

# Lepton Flavour Violating Z Decays in the MSSM

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- Present status of LFV
  - SUSY: parametrization of slepton mass matrices
  - $Z \rightarrow \ell_I \ell_J$  at DESY TESLA GigaZ
  - Bounds from  $\ell_J \rightarrow \ell_I \gamma$
  - LFV and  $(g_\mu - 2)$
  - Summary and conclusions

[hep-ph/0207328]

# Status of LFV

$\text{BR}(\mu \rightarrow e\gamma) < 1.2 \times 10^{-11}$ [MEGA '99]	$\text{BR}(\mu \rightarrow 3e) < 1.0 \times 10^{-12}$ [SINDRUM '88]
$\text{BR}(\tau \rightarrow e\gamma) < 2.7 \times 10^{-6}$ [CLEO '99]	$\text{BR}(\tau \rightarrow 3e) < 2.9 \times 10^{-6}$ [CLEO '98]
$\text{BR}(\tau \rightarrow \mu\gamma) < 1.1 \times 10^{-6}$ [CLEO '00]	$\text{BR}(\tau \rightarrow 3\mu) < 1.9 \times 10^{-6}$ [CLEO '98]
$R(\mu\text{Ti} \rightarrow e\text{Ti}) < 6.1 \times 10^{-13}$ [SINDRUM '98]	
$\text{BR}(Z \rightarrow \mu e) < 1.7 \times 10^{-6}$ [OPAL '95]	$< 2.0 \times 10^{-9}$
$\text{BR}(Z \rightarrow \tau e) < 9.8 \times 10^{-6}$ [OPAL '95]	$\implies < \kappa \times 6.5 \times 10^{-8}$
$\text{BR}(Z \rightarrow \tau \mu) < 1.2 \times 10^{-5}$ [DELPHI '97]	$< \kappa \times 2.2 \times 10^{-8}$
<b>[GigaZ]</b> ( $\kappa = 0.2 - 1.0$ )	

• Neutrino oscillations (neutral LFV)  $\implies$   $\nu$ SM predicts tiny charged LFV

• SUSY models (unbroken R-parity)  $\implies$  slepton mass matrix natural source of LFV

**Aim:  $Z \rightarrow \ell_I \ell_J$     GigaZ potential, independently of origin of SUSY breaking**

with constraints dominantly from  $\ell_J \rightarrow \ell_I \gamma$  since:

$$\text{BR}(\ell_J \rightarrow 3\ell_I) \approx \alpha_{em} \text{BR}(\ell_J \rightarrow \ell_I \gamma), \quad R(\mu\text{Ti} \rightarrow e\text{Ti}) \approx 5 \times 10^{-3} \times \text{BR}(\mu \rightarrow e\gamma)$$

## General Lorentz Structure $V\bar{\ell}_I\ell_J$

- **Amplitudes** for **on-shell** external legs: **one-loop** parametrization [ $\alpha_W = g^2/(4\pi)$ ]

$$\mathcal{M}_Z = -ig \frac{\alpha_W}{4\pi} \varepsilon_Z^\mu \bar{u}_{\ell_I}(p_2) \left[ \gamma_\mu (f_V^Z - f_A^Z \gamma_5) + \frac{\sigma_{\mu\nu} q^\nu}{M_W} (if_M^Z + f_E^Z \gamma_5) \right] u_{\ell_J}(p_1)$$

$$\mathcal{M}_\gamma = -ie \frac{\alpha_W}{4\pi} \varepsilon_\gamma^\mu \bar{u}_{\ell_I}(p_2) \left[ \gamma_\mu \underbrace{f_V^\gamma}_{=0 \text{ for } I \neq J} + \frac{\sigma_{\mu\nu} q^\nu}{m_{\ell_J}} (if_M^\gamma + f_E^\gamma \gamma_5) \right] u_{\ell_J}(p_1)$$

Dipole form factors ( $f_M, f_E$ ) are **chirality flipping** (proportional to a fermion mass)

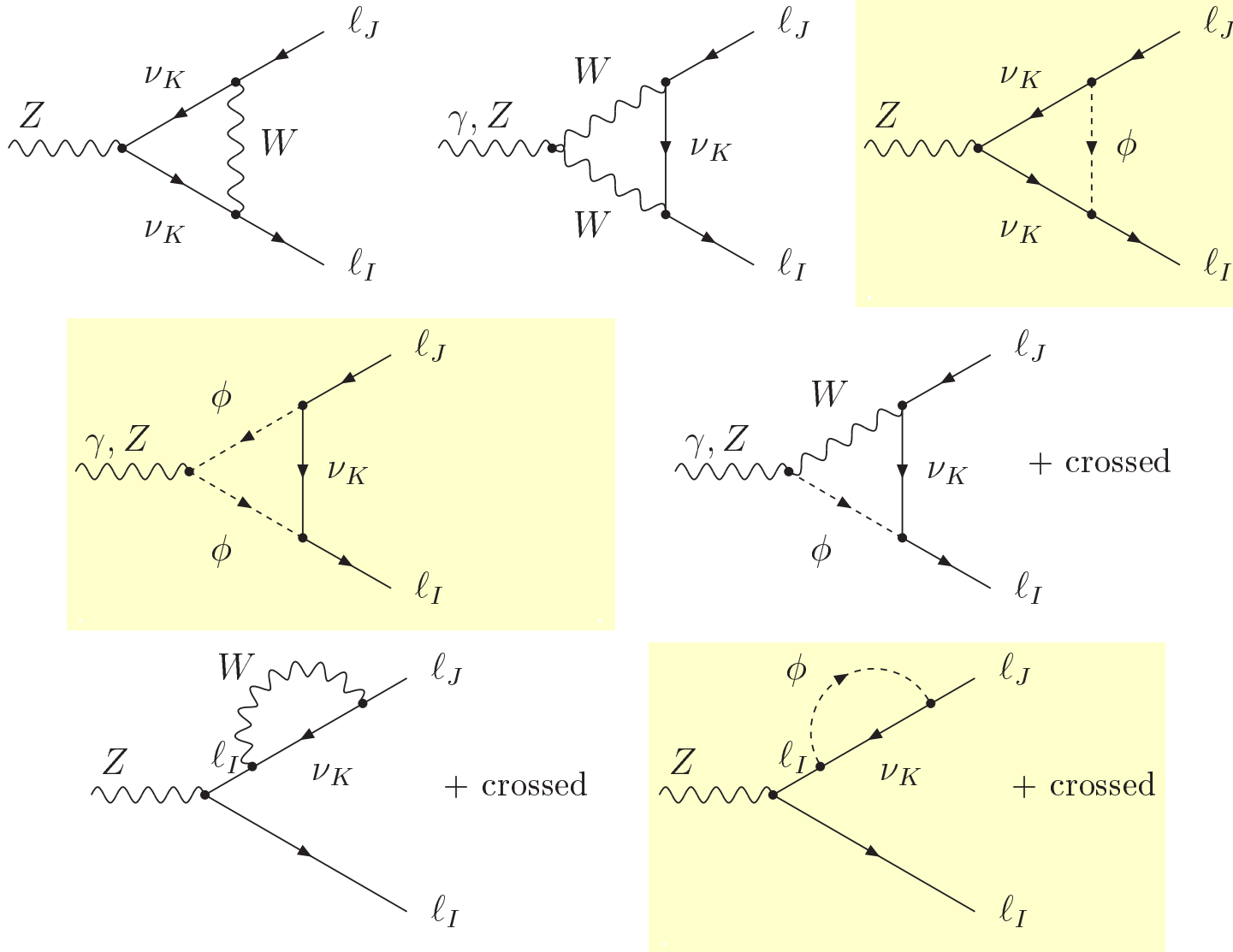
- **Branching ratios** [ $f_{L,R} = f_V \pm f_A$ ]

$$\text{BR}(Z \rightarrow \bar{\ell}_I \ell_J + \ell_I \bar{\ell}_J) = \underbrace{\frac{\alpha_W^3 M_Z}{48\pi^2 \Gamma_Z}}_{\mathcal{O}(10^{-6})} \left[ |f_L^Z|^2 + |f_R^Z|^2 + \frac{1}{c_W^2} (|f_M^Z|^2 + |f_E^Z|^2) \right]$$

$$\text{BR}(\ell_J \rightarrow \ell_I \gamma) = \frac{12\alpha}{\pi} \frac{M_W^4}{m_{\ell_J}^4} (|f_M^\gamma|^2 + |f_E^\gamma|^2) \times \underbrace{\text{BR}(\ell_J \rightarrow \ell_I \nu_J \bar{\nu}_I)}_{=1/0.17/0.17 \text{ for } \mu e/\tau e/\tau \mu}$$

- For  $\ell \equiv \ell_I = \ell_J$ , **anomalous magnetic dipole moment**  $a_\ell = (g_\ell - 2)/2 = \frac{\alpha_W}{4\pi} f_M^\gamma$

# Standard Model + massive neutrinos



$$\mathcal{M} \propto \sin^2 2\theta \frac{\Delta m_{IJ}^2}{M_W^2}$$

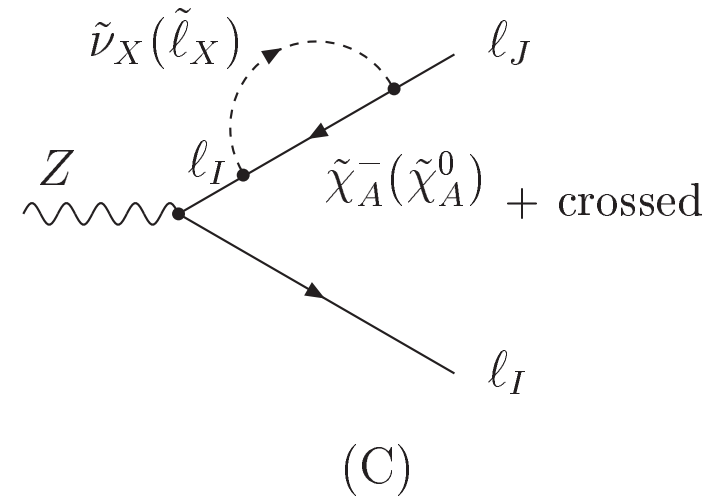
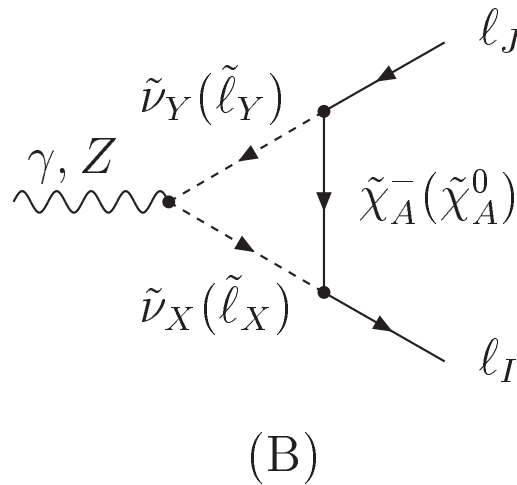
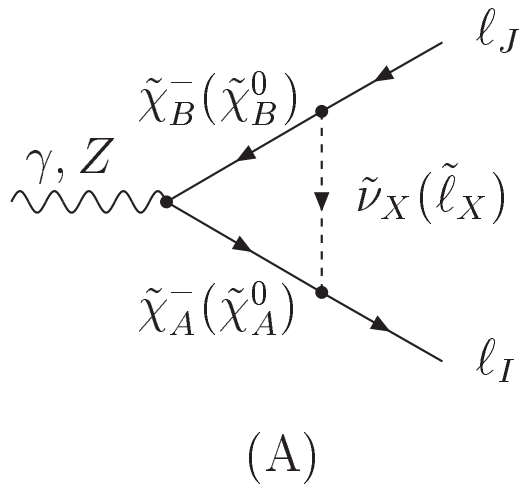
Oscillation data  $\Rightarrow$

$$\text{BR}(l_J \rightarrow l_I \gamma) \lesssim 10^{-48}$$

$$\text{BR}(Z \rightarrow l_I l_J) \lesssim 10^{-54}$$

hopeless!

# MSSM



- **Charginos–sneutrinos**  $(\tilde{\chi}_{A=1,2}^{\pm} ; \tilde{\nu}_{X=1,2,3} \text{ 'L'})$
- **Neutralinos–charged sleptons**  $(\tilde{\chi}_{A=1,2,3,4}^0 ; \tilde{\ell}_{X=1,\dots,6} \text{ 'L' and 'R'})$
- **(A)+(B)+(C): ultraviolet finite and decoupling** of heavy SUSY particles
- (C) no contribution to  $\ell_J \rightarrow \ell_I \gamma$  (no dipole structure)
- **FCNC/LFV** ( $\ell_I \neq \ell_J$ )  $\Leftrightarrow \left\{ \begin{array}{l} \text{slepton family mixing} \\ \text{slepton mass splitting (non-degeneracy)} \end{array} \right. \text{ (LL, RR, LR)}$

# Slepton mass matrices

$\mathbf{M}_{\tilde{\nu}}^2 = (\mathbf{m}_{\tilde{\nu}_L}^2)_{3 \times 3}$ $(\mathbf{m}_{\tilde{\nu}_L}^2)_{IJ} = (\mathbf{m}_L^2)_{IJ} + \frac{1}{2} M_Z^2 c 2\beta \delta_{IJ}$ $(I, J = 1, 2, 3)$	$\mathbf{M}_{\tilde{\ell}}^2 = \begin{pmatrix} \mathbf{m}_{LL}^2 & \mathbf{m}_{LR}^{2T} \\ \mathbf{m}_{LR}^2 & \mathbf{m}_{RR}^2 \end{pmatrix}_{6 \times 6} \quad (\text{symmetric})$ $(\mathbf{m}_{LL}^2)_{IJ} = (\mathbf{m}_L^2)_{IJ} + \left[ m_{\ell_I}^2 + \left( -\frac{1}{2} + s_W^2 \right) M_Z^2 c 2\beta \right] \delta_{IJ}$ $(\mathbf{m}_{RR}^2)_{IJ} = (\mathbf{m}_R^2)_{IJ} + \left[ m_{\ell_I}^2 - M_Z^2 c 2\beta s_W^2 \right] \delta_{IJ}$ $(\mathbf{m}_{LR}^2)_{IJ} = (\mathbf{A}_\ell)_{IJ} v \cos \beta / \sqrt{2} - m_{\ell_I} \mu \tan \beta \delta_{IJ}$
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Assume **two-generation** mixing ( $IJ$ ) and **no alignment** between fermion and scalar fields  $\Rightarrow$

$$\mathbf{M}_{\tilde{\nu}}^2 = \tilde{m}^2 \begin{pmatrix} 1 & \cdot \\ \delta_{LL}^{\tilde{\nu} IJ} & 1 \end{pmatrix}; \quad \mathbf{M}_{\tilde{\ell}}^2 = \tilde{m}^2 \begin{pmatrix} 1 & \cdot & \cdot & \cdot \\ \delta_{LL}^{\tilde{\ell} IJ} & 1 & \cdot & \cdot \\ \delta_{LR}^{\tilde{\ell} II} & \delta_{LR}^{\tilde{\ell} IJ} & 1 & \cdot \\ \delta_{LR}^{\tilde{\ell} JI} & \delta_{LR}^{\tilde{\ell} JJ} & \delta_{RR}^{\tilde{\ell} IJ} & 1 \end{pmatrix} \quad \sin 2\theta = 1$$

and **only one**  $\delta \neq 0$ :  $\tilde{m}_{1,2}^2 = \tilde{m}^2 (\sqrt{1 + \delta^2} \mp \delta) \iff \delta = \frac{\tilde{m}_2^2 - \tilde{m}_1^2}{2\tilde{m}^2}; \quad \tilde{m}^2 = \tilde{m}_1 \tilde{m}_2$

## Calculation

- **Analytical expressions** in terms of **generic couplings** and one-loop **tensor integrals**, implemented with **complete MSSM** Feynman rules (full chargino and neutralino mixings)
- **Constraints on SUSY masses** [PDB]:

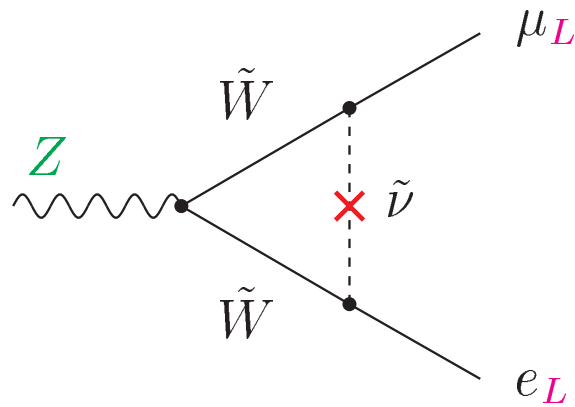
lightest slepton ( $\tilde{m}_1$ )	$m_{\tilde{\nu}} > 45 \text{ GeV}$ $m_{\tilde{\ell}_{L,R}} > 90 \text{ GeV}^*$
lightest chargino	$m_{\tilde{\chi}_1^+} > 75 \text{ GeV}$ , if $m_{\tilde{\nu}} > m_{\tilde{\chi}_1^+}$ $> 45 \text{ GeV}$ , otherwise
lightest neutralino	$m_{\tilde{\chi}_1^0} > 35 \text{ GeV}$

$$* m_{\tilde{\nu}}^2 = m_{\tilde{\ell}_L}^2 + M_Z^2 c_W^2 \cos 2\beta \Rightarrow m_{\tilde{\nu}} > 65 \text{ (40) GeV for } \tan \beta = 2 \text{ (50)}$$

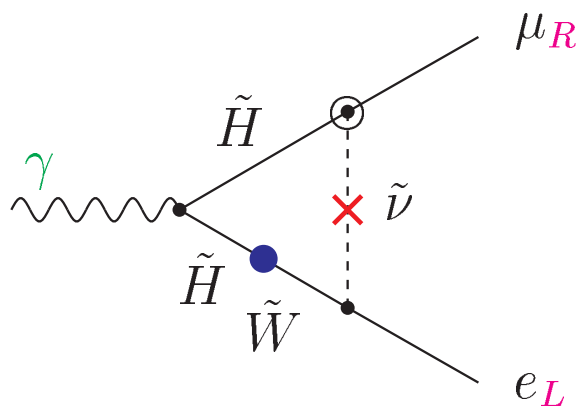
- **Inputs:**  $M_2$ ,  $\mu$ ,  $\tan \beta$  (spectra and couplings of charginos/neutralinos),  $\tilde{m}_1$  (relevant lightest scalar mass),  $\delta$  (LFV parameter)
- **Dominating diagrams** (mass insertion method): proportional to  $\delta$  (if small)

diag

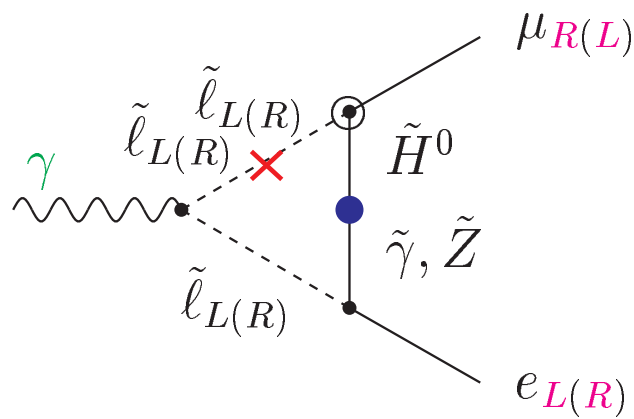
# Dominating diagrams



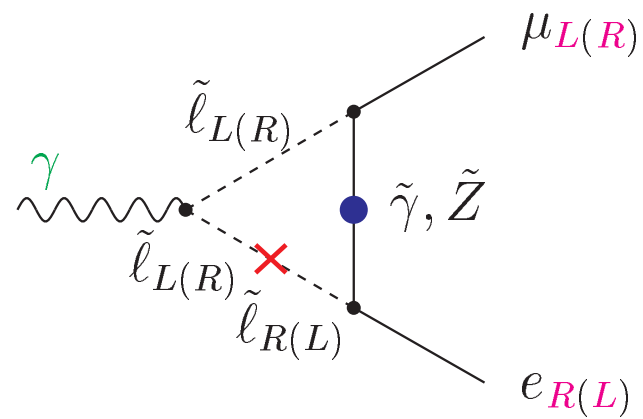
(a)  $\delta_{LL}^{\tilde{\nu} 12}$



(b1)  $\delta_{LL}^{\tilde{\nu} 12} m_\mu \tan \beta$



(b2)  $\delta_{LL(RR)}^{\tilde{\ell} 12} m_\mu \tan \beta$



(b3)  $\delta_{LR}^{\tilde{\ell} 12} m_{\tilde{\gamma}, \tilde{Z}}$



# SUSY predictions for $Z \rightarrow \ell_I \ell_J$

- Ignoring other LFV processes:

- Maximal effects for a decoupled second family ( $\delta^{IJ} \rightarrow \infty$ )

- Chargino–sneutrino diagrams dominate by  $\mathcal{O}(10)$

Results almost independent of lepton masses


Mild dependence on  $\tan \beta$

✓ BR( $Z \rightarrow \ell_I \ell_J$ ) up to  $2.5 \times 10^{-8}$  ( $7.5 \times 10^{-8}$ ) for  $\tan \beta = 2$  (50)

✓ BR( $Z \rightarrow \ell_I \ell_J$ )  $> 2 \times 10^{-9}$  ( $2 \times 10^{-8}$ ) for  $m_{\tilde{\nu}} < 305$  (85) GeV,  $m_{\tilde{\chi}_1^+} < 270$  (105) GeV

- **Cancellations**: contributions from different particles have opposite signs plot

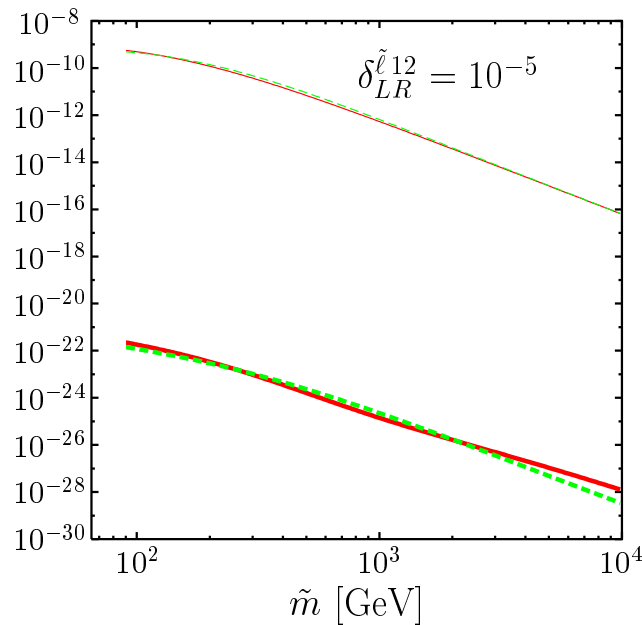
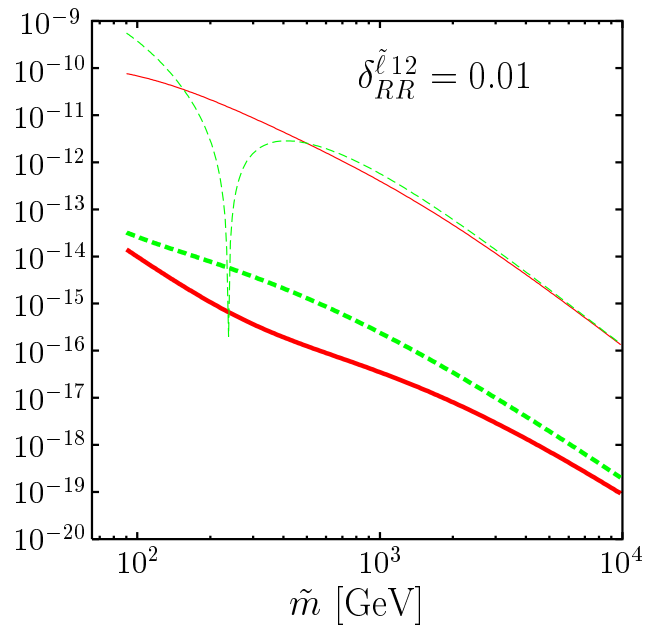
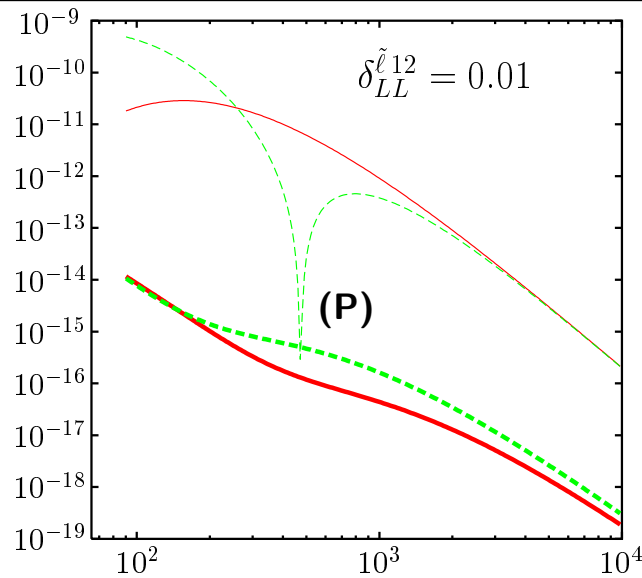
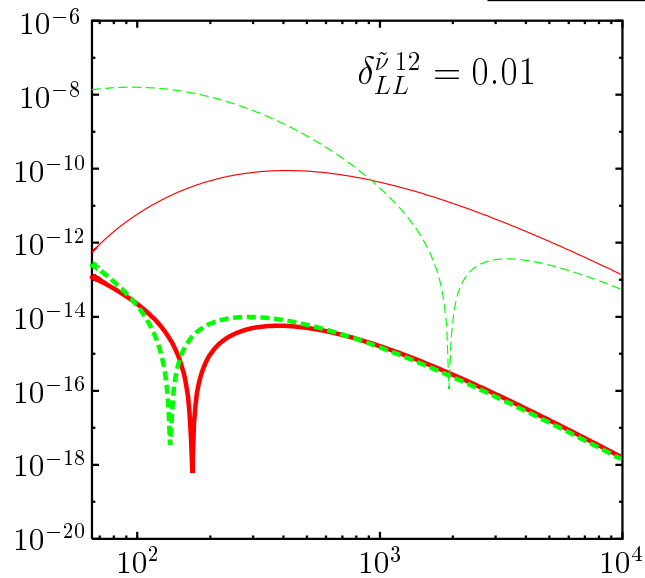
- Correlation with  $\ell_J \rightarrow \ell_I \gamma$ :

 BR( $\mu \rightarrow e \gamma$ )  $< 1.2 \times 10^{-11} \Rightarrow$  BR( $Z \rightarrow \mu e$ )  $< 1.5 \times 10^{-10}$

 BR( $Z \rightarrow \tau e; \tau \mu$ ) **at reach** of best projection of GigaZ ( $\kappa < 1$ )

with BR( $\tau \rightarrow e \gamma; \mu \gamma$ )  $\ll$  present limits plot

# BR( $Z \rightarrow \mu e$ ) & BR( $\mu \rightarrow e \gamma$ )

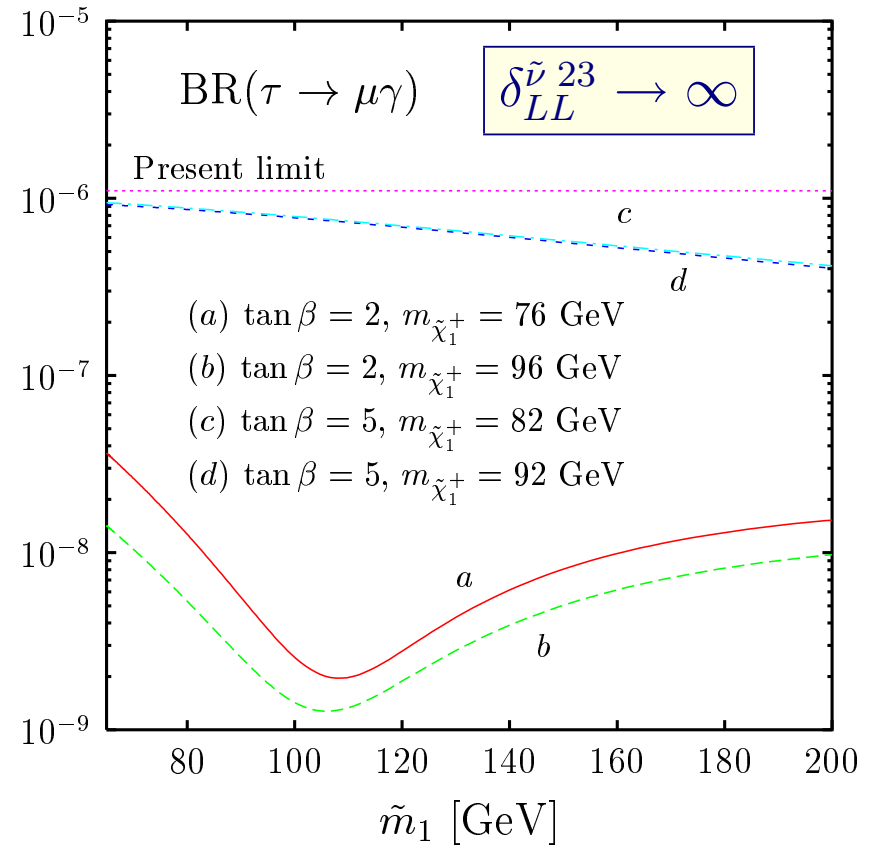
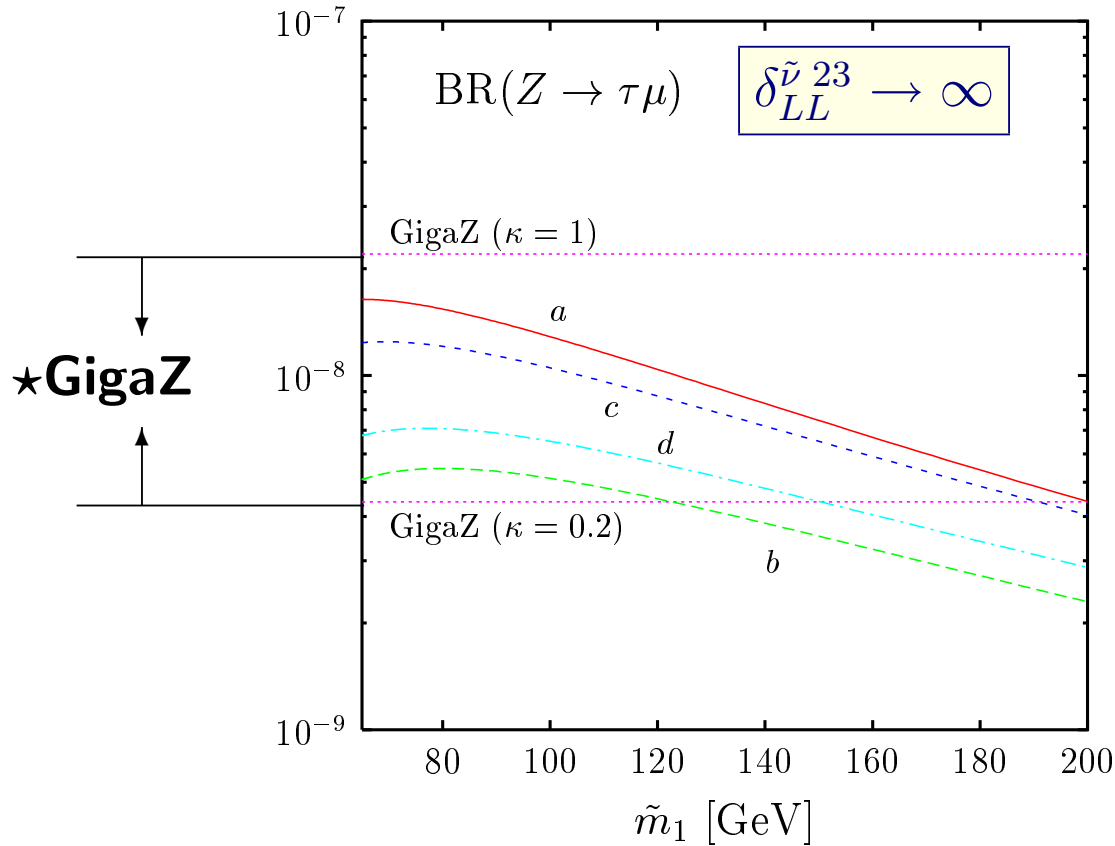


$\tan \beta = 2$   
solid  $M_2 = 150$  GeV  
 $\mu = -500$  GeV  
dashed  $M_2 = \mu = 150$  GeV

asymptotically  $\propto \frac{\delta^2}{\tilde{m}^4}$

(small  $\delta$ )

# BR( $Z \rightarrow \tau\mu$ ) & BR( $\tau \rightarrow \mu\gamma$ )



**(Best) GigaZ reach:**  $55 < m_{\tilde{\nu}} < 215$  GeV

$75 < m_{\tilde{\chi}_1^+} < 100$  GeV

$\tan\beta < 7$

## Bounds from $\ell_J \rightarrow \ell_I \gamma$

- Provide **most stringent constraints** on slepton LFV mass terms (today)
- Establish how severe is the **flavour problem** in the lepton sector of the MSSM
- **Aim:** **update** works by [Masiero *et al* '86–'96] (only photino-mediated diagrams) with **complete** MSSM  $\Rightarrow$  frequent cancellations e.g. plot
- **Bounds** involving the **first two families** are **very restrictive** table

$\delta_{LL}^{\tilde{\nu} 12}$ , $\delta_{LL}^{\tilde{\ell} 12}$ , $\delta_{RR}^{\tilde{\ell} 12}$	:	$\mathcal{O}(10^{-3} \dots 10^{-5})$	stronger for high $\tan \beta$
$\delta_{LR}^{\tilde{\ell} 12}$	:	$\mathcal{O}(10^{-6})$	independent of $\tan \beta$

- **Bounds** involving the **third family** are **much weaker** (or **unexistent!**)

	low $\tan \beta$	high $\tan \beta$
$\delta_{LL}^{\tilde{\nu} I3}$ :	$\infty$	0.03 to 1.3
$\delta_{LL}^{\tilde{\ell} I3}$ :	$\infty$	0.14 to $\infty$
$\delta_{RR}^{\tilde{\ell} I3}$ :	$\infty$	0.11 to $\infty$
$\delta_{LR}^{\tilde{\ell} I3}$ :	0.05 to $\infty$	

# Bounds from $\mu \rightarrow e\gamma < 1.2 \times 10^{-11}$

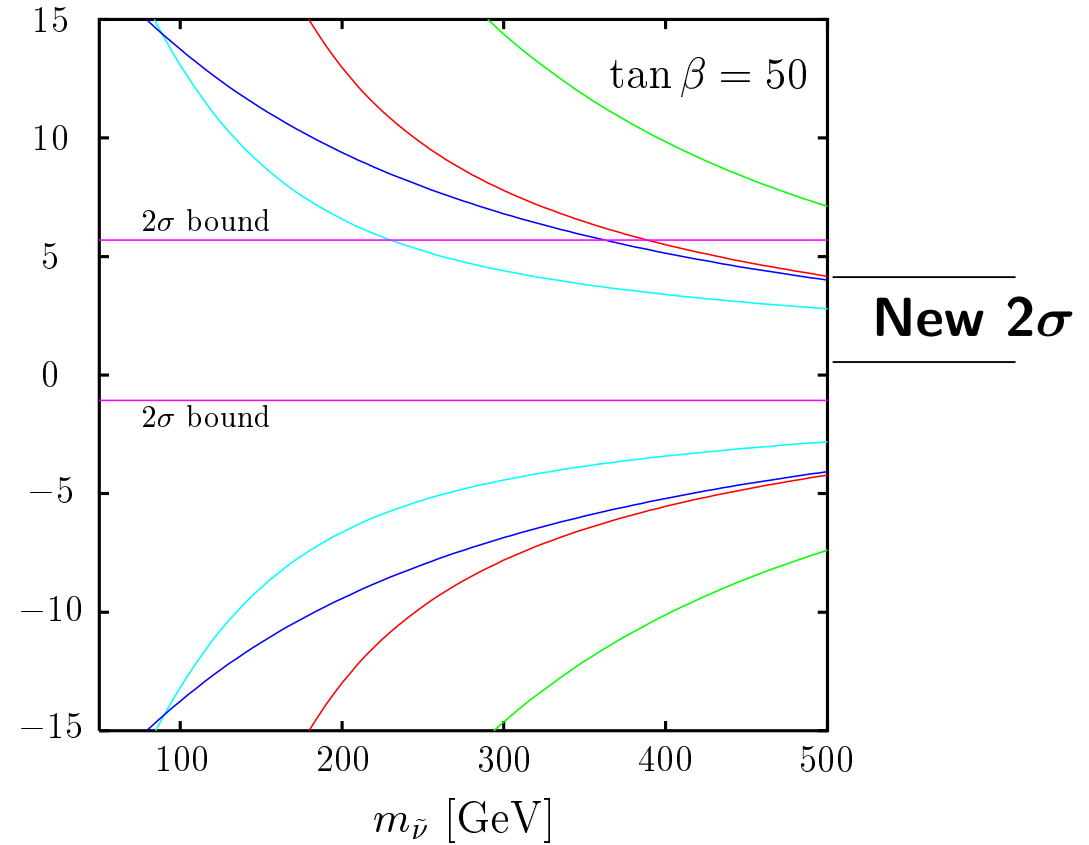
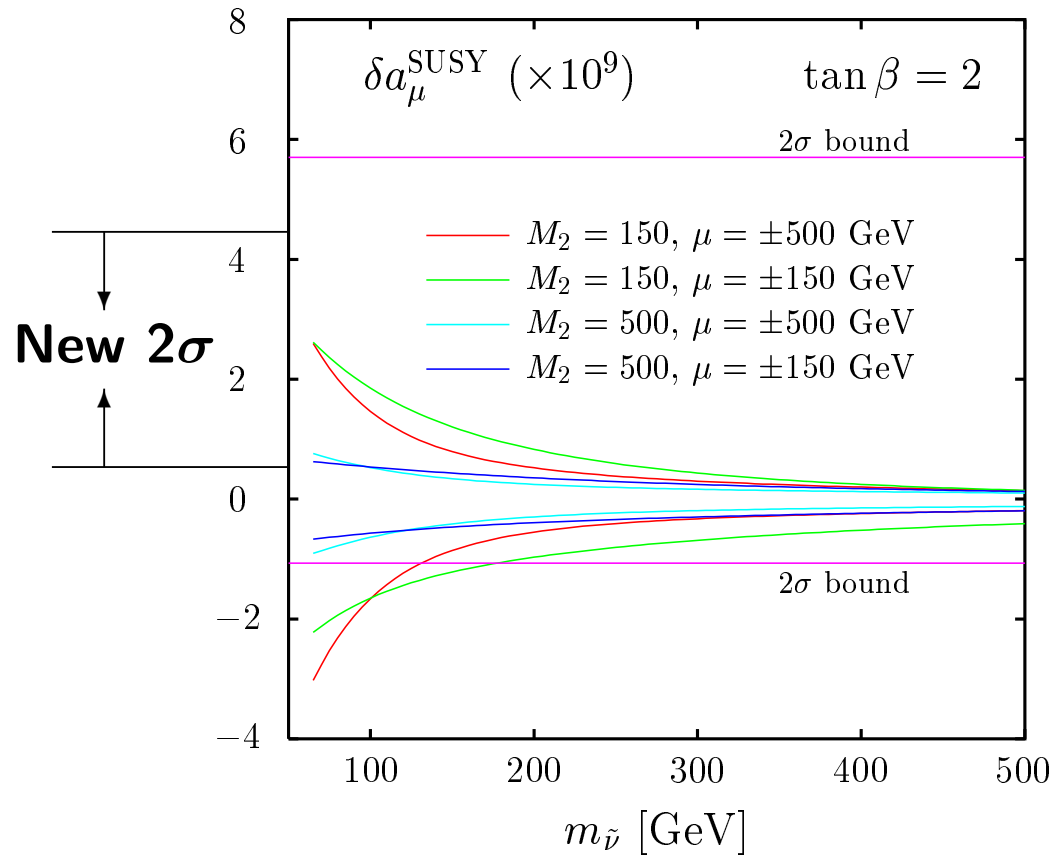
$\tilde{m}_1$	$M_2$	$\delta_{LL}^{\tilde{\nu} 12}$				$\delta_{LL}^{\tilde{\ell} 12}$			
		$\mu = -500$	$\mu = -150$	$\mu = 150$	$\mu = 500$	$\mu = -500$	$\mu = -150$	$\mu = 150$	$\mu = 500$
$\tan \beta = 2$									
100	150	$14 \times 10^{-3}$	$1.0 \times 10^{-3}$	$0.3 \times 10^{-3}$	$1.0 \times 10^{-3}$	$7.5 \times 10^{-3}$	$2.6 \times 10^{-3}$	$1.7 \times 10^{-3}$	$5.0 \times 10^{-3}$
	500	$33 \times 10^{-3}$	$3.0 \times 10^{-3}$	$1.7 \times 10^{-3}$	$13 \times 10^{-3}$	$84 \times 10^{-3}$	$11 \times 10^{-3}$	$7.0 \times 10^{-3}$	$41 \times 10^{-3}$
500	150	$3.7 \times 10^{-3}$	$1.1 \times 10^{-3}$	$1.2 \times 10^{-3}$	$1.7 \times 10^{-3}$	$14 \times 10^{-3}$	$8.5 \times 10^{-3}$	<b>0.12 (P)</b>	$32 \times 10^{-3}$
	500	$7.3 \times 10^{-3}$	$2.6 \times 10^{-3}$	$2.3 \times 10^{-3}$	$4.1 \times 10^{-3}$	$24 \times 10^{-3}$	$20 \times 10^{-3}$	$32 \times 10^{-3}$	$26 \times 10^{-3}$
$\tan \beta = 50$									
100	150	$9.3 \times 10^{-5}$	$2.1 \times 10^{-5}$	$2.0 \times 10^{-5}$	$8.6 \times 10^{-5}$	$2.4 \times 10^{-4}$	$0.8 \times 10^{-4}$	$0.8 \times 10^{-4}$	$2.4 \times 10^{-4}$
	500	$80 \times 10^{-5}$	$9.0 \times 10^{-5}$	$8.8 \times 10^{-5}$	$77 \times 10^{-5}$	$24 \times 10^{-4}$	$3.5 \times 10^{-4}$	$3.4 \times 10^{-4}$	$23 \times 10^{-4}$
500	150	$9.7 \times 10^{-5}$	$4.5 \times 10^{-5}$	$4.5 \times 10^{-5}$	$9.4 \times 10^{-5}$	$7.4 \times 10^{-4}$	$6.4 \times 10^{-4}$	$6.9 \times 10^{-4}$	$7.6 \times 10^{-4}$
	500	$22 \times 10^{-5}$	$9.6 \times 10^{-5}$	$9.5 \times 10^{-5}$	$21 \times 10^{-5}$	$9.7 \times 10^{-4}$	$9.6 \times 10^{-4}$	$9.8 \times 10^{-4}$	$9.7 \times 10^{-4}$
$\tilde{m}_1$	$M_2$	$\delta_{RR}^{\tilde{\ell} 12}$				$\delta_{LR}^{\tilde{\ell} 12}$			
		$\mu = -500$	$\mu = -150$	$\mu = 150$	$\mu = 500$	$\mu = -500$	$\mu = -150$	$\mu = 150$	$\mu = 500$
$\tan \beta = 2$									
100	150	$4.2 \times 10^{-3}$	$1.5 \times 10^{-3}$	$1.8 \times 10^{-3}$	$3.7 \times 10^{-3}$	$1.6 \times 10^{-6}$	$1.5 \times 10^{-6}$	$1.6 \times 10^{-6}$	$1.7 \times 10^{-6}$
	500	$11 \times 10^{-3}$	$3.2 \times 10^{-3}$	$2.5 \times 10^{-3}$	$8.3 \times 10^{-3}$	$4.5 \times 10^{-6}$	$4.4 \times 10^{-6}$	$4.7 \times 10^{-6}$	$4.6 \times 10^{-6}$
500	150	$22 \times 10^{-3}$	$10 \times 10^{-3}$	$22 \times 10^{-3}$	<b>0.33</b>	$1.3 \times 10^{-6}$	$1.2 \times 10^{-6}$	$1.2 \times 10^{-6}$	$1.2 \times 10^{-6}$
	500	$19 \times 10^{-3}$	$12 \times 10^{-3}$	<b>0.33</b>	$35 \times 10^{-3}$	$7.6 \times 10^{-6}$	$7.5 \times 10^{-6}$	$7.6 \times 10^{-6}$	$7.7 \times 10^{-6}$
$\tan \beta = 50$									
100	150	$1.6 \times 10^{-4}$	$0.6 \times 10^{-4}$	$0.6 \times 10^{-4}$	$1.5 \times 10^{-4}$	$1.6 \times 10^{-6}$	$1.5 \times 10^{-6}$	$1.6 \times 10^{-6}$	$1.6 \times 10^{-6}$
	500	$3.8 \times 10^{-4}$	$1.1 \times 10^{-4}$	$1.1 \times 10^{-4}$	$3.8 \times 10^{-4}$	$4.5 \times 10^{-6}$	$4.5 \times 10^{-6}$	$4.6 \times 10^{-6}$	$4.5 \times 10^{-6}$
100	150	$16 \times 10^{-4}$	$13 \times 10^{-4}$	$15 \times 10^{-4}$	$17 \times 10^{-4}$	$1.3 \times 10^{-6}$	$1.2 \times 10^{-6}$	$1.2 \times 10^{-6}$	$1.3 \times 10^{-6}$
	500	$9.4 \times 10^{-4}$	$8.1 \times 10^{-4}$	$8.7 \times 10^{-4}$	$9.6 \times 10^{-4}$	$7.7 \times 10^{-6}$	$7.6 \times 10^{-6}$	$7.6 \times 10^{-6}$	$7.7 \times 10^{-6}$

# LFV and $(g_\mu - 2)$

$$\left. \begin{array}{l}
 \text{World average: } a_\mu^{\text{exp}} = 11\,659\,202.3 (15.1) \times 10^{-10} \\
 \text{including new BNL: } \quad \quad \quad \underbrace{11\,659\,203 (8) \times 10^{-10}} \\
 \text{SM prediction: } a_\mu^{\text{SM}} = 11\,659\,179.2 (9.4) \times 10^{-10}
 \end{array} \right\} \Rightarrow \delta a_\mu = \underbrace{(23.1 \pm 16.9)}_{\text{new: } \substack{24 \\ 10}} \times 10^{-10}$$

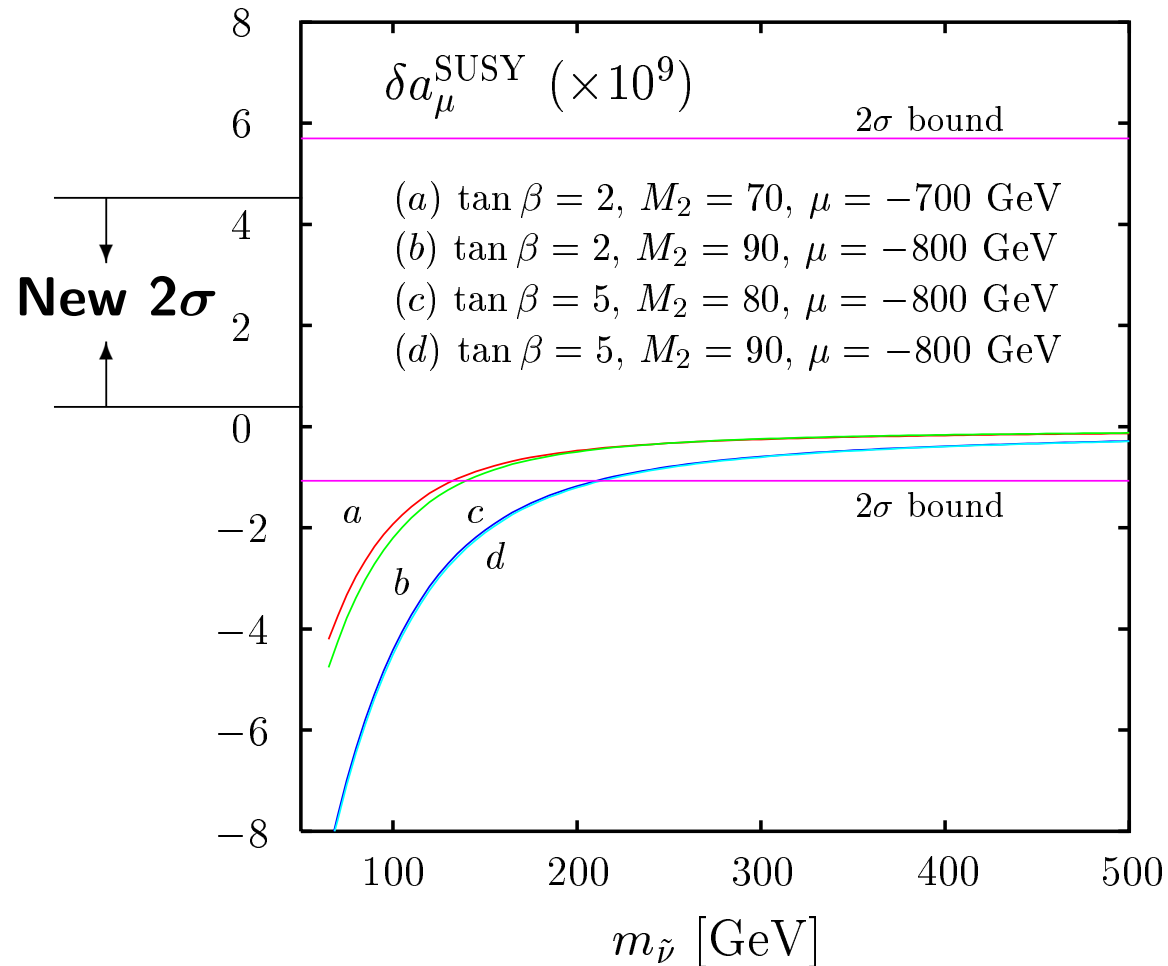
- $(g_\mu - 2) \approx$  ‘normalization’ of LFV lepton decays (small deltas):  
 (similar diagrams  $\Rightarrow$  replace  $\delta_{LR}^{\tilde{\ell} IJ}$  by  $\delta_{LR}^{\tilde{\ell} 22}$  and remove all other delta insertions)
- $\delta a_\mu^{\text{SUSY}} = \delta a_\mu \Rightarrow$  **large uncertainties**
  - SM prediction
  - Extra assumptions:
    - \* take **same** mass scales  $\tilde{m}$  involved in  $(g_\mu - 2)$  as in  $\mu \rightarrow e\gamma$  or  $\tau \rightarrow \mu\gamma$
    - \* soft-breaking terms  $m_L$  and  $m_R$  involved **at the same time** in  $(g_\mu - 2)$
- **But taking present results seriously ...**

# ( $g_\mu - 2$ ) and SUSY



- **Positive** Higgsino mass parameter  $\mu$  preferred (now even more with new BNL data)

# BR( $Z \rightarrow \tau\mu$ ) & ( $g_\mu - 2$ ): very speculative!



- Assuming  $m_L = m_R$ : the favourite scenario for  $Z \rightarrow \tau\mu$  in GigaZ restricted to higher slepton masses (or practically ruled out by new BNL data)



# Summary and conclusions

- **Slepton sector of SUSY** (soft-breaking terms) provides a **natural source of LFV**:  
effects  $\approx$  proportional to slepton mass squared differences (degeneracy  $\delta$ )
- **Our calculation**:
  - $Z \rightarrow \ell_I \ell_J$  realistically correlated with  $\ell_J \rightarrow \ell_I \gamma$  in SUSY models  
(no SUSY-breaking mechanism assumed and exact beyond mass insertion approx.)
  - update to include complete MSSM (several  $\delta$  types) and more recent experiments
- **GigaZ reach**:
  - $Z \rightarrow \mu e$  excluded by  $\mu \rightarrow e \gamma$  not by limits on SUSY masses
  - $Z \rightarrow \tau e; \tau \mu$  possible ( $\kappa < 1$ ) thanks to chargino-sneutrino contributions
- **Best bounds from  $\ell_J \rightarrow \ell_I \gamma \Rightarrow$  SUSY lepton flavour problem?**
  - Non-observation of  $\mu \rightarrow e \gamma \Rightarrow \delta_{LL,RR,LR}^{\tilde{\nu} | \tilde{\ell}}{}^{12} \approx 10^{-3} \dots 10^{-6}$  (large degeneracy!)  
... maybe justifiable by weakness of Yukawas of light lepton sector
  - Non-observation of  $\tau \rightarrow e \gamma; \mu \gamma \Rightarrow$  very weak (even nonexistent) constraints on  $\delta^{I3}s$