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# The Direct Limit on the Higgs Mass & the SM Fit

MC PRL 87:231802, 2001  
PRD 66:073002, 2002

Updated:

- 1) LEPEWWG/2002-02 12/02
- 2) 2-loop  $m_W$  Freitas, Heinemeyer, Weiglein 12/02
- 3) **New  $m_W$**  Aleph Today!

## New at this meeting:

$$m_W = 80.426 (34)$$

Previously:  $80.449 (34)$

From 79 MeV shift in Aleph measurement,  
due to systematic effect.

Improves betting odds on SM fit relative to Summer 02:  
improves global fit & raises  $m_H$  prediction.

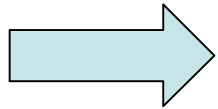
This talk: initially prepared results (based on Summer 02)  
+ revised results (with new  $m_W$ )

## LEP direct limit

(New  $m_W$  results in green (...)’s )

$$m_H > 114 \text{ GeV} \quad (95\%) \quad \text{N.B., } \text{CL}(m < 114) \ll 5\%$$

imposes important constraint on interpretation of EW data.



To test SM, consider

$$\text{Fit:} \quad \text{CL}(\chi^2) = 0.01 \quad (0.02) \quad (\text{No APV, } \chi_W)$$

AND

$$m_H > 114: \quad \text{CL}(m) = 0.3 \quad (0.35) \quad \left\{ \begin{array}{l} \text{CL from SM fit} \\ \text{that } m_H > 114 \end{array} \right.$$

Useful to consider **TOTAL CL**:

$$\text{CL}_T = \text{CL}(\chi^2) \parallel \text{CL}(m) = 0.003 \quad (0.0066)$$

Poor CL( $\chi^2$ ) due to two 3 $\chi$  anomalies:

$$\begin{array}{l}
 1) \quad x_W^{\chi^2 N} \\
 2) \quad x[A_L] = 0.23113 \quad (21) \\
 \\
 x[A_H] = 0.23217 \quad (29)
 \end{array}
 \left\{ \begin{array}{l}
 A_{LR}, A_{FB}^1, A(P_\chi) \\
 \chi^2 = 1.7/2 \\
 \\
 A_{FB}^b, A_{FB}^c, Q_{FB} \\
 \chi^2 = 0.05/2
 \end{array} \right.$$

dominated by  $\left\{ \begin{array}{l} A_{LR} = 0.23098 \quad (26) \\ A_{FB}^b = 0.23217 \quad (31) \end{array} \right.$

Generic Explanations:

New physics -- **certainly possible**

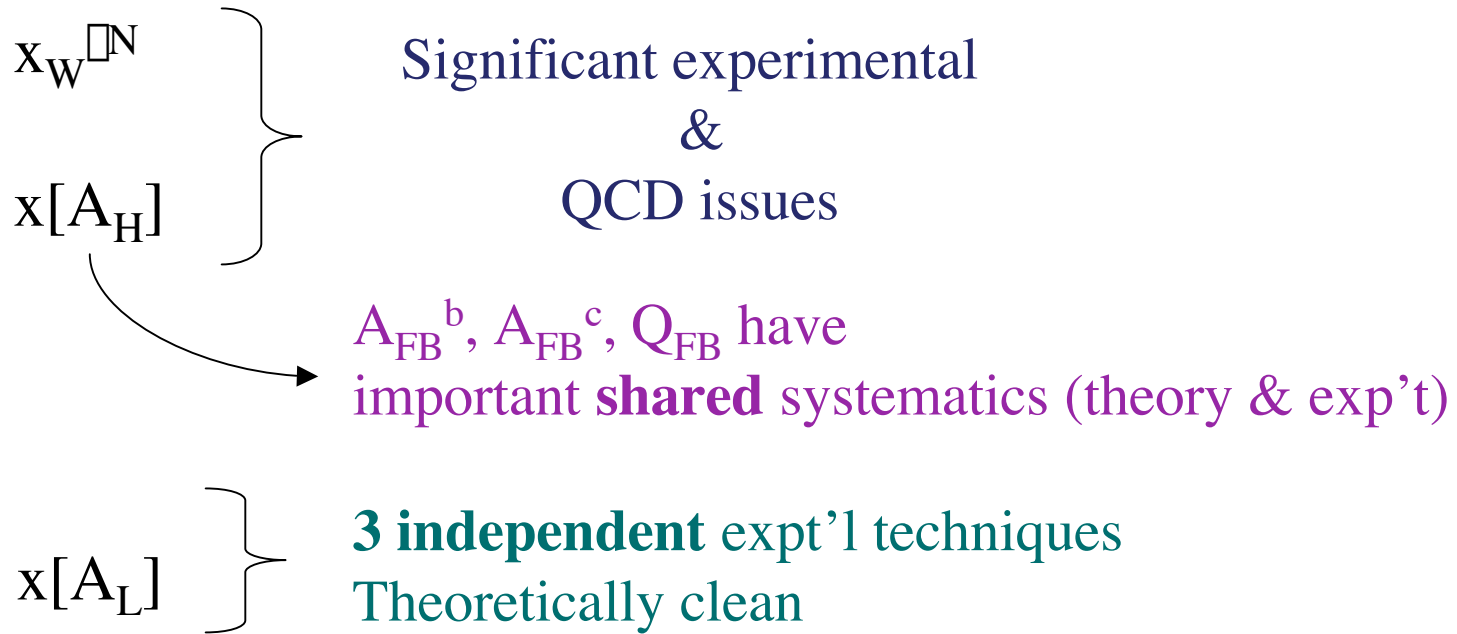
Statistical fluctuation -- **fairly valued**

Underestimated systematic uncertainty

Focus on sys. uncertainty **not** because it is more likely **but** to see if it can rescue the SM.

$m_H$  direct lower limit is central to the analysis.

# Systematics



➡ underestimated sys. uncertainty unlikely for  $x[A_L]$ ,  
more conceivable for  $x_W^N$  &  $x[A_H]$ ,  
and  $x_W^N, A_{FB}^b$  contribute largest pulls, causing poor  $CL(\chi^2)$ .

**Q:** Could underestimated sys. uncertainty in  $x_W^N, x[A_H]$  improve SM fit?

## SM Fits

SM EWRC from ZFITTER 6.30 + 2-loop  $m_W$  FHW  $\rightarrow$  - (5 - 10) MeV

$$m_Z, m_t, \alpha_5, \alpha_S, m_H \quad \longrightarrow \quad O_{Z\text{-Pole}} + m_W + x_W \alpha^2 + \dots$$

Good agreement with EWWG.

$\alpha_5$  from BP (BES) -- EWWG default

Stronger constraints from “theory driven.” cf MC PRL87:231802, 2001

$\chi^2$  and “Bayesian” likelihood fits

- Vary  $m_t, \alpha_5, \alpha_S, m_H$
- Fit  $m_t, \alpha_5$  + all/some of  $\{13 O_{Z\text{-Pole}}, m_W, x_W \alpha^2\}$
- Correlations alla EWWG

(constrain  $\alpha_S = 0.118(3)$  if  $\alpha_Z, R_1,$  or  $\alpha_H$  not in fit)

Global Fits      (Summer 02  $m_W$ )

A)      **All**  
           27.7/13  
           CL( $\chi^2$ ) = **0.010**  
**P( $\geq 1.3\sigma$  &  $\geq 1.26\sigma$ , N = 13) = **0.005****

B)      -  $\chi_W^2$   
           18.3/12  
           CL( $\chi^2$ ) = **0.11**  
**P( $\geq 1.28\sigma$ , N = 12) = **0.09****

C)      -  $\chi[A_H]$   
           17.7/10  
           CL( $\chi^2$ ) = **0.060**  
**P( $\geq 1.32\sigma$ , N = 10) = **0.013****

D)      -  $\chi_W^2$  -  $\chi[A_H]$   
           7.1/9  
           CL( $\chi^2$ ) = **0.63**

Global CL's correctly reflect probability for outliers relative to sample size.  
 The statistical ensemble is multiple replays of the 90's @ LEP, SLC, TeVatron.

**Likelihoods for statistical fluctuation are fairly valued.**

# $m_W$ : Shapes SM Global Fit

'98 80.410 (90)

'02 80.449 (34)

Heavy, precise  $m_W$  contributes to marginalization of SM fit:

\*Favors  $x[A_L]/A_{LR}$ , forces  $x[A_H]/A_{FB}^b$  to large pull.

\*IF  $m_W$  were  $\sim 2\%$  lighter or  $\sim 3$  times less precise

-  $CL(\chi^2)$  increases by  $\sim 2$  (fit B: 0.1  $\rightarrow$  0.2)

- no preference for  $x[A_L]$  over  $x[A_H]$

$$m_W = 80.449$$



$$x_W^1 = \begin{cases} 0.23081 & \text{ZFITTER} \\ 0.23071 & \text{2-loop} \end{cases}$$

$80.426 \rightarrow 0.23095$

$$x[A_{LR}] = 0.23098$$

$$x[A_{FB}^b] = 0.23217$$

# $m_H$ -sensitive observables: $CL(\chi^2)$

New  $m_W$  in (green)

Non-asymmetry:  $m_W, \chi_\chi, R_1$   
 $x[A_L]: A_{LR}, A_{FB}^1, A[P_\chi]$   
 $x[A_H]: A_{FB}^b, A_{FB}^c, Q_{FB}$

$$CL(\chi^2) = 0.035 \quad (0.067)$$

+  $x_W \chi^N$

$$CL(\chi^2) = 0.0019 \quad (0.0046)$$

Highest precision:

$A_{LR}, A_{FB}^b, m_W$

$$CL(\chi^2) = 0.0027 \quad (0.006)$$



**Problems of global fit are concentrated in  $m_H$ -sensitive sector.**

Non  $m_H$ -sensitive:

$\chi_\chi, R_b, A_b, R_c, A_c$

$$CL(\chi^2) = 0.68$$

# $m_H$ -sensitive observables: $m_H$ predictions

New  $m_W$  in (green)

<u>Highest Precision</u>	$m_H$	95%	CL( $m_H > 114$ )
$A_{LR}$	39	$< 122$	0.062
$A_{FB}^b$	410	$130 < m < 1200$	0.97
$m_W$	19 (35)	$< 111$	0.047 (0.12)

<u>Aggregates</u>	$m_H$	95%	CL( $m_H > 114$ )
$x[A_L]$	55	$< 143$	0.10
$x[A_H]$	410	$140 < m < 1200$	0.97
$m_W, \square_Z, R_1$	14 (17)	$< 61$	0.016 (0.057)

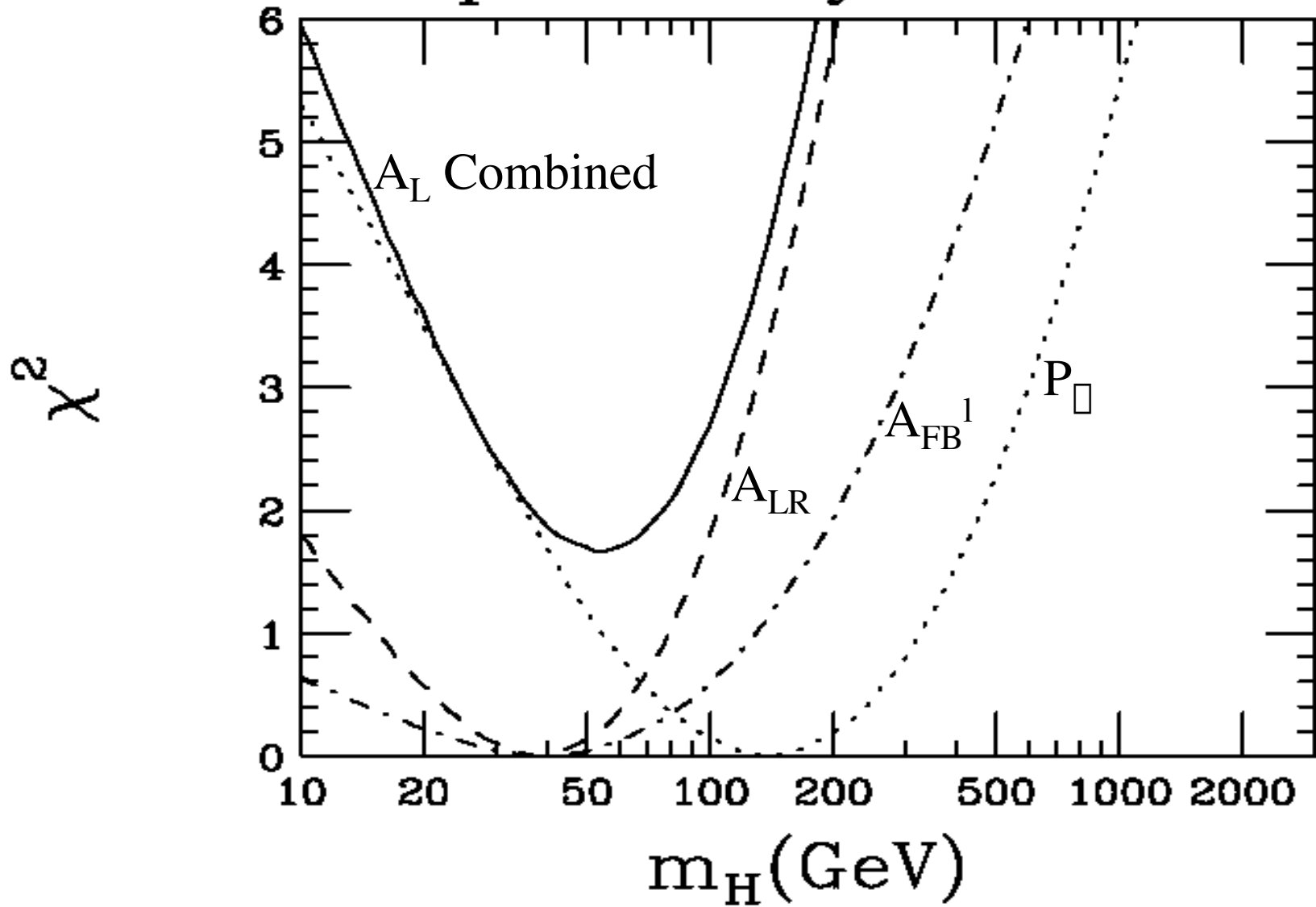


→ Non-asymmetry observables

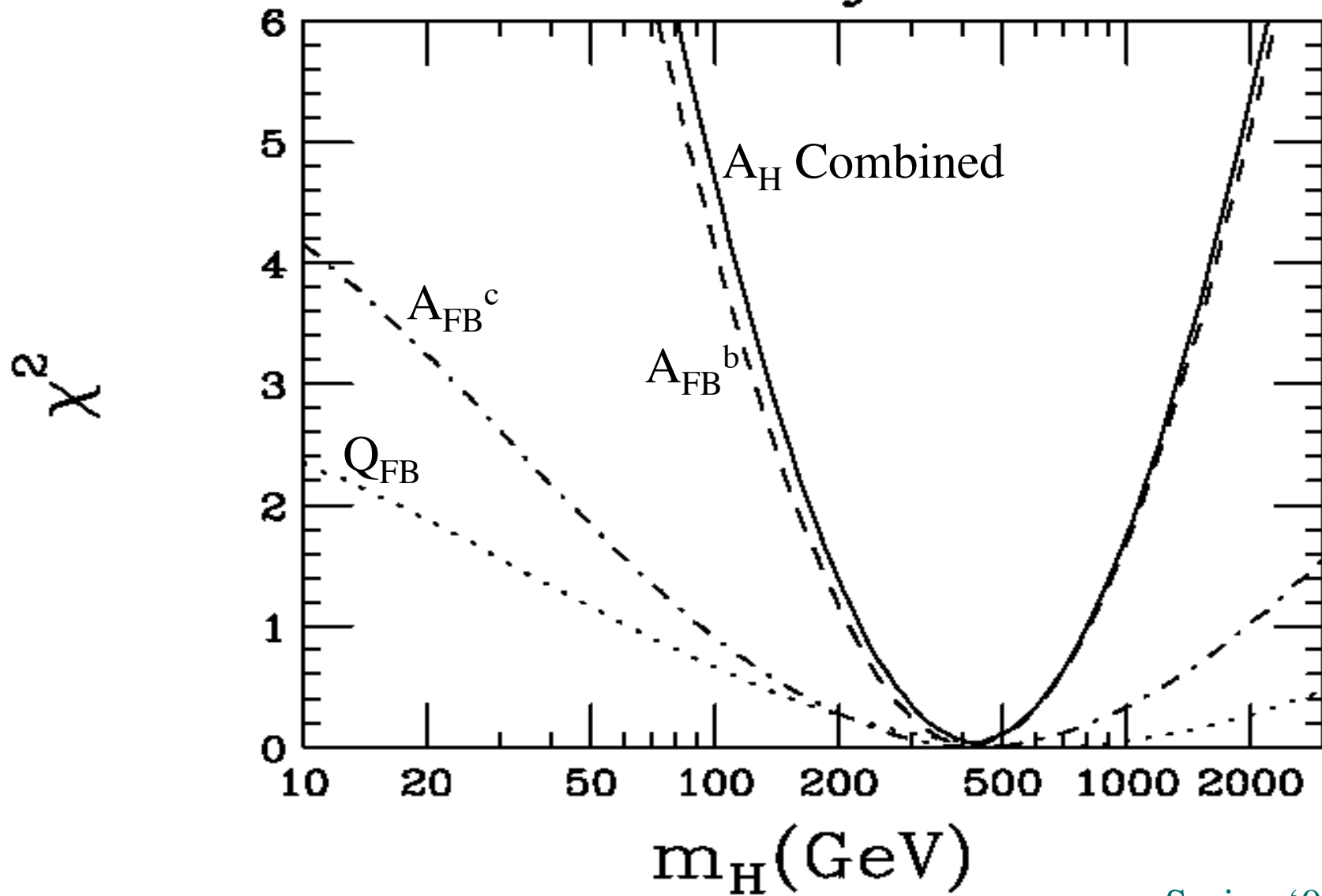


Support for  $m_H > 114$  only from  $x[A_H]$  (+  $x_W \square^N$ )

# Leptonic Asymmetries

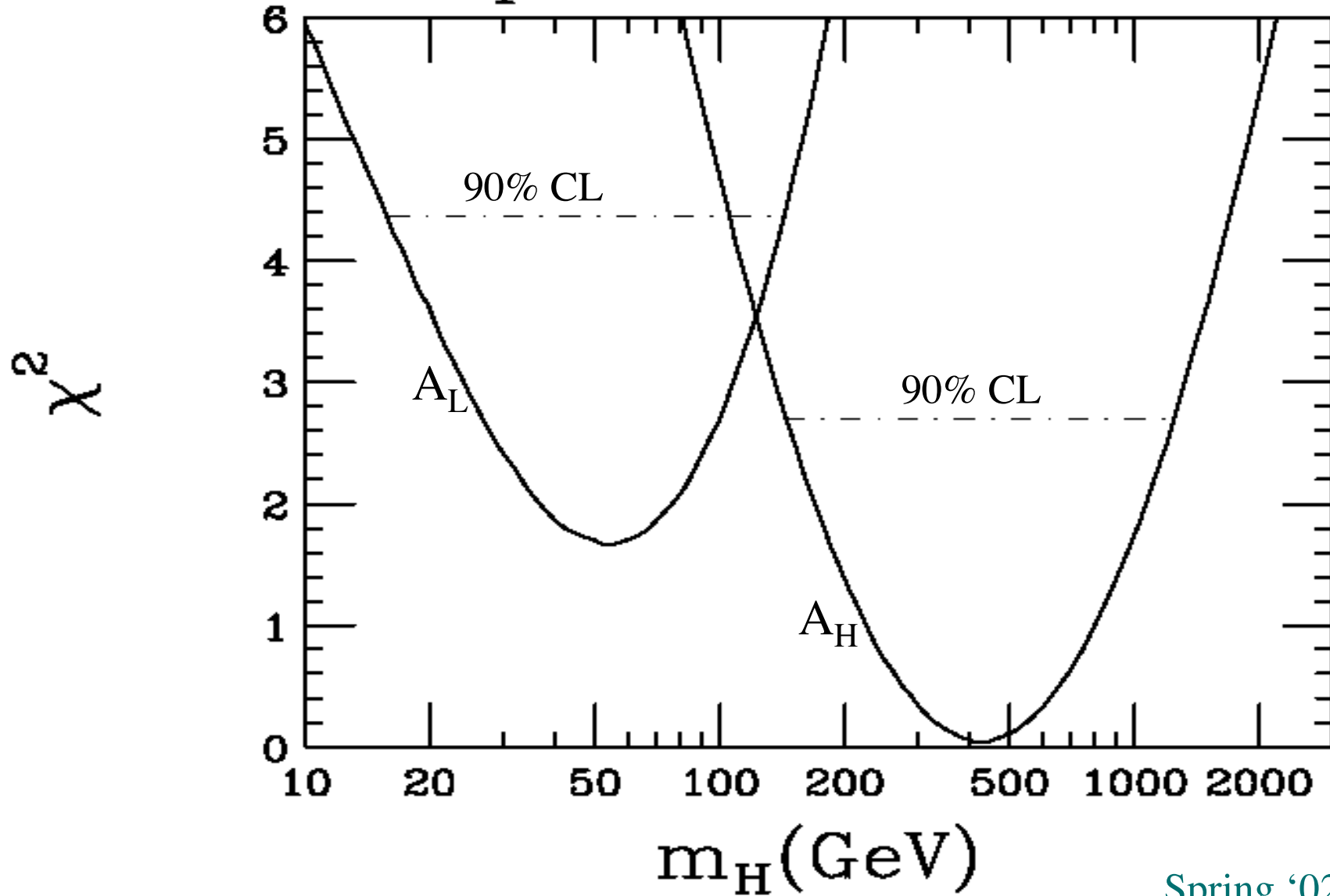


# Hadronic Asymmetries

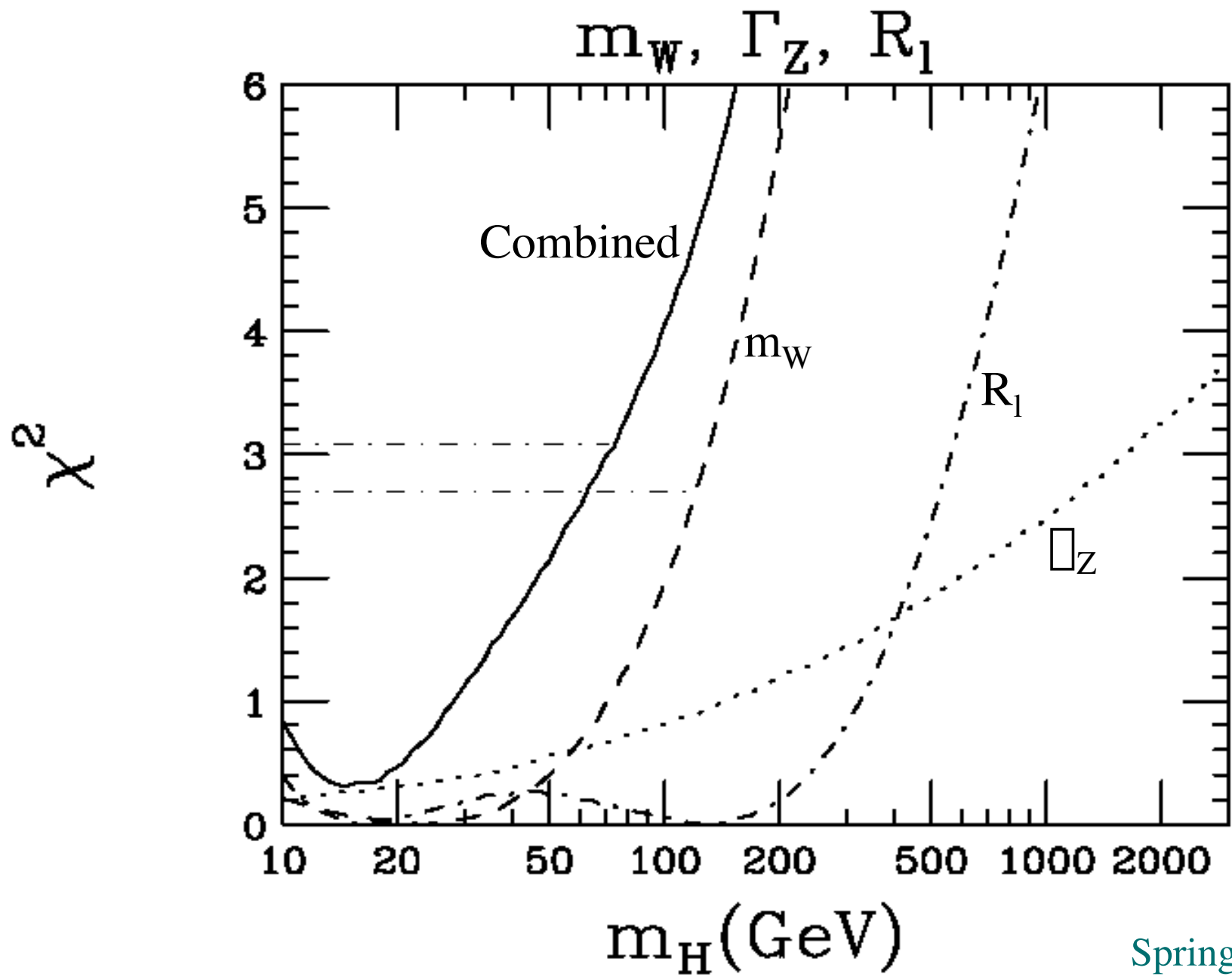


Spring '02

# Leptonic & Hadronic



Spring '02



Spring '02

Global Fits:  $CL_T = CL(\square) \parallel CL(m_H > 114)$

Summer 02  $m_W$

A) All  
27.7/13  
 $m_H = 90$   
 $CL_T = 0.010 \parallel 0.30 = 0.0030$

B)  $-x_W^{\square N}$   
18.3/12  
 $m_H = 81$   
 $CL_T = 0.107 \parallel 0.247 = 0.026$

C)  $-x[A_H]$   
17.7/10  
 $m_H = 45$   
 $CL_T = 0.060 \parallel 0.048 = 0.0029$

D)  $-x_W^{\square N} - x[A_H]$   
7.1/9  
 $m_H = 43$   
 $CL_T = 0.63 \parallel 0.035 = 0.022$

**$CL_T \sim$  invariant under removal of  $x[A_H]$ .**

**Without  $x[A_H]$ , SM fit is inconsistent with search limit.**

Global Fits:  $CL_T = CL(\square) \parallel CL(m_H > 114)$

New  $m_W$

A) All  
25.7/13  
 $m_H = 89$   
 $CL_T = 0.019 \parallel 0.35 = 0.0066$

B)  $-x_W^{\square N}$   
16.5/12  
 $m_H = 89$   
 $CL_T = 0.17 \parallel 0.29 = 0.049$

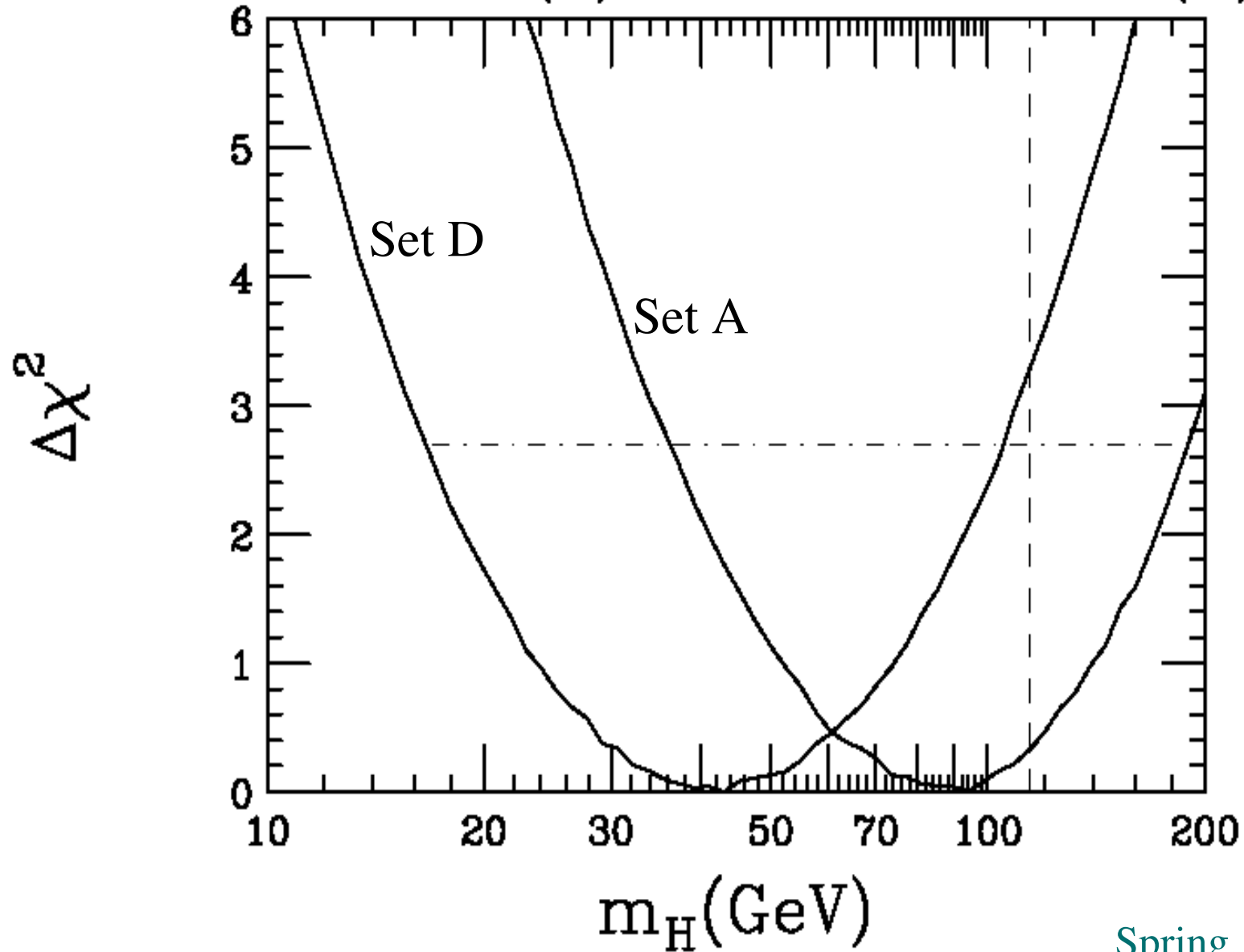
C)  $-x[A_H]$   
16.7/10  
 $m_H = 45$   
 $CL_T = 0.081 \parallel 0.068 = 0.0055$

D)  $-x_W^{\square N} - x[A_H]$   
6.3/9  
 $m_H = 45$   
 $CL_T = 0.71 \parallel 0.049 = 0.035$

**$CL_T \sim$  invariant under removal of  $x[A_H]$ .**

**Without  $x[A_H]$ , SM fit is inconsistent with search limit.**

# All-Data (A) & Minimal Set (D)



Spring '02

# “Bayesian” Likelihood

Summer 02

Instead of  $\chi^2$ , get  $CL(m_H > 114)$  from likelihood.

“Bayesian:”

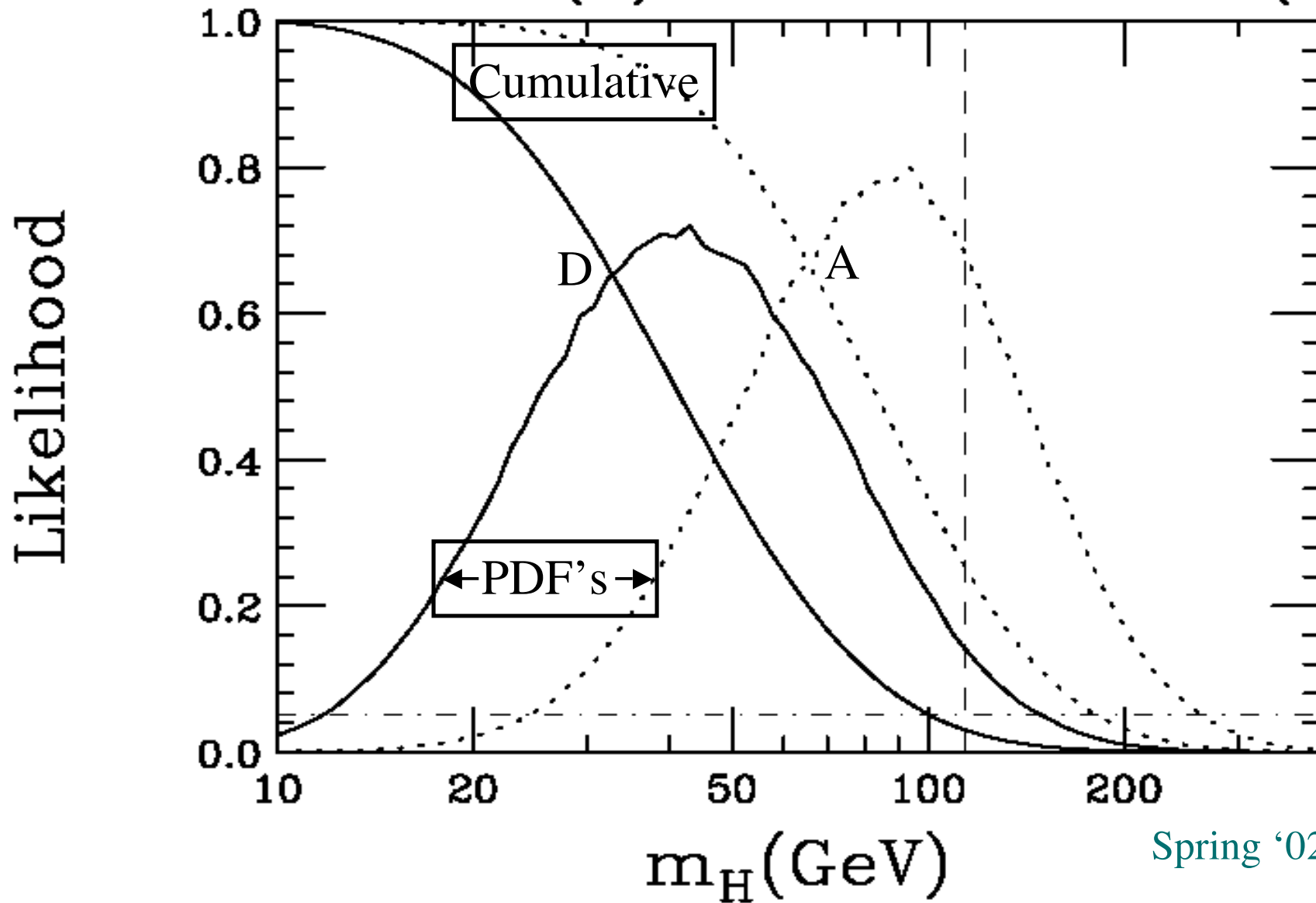
- Probability measure  $d \ln(m_H)$
- Normalize pdf to  $10 < m_H < 3000 \text{ GeV}$

Results agree with  $\chi^2$ :

E.g., ‘Fit D’ = All -  $x[A_H]$  -  $x_W^N$

	PDF	$\chi^2$
$m_H$	43	43
95%	< 102	< 105
CL( $m > 114$ )	0.032	0.035

# All Data (A) & Minimal Set (D)

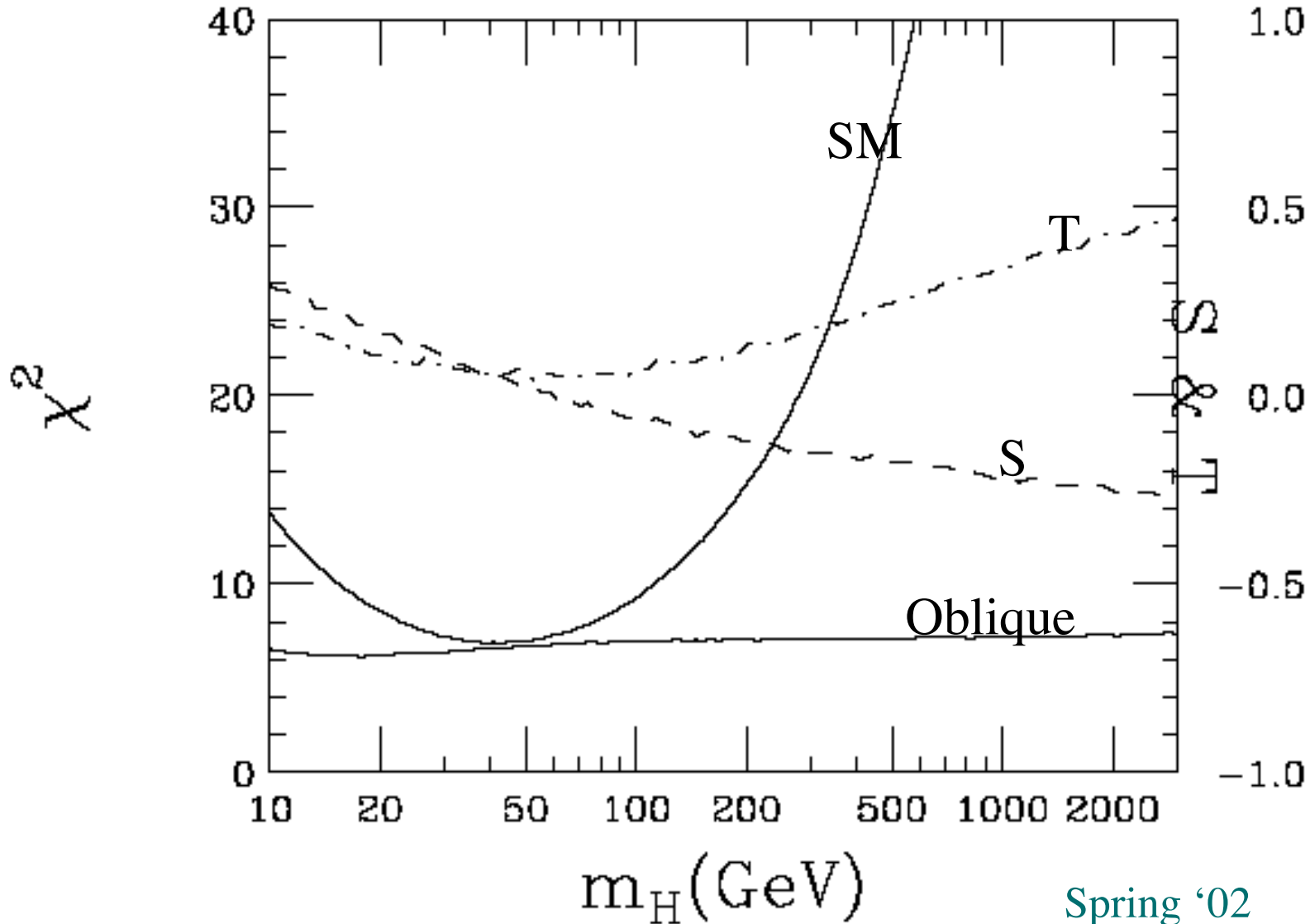


## New Physics alla oblique

- S,T alone do not improve fits with  $\chi^2[A_H]$ ,  $\chi^2_W$
- For fits with  $\chi^2[A_H]$  excluded, S,T can raise  $m_H$  arbitrarily

	<u>SM</u>	<u>S,T <math>\neq 0</math></u>
$\chi^2 =$	7.1/9	6.1/7
CL =	0.63	0.53
$m_H =$	43	All $m_H$ allowed

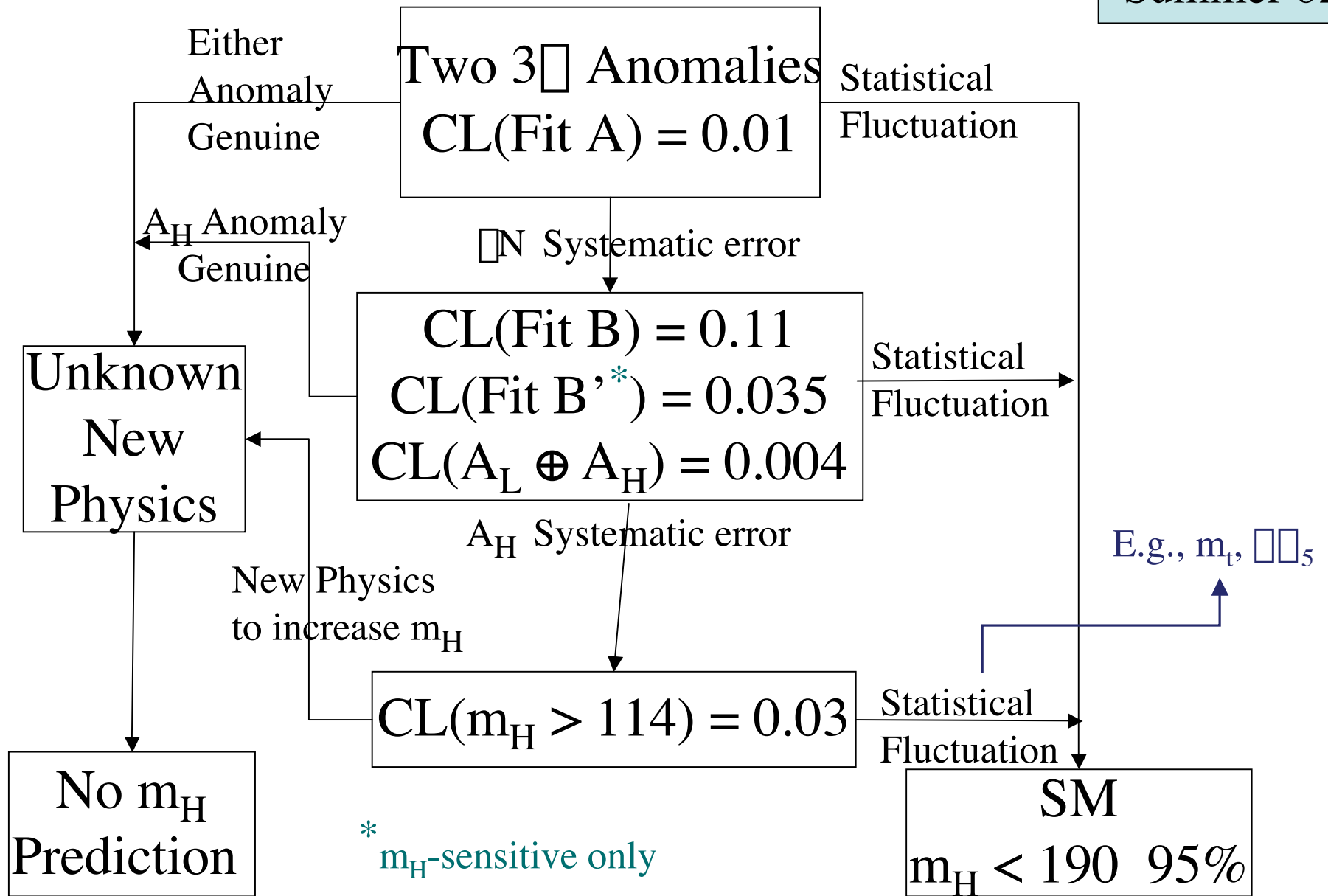
# Minimal Data Set



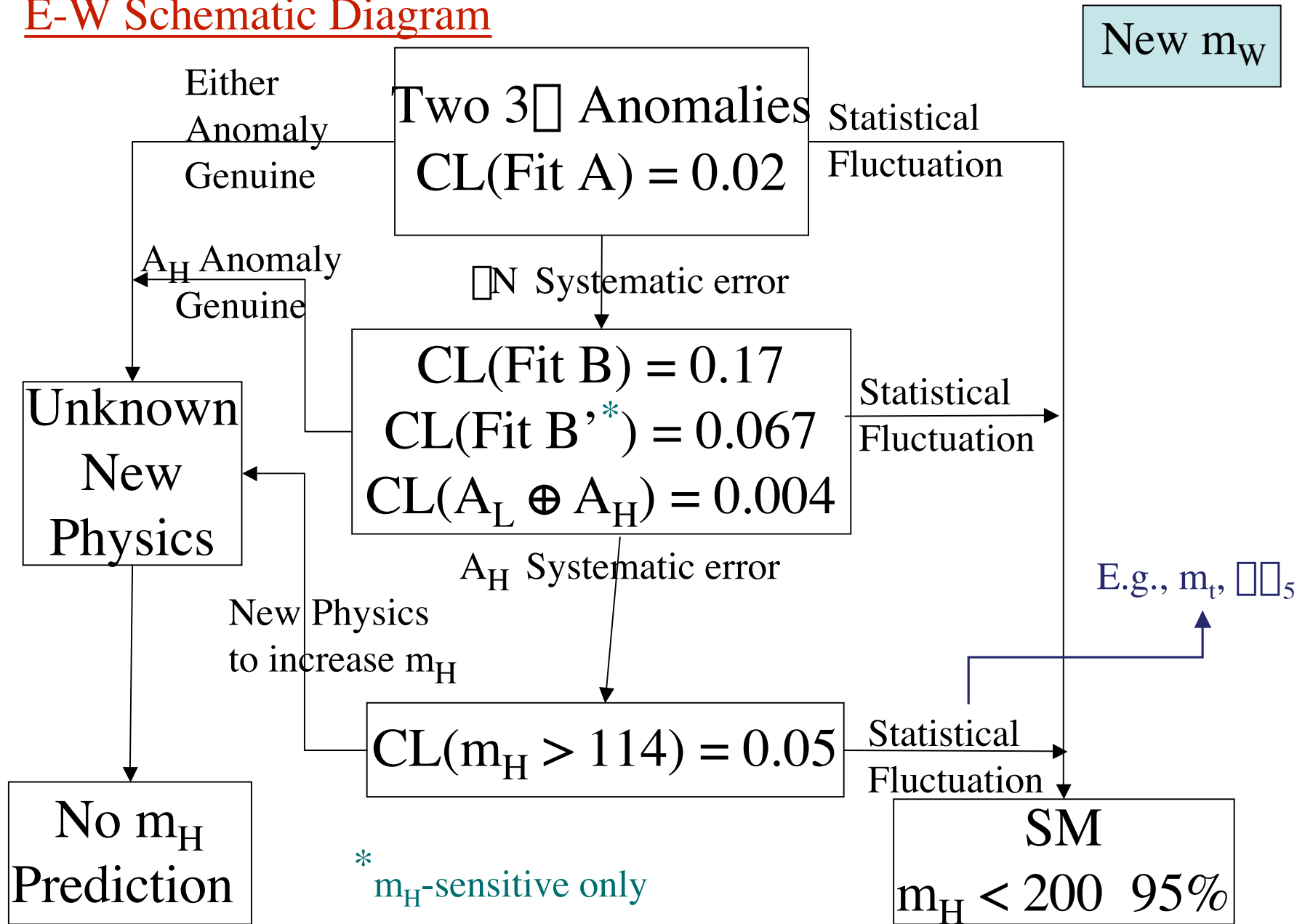
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# E-W Schematic Diagram

Summer 02



# E-W Schematic Diagram



With new  $m_W$ , betting odds that data reflects SM statistical fluctuation are improved, but both  $3\sigma$  anomalies remain.


Still not clear if anomalies result from new physics, systematic effects, or statistical fluctuation.

Systematic error hypothesis for both  $3\sigma$  anomalies now implies  $m_H < 114$  GeV at 95% CL, in conflict with the direct lower limit,  $m_H > 114$  GeV.

The E-W decade @LEP/SLC/FNAL has verified the SM at quantum loop level for non  $m_H$ -sensitive observables.

However,  $x[A_{FB}^b] - x[A_{LR}]$  discrepancy has proven to be a stubborn problem that refuses to go away.

LEP II limit on  $m_H$  makes problem even more stubborn:

- New physics is favored whether  $A_{FB}^b$  attributed to sys. error or not  
     no prediction for  $m_H$  until new physics is known.
- SM & usual  $m_H$  prediction require statistical fluctuations of both anomalous and non-anomalous measurements.

**What's it all mean?**

**Beats me --- a great puzzle!**

- The answer could begin to emerge at the TeVatron and will surely begin to emerge at LHC.
- Final clarity will probably require revisiting the Z boson with even greater precision, as for instance at Giga-Z.