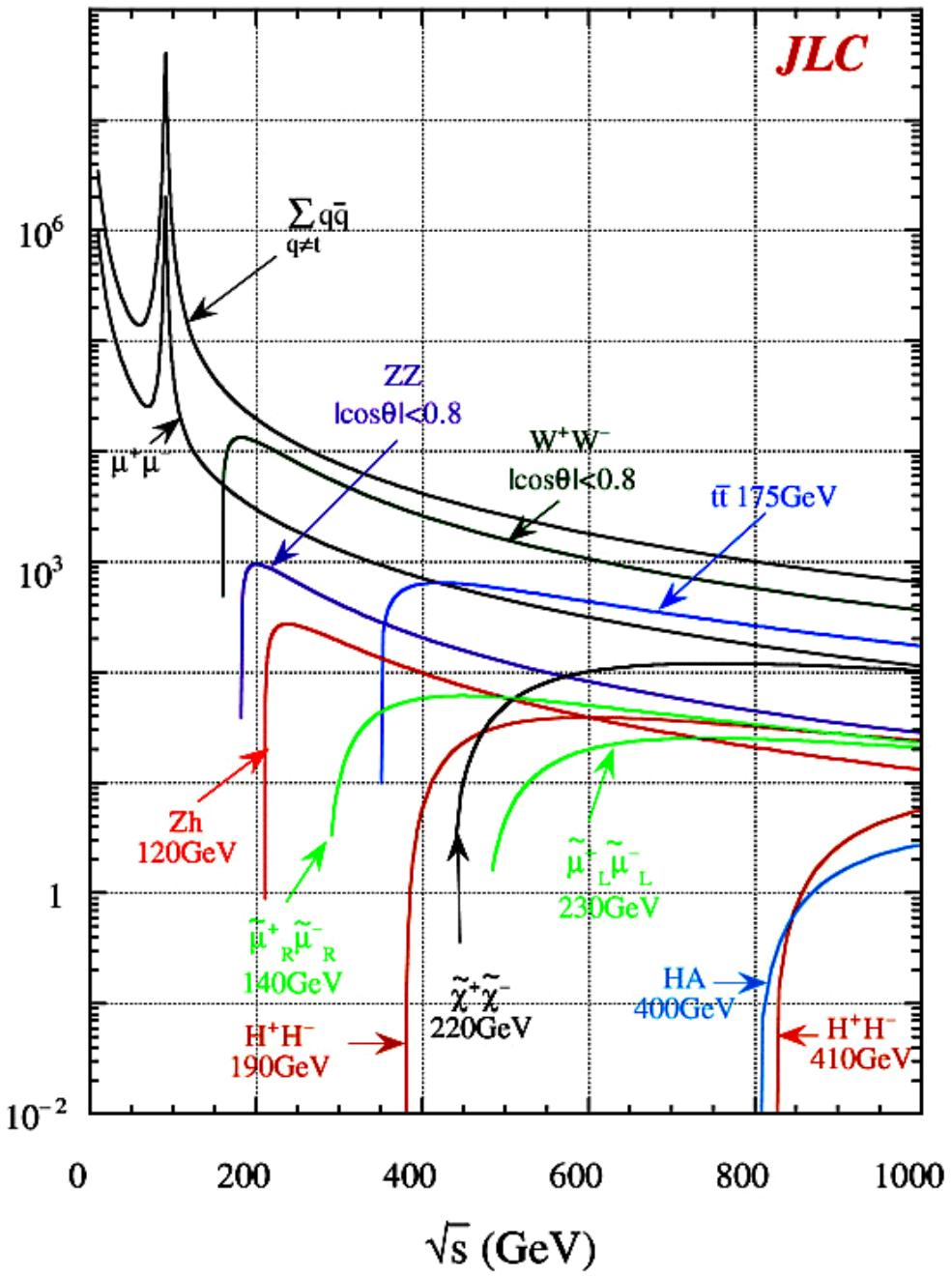
decays:

- missing E_T (LSP)
- jets (squarks/gluino)
- leptons (sleptons)

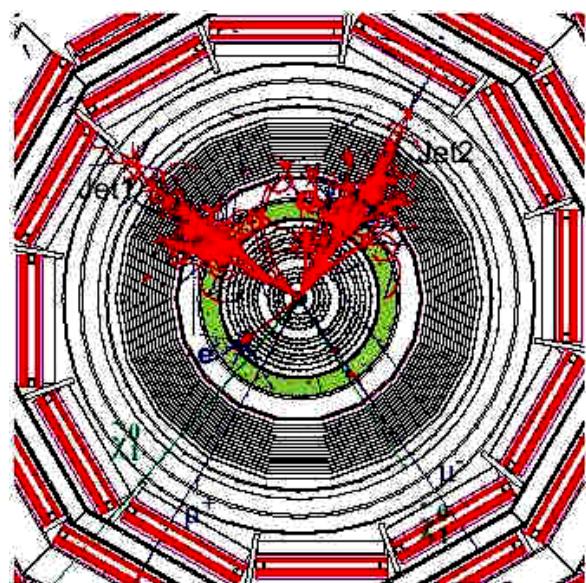
spin:

- J (neutralino) = $1/2$?
- J (squark,slepton) = 0 ?

couplings: universal ?



SUSY @ ILC



SUSY event
with 3 leptons + 2 jets

large missing E_T
due to neutralinos

SUSY: rare decays

$\mu^+ \rightarrow e^+ \gamma$

current limit: $< 1.2 \cdot 10^{-11}$

MEG @ PSI: $< 10^{-14}$

$\mu^+ \rightarrow e^+ e^+ e^-$

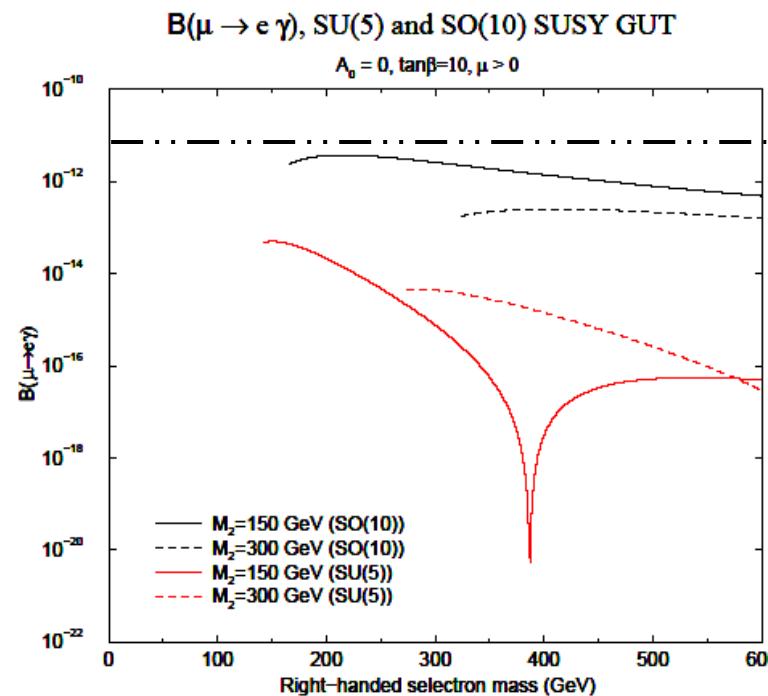
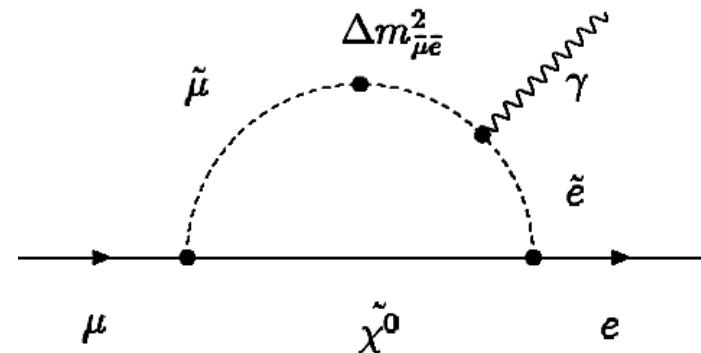
current limit: $< 10^{-12}$

$\mu^- A \rightarrow e^- A$

current limit: $< 6 \cdot 10^{-13}$

MECO: $< 10^{-16}$

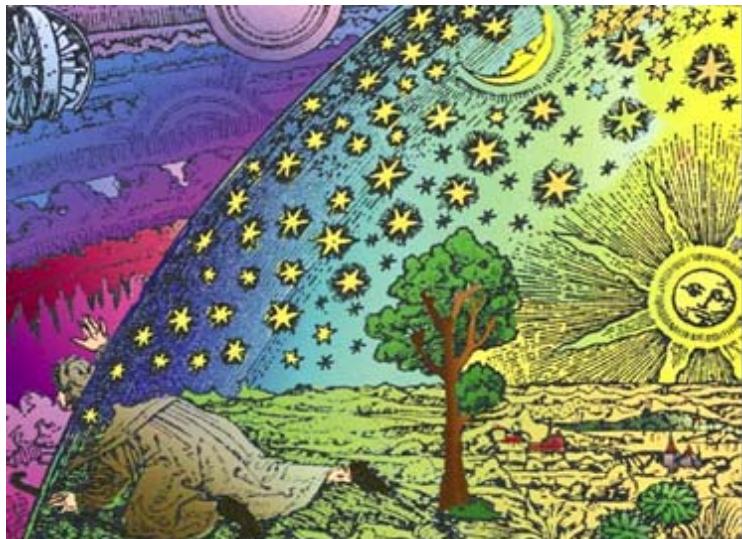
PRISM@JPARC: $< 10^{-19}$



Superstrings

Is the World

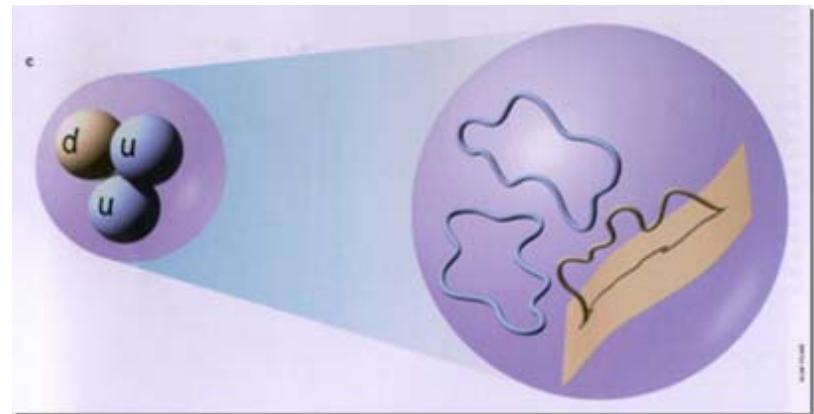
2-dimensional ...



3-dimensional ...

10-dimensional ?

Only our **4** dimensions
expanded after the Big Bang.
The other **6** stayed **compact**.

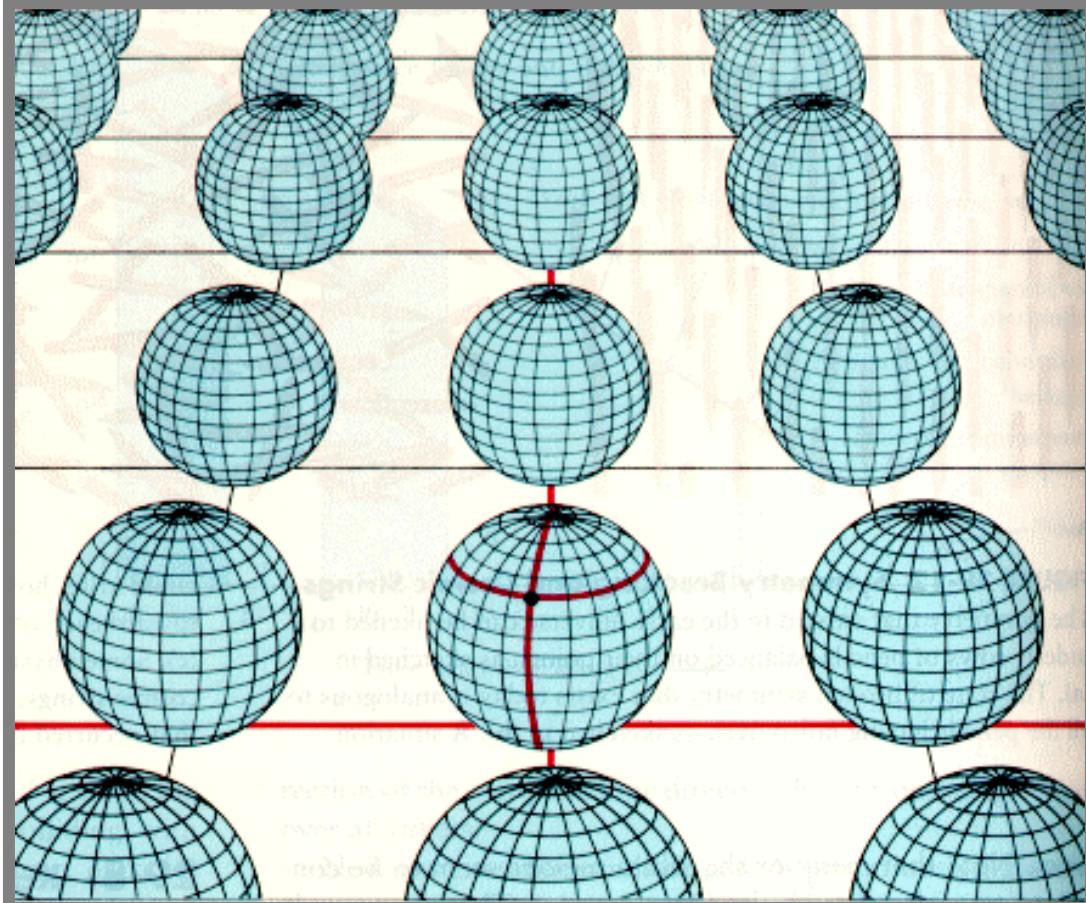


Elementary particles =
string excitations ?

string radius ?

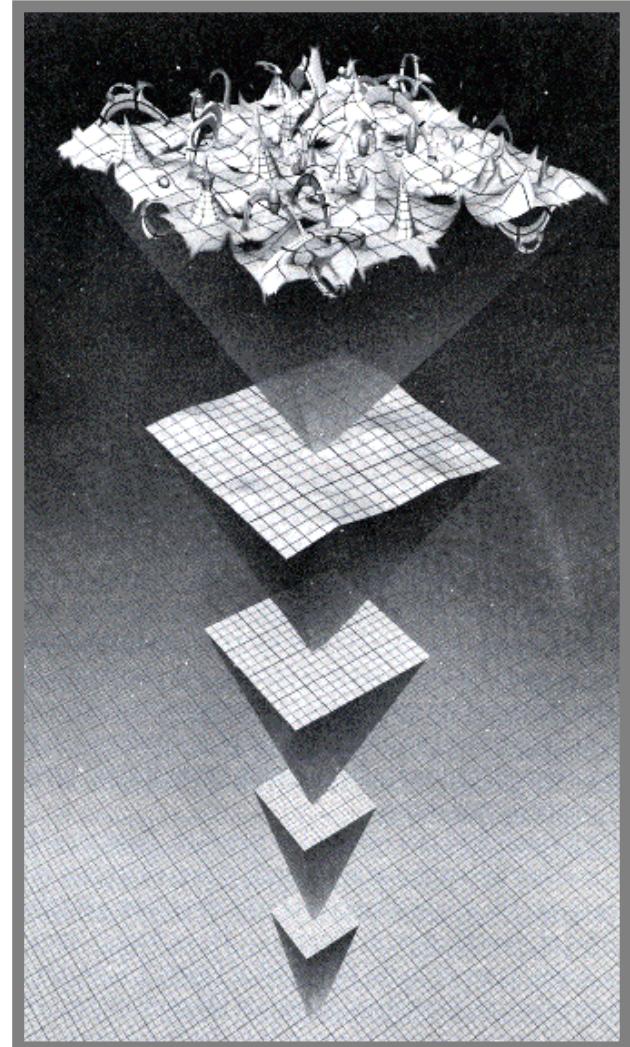
Planck length = $(\hbar G_N/c^3)^{1/2}$
= 10^{-35} m ?

Superstrings

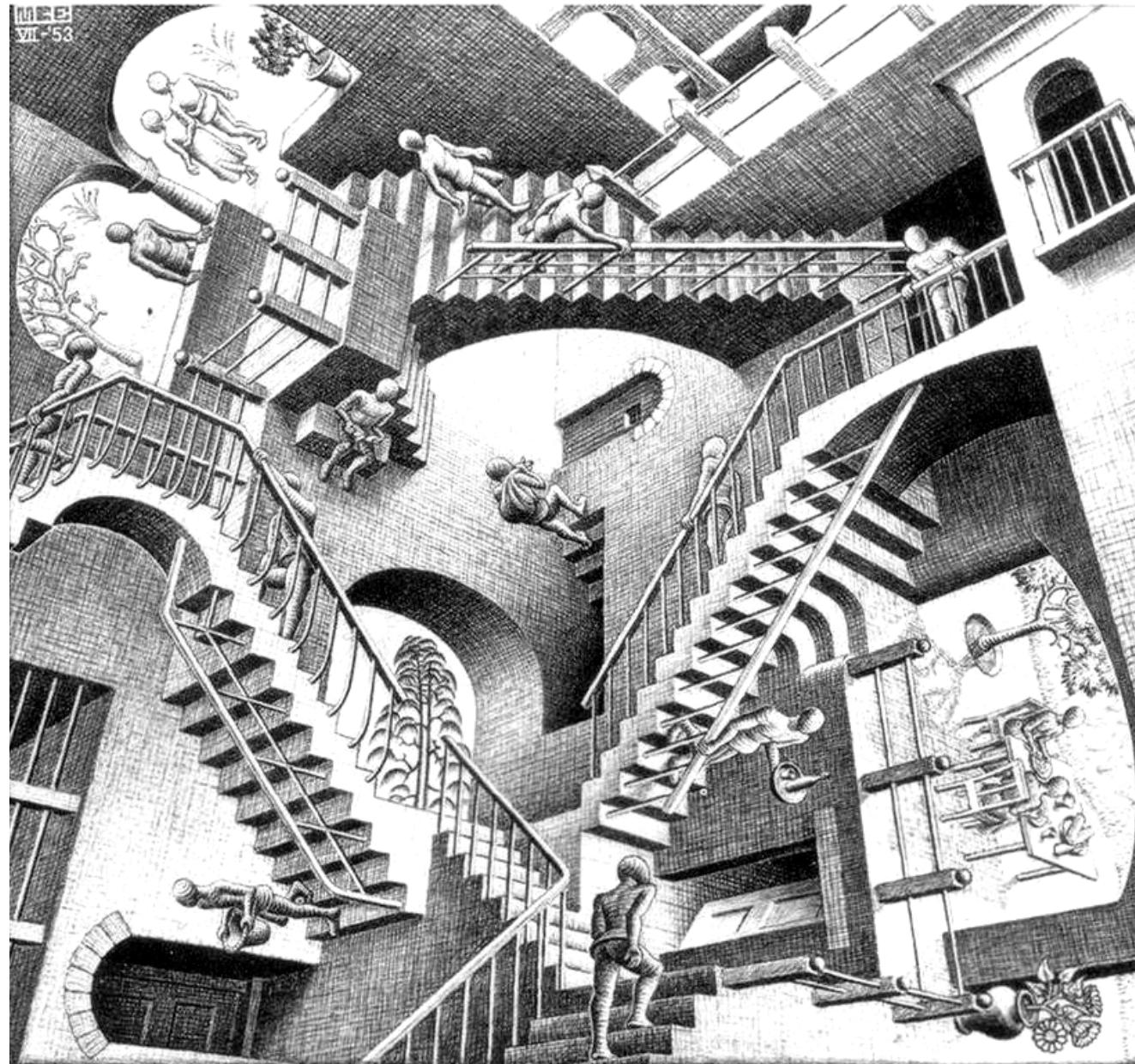


curved extra dimensions
at Planck length = $(\hbar G_N/c^3)^{1/2} = 10^{-35}$ m ?

foamy space-time
at Planck length ?



Extra Dimensions

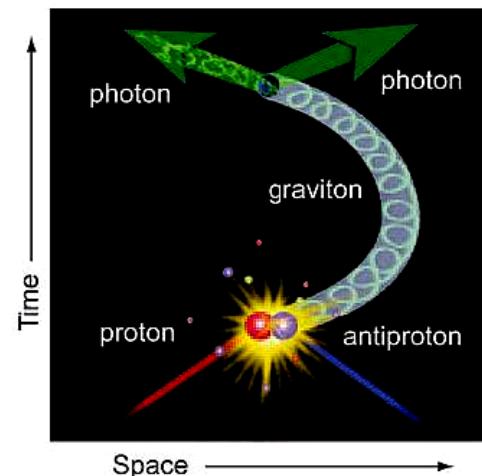


Hierarchy problem:

$$M_{\text{ew}} / M_{\text{Planck}} \sim 10^{-17}$$

lower gravity scale
from $M_{\text{Planck}} \sim 10^{19} \text{ GeV}$
to $M_D \sim 1 \text{ TeV}$

possible if gravity
propagates in
4+n dimensions.



Extra Dimensions

Brane Theories (Arkani-Hamed, Dimopoulos, Dvali, 1998,99)

- elm., strong + weak force act on (3+1) brane
- N extra dimensions compactified with radius R
- only gravitation acts in extra dimensions
 - modification of Newton's law:

$$F \sim G_N / r^2 \sim 1 / M_{\text{Pl}}^2 r^2 \sim 1 / M_{\text{KK}}^{2+N} r^{2+N}$$

- G_N , M_{Pl} ... only effective constants
- R , M_{KK} ... new, fundamental Planck length, mass

$$M_{\text{Pl}}^2 \sim G_N^{-1} \sim (10^{19} \text{ GeV})^2 \sim R^N M_{\text{KK}}^{N+2}$$

- Grand Unification = gravity becomes strong already at electroweak scale : $M_{\text{KK}} = 1 \text{ TeV}$:

$$N=1 \quad R \sim 10^9 \text{ km}$$

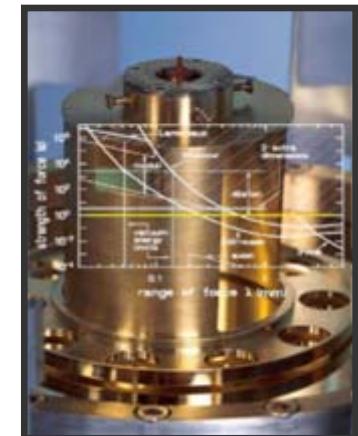
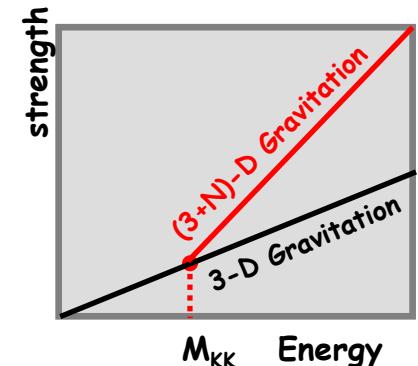
$$N=2 \quad R \sim 1 \text{ mm}$$

Eöt-Wash: $1/r$ ok to $50 \mu\text{m}$

...

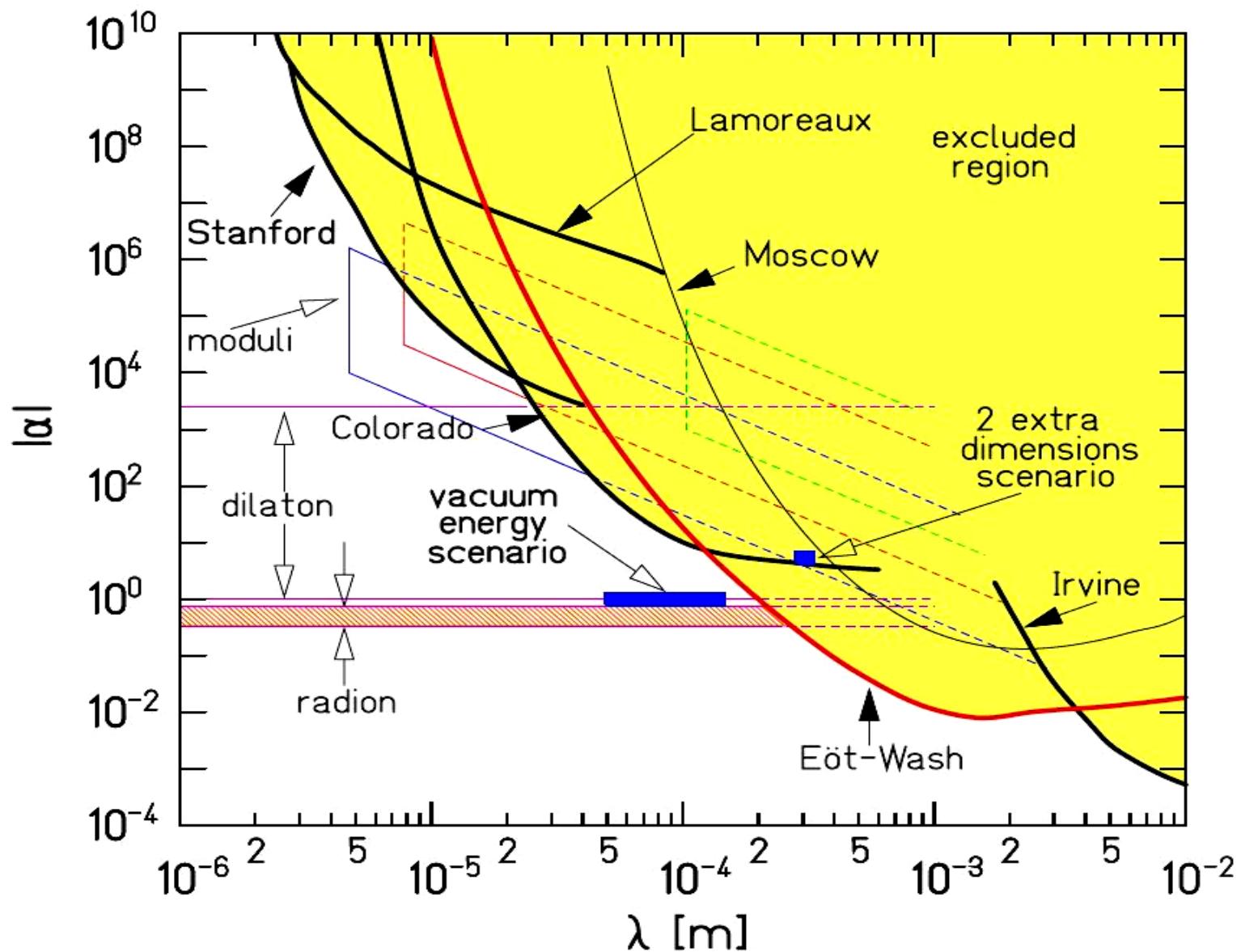
$$N=6,7 \quad R \sim 10 \text{ fm}$$

new Planck length



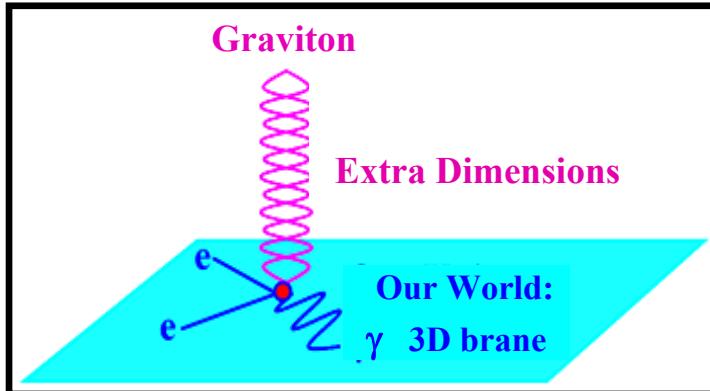
Torsion pendulum
Eöt-Wash Group,
Univ. Washington, Seattle

Extra Dimensions

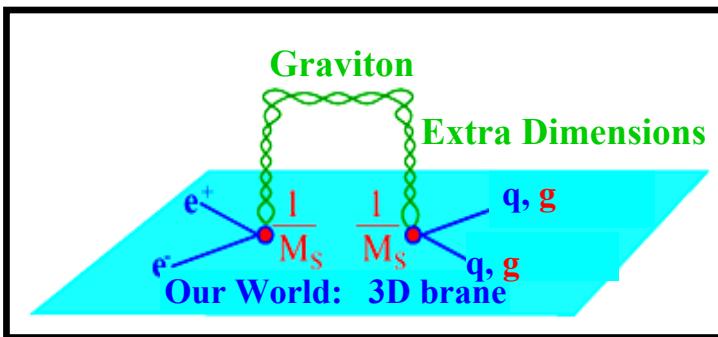


Extra Dimensions

Emission of Kaluza-Klein gravitons in Extra Dimensions



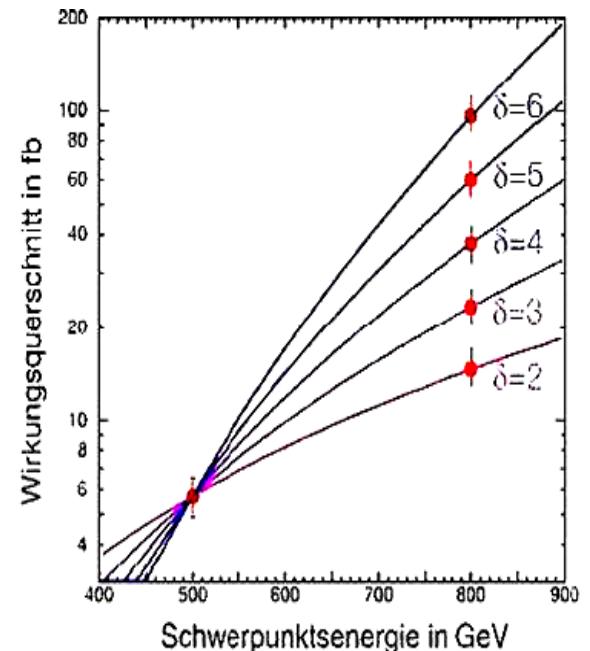
real emission: missing momentum



virtual exchange: larger rates

Graviton: Spin-2 tensor force: angular distribution

$\sigma (e^+ e^- \rightarrow G X)$ for δ extra dim.

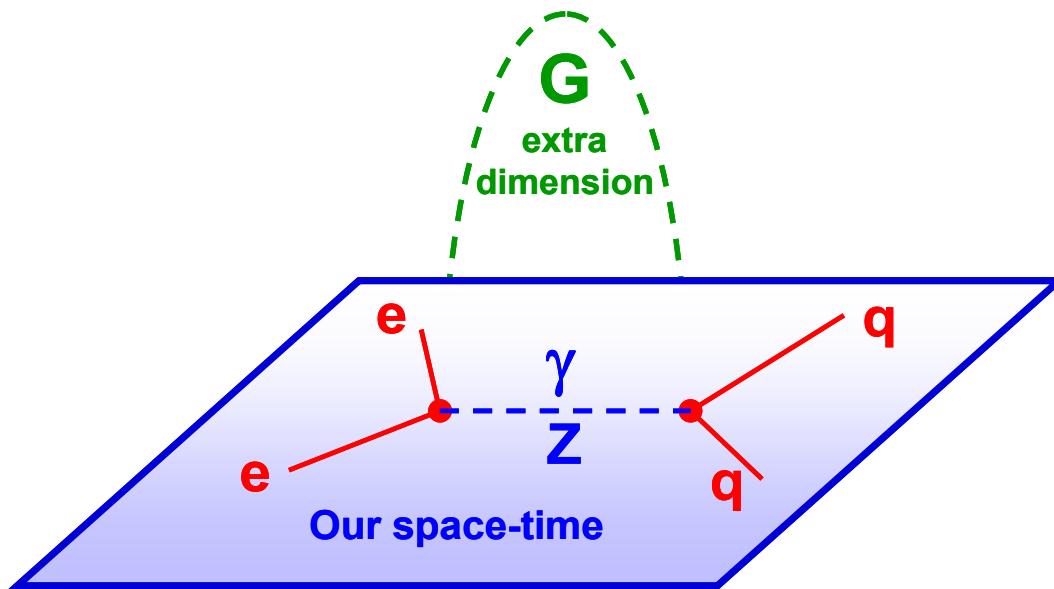


LHC 100 fb^{-1} :

$M_{KK} > 9, 7, 6 \text{ TeV}$ for
 $\delta = 2, 3, 4$

Extra Dimensions

Emission of Kaluza-Klein gravitons



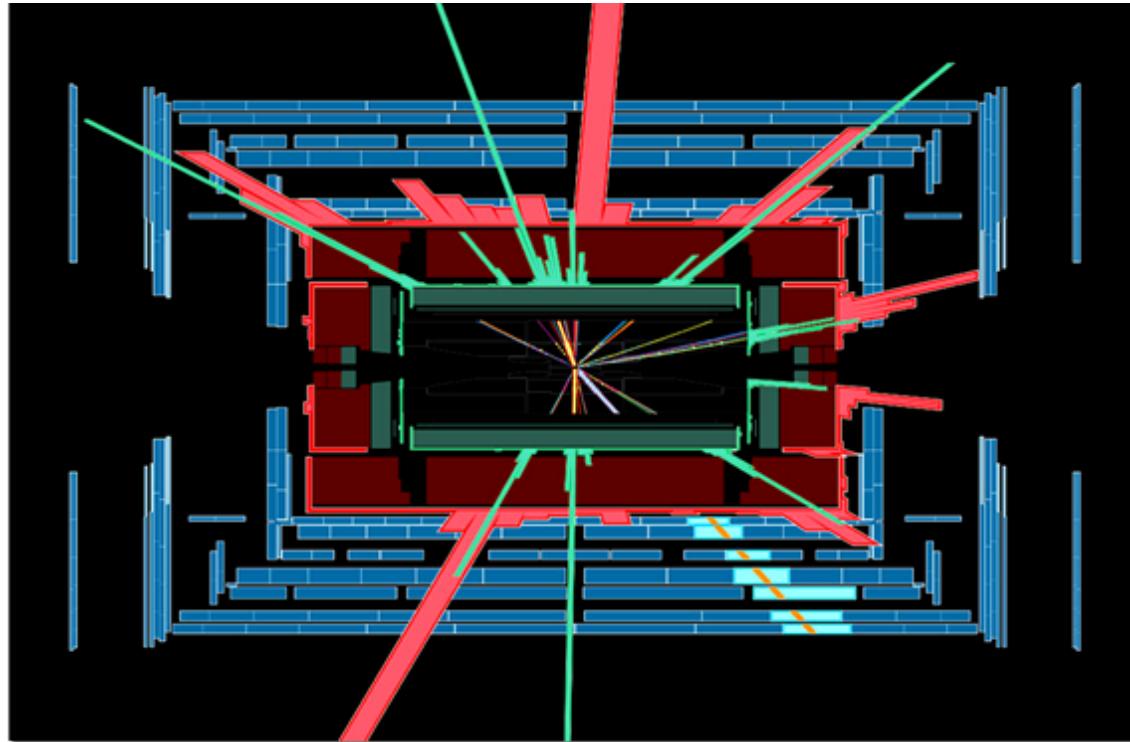
real emission: missing momentum

virtual exchange: larger rates

Graviton: Spin-2 tensor force: angular distribution

Mini Black Holes

- **gravitative resonance:** decay in SUSY + other particles
- evaporate by Hawking radiation
- quantum theory of black holes in the lab !
- detect in **LHC** + cosmic radiation (Auger, ICECUBE)



simulated decay of a Mini Black Hole
in the ATLAS detector at the LHC at CERN

LHC discovery potential

Scenario	LHC Reach
$Z', W' \rightarrow ll, l\nu$	~ 5-6 TeV
Leptoquarks	~ 1.5 TeV
Excited quarks	~ 6.5 TeV
Compositeness scale	~ 40 TeV
Extra dimensions gravity scale	~ 9 TeV for 2 extra dimensions
Magnetic monopoles	~ 20 TeV

